

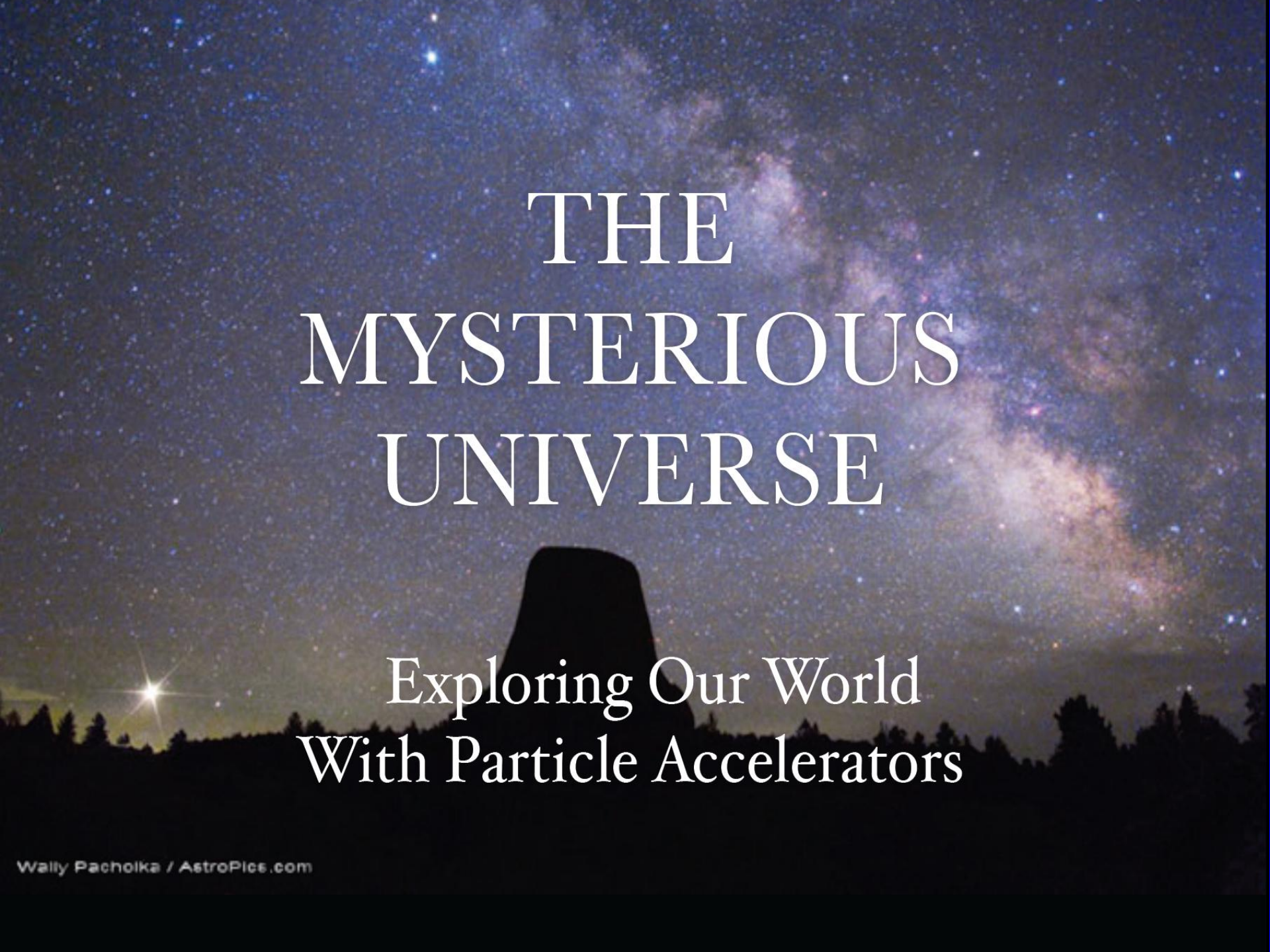


International Linear Collider (ILC)
Superconducting Radio Frequency (SRF) Acceleration
and
ILC High Gradient SRF Cavity R&D at JLab

Rong-Li Geng
Jefferson Lab

JLab Graduate Student Lunch Seminar, September 21, 2011



The background of the slide is a night sky featuring the Milky Way galaxy. The galaxy's band of stars and dust is visible, curving from the upper right towards the center. The sky is filled with numerous stars of varying brightness. In the lower foreground, the dark silhouette of a rock formation, likely a butte or mesa, is visible against the horizon. The overall scene is a deep blue and black, with the light from the stars and the Milky Way providing the primary illumination.

THE MYSTERIOUS UNIVERSE

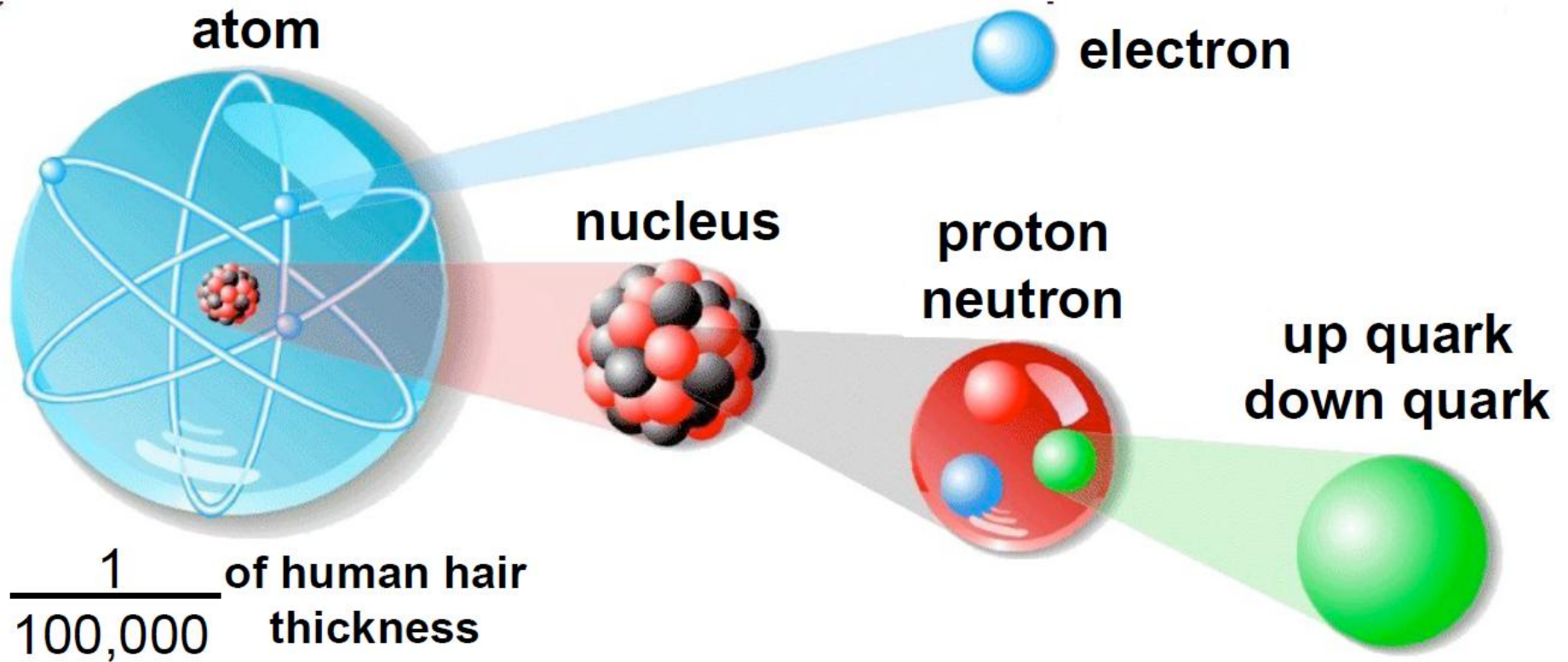
Exploring Our World
With Particle Accelerators

~90 years ago

~60 years ago

~40 years ago

Present



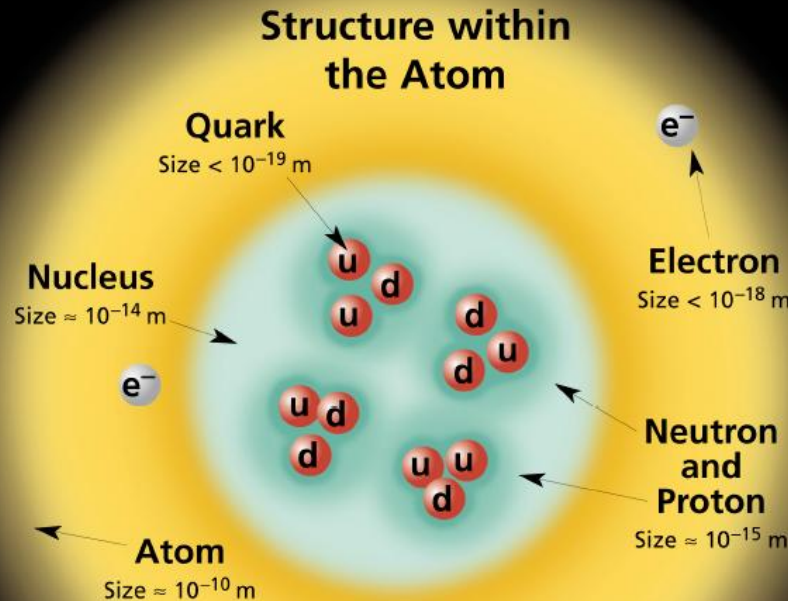
$\frac{1}{10,000}$

$\frac{1}{10}$

$\frac{1}{100,000}$

The Nature of Matter

Could there be more quarks?
Or something smaller?

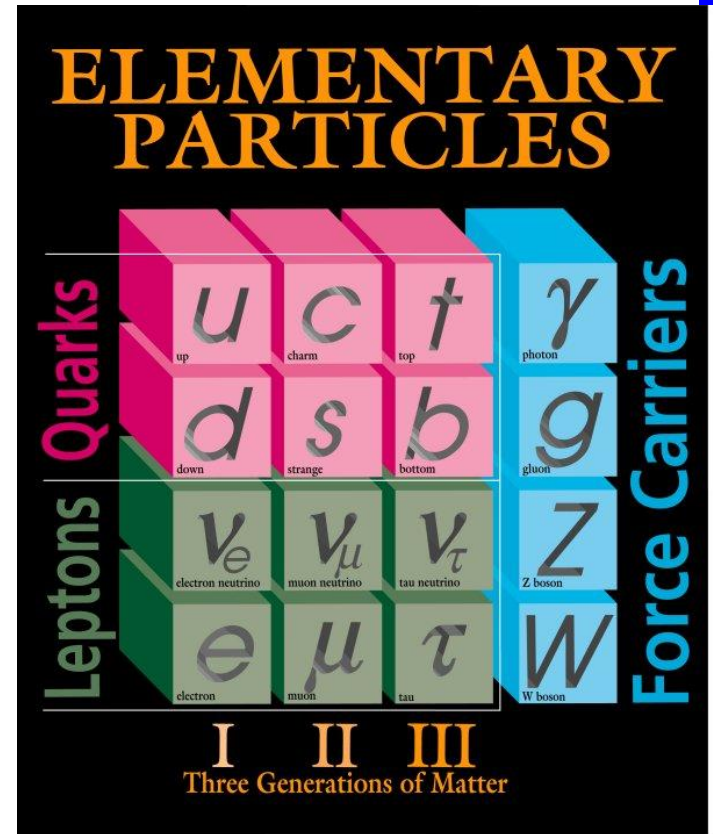


If the protons and neutrons in this picture were 10 cm across, then the quarks and electrons would be less than 0.1 mm in size and the entire atom would be about 10 km across.

**Atoms as we know
them today**

The fundamental questions

- What is the nature of the universe and what is it made of?
- What are matter, energy, space and time?
- How did we get here and where are we going?



