

# Surface Properties of Strange Superheavy Nuclei

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- Motivation – in principle, neutral, strange, superheavy nuclei stable against strong decay can be arbitrarily large
- For some critical number of  $\Lambda$ 's in the nuclear medium!



- Include  $\Lambda$ 's and  $\Xi = \begin{pmatrix} \Xi^0 \\ \Xi^- \end{pmatrix}$
- Construct an effective lagrangian - limit to scalar and vector mesons only
- Impose the conditions  $Q = 0$  and  $|S|/B = 1$
- The resulting SEMF is

$$\frac{E}{B} = a_1 + a_2 B^{-1/3}$$

1 J. Schaffner, C. Dover, A. Gal, C. Greiner, and H. Stocker. Phys. Rev. Lett. 71, 1328 (1993)

## Modern approach to nuclear structure

- Construct an effective lagrangian using hadronic degrees of freedom and incorporate the following:
  - Special Relativity and Quantum Mechanics
  - The underlying symmetry structure of QCD
  - Nonlinear realization of spontaneously broken chiral symmetry
- The fields are small compared to  $M$
- Now a controlled expansion is performed in the effective lagrangian

$$\frac{\Phi}{M}, \frac{W}{M} \approx \frac{1}{3}, \quad \frac{k_F}{M} \approx \frac{1}{4}$$

- If we assume that the lagrangian provides an appropriate energy functional, then Density Functional Theory states that this lagrangian can be used to reproduce the exact ground-state scalar and vector densities, energy, and chemical potential

- Compute the density functional by assuming that the system is a local Fermi gas filled up to some  $k_F(r)$
- Make the simplest approximation for both the nucleons and hyperons
- Retain a couple of nonlinear scalar field self couplings

- The energy density

$$\begin{aligned} \mathcal{E}[\phi_0, V_0, \rho_B] = & \frac{4}{(2\pi)^3} \int_0^{k_F} d^3k \sqrt{(k^2 + M^*)^2} + \frac{4}{(2\pi)^3} \int_0^{k_F} d^3k \sqrt{(k^2 + M_\Xi^*)^2} \\ & + \frac{2}{(2\pi)^3} \int_0^{k_{F\Lambda}} d^3k \sqrt{(k^2 + M_\Lambda^*)^2} + \frac{1}{2} [(\nabla\phi)^2 + m_S^2\phi_0^2] + \frac{\kappa_3 m_S^2}{3!} \frac{g_S \phi_0^3}{M} \\ & + \frac{\kappa_4 m_S^2}{4!} \frac{g_S^2 \phi_0^4}{M^2} + g_V \rho_B V_0 - \frac{1}{2} [(\nabla V_0)^2 + m_V^2 V_0^2] \end{aligned}$$

- Effective mass

$$M^* \equiv M - g_S \phi_0$$

- Nucleon coupling constants<sup>1</sup>

<sup>1</sup> B. Serot and J. D. Walecka Inter. J. of Mod. Phys. E6 515 (1997)

- Hyperon coupling constants
  - Fit scalar couplings to experimental data where it exists
  - Assume a universal vector coupling to the conserved baryon current
- The meson equations are
 
$$(\nabla^2 - m_S^2)\phi_0 = -g_S \rho_S + \frac{\kappa_3 m_S^2}{2} \frac{g_S \phi_0^2}{M} + \frac{\kappa_4 m_S^2}{6} \frac{g_S^2 \phi_0^3}{M^2}$$

