

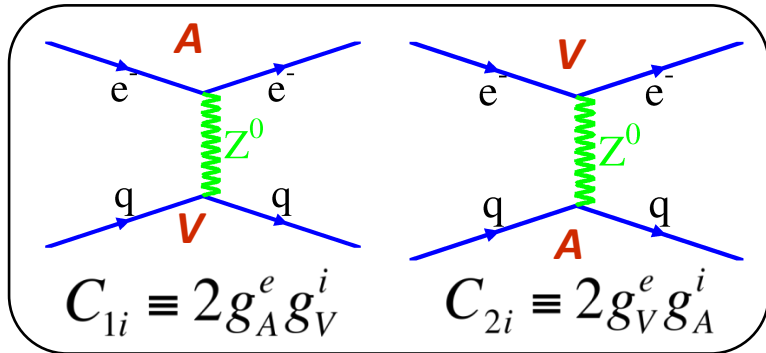
# PVDIS with SoLID

and radiative corrections

# PVDIS Theory

$$A_{\text{iso}} = \frac{\sigma^l - \sigma^r}{\sigma^l + \sigma^r}$$

$$= - \left( \frac{3G_F Q^2}{\pi\alpha 2\sqrt{2}} \right) \frac{2C_{1u} - C_{1d} (1 + R_s) + Y (2C_{2u} - C_{2d}) R_v}{5 + R_s}$$



$$Y = \frac{1 - (1 - y)^2}{1 + (1 - y)^2 - y^2 \frac{R}{R+1}}$$

$$R(x, Q^2) = \sigma^l / \sigma^r \approx 0.2$$

$$C_{1u} = -\frac{1}{2} + \frac{4}{3} \sin^2 \theta_W \approx -0.19$$

$$C_{1d} = \frac{1}{2} - \frac{2}{3} \sin^2 \theta_W \approx 0.35$$

$$C_{2u} = -\frac{1}{2} + 2 \sin^2 \theta_W \approx -0.04$$

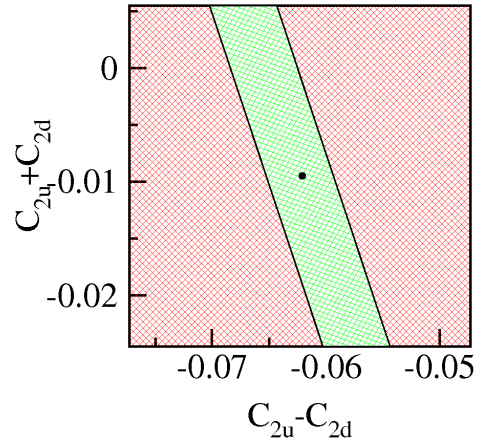
$$C_{2d} = \frac{1}{2} - 2 \sin^2 \theta_W \approx 0.04$$

Cahn and Gilman, PRD 17  
1313 (1978) polarized  
electrons on deuterium

# Motivation for PVDIS

## Standard Model

$$b(x) = \frac{\sum_i C_{2i} Q_i f_i^-(x)}{\sum_i Q_i^2 f_i^+(x)}$$



$$A_{PV} = \frac{G_F Q^2}{\sqrt{2}\pi\alpha} [a(x) + Y(y) b(x)]$$

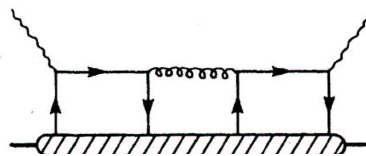
## CSV at Quark Level

$$\delta u(x) = u^p(x) - d^p(x)$$

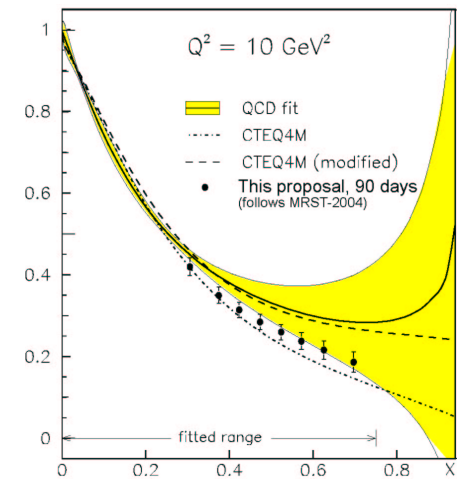
$$\delta d(x) = d^p(x) - u^p(x)$$

$$R_{CSV} = \frac{\delta A_{PV}(x)}{A_{PV}(x)} = 0.28 \frac{\delta u(x) - \delta d(x)}{u(x) + d(x)}$$