Today's biggest question

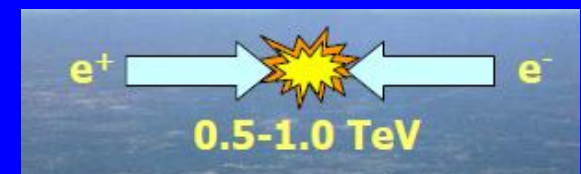
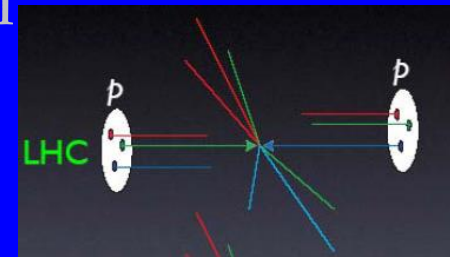
What's beyond the Standard Model?

1. Are there undiscovered principles of nature:
New symmetries, new physical laws?
2. How can we solve the mystery of dark energy?
3. Are there extra dimensions of space?
4. Do all the forces become one?
5. Why are there so many kinds of particles?
6. What is dark matter?
How can we make it in the laboratory?
7. What are neutrinos telling us?
8. How did the universe come to be?
9. What happened to the antimatter?

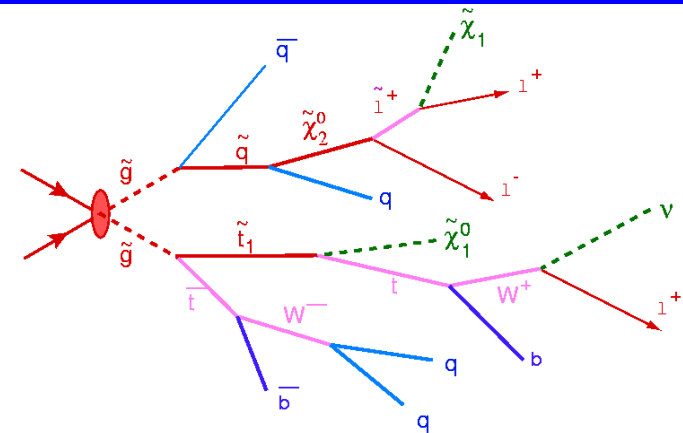
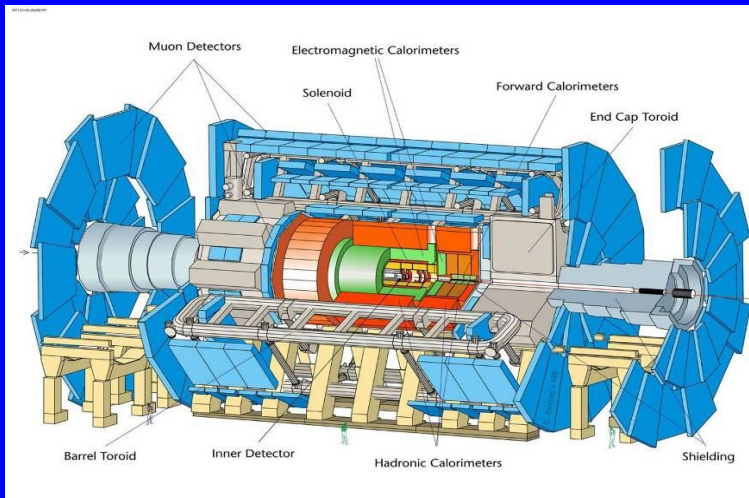
from the Quantum Universe

Addressing the Questions

- **Neutrinos**
 - Particle physics and astrophysics using a weakly interacting probe
- **Particle Astrophysics/Cosmology**
 - Dark Matter; Cosmic Microwave, etc
- **High Energy pp Colliders**
 - Opening up a new energy frontier (~ 1 TeV scale)
- **High Energy e^+e^- Colliders**
 - Precision Physics at the new energy frontier



Megascience project --- LHC



3 isolated leptons
 + 2 b-jets
 + 4 jets
 + E_t^{miss}

Exploring the Terascale

the tools

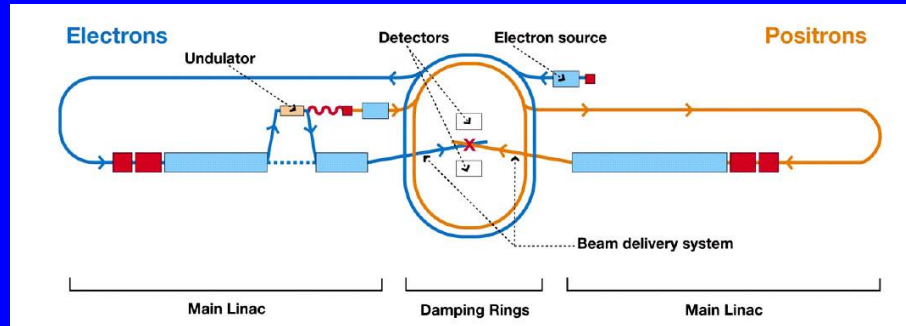
- **The LHC**
 - It will lead the way and has large reach
 - Quark-quark, quark-gluon and gluon-gluon collisions at 0.5 - 5 TeV
 - Broadband initial state
- **The ILC**
 - A second view with high precision
 - Electron-positron collisions with fixed energies, adjustable between 0.1 and 1.0 TeV
 - Well defined initial state
- **Together, these are our tools for the terascale**

Why a TeV Scale e^+e^- Accelerator?

- Two parallel developments over the past few years (**the science & the technology**)
 - The precision information from LEP and other data have pointed to a low mass Higgs; Understanding electroweak symmetry breaking, whether supersymmetry or an alternative, will require precision measurements.
 - There are strong arguments for the complementarity between a $\sim 0.5\text{-}1.0$ TeV ILC and the LHC science.

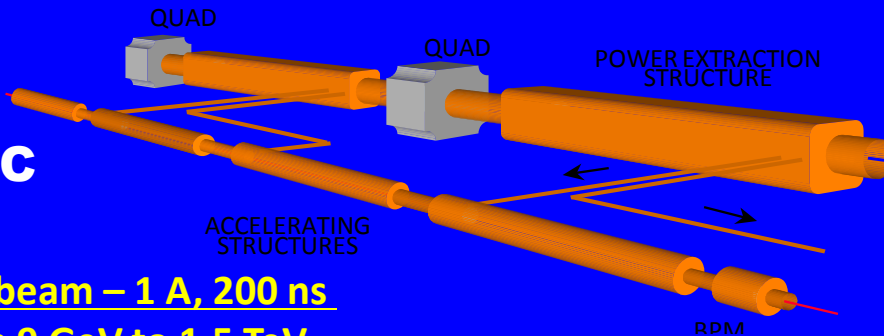
Possible TeV Scale Lepton Colliders

ILC



ILC < 1 TeV
Technically possible
~ 2019

CLIC

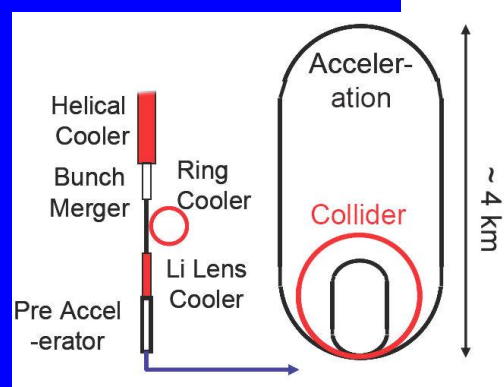


Drive beam - 95 A, 300 ns
from 2.4 GeV to 240 MeV

Main beam – 1 A, 200 ns
from 9 GeV to 1.5 TeV

CLIC < 3 TeV
Feasibility?
ILC + 5-10 yrs

Muon Collider



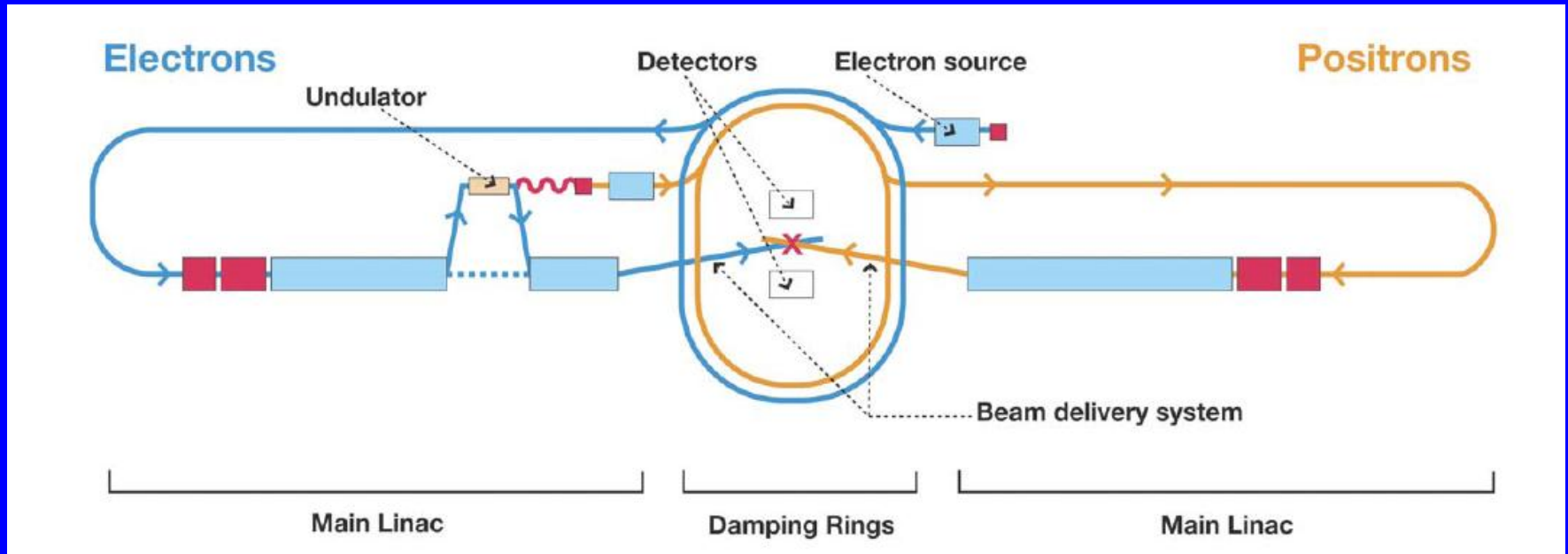
Muon Collider
< 4 TeV
FEASIBILITY??
ILC + 15 yrs?