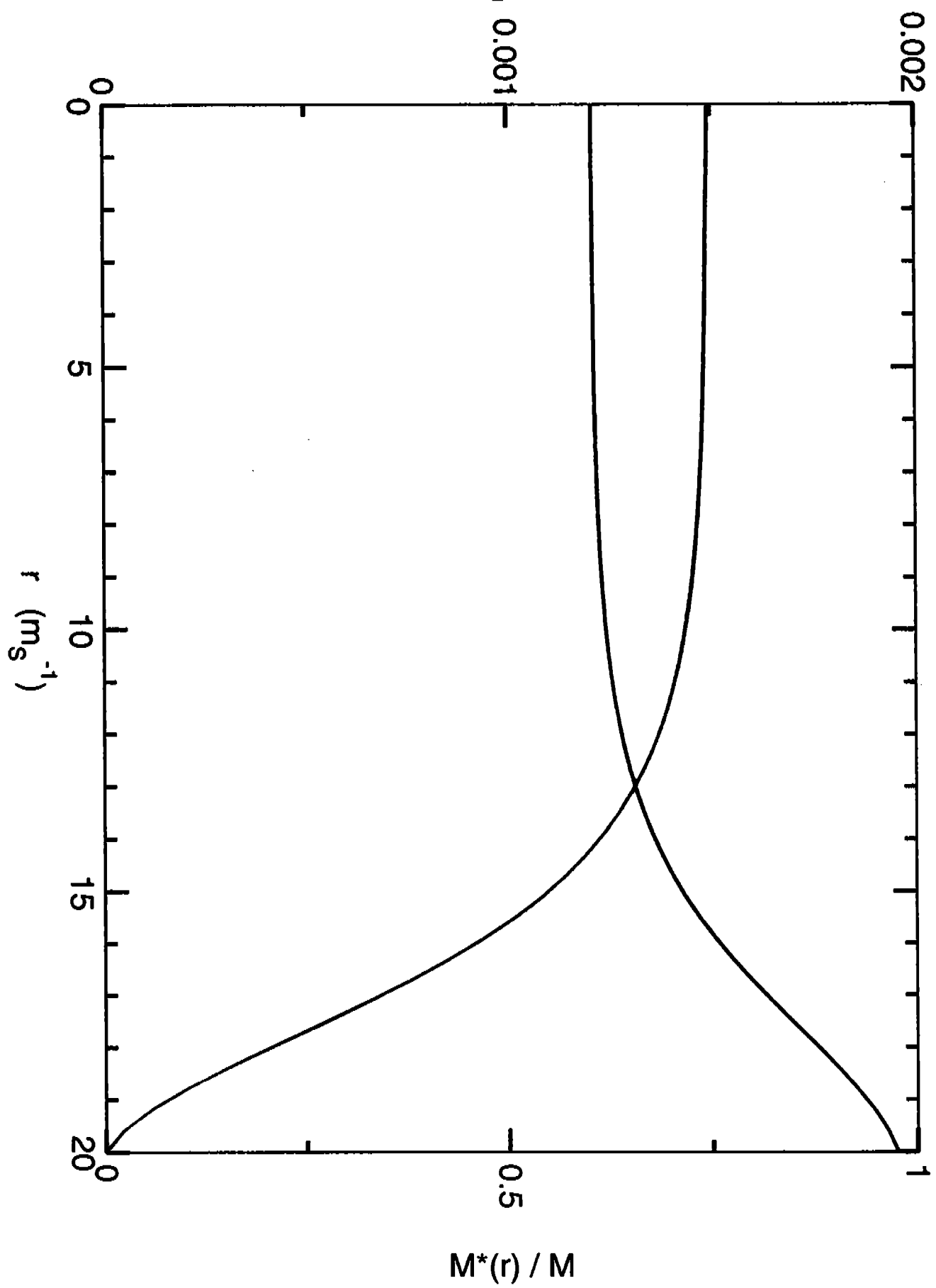
$$(\nabla^2 - m_V^2)V_0 = -g_V \rho_B$$
- The constraint equations - to be solved for  $k_F(r)$ 

$$\mu = g_V V_0 + \frac{1}{2} \left[ (k_F^2 + M^{*2})^{1/2} + (k_F^2 + M_\Xi^{*2})^{1/2} \right]$$