## Di-quarks in the nucleon (Q<sup>2</sup> Dependence)

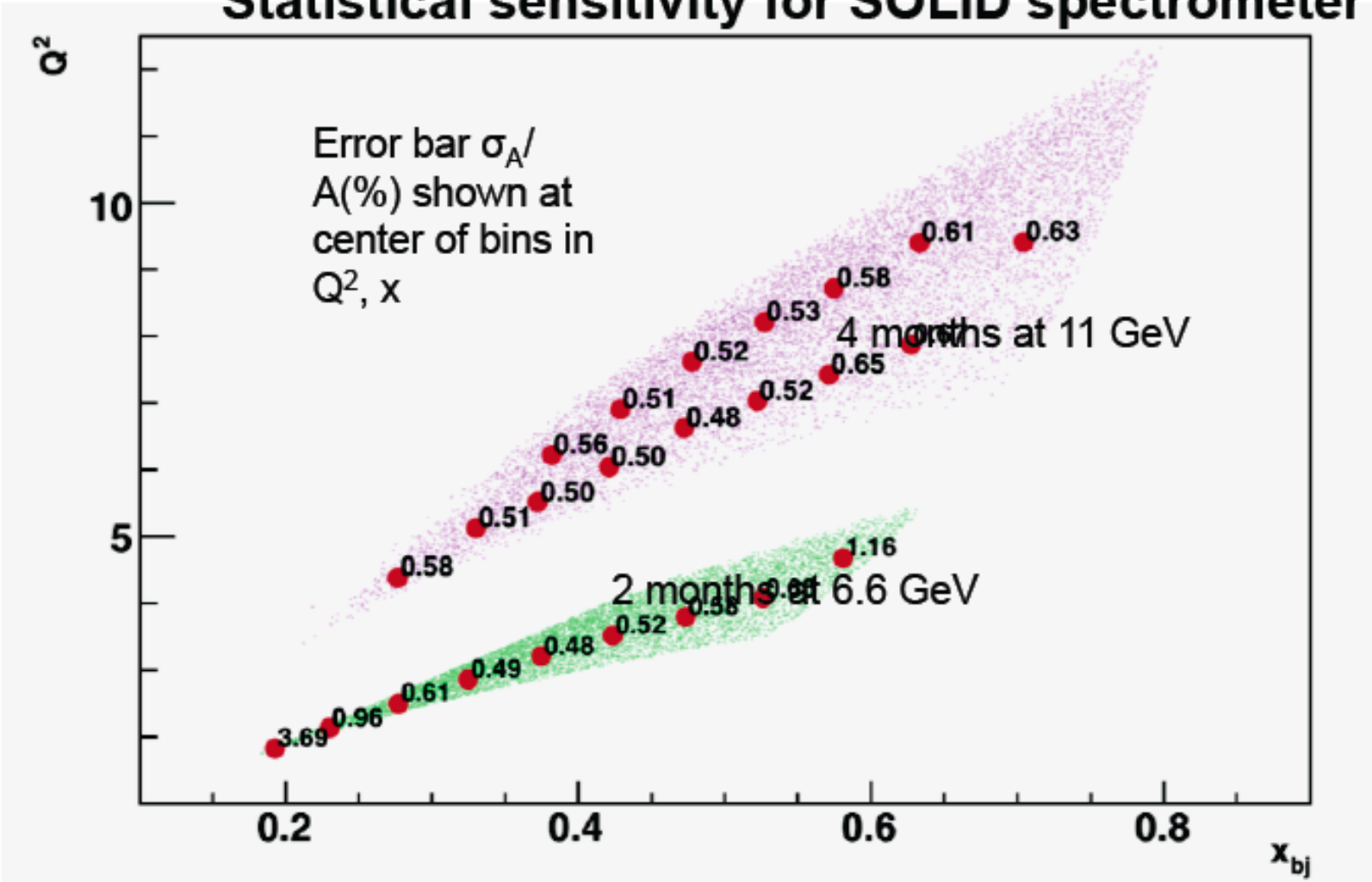