Much R&D Needed

- Neutrino Factory R&D +
- bunch merging
- much more cooling
- etc

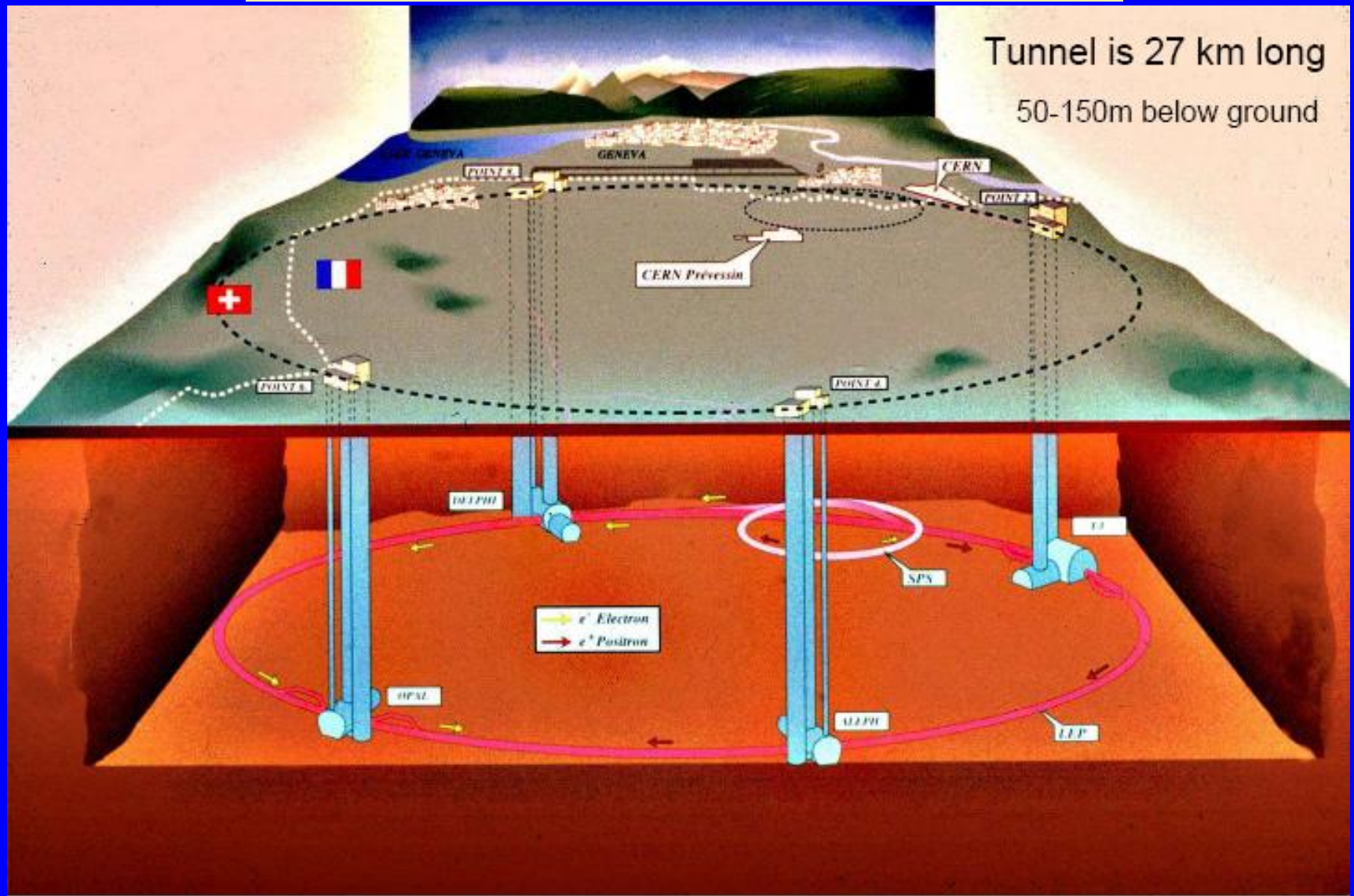
Rongli Geng

The ILC



- Two linear accelerators, with tiny intense beams of electrons and positrons colliding head-on-head
- Total length ~ 30 km long (comparable scale to LHC)
- COM energy = 500 GeV, upgradeable to 1 TeV

LHC --- Deep Underground

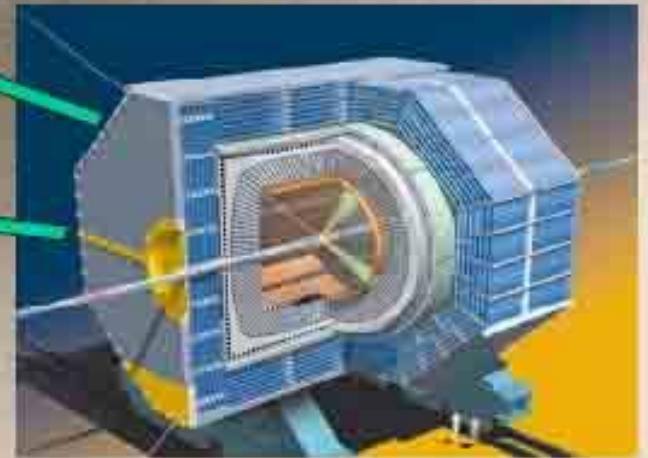


ILC --- Deep Underground

Main Research Center

Particle Detector

~30 km long tunnel



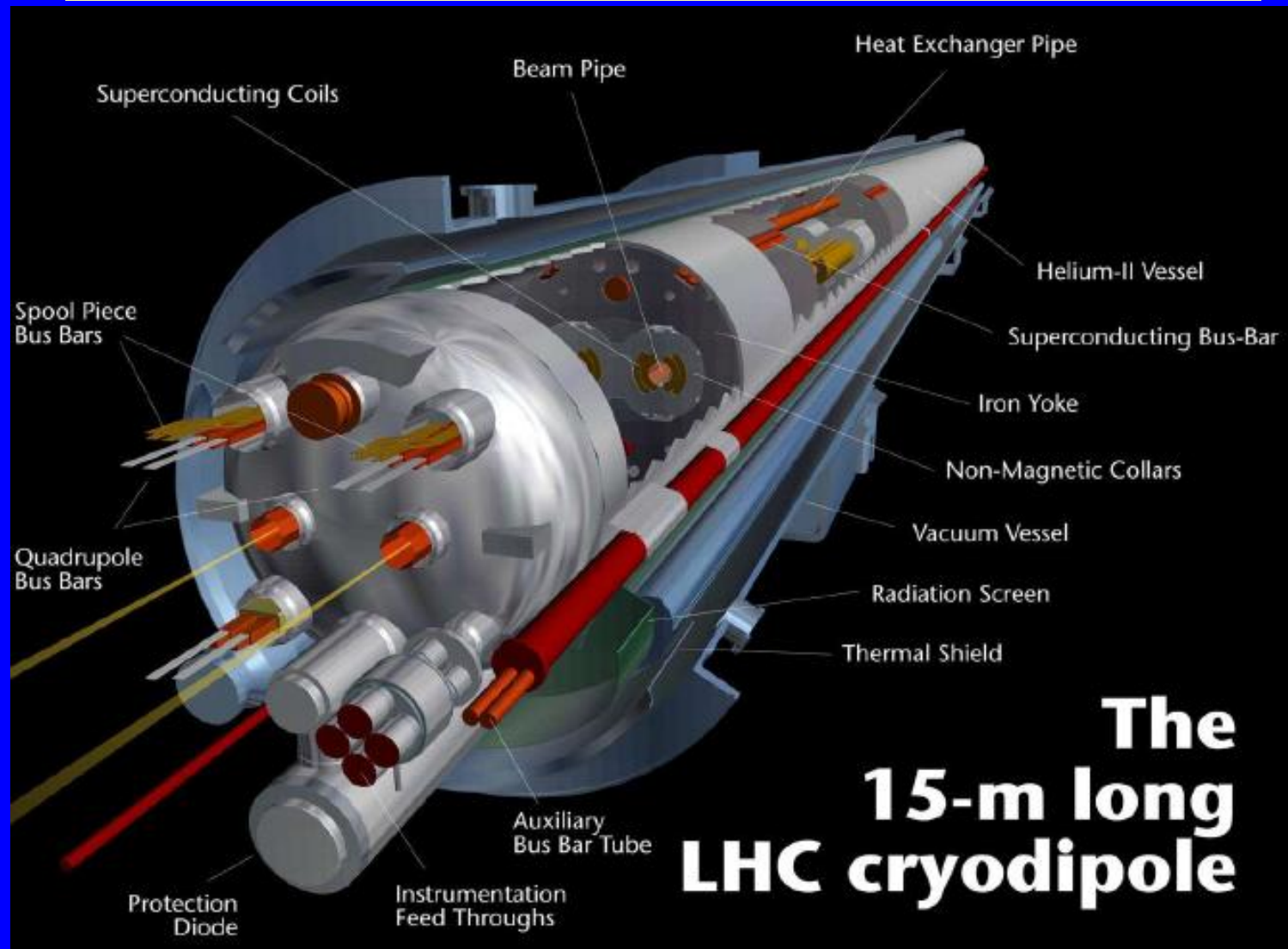
Two tunnels

- accelerator units
- other for services - RF power

Comparison: ILC and LHC

	ILC	LHC
Beam Particle :	Electron x Positron	Proton x Proton
CMS Energy :	0.5 – 1 TeV	14 TeV
Luminosity Goal :	$2 \times 10^{34} \text{ /cm}^2\text{/sec}$	$1 \times 10^{34} \text{ /cm}^2\text{/sec}$
Accelerator Type :	Linear	Circular Storage Rings
Technology :	Supercond. RF	Supercond. Magnet

LHC --- Superconducting Magnet



ILC - Superconducting RF Cryomodule



Linear Collider Conceptual Scheme



ILC Subsystems

- **Electron source**

To produce electrons, light from a titanium-sapphire laser hit a target and knock out electrons. The laser emits 2-ns "flashes," each creating billions of electrons. An electric field "sucks" each bunch of particles into a 250-meter-long linear accelerator that speeds up the particles to 5 GeV.

- **Positron source**

To produce positron, electron beam go through an undulator. Then, photons, produced in an undulator, hit a titanium alloy target to generate positrons. A 5-GeV accelerator shoots the positrons to the first of two positron damping rings.

- **Damping Ring for electron beam**

In the 6-kilometer-long damping ring, the electron bunches traverse a wiggler leading to a more uniform, compact spatial distribution of particles. Each bunch spends roughly 0.2 sec in the ring, making about 10,000 turns before being kicked out. Exiting the damping ring, the bunches are about 6 mm long and thinner than a human hair.

- **Damping Ring for positron beam**

To minimize the "electron cloud effects," positron bunches are injected alternately into either one of two identical positron damping rings with 6-kilometer circumference.

- **Main Linac**

Two main linear accelerators, one for electrons and one for positrons, accelerate bunches of particles up to 250 GeV with 8000 superconducting cavities nestled within cryomodules. The modules use liquid helium to cool the cavities to -2°K . Two 12-km-long tunnel segments, about 100 meters below ground, house the two accelerators. An adjacent tunnel provides space for support instrumentation, allowing for the maintenance of equipment while the accelerator is running. Superconducting RF system accelerate electrons and positrons up to 250 GeV.