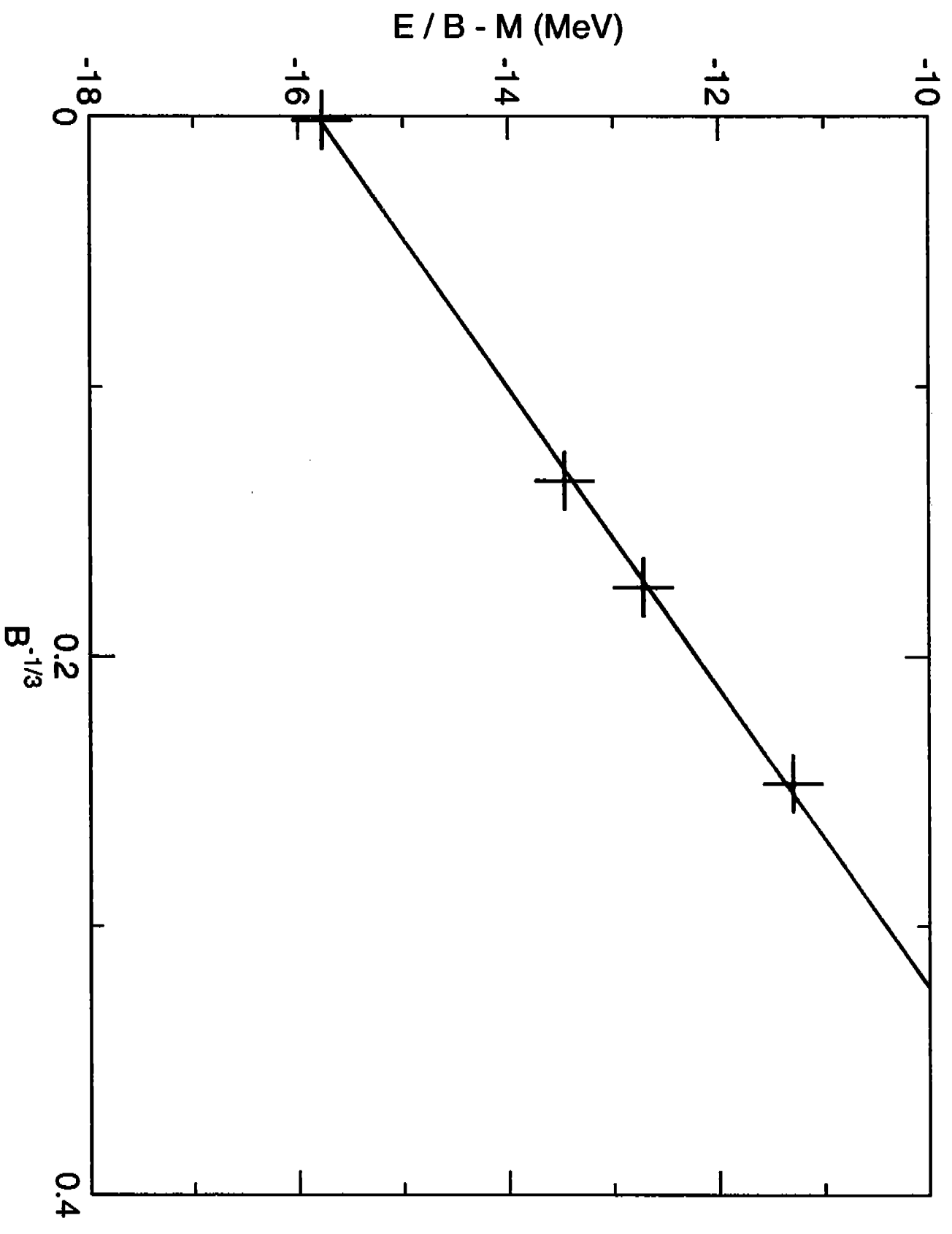
$$\mu = g_V V_0 + (k_{F\Lambda}^2 + M_\Lambda^{*2})^{1/2}$$
- First reproduce nucleon matter with  $N = Z$

# Nucleon Matter with $N=Z$

$$n_B(r) = \rho_B(r) / M^3$$



# Nucleon Matter with $N=Z$



## Surface energies for nucleon matter

	a2 (MeV)
L2 <sup>1</sup>	26.51
NLC <sup>1</sup>	18.01
Q1 <sup>1</sup>	19.11
Expt.	17.8