## d/u for Hydrogen

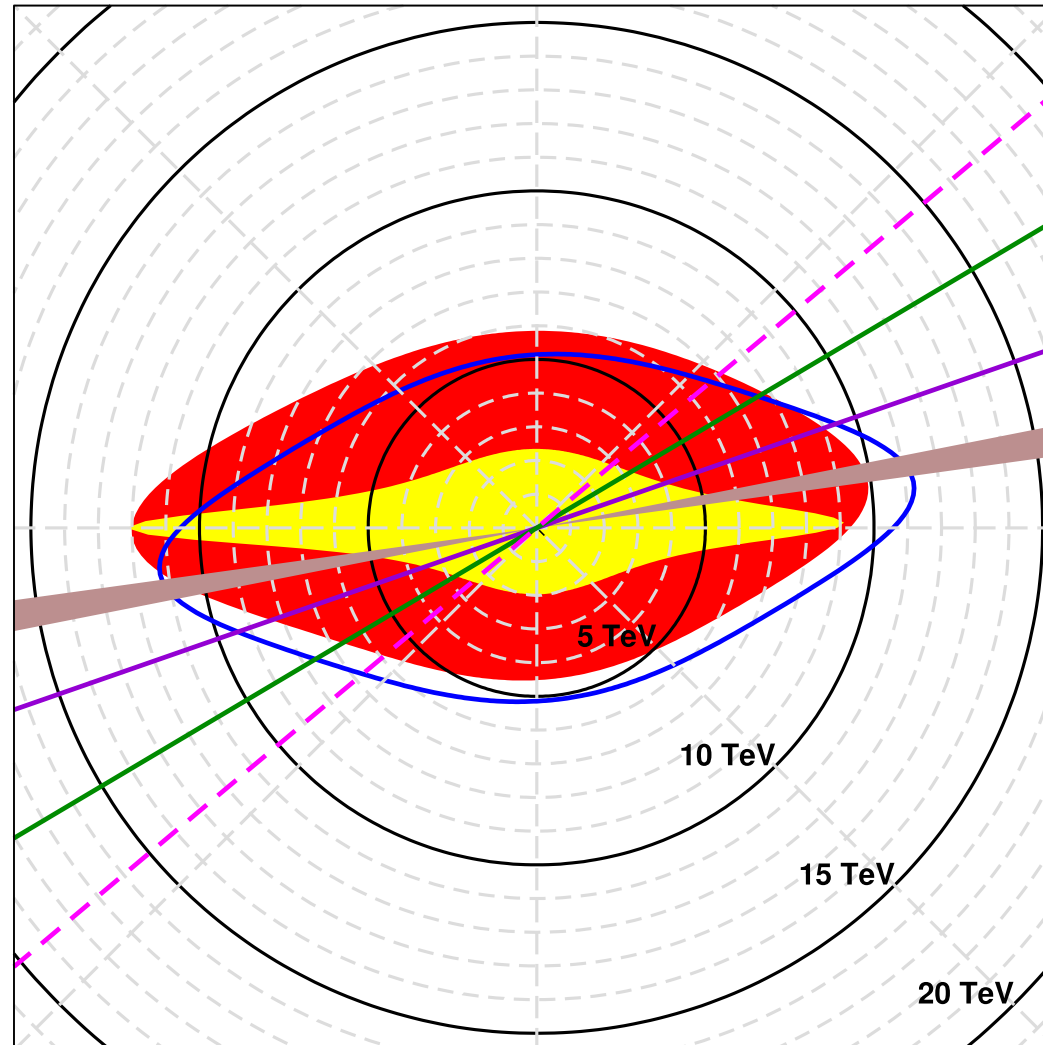


# Statistical Errors (%) vs. Kinematics