- **Beam Delivery System**

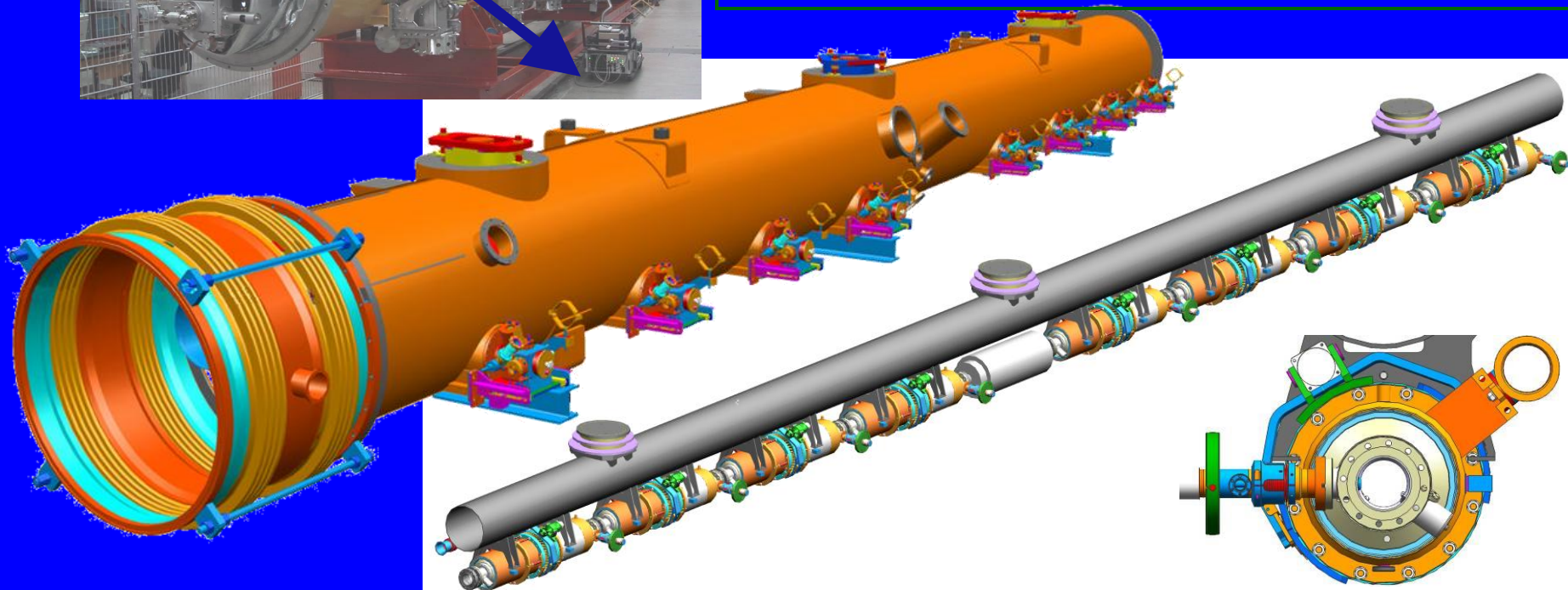
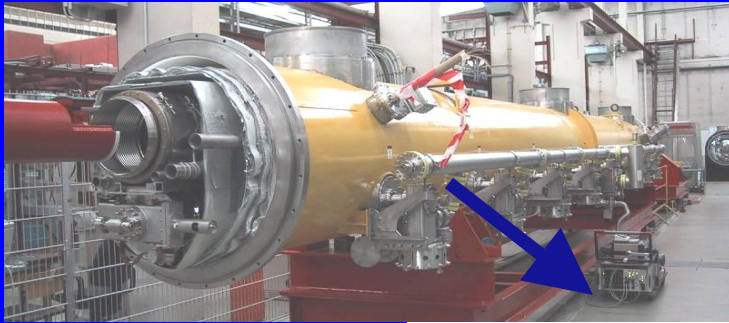
Traveling toward each other, electron and positron bunches collide at 500 GeV. The baseline configuration of the ILC provides for two collision points, offering space for two detectors.

RDR Design Parameters

Max. Center-of-mass energy	500	GeV
Peak Luminosity	$\sim 2 \times 10^{34}$	1/cm ² s
Beam Current	9.0	mA
Repetition rate	5	Hz
Average accelerating gradient	31.5	MV/m
Beam pulse length	0.95	ms
Total Site Length	31	km
Total AC Power Consumption	~ 230	MW

ILC Reference Cryomodule

- Developed by INFN for TTF-TESLA
- 3rd generation of improvements
- Many years of successful operation
- Baseline for XFEL and ILC
- Reference for others (Project X, etc)

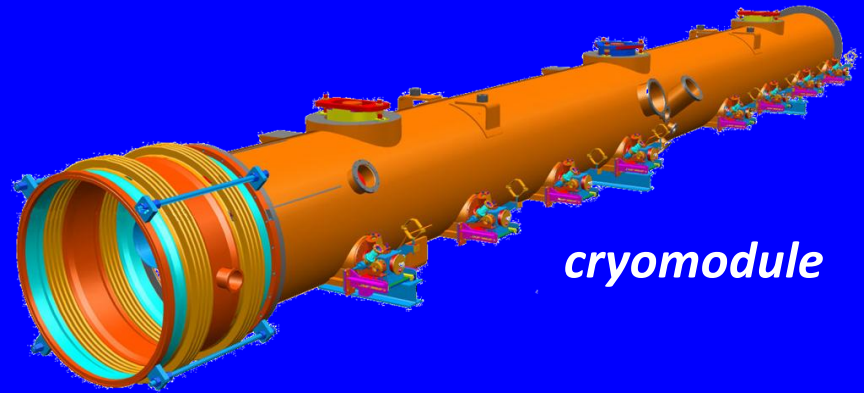


Superconducting RF Linac Technology

cavity



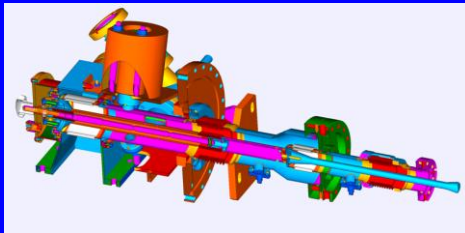
cryomodule



SCRF Linac Technology



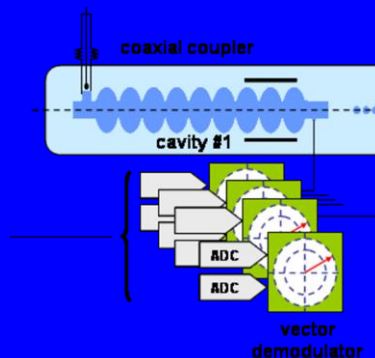
coupler



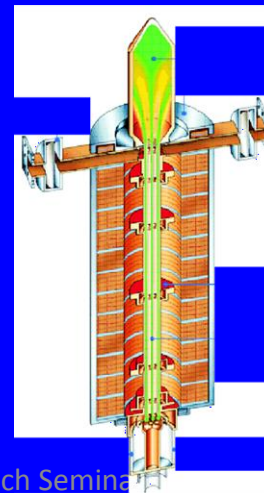
tuner



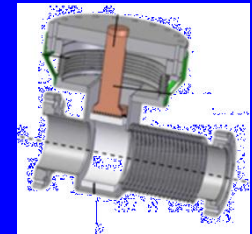
LLRF



RF



HOMs



Milestones that led to accelerators based on SRF

Superconductivity

1908: Heike Kamerlingh Onnes (Holland)

Liquefied Helium for the first time.

1911: Heike Kamerlingh Onnes

Discovered Superconductivity.

1928-34: Walther Meissner (Germany)

Discovered Superconductivity of Ta, V, Ti and Nb.

RF Acceleration

1924: Gustaf Ising (Sweden)

The First Publication on RF Acceleration
Arkiv för Matematik, Astronomi och Fysik.

1928: Rolf Wideröe (Norway, Germany)

Built the first RF Accelerator,
Arch. für Elektrotechnik 21, vol.18.

1947: Luis Alvarez (USA)

Built first DTL (32 MeV protons).

1947: W. Hansen (USA)

Built first 6 MeV e-accelerator, Mark I
(TW- structure).

↓

1961: W. Fairbank (Stanford Univ.)
Presented the first proposal for a superconducting accelerator for electrons
A. Banford and G. Stafford (Rutherford Appleton Lab.)
Presented the first proposal for a superconducting accelerator for protons

↓

1964: W. Fairbank, A. Schwettman, P. Wilson (Stanford Univ.)
First acceleration of electrons with sc lead cavity

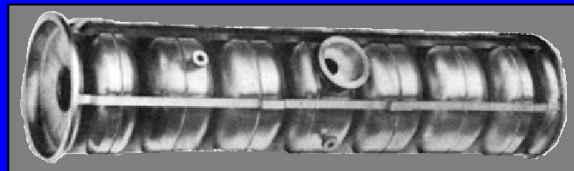
↓

1970: J. Turneaure (Stanford Univ.)
 $E_{\text{peak}} = 70 \text{ MV/m}$ and $Q \sim 10^{10}$ in 8.5 GHz cavity !

↓

1968-1981: M. McAshan, A. Schwettman, T. Smith, J. Turneaure, P. Wilson
(Stanford Univ.)

Developed and Constructed the Superconducting Accelerator SCA

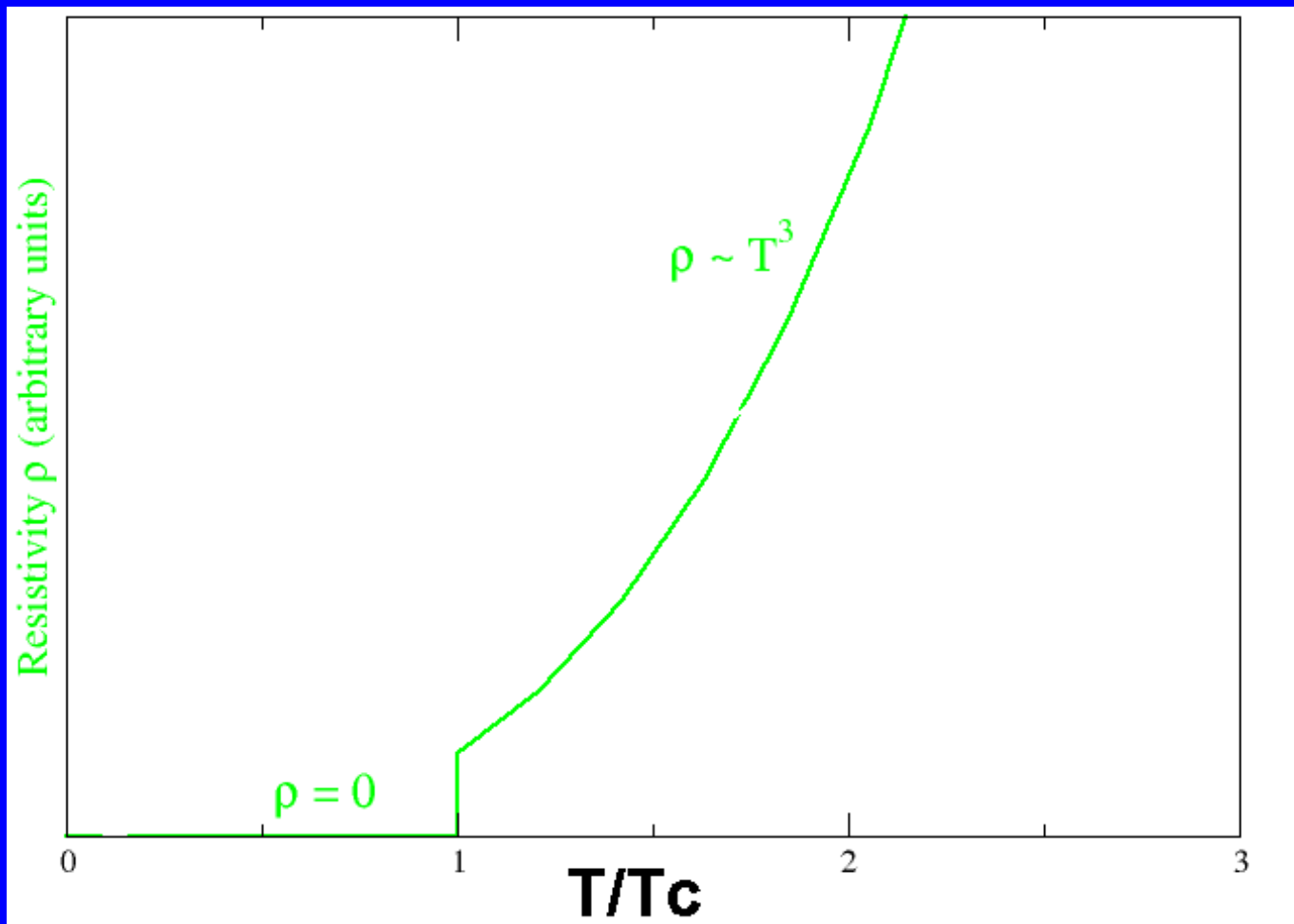


Since then, many sc accelerators were built and we are constructing and making plans for many new facilities.