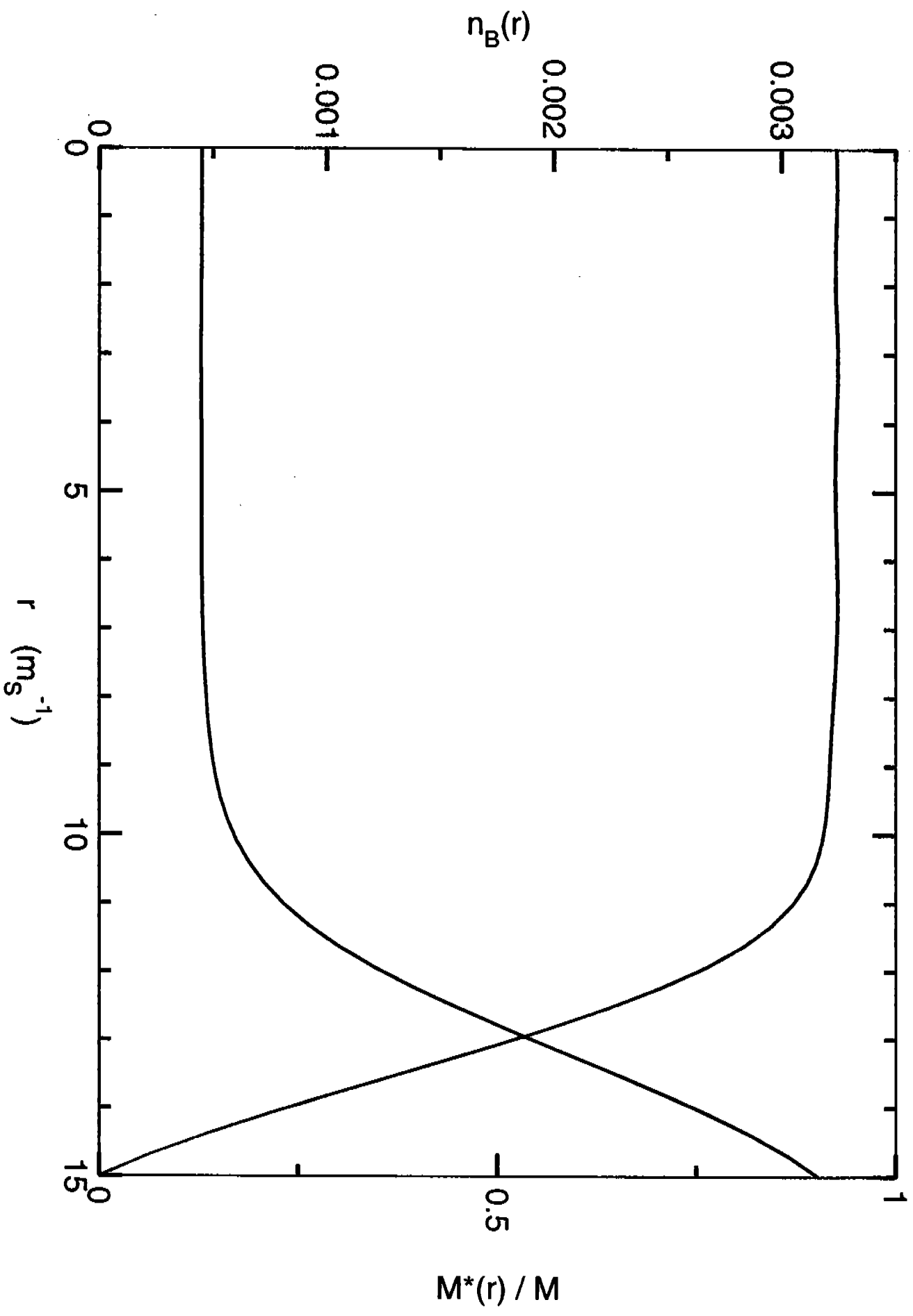


– This calibrates the approach

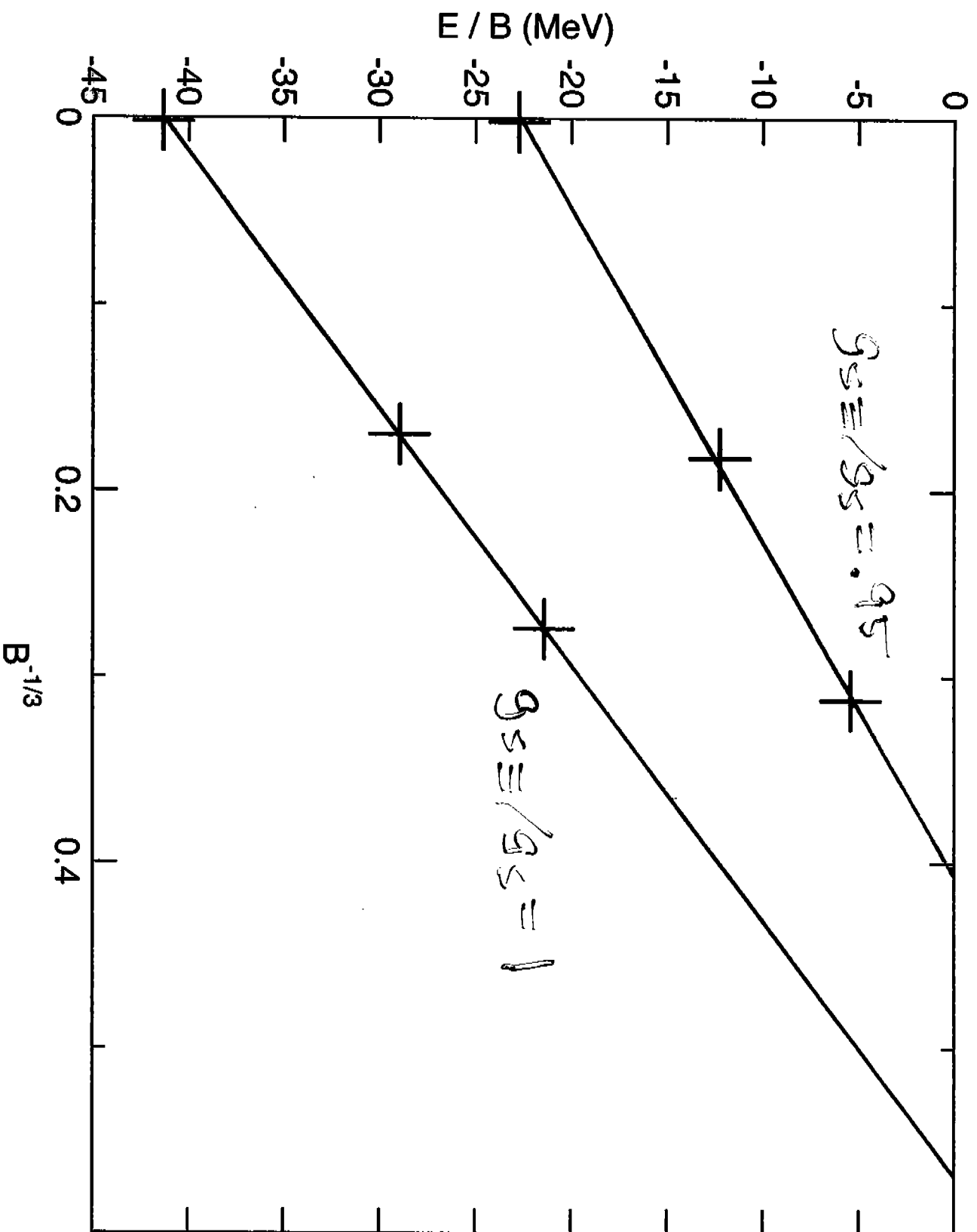
<sup>1</sup> B. Serot and J. D. Walecka Inter. J. of Mod. Phys. E6 515 (1997)