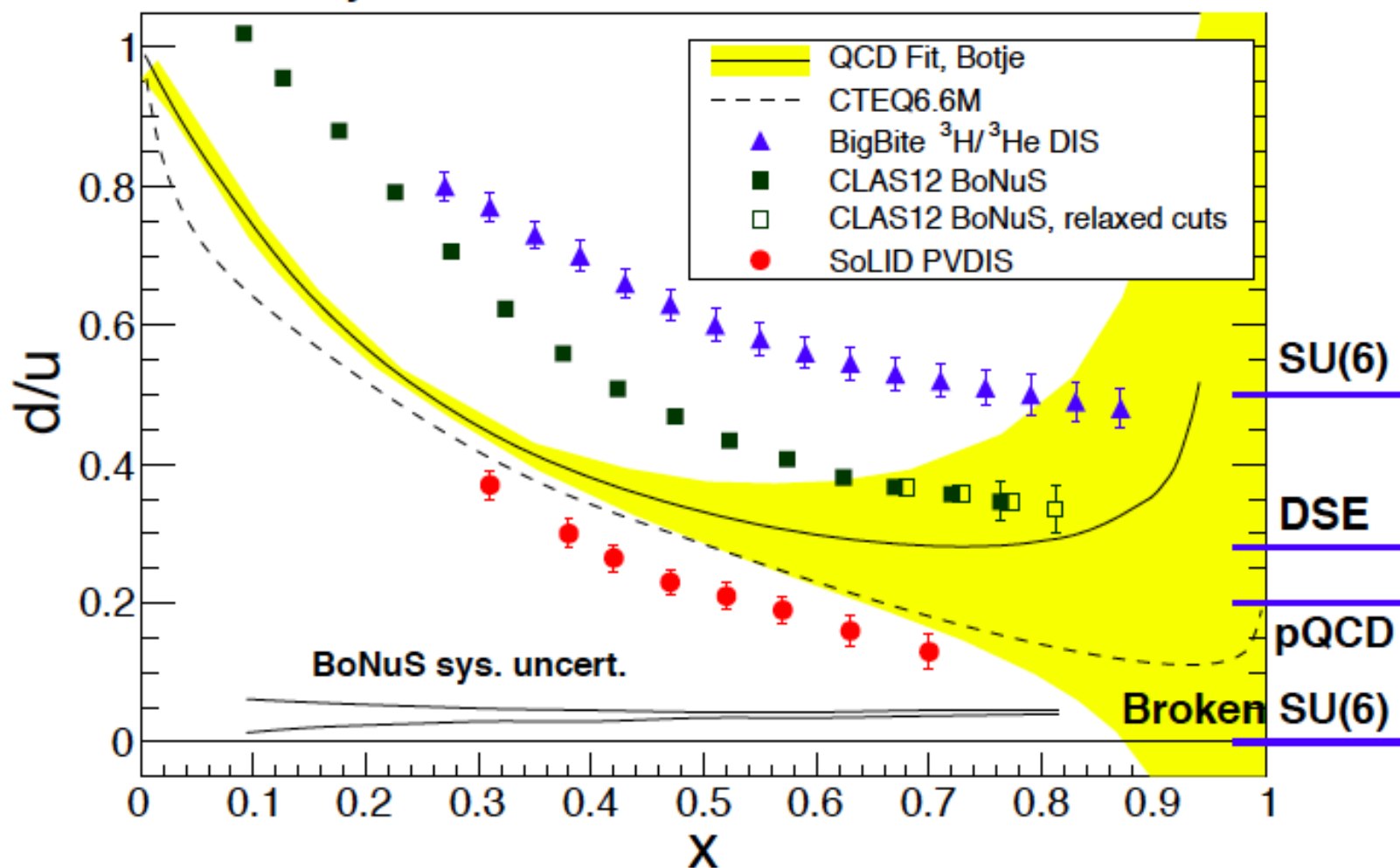
## Statistical sensitivity for SOLID spectrometer



