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# Approximations to the CLAS Magnetic Field M.V. Kossov, CNU/CEBAF/ITEP

#### 1. Introduction

Historically, the magnetic field of the CLAS detector has been calculated by the subroutine TORUS which subdivides conductors into small straight elements. The goal of this note is to compare two approximation methods for the magnetic field. Both approximation algorithms use a data base consisting of the node values of the magnetic field, calculated by the TORUS subroutine. In the SDA package, a 2nd order approximation algorithm is used, and in the IFRAME package a linear approximation algorithm is used. These two algorithms will be compared to make a decision which algorithm should be used in CLAS GEANT simulation.

For the comparison three parameters were used:

- 1. The number of nodes in the data base for the approximation (the size);
- 2. The speed of the field approximation;
- 3. The accuracy of the  $\int Bdl$  calculation in comparison with exact TORUS values.

### 2. 2-nd order approximation of CLAS magnetic field in SDA.

The SDA package uses the subroutine SNAKE, which performs a 3-dimensional 2nd order interpolation for each component of the magnetic field using Taylor expansion. For the data base the node's values of the cubic 3-dimensional table are:

- 1. 61 nodes along the X-axis (along a mid.plane perpendicular to the beam direction) with 7 cm step ( $X_{min} = -7 \text{ cm}, X_{max} = 420 \text{ cm}$ );
- 2. 41 nodes along the Y-axis (along the wires of the CLAS Drift Chamber) with 7 cm step ( $Y_{min} = -7$  cm,  $Y_{max} = 280$  cm);
- 3. 87 nodes along the Z-axis (along the ideal beam direction) with 10 cm step (  $Z_{min} = -350 \text{ cm}, Z_{max} = 520 \text{ cm}$ ).

The SNAKE algorithm uses 8 points, which define a cube around the space point, where the field value is needed: (i,i,i), (i+1,i,i), (i,i+1,i), (i,i,i+1), (i,i+1,i+1), (i+1,i+1,i), (i+1,i+1,i+1). In addition, three upstream points (i-1,i,i), (i,i-1,i), (i,i,i-1) are used. Because of these upstream points the data base needs additional points with negative values of X and Y. The algorithm is as follows:

```
xx=x-x_{i,i,i}
yy=y-y_{i,i,i}
zz=z-z_{i,i,i}
bx = f_{i+1,i,i} - f_{i-1,i,i}
by = f_{i,i+1,i} - f_{i,i-1,i}
bz = f_{i,i,i+1} - f_{i,i,i-1}
ff = f_{i,i,i} + f_{i,i,i}
cx = f_{i+1,i,i} + f_{i-1,i,i}-ff
cy = f_{i,i+1,i} + f_{i,i-1,i}-ff
cz = f_{i,i,i+1} + f_{i,i,i-1}-ff
xy = f_{i,i,i} - f_{i+1,i,i} + f_{i+1,i+1,i} - f_{i,i+1,i}
yz = f_{i,i,i} - f_{i,i+1,i} + f_{i,i+1,i+1} - f_{i,i,i+1}
zx = f_{i,i,i} - f_{i,i,i+1} + f_{i+1,i,i+1} - f_{i+1,i,i}
dd = f_{i+1,i+1,i+1} - f_{i+1,i,i+1} + f_{i,i,i+1} - f_{i,i+1,i+1} - f_{i,i,i} + f_{i+1,i,i} - f_{i+1,i+1,i} + f_{i,i+1,i}
f = f_{i,i,i} + (xx \cdot (bx + xx \cdot cx) + yy \cdot (by + yy \cdot cy) + zz \cdot (bz + zz \cdot cz) + xx \cdot yy \cdot xy + yy \cdot zz \cdot yz + zz \cdot zz
xx \cdot zx + xx \cdot yy \cdot zz \cdot dd)/2
```

An obvious disadvantage of this algorithm is the large number of dummy nodes in case of the 60 degrees sector. The cross section of the grid in a plane perpendicular to the beam is a rectangle, while a cross section of the sector is a triangle. That is why at least half of nodes are dummy.

### 3. Linear approximation of CLAS magnetic field in IFRAME.

The IFRAME package uses the algorithm of the FINT function (CERNLIB). The existing FINT algorithm is optimized for the constant step (it is arbitrary in the original function) and for 3 dimensions (the original function can work with up to 5 dimensions). The FINT algorithm performs a simple first order interpolation. For the data base the following nodes were used:

- 1. 17 nodes for the  $tg(\phi)$  subdivision with a step 0.03625 (  $tg(\phi)_{min} = 0$ ,  $tg(\phi)_{max} = 0.58$ );
- 2. 129 nodes along the Z-axis (along the ideal beam direction) with 7 cm step (  $Z_{min} = -360 \ cm, Z_{max} = 535 \ cm$ ).
- 3. 65 nodes along the X-axis (along a middle plane perpendicular to the beam direction) with 7 cm step ( $X_{min} = 0$  cm,  $X_{max} = 448$  cm).

The IFRAME data base covers a 15% larger volume than the SDA data base, while the number of nodes is 1.53 times smaller than in SDA data base.

