

Fully Coherent Hard X-ray Free-Electron Lasers

Ryan R. Lindberg Advanced Photon Source, Argonne National Laboratory

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Outline

- Intro: two paths to completely coherent hard x-rays
 - Apply monochromator to self-amplified spontaneous emission and amplify the result in downstream FEL (self-seeding)
 - Form x-ray cavity using Bragg mirrors for x-ray FEL oscillator
- Operating principles of self-seeding
- "Wake monochromator" as a simple self-seeding scheme
- An alternative approach: the x-ray FEL oscillator (XFELO)
- XFELO performance and scaling
- Stability concerns
- Physics of Bragg scattering for oscillators and wake monochromators
- Performance of high-quality diamond crystals
- Conclusions

FELs based on self-amplified spontaneous emission (SASE) are now producing hard x-rays for science

LCLS pioneered bright, hard x-ray sources based on SASE, with many more projects underway

SPring8 SACLA in operation, Euro XFEL under construction, LCLS II under development, Swiss FEL designed, PAL XFEL (Korea) in planning stages, ...



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- High pulse intensity (~ 10¹² photons/pulse)
- Temporally chaotic
 Normalized bandwidth $\Delta\omega/\omega \sim 10^{-3}$
- Sub-femtosecond pulses possible



Flagship applications include:

Single shot x-ray imaging, Nonlinear physics, and Atto-second dynamics

Brightness limits of SASE

- Because the FEL gain is initiated by electron beam shot noise, the SASE radiation comprises many longitudinal modes (temporally chaotic)
- In the frequency domain, the shot noise seeds radiation over the entire FEL bandwidth, so that $\Delta\omega/\omega\approx\rho$

Either the pulse is not Fourier limited/coherent OR The x-ray pulse is short $(\sim\lambda/c\rho)$ with $\sim100\%$ shot-to-shot fluctuations in energy ("single spike")¹

 Longitudinal coherence/spectral brightness can be improved by initializing the interaction with a coherent signal at the wavelength of interest

Harmonics of induced current modulation (High Gain Harmonic Generation², Echo-Enabled Harmonic Generation³) Coherent radiation sources (using, e.g., High Harmonic Generation)

SASE + monochromator + radiator FEL ("self-seeding") Oscillator using Bragg crystal mirrors

1. R. Bonifacio, L. De Salvo, N. Piovella, and C. Pellegrini, *Phys. Rev. Lett.* 73, 70 (1994).

2. L. H. Yu, Phys. Rev. A 44, 5178 (1991).

3. G. Stupakov, Phys. Rev. Lett. 102, 074801 (2009).

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SASE + monochromator + radiator FEL ("self-seeding")

Oscillator using Bragg crystal mirrors

Appear to be limited to soft x-rays

Applicable to hard x-rays

1. R. Bonifacio, L. De Salvo, N. Piovella, and C. Pellegrini, Phys. Rev. Lett. 73, 70 (1994).

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Self-seeding schematic





- Power large enough so that monochromatic signal dominates the seeding of SASE by electron beam shot noise
- FEL-induced energy spread on beam must remain small (well before nonlinear regime)

Example uses LCLS-type parameters with ρ = 5×10⁻⁴, $\sigma_e \approx$ 12 fs and $(\Delta \omega / \omega)_{mono}$ = 2×10⁻⁵



- Monochromator selects narrow bandwidth seed whose energy fluctuates by ~100%
- Electron beam must be delayed to overlap and amplify radiation in downstream undulator

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The monochromator and e-beam bypass (chicane)

Standard 4-bounce monochromator delays monochromatic seed ~30 ps



Chicane delays radiation for overlap and washes out the SASE-induced microbunching



- 1. E.L. Saldin, E.A. Schneidmillera, Yu.V. Shvyd'ko, and M.V. Yurkov, Nucl. Instrum. Methods Phys. Res., Sect. A 475, 357 (2001)
- 2. Y. Ding, Z. Huang, and R. D. Ruth, Phys. Rev. ST Accel. Beams 13, 060703 (2010)

The "wake" monochromator for small delays from temporally short beams

 Uses time dependence of forward Bragg diffraction from a single crystal, i.e., time response of the crystal transmission function¹



1. G. Geloni, V. Kocharyan , and E.L. Saldin, J. Modern Optics 58, 1391 (2011)

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 Uses time dependence of forward Bragg diffraction from a single crystal, i.e., time response of the crystal transmission function¹



For broad initial spectrum and optimal seeding overlap, best suited to short ~fs pulses