# 6 GeV PVDIS Result: Contact Interactions

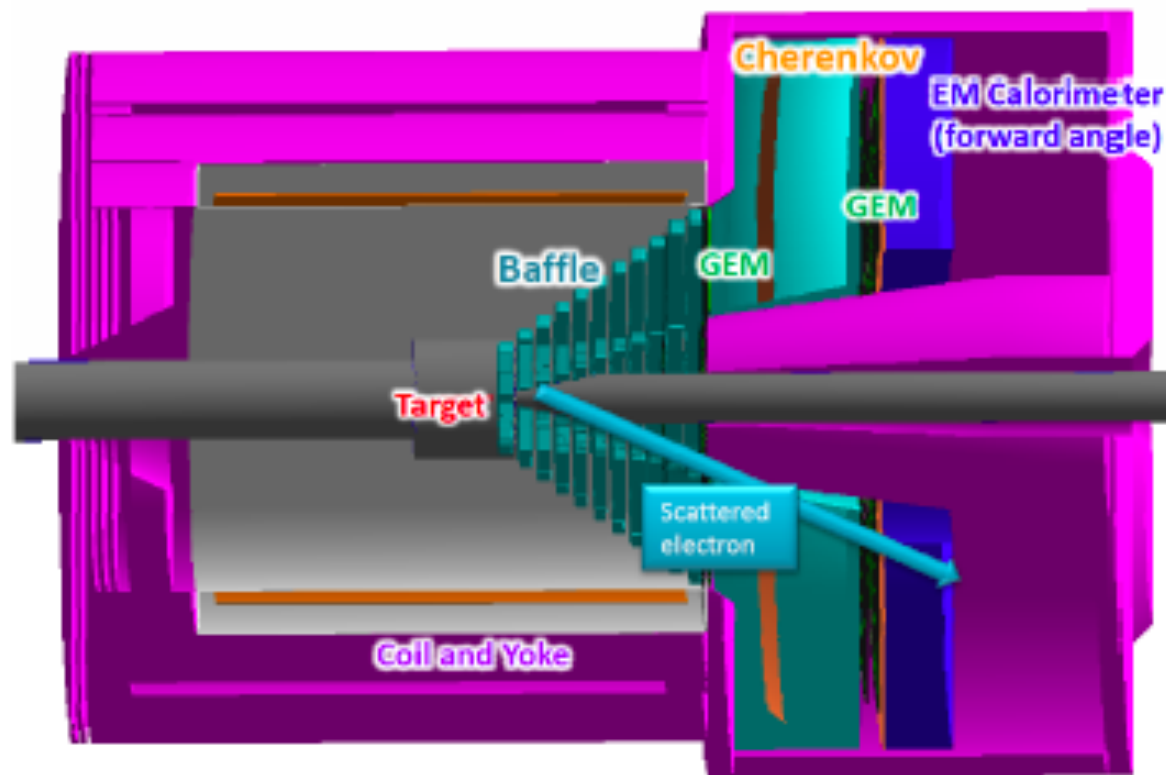


# Projected 12 GeV d/u Extractions

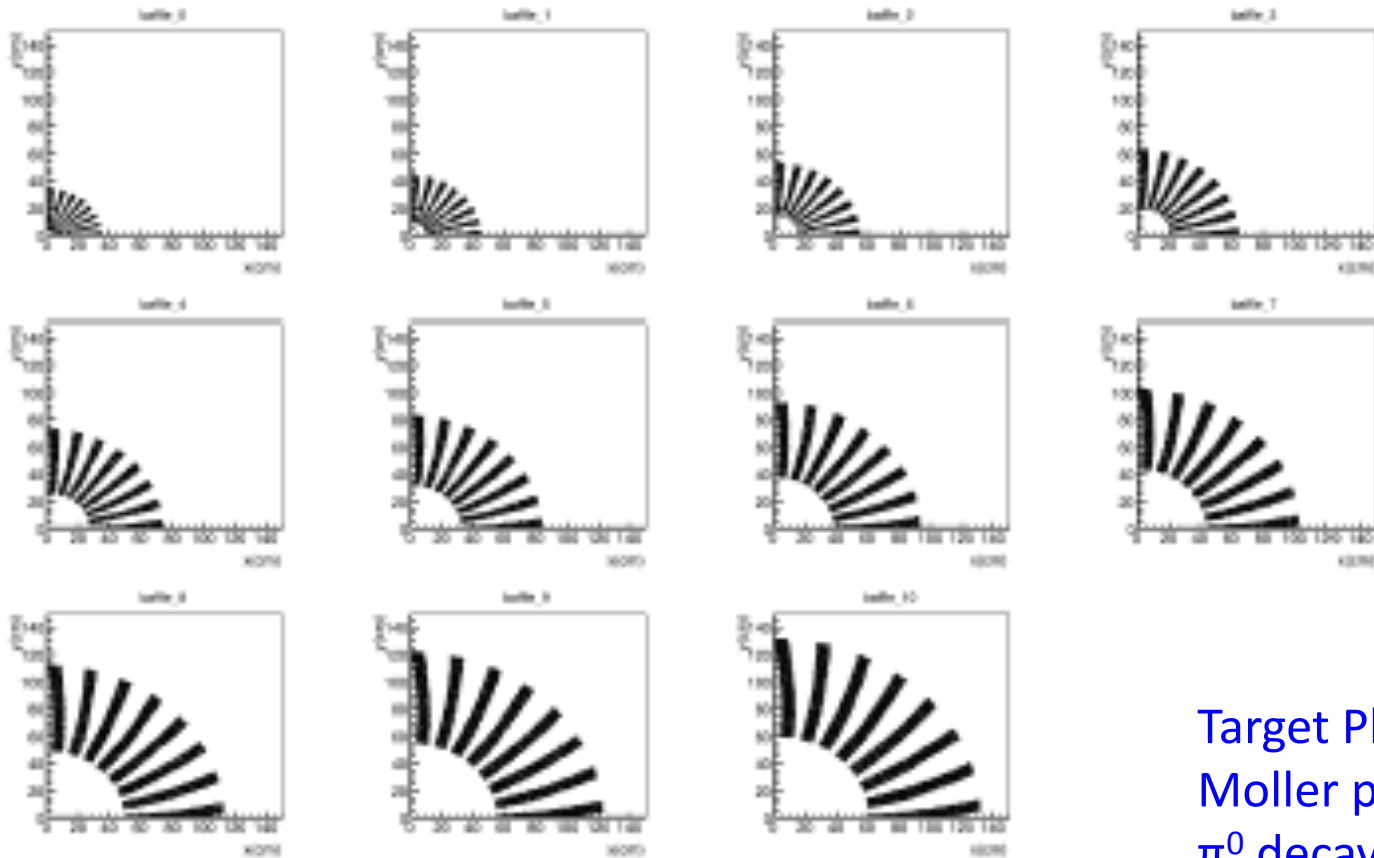


# SoLID PVDIS Apparatus

SoLID CLEO PVDIS