# Cascade Nucleon Matter



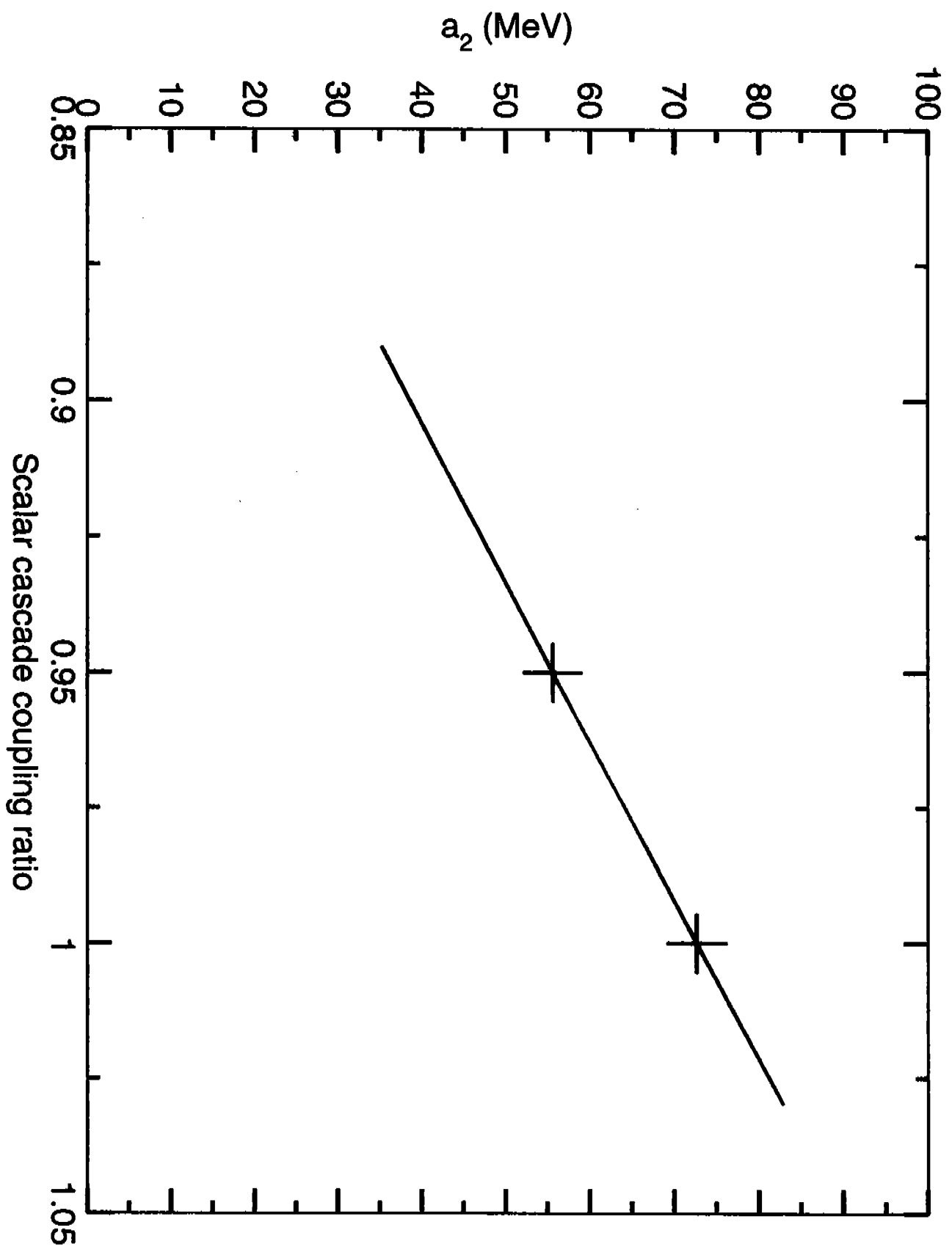
# Cascade Nucleon Matter



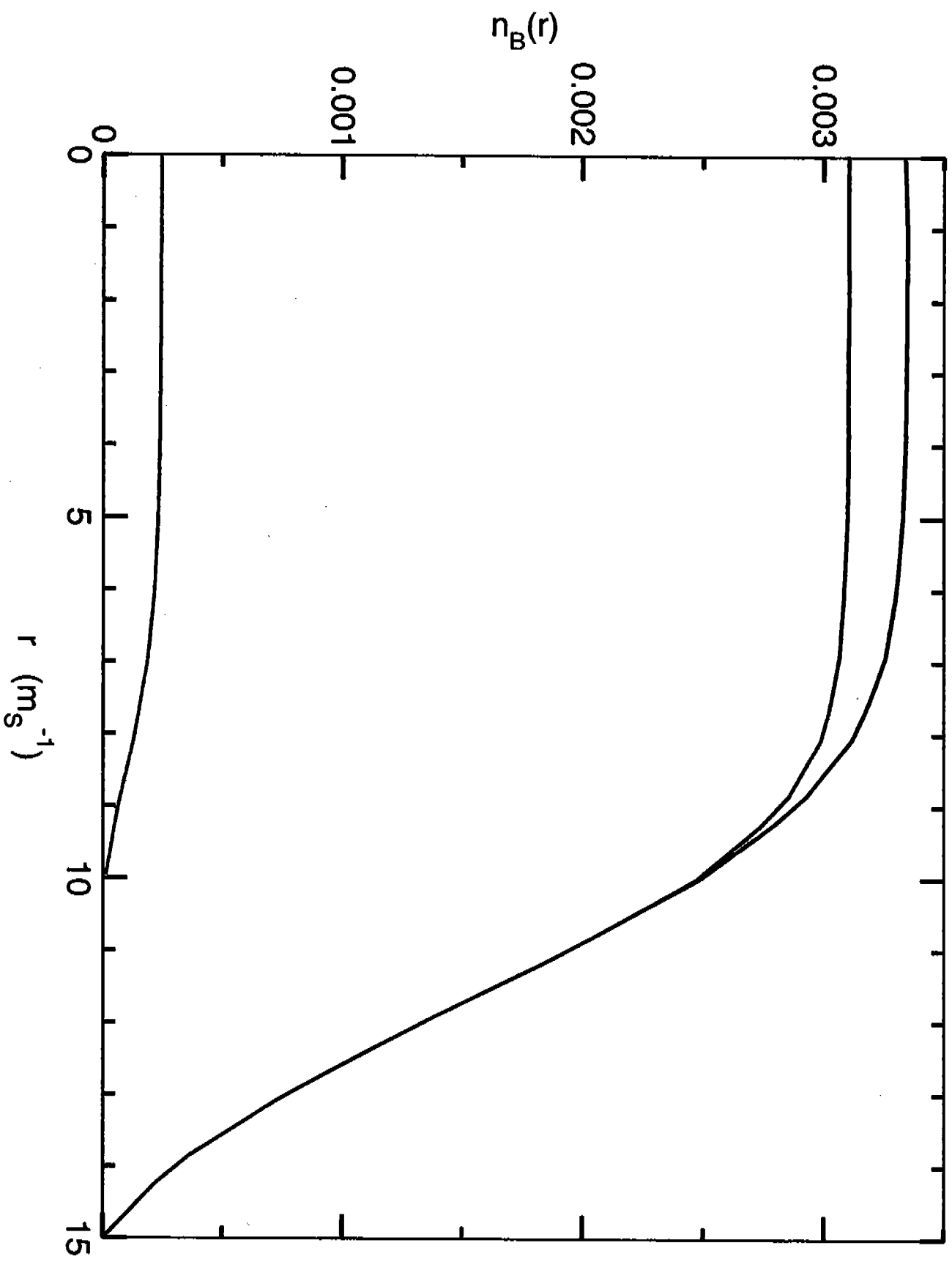
# Surface energies for cascade-nucleon matter<sup>1</sup>

	$g_{SE} / g_S$	min B	a2 (MeV)
NLC	1	5.4	72.69
	0.95	14.6	55.64

<sup>1</sup> J. McIntire. Phys. Rev. C66 064319 (2002)



# Lambda Cascade Nucleon Matter



## Conclusions

- Reproduced the surface energy for ordinary nuclei
- Calculated the surface energy and structure for strange superheavy nuclei