1. G. Geloni, V. Kocharyan ,and E.L. Saldin, J. Modern Optics 58, 1391 (2011

Post-saturation taper of undulator can result in significant gains in field energy

- Tapering the undulator strength (or period) lowers the ponderomotive potential, so that particles can continue to lose energy to the field
- In SASE, different coherent regions of the electron beam do different things



In seeded FEL, the average phase of particles is uniform across the bunch

Single taper can be used to extract more energy across entire bunch

Significant increase in FEL power and efficiency



SLAC-ANL-TISNCM collaboration has demonstrated a reduction of bandwidth ~10 to 40 using the wake monochromator at the LCLS

- Configuration proposed by scientists at DESY¹
- <u>SLAC</u>: project lead; designed and built chicane; built monochromator; prepared control systems; installed hardware; designed and built spectrometer; ...
- <u>ANL</u>: designed and built the vacuum chamber and YAG diagnostic; designed monochromator tank; procured and tested diamond crystals
- **<u>TISNCM</u>**: grew high-quality diamond crystals of 100 and 150 μm thickness
- Data still being analyzed and prepared for publication

1. G. Geloni, V. Kocharyan ,and E.L. Saldin, "Cost-effective way to enhance the capabilities of the LCLS baseline", DESY 10-133, arXiv:1008.3036 (2010)

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Basic idea has been confirmed Significant reduction in the bandwidth has been measured Time delay scan agrees with theory

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Decrease in fluctuations yet to be observed Additional tapering has not improved performance to date Nonlinear saturation not yet reached

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First proposed by Colella and Luccio using Silicon mirrors, but without detailed understanding of the FEL physics the idea went dormant as SASE came into prominence¹

Interest in the oscillator was renewed recently when concrete, realizable parameters for the undulator, the electron beam and the x-ray optical cavity were presented by Kim, Shvyd'ko, and Reiche²



2. K-J. Kim, Yu. Shvyd'ko, and S. Reiche, Phys. Rev. Lett. 100, 244802 (2008)

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Parameters

Cavity length

~ 100 m	→	Repetition rate \sim 1 MHz
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Peak Currents $\sim 10 - 100$ A Bunch Lengths $\sim 2 - 0.1$ ps

• Low charge $\sim 10 - 50 \text{ pC}$

High quality electron beam: $\epsilon_{x,n} \sim 0.1 - 0.3 \text{ mm-mrad}, \Delta \gamma / \gamma \sim 0.01 - 0.03\%$ Low single pass gain, losses: $G \sim 0.3 - 1.0, R_{tot} \sim 0.85 - 0.5 > (1 + G)^{-1}$

Performance

Covity power 10 200 MM	Output power	Example @ 1 A	
Cavity power $\sim 10 - 200$ Miv \checkmark	$\sim 0.5 - 10 \text{ MW}$	Pout	1.7 MW
Output ~ $10^9 - 5 \times 10^9$ photons \rightarrow	Spectral brightness $\sim 10^{31}$ – 10^{34}	Photons/ pulse	1.1×10 ⁹
	$\Lambda E = 1 = 10 \text{ moV}$	ΔΕ	1.29 meV
Nearly Fourier limited pulses \rightarrow	$\Delta \omega / \omega \sim 10^{-6} - 10^{-8}$	Δt	0.51 ps
	,		

R.R. Lindberg, K-J. Kim, Yu. Shvyd'ko, and W.M. Fawley, Phys. Rev. ST-AB. 14, 010701 (2011)

The Advanced Photon Source is an Office of Science User Facility operated for the U.S. Department of Energy Office of Science by Argonne National Laboratory

Example @ 1 Å

E _{beam}	7 GeV	
I _{peak}	10 A	
ε _{<i>x</i>,<i>n</i>}	0.2 mm-mrad	
$\Delta \gamma / \gamma$	0.02%	
L _{und}	52 m	
G	0.36	
<i>R_{tot}</i> 0.85		
crystal C(4 4 4)		

Evample @ 1 Å

Temporal and spectral evolution of XFELO



R.R. Lindberg, K-J. Kim, Yu. Shvyd'ko, and W.M. Fawley, Phys. Rev. ST-AB. 14, 010701 (2011)

Tunable source of hard x-rays



By varying the incidence angle Θ , one can obtain a wide range of photon energies that satisfy Bragg's law $E = E_H \cos \Theta$

Tunability allows one to pick a single crystal for all wavelengths of interest

R.M.J. Cotterill , *Appl. Phys. Lett.* **12**, 403 (1968) K.-J. Kim and Yu. Shvyd'ko, *Phys. Rev. ST-AB* **12**, 030703 (2009)

