Role of Hyperons on the Hadron Star to Quark Star Conversion Model

I. Bombaci, I. Parenti and I. Vidaña





Compact Objects

Hadronic Stars (HS)

- Traditional NS
- Hyperonic Stars





Scenario for HS to HyS(SS) conversion

Central Pressure increase due to spin down or mass accretion

Formation of drop of QM



QM drop formation



After a time τ the virtual QM drop tunnels the potential well becoming a real QM drop



$$d
ightarrow u + e^- + ar{
u}_{e}$$

 $u + e^-
ightarrow d +
u_{e}$
 $s
ightarrow u + e^- + ar{
u}_{e}$
 $u + e^-
ightarrow s +
u_{e}$

Weak interactions acts changing the quark flavor fraction of deconfined QM drop to lower its energy and a drop of β -stable SQM is formed

Lifshitz-Kagan quantum nucleation theory

Quantum fluctuation of a virtual drop of QM in HM

$$L = -M(R)\sqrt{1 - \left(\frac{dR}{dt}\right)^2} + M(R) - U(R)$$

$$M(R) = 4\pi\rho_H \left(1 - \frac{n_Q}{n_H}\right)^2 R^3$$
QM drop
Hadronic
Matter



$$U(R) = \frac{4}{3}\pi n_{Q} (\mu_{Q} - \mu_{H}) R^{3} + 4\pi\sigma R^{2}$$

$$\sigma \sim 10 - 50 MeV / fm^{2}$$

Nucleation Time

Oscillation frequency of the virtual drop inside the potential well and Penetrability of the potential barrier (WKB)

$$v_0 = \left(\frac{dI}{dE}\right)^{-1} at E = E_0$$
$$p_0 = \exp\left[-\frac{A(E_0)}{\hbar}\right]$$

$$I(E) = 2 \int_{R_1}^{R_1} dR \sqrt{[2M(R) + E - U(R)][E - U(R)]]}$$

$$A(E) = 2 \int_{R_1}^{R_2} dR \sqrt{[2M(R) + E - U(R)][U(R) - E]]}$$

Action over and under the barrier

Nucleation time
$$\tau = (v_0 p_0 N_c)^{-1}$$
, $N_c \sim 10^{48}$

The EoS of dense matter

GM1: K=300 MeV, m^{*}/m=0.7 GM3: K=240 MeV, m^{*}/m=0.78

B/A=-16.3 MeV, ρ =0.153 a_{sym} =32.5 MeV



Hadronic phase: RelativisticMean Field Theory of hadronsinteracting meson exchange.[Glendenning & Moszkowsky,Phys. Rev. Lett. 67 (1991)]

Mixed phase: Gibbs construction for a multicomponent system with two conserved "charges". [Glendenning, Phys. Rev. D 46 (1992)]

Quark phase: EoS based on the MIT Bag Model for hadrons. [Farhi & Jaffe, Phys. Rev. D 46 (1992)]

The Equation of State

Without hyperons

With hyperons



Chemical potentials



Potential barrier



Hadronic Star mean-life time



Conversion of a HS into a HyS(SS)

Without hyperons With hyperons



Summary and Conclusions

- I-. Existence of two different families of compact objects: Hadronic Stars and Hybrid or Strange Stars.
- II-. Hadronic Stars are metastable to "decay" to Hybrid Stars or to Strange Stars. The HS mean-life time range within $\tau >>$ age of the universe to few years (days) and depends dramatically on the central pressure.
- III-. The conversion mechanism is favoured by the presence of hyperons.