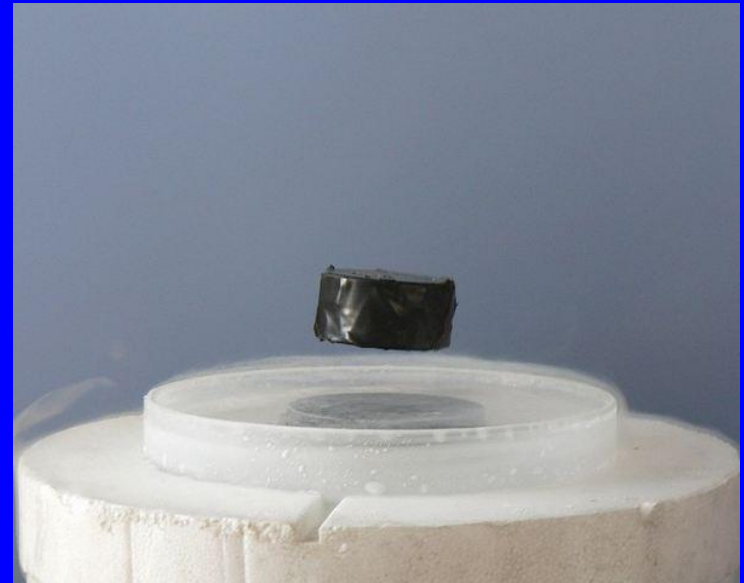
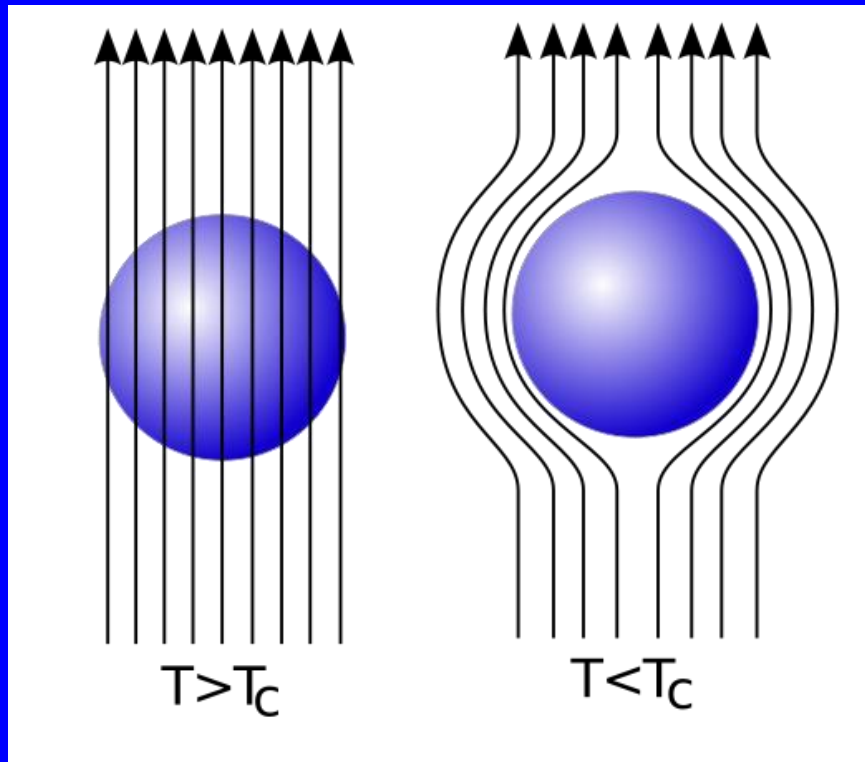
Superconductivity – Zero DC Resistance

Heike Kammerlingh-Onnes, 1911, discovery of SC in mercury



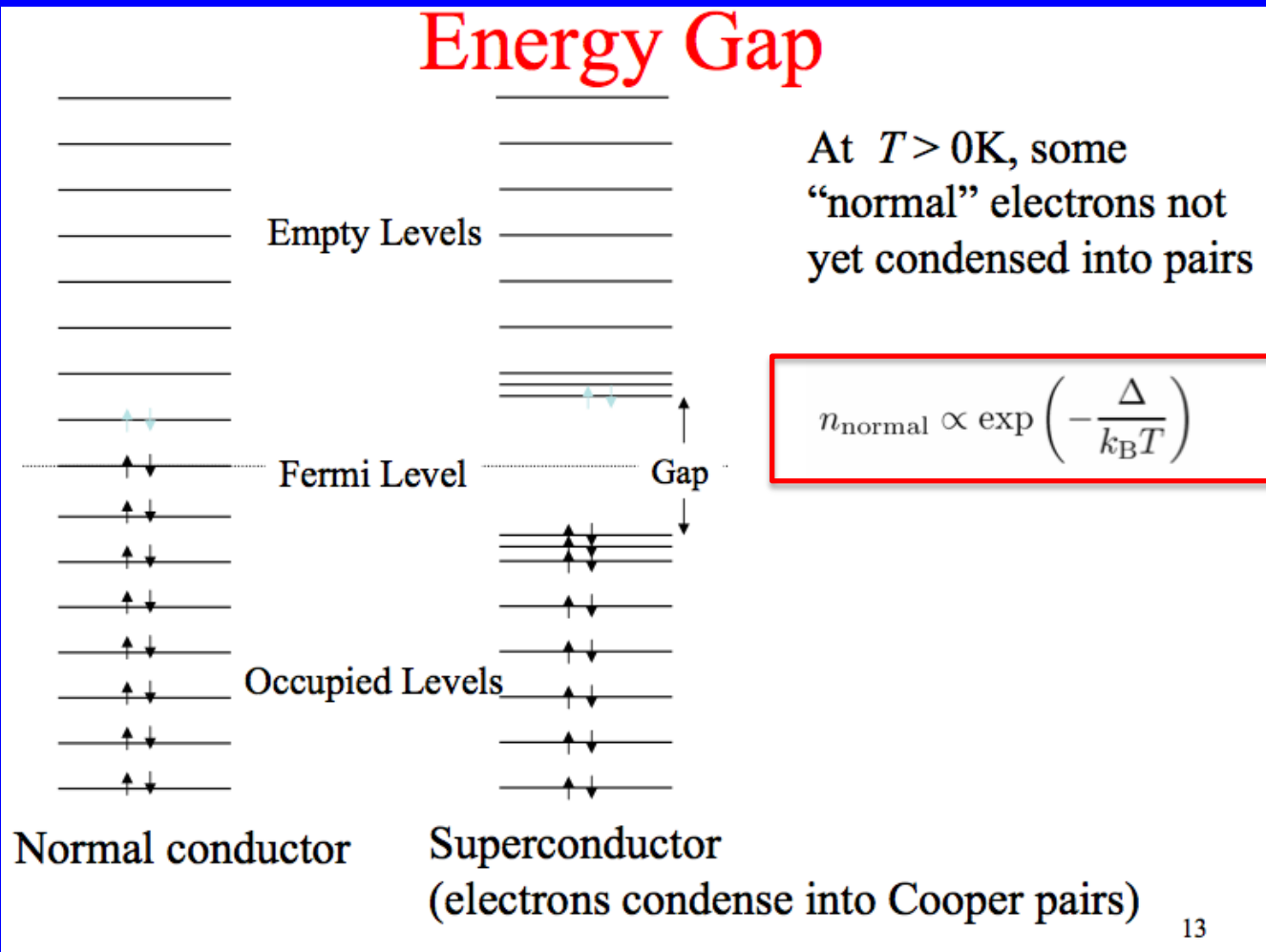
Superconductivity – Meissner Effect

Magnetic field is expelled from a superconductor



Complete magnetic shielding by circulating surface supercurrents

Energy Gap and Two-Fluid Model



- Two fluid model
 - SC electrons
 - Cooper pairs
 - Below T_c , **Cooper pairs** are formed with an energy gap 2Δ
 - Normal electrons
- DC case
 - Cooper pairs short out field
 - Normal electrons not accelerated
 - SC is Lossless even at $T > 0\text{ K}$

Losses in Superconductor

- Now look at the RF case
- Cooper pairs have inertia
 - They can not follow the AC field instantly
 - Thus do not shield AC field perfectly
 - A residual field remains
 - The normal electrons are accelerated
 - Thus dissipate power
- Scaling of RF surface resistance
 - The faster the field oscillates the less perfect the shielding
 - RF surface resistance increases with frequency
 - The more normal electrons, the lossier the material
 - RF surface resistance decreases with temperature below T_c

Figure of Merit

Surface current ($\propto H$) results in power dissipation proportional to the surface resistance (R_s)

$$\frac{dP_c}{ds} = \frac{1}{2} R_s |H|^2$$

Total power dissipation in cavity wall $P_c = \frac{1}{2} \int_s R_s |\mathbf{H}|^2 ds$

Stored energy in cavity $U = \frac{1}{2} \mu_0 \int_V |\mathbf{H}|^2 dv$

Cavity quality factor $Q_0 = \frac{\omega_0 U}{P_c} = \frac{\omega_0 \mu_0}{R_s} \frac{\int_V |\mathbf{H}|^2 dv}{\int_s |\mathbf{H}|^2 ds}$

$$\begin{aligned} &= 10^4 \text{ for n.c.} \\ &= 10^{10} \text{ for s.c.} \end{aligned}$$

Measure of how lossy the cavity material is

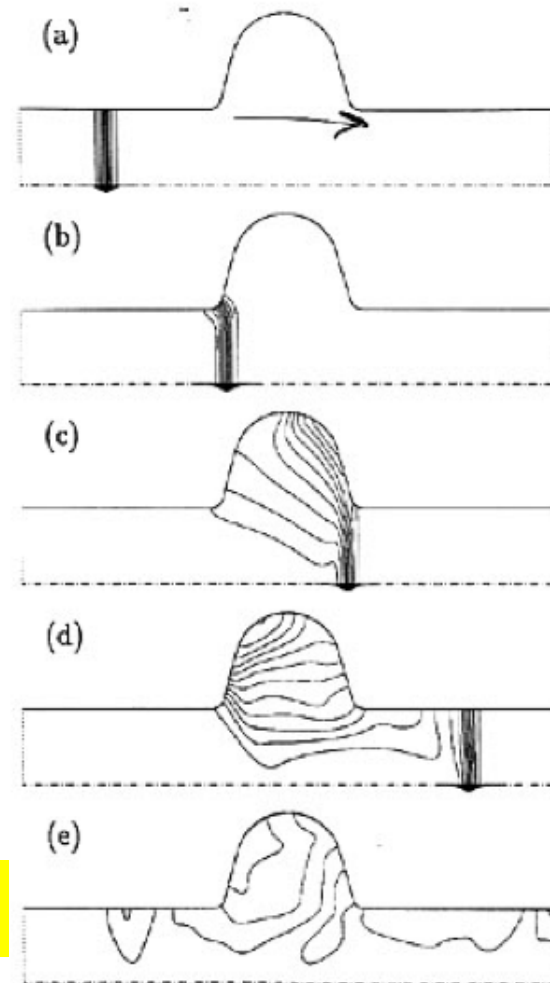
Features of SRF Cavity

- Low power dissipation
 - allows high gradient in CW or long-pulsed operation
 - Less number of cells
 - Less disruption to beam
 - Shorter linac and tunnel length
 - Cost saving
 - allows cavity design with large beam tube
 - Many benefits (next slide)

Features of SRF Cavity

Large beam tube & Fewer cells

- Reduces the interaction of the beam with the cavity (scales as size^3) \rightarrow
- The beam quality is better preserved (important for, e.g., FELs).
- HOMs are removed easily \rightarrow better beam stability \rightarrow more current accelerated (important for, e.g., B-factories)
- Reduce the amount of beam scraping \rightarrow less activation in, e.g., proton machines (important for, e.g., SNS, Neutrino factory)