X-ray FEL oscillator: a complementary hard x-ray source

Characteristic	SASE	XFELO
Pulse duration (fs)	1 to 100	100 to 1000
Photons/pulse	1012	10 ⁹
ΔE (eV)	10	10 ⁻³
Coherence	Transverse	Fully

Science:

Inelastic x-ray scattering, Nuclear resonant scattering, X-ray photoemission spectroscopy, Hard x-ray imaging, X-ray photon correlation spectroscopy Will revolutionize techniques pioneered at 3rd generation light sources, and complement the science of SASE FELs



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Science Opportunities with an XFELO Workshop, APS, May 5th, 2010

XFELO scales well to extremely low electron beam charge and emittance

As an extreme example, consider a suitably modified version of the parameters first proposed for the ultra-short, "single spike" regime of high-gain FELs¹:

Q = 1 pC $\varepsilon_{xn} = 0.062 \text{ mm} \cdot \text{mrad}$ $\Delta E = 250 \text{ keV}$ $\sigma_e = 250 \text{ fs} \Rightarrow I = 1.6 \text{ A}$

Using the same undulator (N_u = 3000) and optical cavity, G = 74%

$$\begin{split} R &= 85\% \twoheadrightarrow P_{\rm out} = 9.5 \; {\rm MW}, N_{\rm ph} = 10^8 \\ R &= 50\% \twoheadrightarrow P_{\rm out} = 600 \; {\rm kW}, N_{\rm ph} = 5 \times 10^6 \end{split} \qquad \Delta \omega / \omega \approx 2 \times 10^{-6} \quad \hbar \Delta \omega \approx 3 \; {\rm meV} \end{split}$$

Reduced heat load on crystals → Increase repetition rate

1. J.B. Rosenzweig, et al.,, Nucl. Instrum. Methods A. 593, 39 (2008)

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 $\Delta\omega/\omega \approx 2 \times 10^{-6}$ $\hbar\Delta\omega \approx 3 \text{ meV}$

Reduced heat load on crystals → Increase repetition rate up to ~GHz



Stability

- Self-seeding
 - Intrinsic: 100% fluctuations in linear regime ~20% fluctuations at/after saturation
- Oscillator
 - Intrinsic: Ratio of spontaneous to saturated power $\sim 10^{-5} 10^{-6}$
 - Stability likely dominated by crystals + mirrors angular tolerances ~10 nrad
- Output of both are affected by
 - Variations of beam energy $\sim 10^{-4}$
 - Energy/current uniformity of beam
 (*E*/*I* modulations create sidebands that can be amplified)
 - Heating issues of crystals
 - Etc.
- Long term drifts

 Compensated using feedback

Fast time scale fluctuations → Directly mapped onto seeded output

Reduced in XFELO by the effective quality factor of the cavity

"Favored" parameters/performance of fully coherent hard x-ray FELs

Self-seeding with wake monochromator $\mathcal{B}_{sat} \approx \rho \frac{\gamma m c^2}{\hbar \omega \mathcal{V}_{coh}} \frac{Q}{e}$

Constraints: $T_{\rm FWHM} \sim 5$ fs ($T_{\rm FWHM} >> T_{\rm coh}$ SASE) – 50 fs (to maintain overlap in radiator) $Q \sim 20 - 150$ pC (to produce necessary current)

 $\begin{array}{ll} \mbox{Output:} \ P_{\rm FWHM} \sim 20 - 1000? \ \mbox{GW} & \mbox{(depending on} \\ N_{\rm photons} \sim 10^{10} - 10^{14}? & \mbox{length of taper)} \end{array}$

 $\Delta\omega/\omega\sim 10^{-4}-10^{-6}$

X-ray FEL oscillator $\mathcal{B}_{\text{sat}} \approx \frac{T}{1-R} \frac{1}{2N_u} \frac{\gamma mc^2}{\hbar\omega \mathcal{V}_{\text{coh}}} \frac{Q}{e}$

Constraints: $T_{\rm FWHM} \sim 0.1 \text{ ps} (T > /\Delta E_{\rm refl}) - 2 + \text{ ps}$

 $Q \sim 1-50+$ pC (to produce necessary gain, depending on ε_x)

Output:
$$P_{\rm FWHM}$$
 ~ 1– 10 MW $N_{\rm photons}$ ~ 10⁸ – 10¹⁰ $\Delta \omega / \omega \sim$ 10⁻⁶ – 10⁻⁸ and below

Bragg crystals in reflection for XFELO

Bragg's law defines central energy of reflection $\lambda = \lambda_B \sin \theta$