The FINT algorithm uses only 8 points, which define a cube around the space point, where the field value is needed. The algorithm for the calculation is as follows:

```
xx=x-x_{i,i,i}

yy=y-y_{i,i,i}

zz=z-z_{i,i,i}
```

```
d\mathbf{x} = x_{i+1,i,i} - x_{i,i,i}
dy=y_{i,i+1,i}-y_{i,i,i}
dz=z_{i,i,i+1}-z_{i,i,i}
xn=xx/dx
yn=yy/dy
zn=zz/dz
xm=1.-xn
ym=1.-yn
zm=1.-zn
fy_{i,i}=f_{i,i,i}\cdot xm+f_{i+1,i,i}\cdot xn
f_{y_{i,i+1}} = f_{i,i,i+1} \cdot xm + f_{i+1,i,i+1} \cdot xn
f_{y_{i+1,i}} = f_{i,i+1,i} \cdot xm + f_{i+1,i+1,i} \cdot xn
f_{i+1,i+1} = f_{i,i+1,i+1} \cdot xm + f_{i+1,i+1,i+1} \cdot xn
fz_i = fy_{i,i} \cdot ym + fy_{i+1,i} \cdot yn
fz_{i+1}=fy_{i,i+1}\cdot ym+fy_{i+1,i+1}\cdot yn
f=fz_i\cdot zm+fz_{i+1}\cdot zn
```

#### 3. Speed and accuracy.

The linear IFRAME algorithm is obviously faster than the 2-nd order SDA algorithm. Measurements on a VAX computer show that the field extraction by the SNAKE subroutine requires 2 msec while the extraction by the modified FINT subroutine requires 1.25 msec. That is a linear algorithm is factor of 1.6 faster than the 2nd order algorithm.

The accuracy of the calculations was compared using  $\int Bdl$  calculations for the exact TORUS magnetic field values and for approximated values. The  $\int Bdl$  integrations were made along directions with random  $\theta$  and  $\phi$ . For most of the points (67%) the accuracy of the  $\int Bdl$  integration is better then 1% for both algorithms. In case the accuracy is worse than 1% for at least one projection, the resulting relative differences are listed in the table.

One can see, that only at very small angles ( $\theta$ =11.48 in Table 1) and at very large angles ( $\theta$ =137.31 in Table 1) IFRAME algorithm gives worse results. In all other cases it seems to be even more accurate than the SDA algorithm.

It should be noted, that if the same magnetic field approximation is used in a simulation and in a reconstruction, the relative systematic errors in the acceptances will be much smaller than the relative errors in  $\int Bdl$  integrals.

Table 1. Fractional accuracies of  $\int Bdl$  for the SDA and IFRAME methods

$\theta(\deg)$	$\phi(\deg)$	Bx(SDA)	Bx(IFR)	By(SDA)	By(IFR)	Bz(SDA)	Bz(IFR)
11.48	7.97	.0601	.0701	.0008	0073	0031	0090
11.48	9.03	.0426	.0714	.0005	0072	0007	0095
11.48	9.54	.0282	.0780	.0004	0072	.0020	0096
11.48	12.13	.1029	.1748	.0020	0077	.0084	0101
11.48	24.18	.1435	.0211	0058	0091	.0124	.0119
62.23	1.89	0073	.0001	0012	.0002	.0177	.0001
62.23	2.41	0063	0002	0012	.0001	.0172	0002
71.88	7.46	0218	0069	0022	0006	.0166	.0050
71.88	10.98	0176	0057	0008	0002	.0139	.0042
71.88	14.57	0122	0040	.0008	.0004	.0102	.0033
72.72	6.28	0232	0082	0027	0008	.0154	.0060
72.72	15.70	0110	0043	.0011	.0004	.0085	.0031
73.20	2.86	0247	0096	0037	0013	.0159	.0055
73.20	11.86	0166	0060	0005	0001	.0109	0039
73.20	12.02	0163	0058	0004	0001	.0108	.0039
79.28	12.73	0197	0080	0001	.0000	.0027	.0013
79.28	16.49	0125	0050	.0016	.0008	.0017	.0010
79.28	16.75	0120	0050	.0017	.0008	.0017	.0008
90.00	1.03	0397	0173	0052	0022	0159	0063
90.00	7.24	0333	0147	0032	0013	0132	0054
90.00	12.19	0237	0099	0004	.0002	0100	0036
94.30	0.92	0370	0165	0047	0021	0188	0085
94.30	3.26	0358	0160	0044	0020	0186	0082
94.30	14.09	0185	0081	.0007	.0003	0111	0042
94.30	15.11	0163	0077	.0011	.0004	0101	0043
95.62	16.86	0124	0047	.0018	.0007	0094	0032
96.26	4.69	0295	0127	0034	0014	- 0174	0073
96.26	12.79	0177	0078	0001	.0001	0120	0049
96.26	15.16	0136	0062	.0009	.0005	0099	0042
96.89	13.98	0162	0071	.0004	.0003	0113	0045
110.18	8.87	0105	0032	0007	0003	0096	0020
111.96	26.89	.0390	0029	.0358	0002	.0066	.0046
122.28	2.12	0150	0052	.0000	0005	.0016	.0017
132.22	17.22	.0122	0083	0011	0024	.0038	0014
132.22	20.41	.0218	0005	0003	0043	.0066	0003
132.22	25.69	0863	0522	.0154	.0051	.0154	.0151
137.31	25.64	.0098	0164	0018	0123	0035	0154