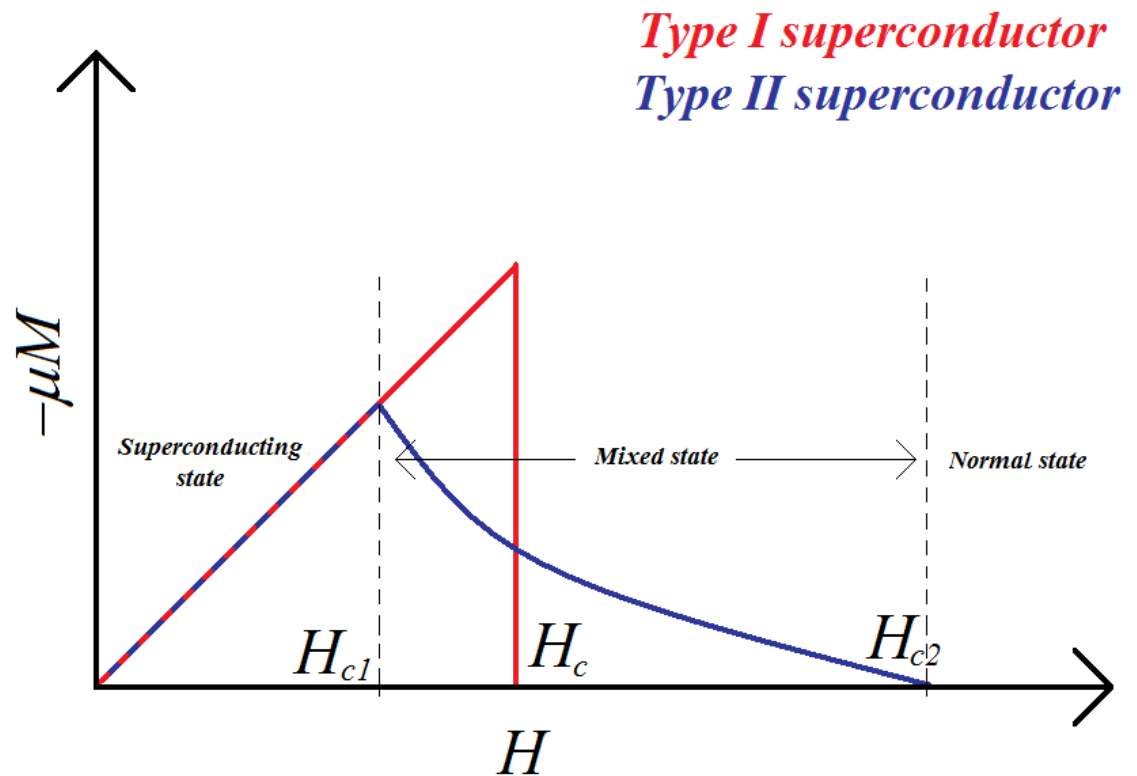
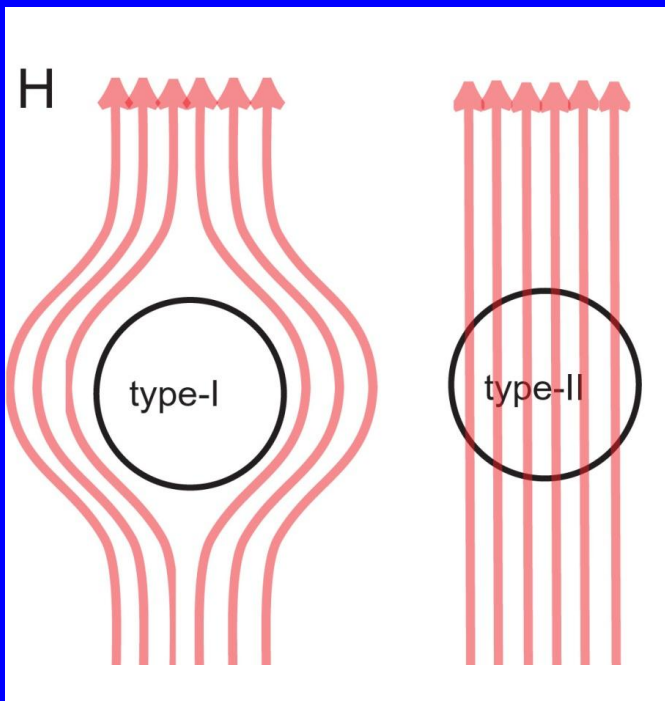


Large aperture of SRF cavity relaxes wakefields

Type-I and Type-II Superconductor

- Two types of superconductors defined by Ginsburg-Landau

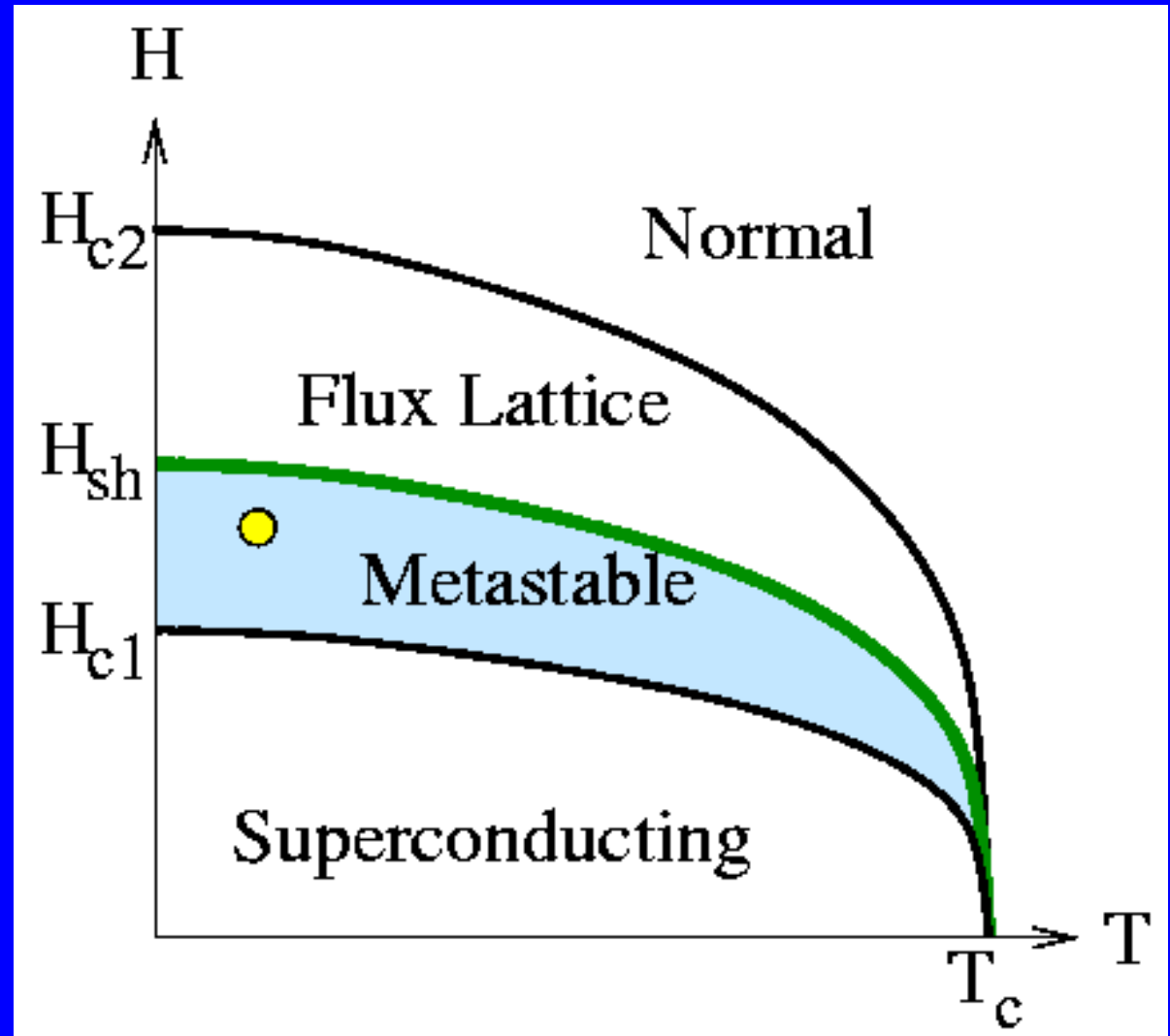
$$\kappa = \lambda(T) / \xi(T)$$



- $\kappa < 1/\sqrt{2}$ and $\kappa > 1/\sqrt{2}$

DC and RF Critical Field of Superconductor

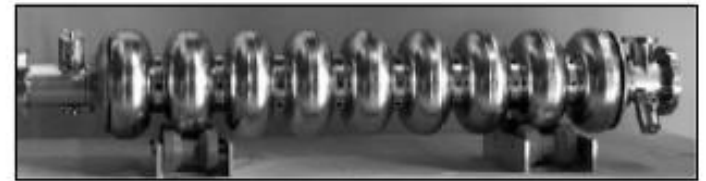
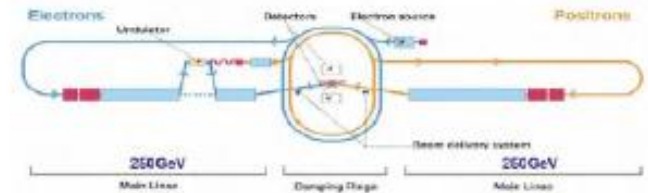
Superheating field
due to surface barrier
to vortex penetration





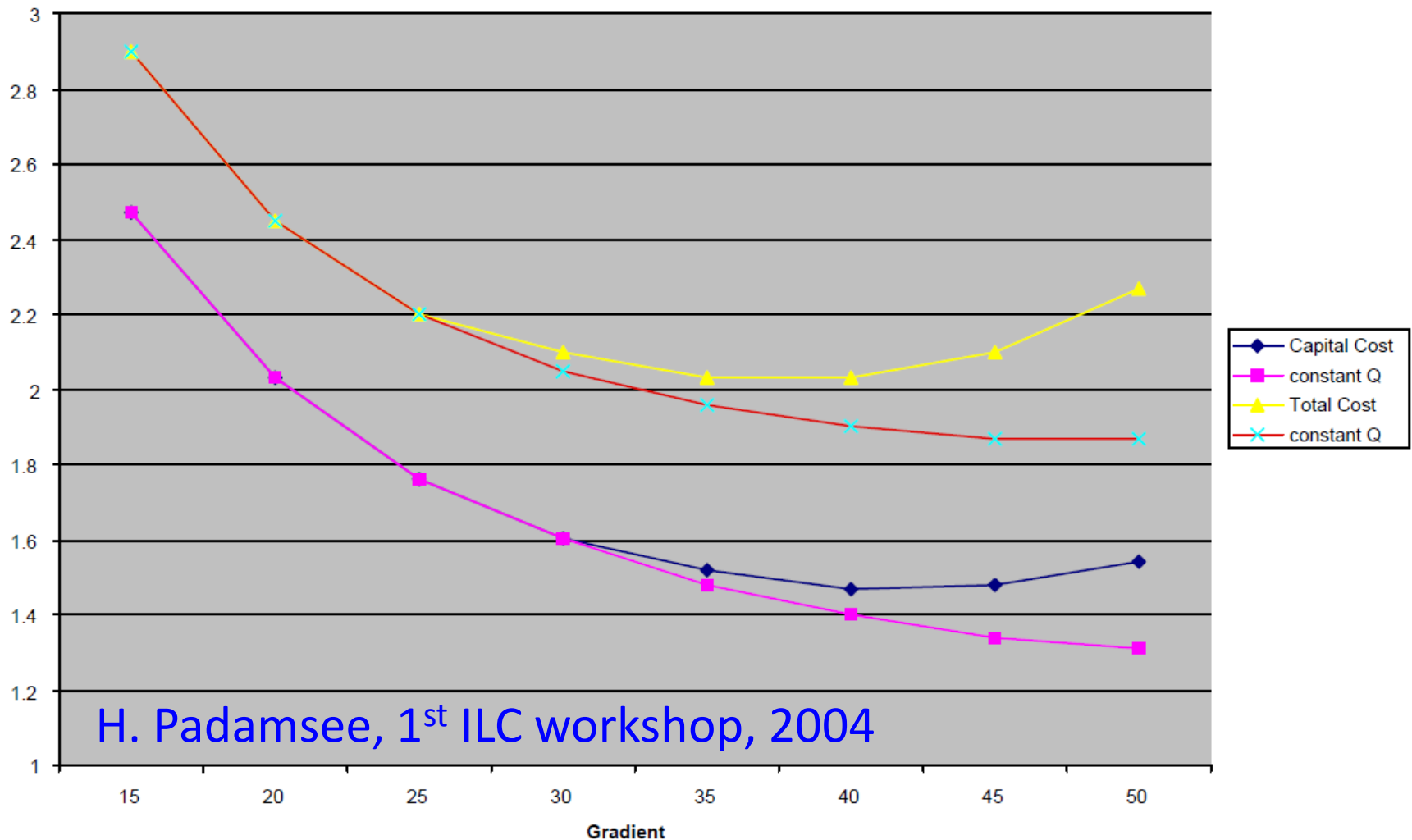
SCRF Technology Required

Parameter	Value
C.M. Energy	500 GeV
Peak luminosity	$2 \times 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$
Beam Rep. rate	5 Hz
Pulse time duration	1 ms
Average beam current	9 mA (in pulse)
Av. field gradient	31.5 MV/m
# 9-cell cavity	14,560
# cryomodule	1,680
# RF units	560