Reflection $|R(E)|^2$





 Extinction length A gives depth into crystal over which x-rays near Bragg's law are reflected

 $\Delta E_{\rm refl} \sim \frac{\hbar c}{\Lambda}$

• XFELO radiation has $\hbar/T_{\rm rad} \lesssim \Delta E_{\rm refl}$ to minimize reflection losses

 \rightarrow Narrow region of *R*, *T* contributes

 Decreasing thickness to few Λ allows for transmission to users

Bragg crystals in transmission for "wake" self-seeding







• Radiation has $\hbar/T_{
m rad}\gg\Delta E_{
m refl}$

 \rightarrow Large Region of *T* contributes

Seeding "wake" is determined by the difference T(E) − T(∞)

R.R. Lindberg and Yu. Shvyd'ko, submitted to Phys. Rev. ST-AB., eprint arXiv:1202.1472 (2011)

Bragg crystals in transmission for "wake" self-seeding

Bragg's law defines central energy of reflection $\lambda = \lambda_B \sin \theta$





• Radiation has $\hbar/T_{
m rad}\gg\Delta E_{
m refl}$

 \rightarrow Large Region of *T* contributes

- Seeding "wake" is determined by the difference $T(E) T(\infty)$
- Relevant spectral components interact over entire crystal thickness

$$\Delta E_{\rm trans} \sim \frac{d}{\Lambda \sin \theta} \Delta E_{\rm refl}$$

$$\Delta t_{\rm trans} \sim \frac{\Lambda^2 \sin \theta}{cd}$$

R.R. Lindberg and Yu. Shvyd'ko, submitted to Phys. Rev. ST-AB., eprint arXiv:1202.1472 (2011)

Time domain picture of wake monochromator



Position of first peak proportional to interaction area: $\Delta t d = \text{constant}$ Δt scales inversely with crystal thickness

Seed power scales quadratically with crystal thickness

Diamond crystals have superlative material properties



+ Technological Institute for Superhard Novel Carbon Materials, Russia



Yu. Shvyd'ko, S. Stoupin, V. Blank, and S. Terentyev, Nature Photonics 5, 539 (2011)

+ Technological Institute for Superhard Novel Carbon Materials, Russia



There exist regions whose area > 1 mm^2 that are free of dislocations and stacking faults.

Suitable for FEL applications

Yu. Shvyd'ko, S. Stoupin, V. Blank, and S. Terentyev, Nature Photonics 5, 539 (2011)

+ Technological Institute for Superhard Novel Carbon Materials, Russia



Reflectivity curve width and maximum match that predicted by theory

Yu. Shvyd'ko, S. Stoupin, V. Blank, and S. Terentyev, Nature Photonics 5, 539 (2011)

Crystal heating

Crystal temperature rise $\delta T \rightarrow Crystal expansion \delta a = \beta a \delta T$

→ Variation in atomic spacing → Variation in Bragg's law $\delta E/E = \beta \delta T$

How does the variation δE compare to the bandwidth ~ meV?



Crystal Damage

- There are significant unknowns regarding exact damage limits of x-ray crystals
- Present thinking is that time averaged photon intensity is main concern
 - Photoabsorption can produce high energy ``free'' electron and secondary Auger electron which can lead to lattice deformations/damage
- Graphitization of diamond crystals seen at synchrotrons



Graphitization seen after a few days, but no significant performance degradation after one year APS undulator power flux ~0.15 kW/mm²

XFELO with 50 pC @ 1 MHz provides ~4 kW/mm² on crystal (scales with charge)

Self-seeding typically has ~0.5 kW/mm² × f[MHz] × $T_{\rm FWHM}$ [fs]

There is evidence that diamond exposed to 3.5 kW/mm² survives¹

1. J. Als-Niesen, A.K Freund, and S. Terentyev, Nucl. Instrum. Methods Res. Sec. B 94, 348 (1994)

Conclusions

- Fully coherent hard x-ray FELs are possible
- Self-seeding with the wake monochromator can dramatically improve SASE
 - Route to Fourier limited short pulses (< 50 fs or so)
 - Increase power with efficient nonlinear undulator tapering
 - Early experimental results are quite promising
- X-ray FEL oscillator offers complementary approach to small linewidths
 - Revolutionize techniques developed at 3rd generation light sources
 - Scales quite favorably to low charge
 - Natural fit in ERL based light source
- Diamond Bragg crystals are the future of high-intensity hard x-ray optics
 - Low absorption, high reflectivity, narrow bandwidth, low thermal expansion, high thermal conductivity, high damage threshold, ...
 - Near perfect crystals are now routinely grown at a few companies/institutes