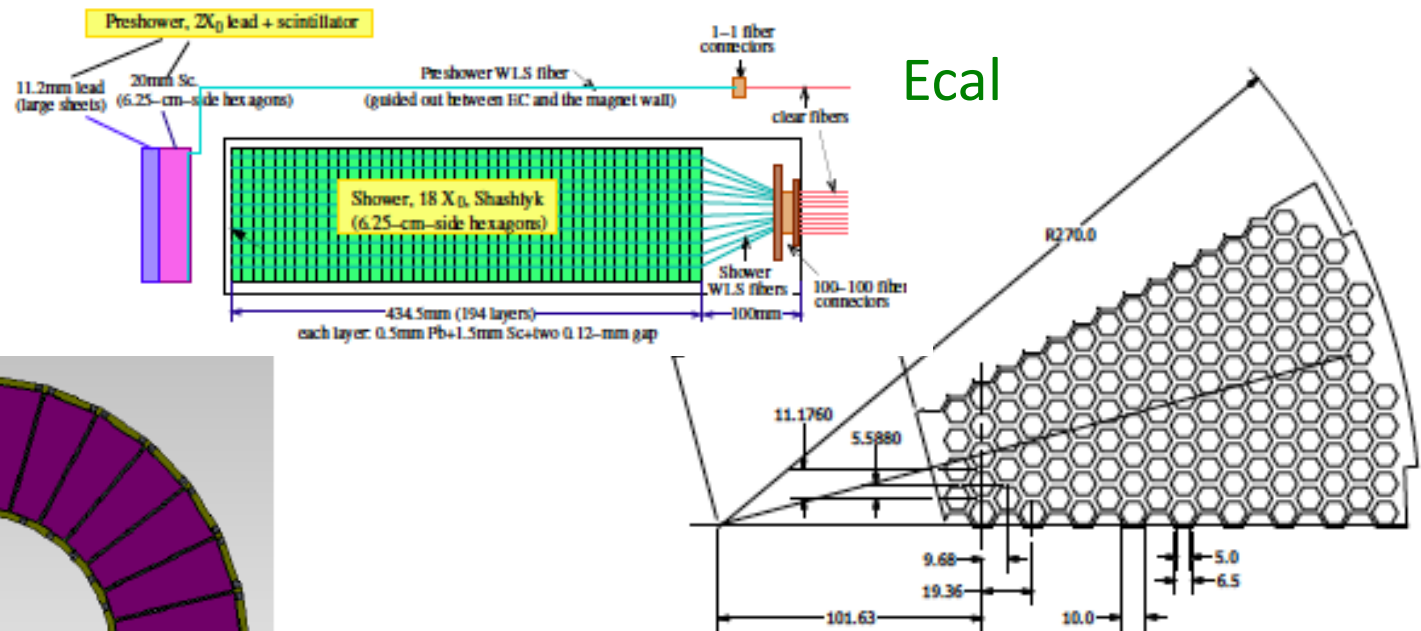
# Baffles Reduce Backgrounds



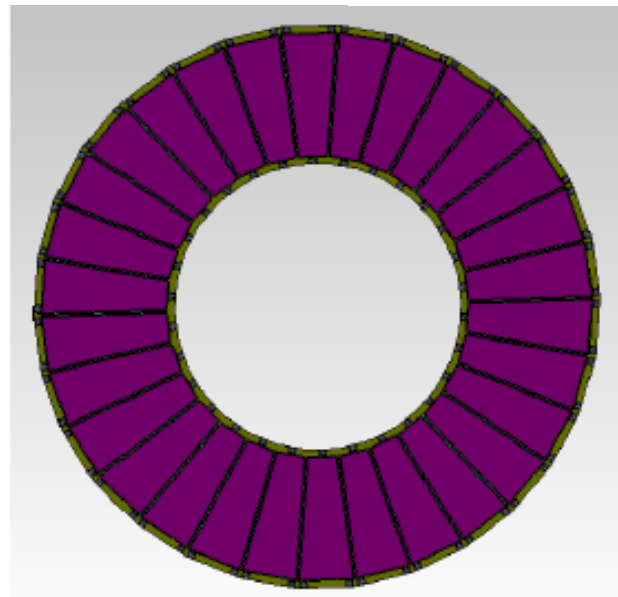
Target Photons  
Moller photons  
 $\pi^0$  decay photons  
Charged hadrons