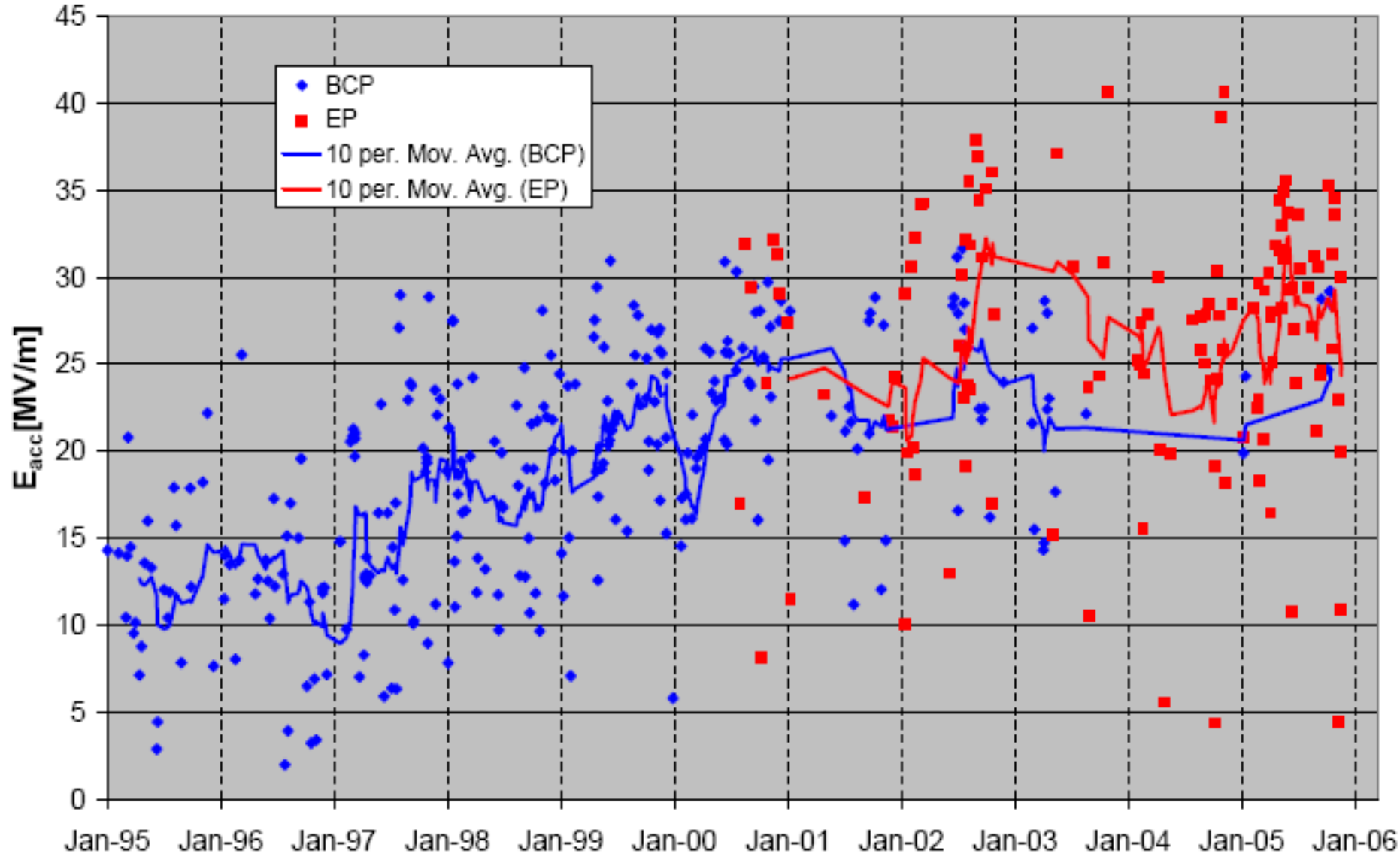


Gradient a Major Cost Driver for ILC

Cost vs Gradient



H. Padamsee, 1st ILC workshop, 2004



Lutz Lilje DESY -MPY-



28.07.2006

Global Plan for ILC Gradient R&D

Year	07	2008	2009	2010	2011	2012
Phase	TDP-1			TDP-2		
Cavity Gradient in v. test to reach 35 MV/m	→ Yield 50%			→ Yield 90%		
Cavity-string to reach 31.5 MV/m, with one-cryomodule		Global effort for string assembly and test (DESY, FNAL, INFN, KEK)				
System Test with beam acceleration			FLASH (DESY) , NML (FNAL) STF2 (KEK, test start in 2013)			
Preparation for Industrialization				Production Technology R&D		

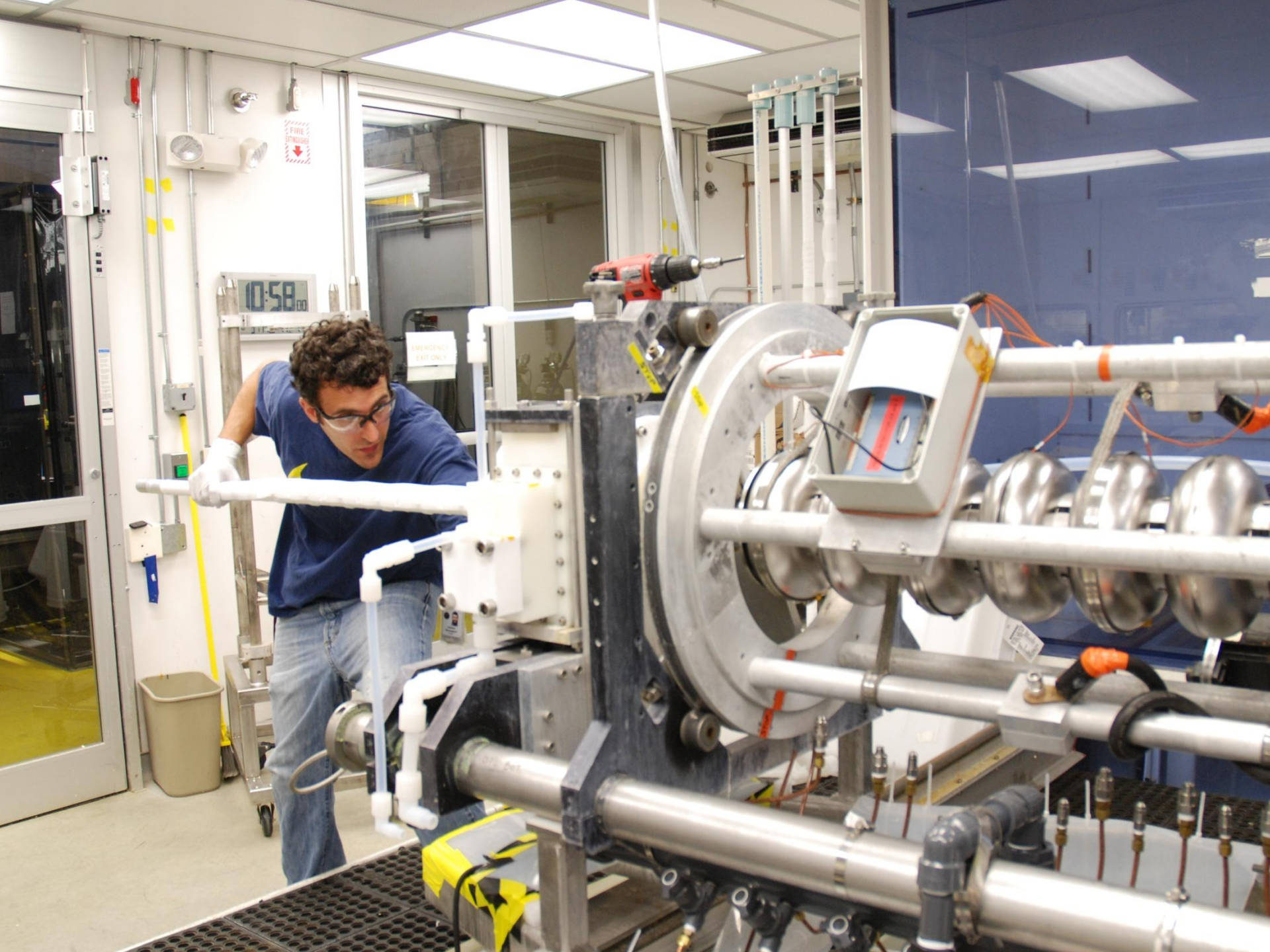
New baseline gradient:

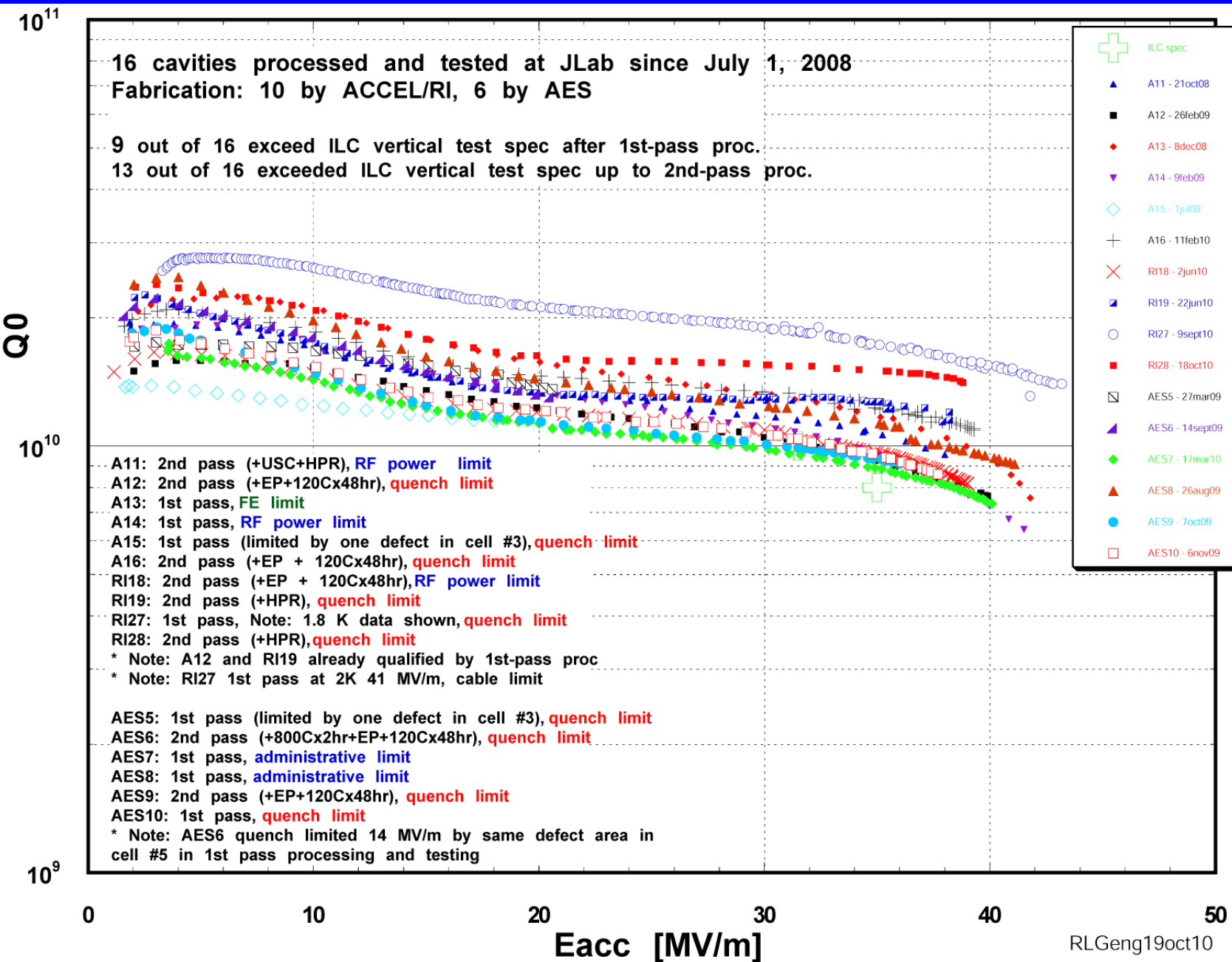
Vertical acceptance: 35 MV/m average, allowing $\pm 20\%$ spread (28-42 MV/m)