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- Sasha Zholents
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- Harald Sinn (Euro XFEL)

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W. Berg	A. Lutman
FJ. Decker	HD. Nuhn
Y. Ding	D. Ratner
P. Emma	J. Rzepiela
Y. Feng	S. Spampinati
J. Frisch	E. Trakhtenberg
D. Fritz	D. Walz
J. Hastings	J. Welch

J. Wu

- Z. Huang
- J. Krzywinski

Extra slides



Example uses LCLS-type parameters with ρ = 5×10⁻⁴, $\sigma_e \approx$ 12 fs and ($\Delta \omega / \omega$)_{mono} = 2×10⁻⁵

Parameters

Cavity length $\sim 100~m~$ $\Longrightarrow~$ Repetition rate $\sim 1~MHz$

 $\label{eq:Low charge} \ensuremath{\mathsf{Low charge}} \sim 1 \ensuremath{\mathsf{pC}} \\ \ensuremath{\mathsf{Bunch Lengths}} \sim 0.25 \ensuremath{\,\mathsf{ps}} \\ \ensuremath{\mathsf{s}} \sim 0.25 \ensuremath{\,\mathsf{s}} \\ \ensuremath{\mathsf{s}} \sim 0.25 \ensuremath{\,\mathsf{ps}} \\ \ensuremath{\mathsf{s}} \sim 0.25 \ensuremath{\,\mathsf{s}} \\ \ensuremath{\mathsf{ps}} \simeq 0.25 \ensuremath{\,\mathsf{s}} \\ \ensuremath{\mathsf{s}} \simeq 0.25 \ensuremath{\,\mathsf{s}} \\ \ensuremath{\s} \simeq 0.25 \ensuremath{\,\mathsf{s}} \\ \ensuremath{\s}$

 \rightarrow Peak currents ~ 1.6 A

High quality electron beam: $\epsilon_{x,n} \sim 0.082 \text{ mm-mrad}, \Delta E \sim 250 \text{ keV}$ Low single pass gain, losses: (Nu = 2000) $G = 0.4, R_{tot} \sim 0.85$

Example @ 1 Å

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crystal	C(4 4 4)	

. . .

Performance

Output power $\sim 0.6 \text{ MW}$	Example @ 1 A	
	Pout	1.7 MW
Spectral brightness $\sim 10^{31}$ – 10^{34}	Photons/ pulse	1.1×10 ⁹
	ΔΕ	1.29 meV
$\Delta \omega \sim 10^{-6} - 10^{-8}$	Δt	0.51 ps
	Output power $\sim 0.6 \text{ MW}$ Spectral brightness $\sim 10^{31} - 10^{34}$ $\Delta E \sim 5.0 \text{ meV FWHM}$ $\Delta \omega / \omega \sim 10^{-6} - 10^{-8}$	Output powerExam $\sim 0.6 \text{ MW}$ P_{out} Spectral brightnessPhotons/ $\sim 10^{31} - 10^{34}$ Photons/ $\Delta E \sim 5.0 \text{ meV FWHM}$ ΔE $\Delta \omega/\omega \sim 10^{-6} - 10^{-8}$ Δt

9. R.R. Lindberg, K-J. Kim, Yu. Shvyd'ko, and W.M. Fawley, Phys. Rev. ST-AB. 14, 010701 (2011)

Cavity tolerances and stability

To preserve radiation-electron beam overlap and FEL gain, we require:

Cavity length stability δL < 3 μm rms Crystal angular stability $\delta \theta$ < 10 nrad rms



Null detection feedback (similar to that used at LIGO)

Variation in output signal is proportional to deviation from maximum Feedback correction signal is extracted using a lock-in amplifier

Stoupin, Lenkszus, Laird, Goetze, Kim, and Shvyd'ko, Rev. Sci. Instrum. 81, 055108 (2010)

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Null detection feedback: proof of principle experiment @ APS



Crystal stability of \sim 15 nrad rms was shown at the APS HERIX monochromator

Stoupin, Lenkszus, Laird, Goetze, Kim, and Shvyd'ko, Rev. Sci. Instrum. 81, 055108 (2010)

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Crystal Heating

Crystal temperature rise $\delta T \rightarrow Crystal expansion \delta a = \beta a \delta T$

→ Variation in atomic spacing → Variation in Bragg's law $\delta E/E = \beta \delta T$

Two heatings:



Hard x-ray optical cavity



Total external reflection of grazing incidence mirrors require $\phi \ll 1$

 \rightarrow Requires $\Theta << 1$ which leads to a limited tuning range