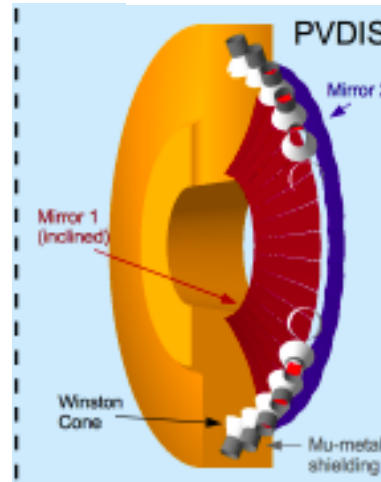
# Detectors



Ecal



Gem Tracking Detector



Cerenkov

Ecal Geometry

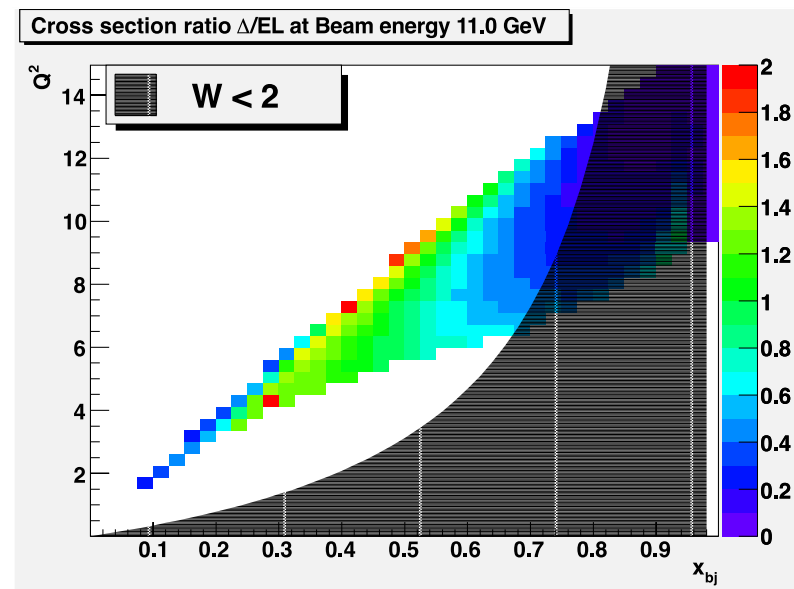
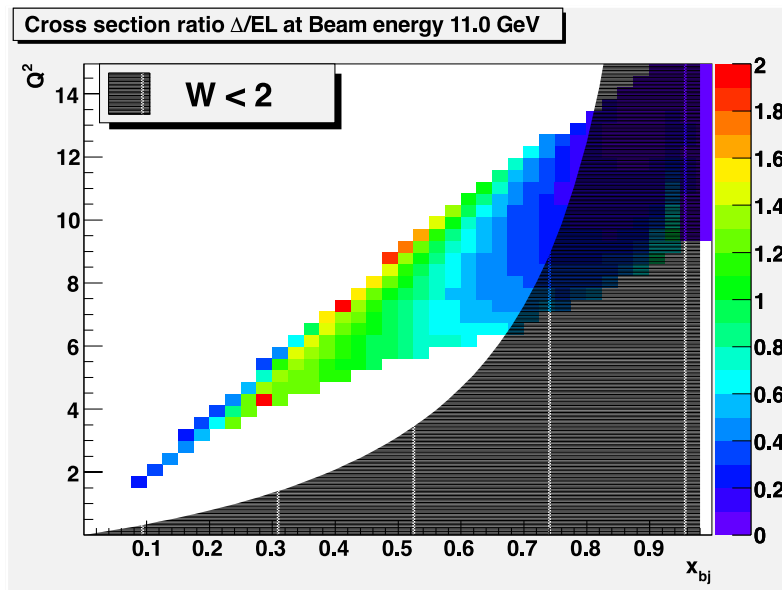


# Issues (Quote JFK)

- What you can do for us:
  - Help with theory of radiative corrections
  - Tell what data we need for radiative corrections
- What we can do for you:
  - Obtain elastic data over a wide kinematic range

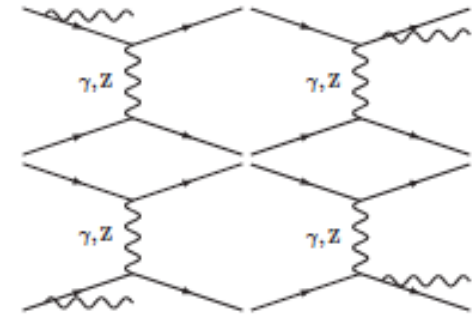
# Radiative DIS Contamination (Peaking Approximation)

Most of our events are DIS,  
not elastic or resonance events



# Divergent Radiative Corrections

Zygunov , Suarez, Kolomensky, Tweedy (Moller)



$$\sigma^R = -\frac{\alpha^3}{4s\pi} \int_0^{v^{\max}} dv \int \frac{d^3k}{k_0} \delta[(k_1 + p_1 - k_2 - k)^2 - m^2] \sum_{j,t=1,4} M_{ij}^R (-1)^{t+j} ,$$

$$\begin{aligned} (M_{ij}^R)_{zz} &= \text{Tr}[G_1^{\mu\alpha} \rho^{ij}(k_1) G_1^{\nu\alpha T} \Lambda(k_2)] \text{Tr}[\gamma_\mu \rho^{ij}(p_1) \gamma_\nu \Lambda(p_2)] D^{it_1} D^{jt_1} , \\ (M_{ij}^R)_{zv} &= \text{Tr}[G_1^{\mu\alpha} \rho^{ij}(k_1) \gamma_\nu \Lambda(k_2)] \text{Tr}[\gamma_\mu \rho^{ij}(p_1) G_2^{\nu\alpha T} \Lambda(p_2)] D^{it_1} D^{jt} , \\ (M_{ij}^R)_{vz} &= \text{Tr}[G_2^{\mu\alpha} \rho^{ij}(p_1) \gamma_\nu \Lambda(p_2)] \text{Tr}[\gamma_\mu \rho^{ij}(k_1) G_1^{\nu\alpha T} \Lambda(k_2)] D^{it} D^{jt_1} , \\ (M_{ij}^R)_{vv} &= \text{Tr}[G_2^{\mu\alpha} \rho^{ij}(p_1) G_2^{\nu\alpha T} \Lambda(p_2)] \text{Tr}[\gamma_\mu \rho^{ij}(k_1) \gamma_\nu \Lambda(k_2)] D^{it} D^{jt} , \\ (M_{ij}^R)_f &= \text{Tr}[G_1^{\mu\alpha} \rho^{ij}(k_1) G_3^{\nu\alpha} \Lambda(p_2) \gamma_\mu \rho^{ij}(p_1) \gamma_\nu \Lambda(k_2)] D^{it_1} D^{jt_u} , \\ (M_{ij}^R)_l &= \text{Tr}[G_1^{\mu\alpha} \rho^{ij}(k_1) \gamma_\nu \Lambda(p_2) \gamma_\mu \rho^{ij}(p_1) G_4^{\nu\alpha} \Lambda(k_2)] D^{it_1} D^{jt_{z_2}} , \\ (M_{ij}^R)_{lu} &= \text{Tr}[\gamma_\mu \rho^{ij}(k_1) G_3^{\nu\alpha} \Lambda(p_2) G_2^{\mu\alpha} \rho^{ij}(p_1) \gamma_\nu \Lambda(k_2)] D^{it} D^{jt_u} , \\ (M_{ij}^R)_s &= \text{Tr}[\gamma_\mu \rho^{ij}(k_1) \gamma_\nu \Lambda(p_2) G_2^{\mu\alpha} \rho^{ij}(p_1) G_4^{\nu\alpha} \Lambda(k_2)] D^{it} D^{jt_{z_2}} , \end{aligned}$$

For PVDIS,  
QCD effects  
are also important.

Starting point:  
HERA code

Compare to Moller: PVDIS 5x less sensitive, but quarks interact

# Accessible Non-DIS Data

Stat Precision [%] - SoLID D<sub>2</sub>, 11 GeV 60 days, 6.6 GeV 30 days