Operational: 31.5 MV/m average, allowing $\pm 20\%$ spread (25-38 MV/m)

Baseline ILC Nb Cavity Proc. Procedure

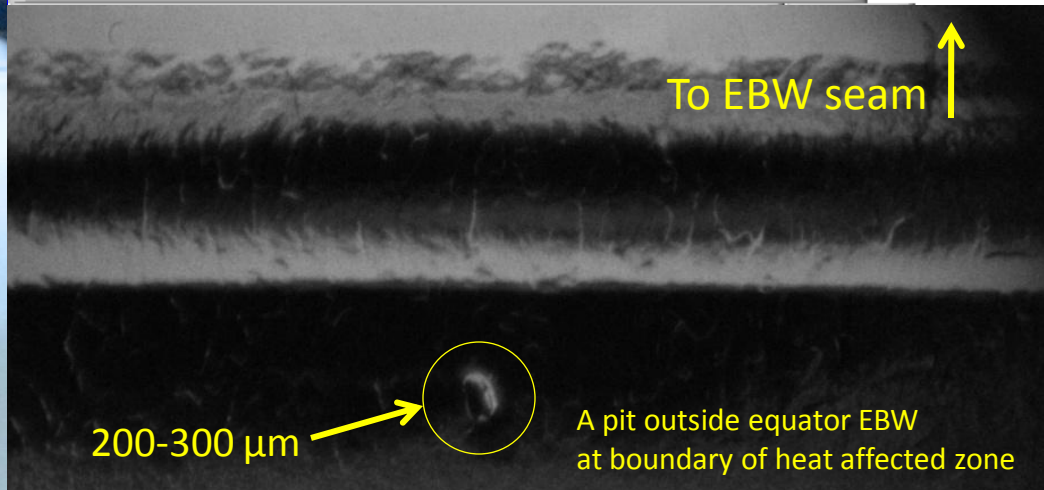
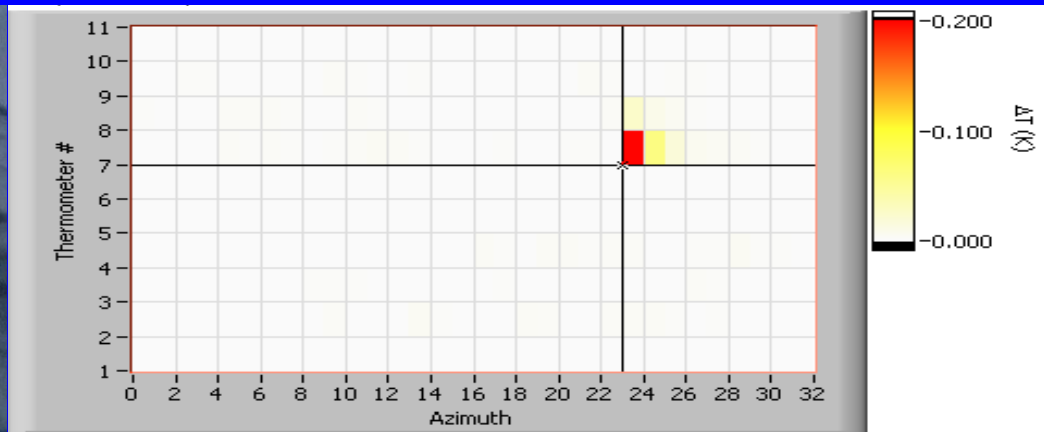
- Cavity manufacture (EBW) using RRR 300 Nb
- Initial light chemistry 5-30 μm (BCP)
- Heavy chemistry 80-150 μm (EP)
- Post-EP cleaning
- Vacuum furnace heat treatment 750-800 $^{\circ}\text{C}$
- Light chemistry 20-50 μm (EP)
- Post-EP cleaning (ER/USC+HOM coupler brushing)
- Initial HPR
- Clean room assembly
- Final HPR
- Pump down
- 120 $^{\circ}\text{C}$ x 48hr bake-out





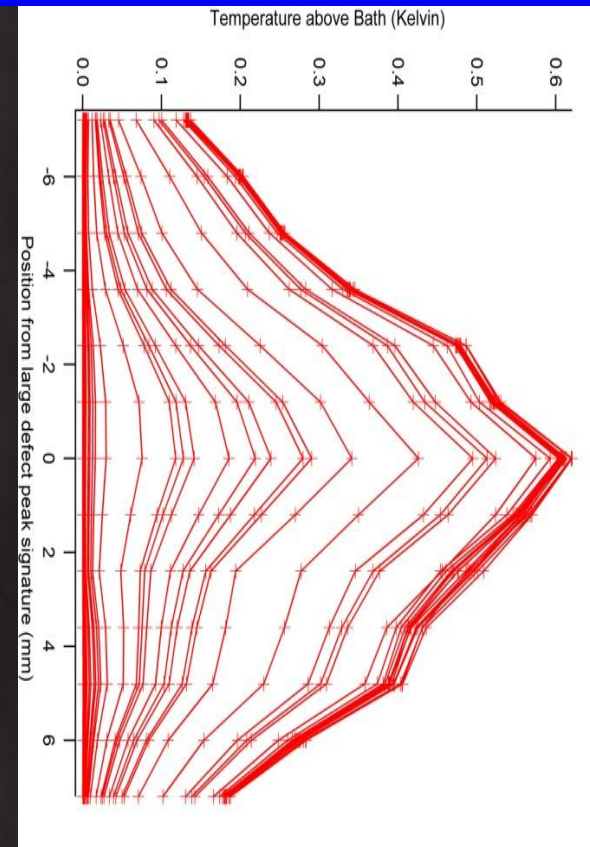
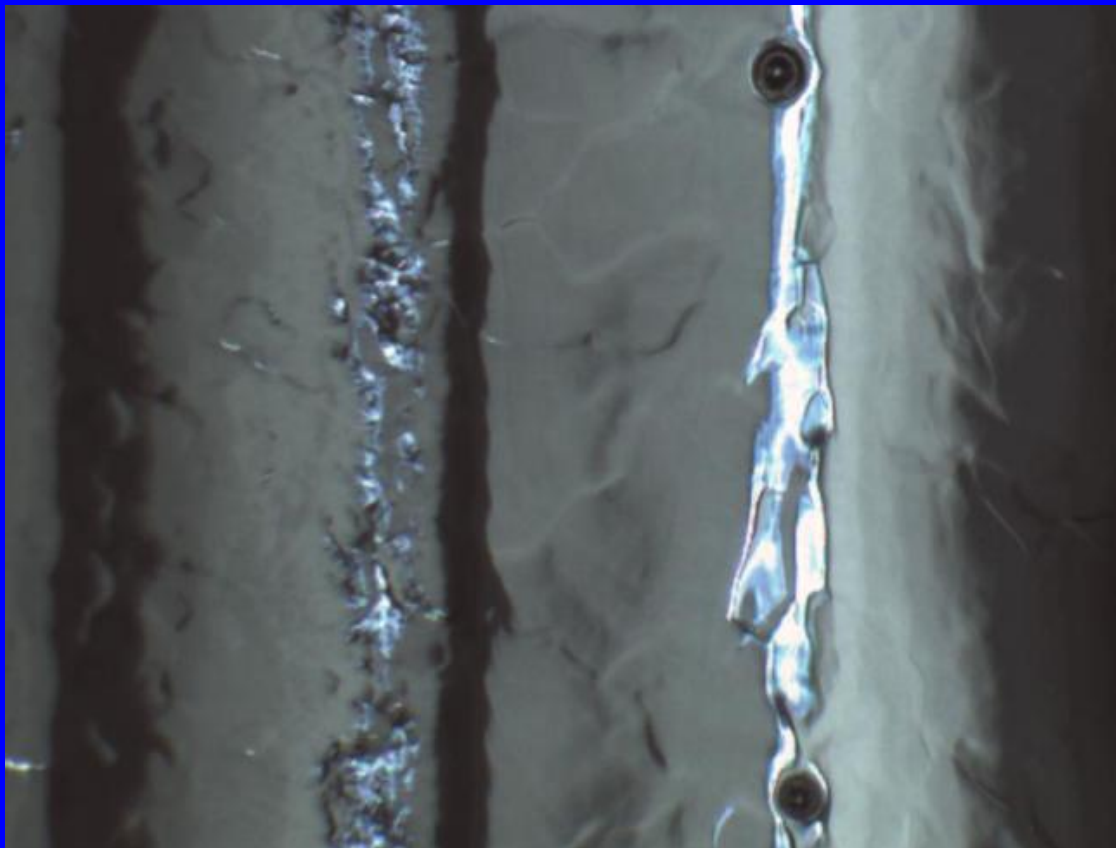
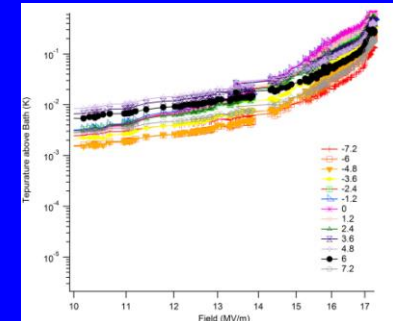
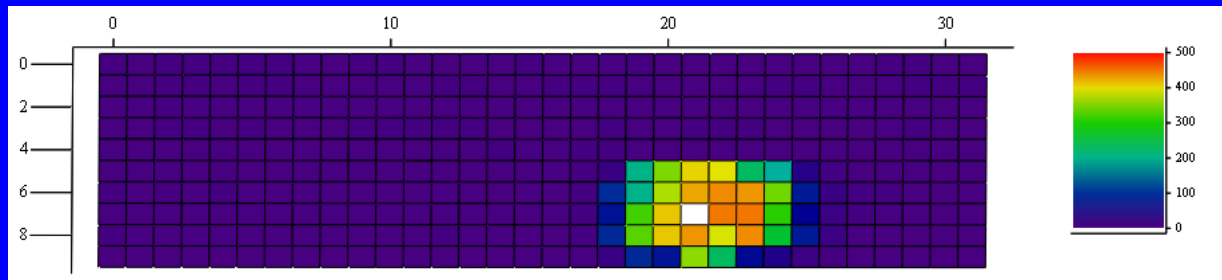
A15 gradient limit at 19 MV/m: T-mapping found a hot spot correlated to quench

Long distance microscope identified a defect near hot spot

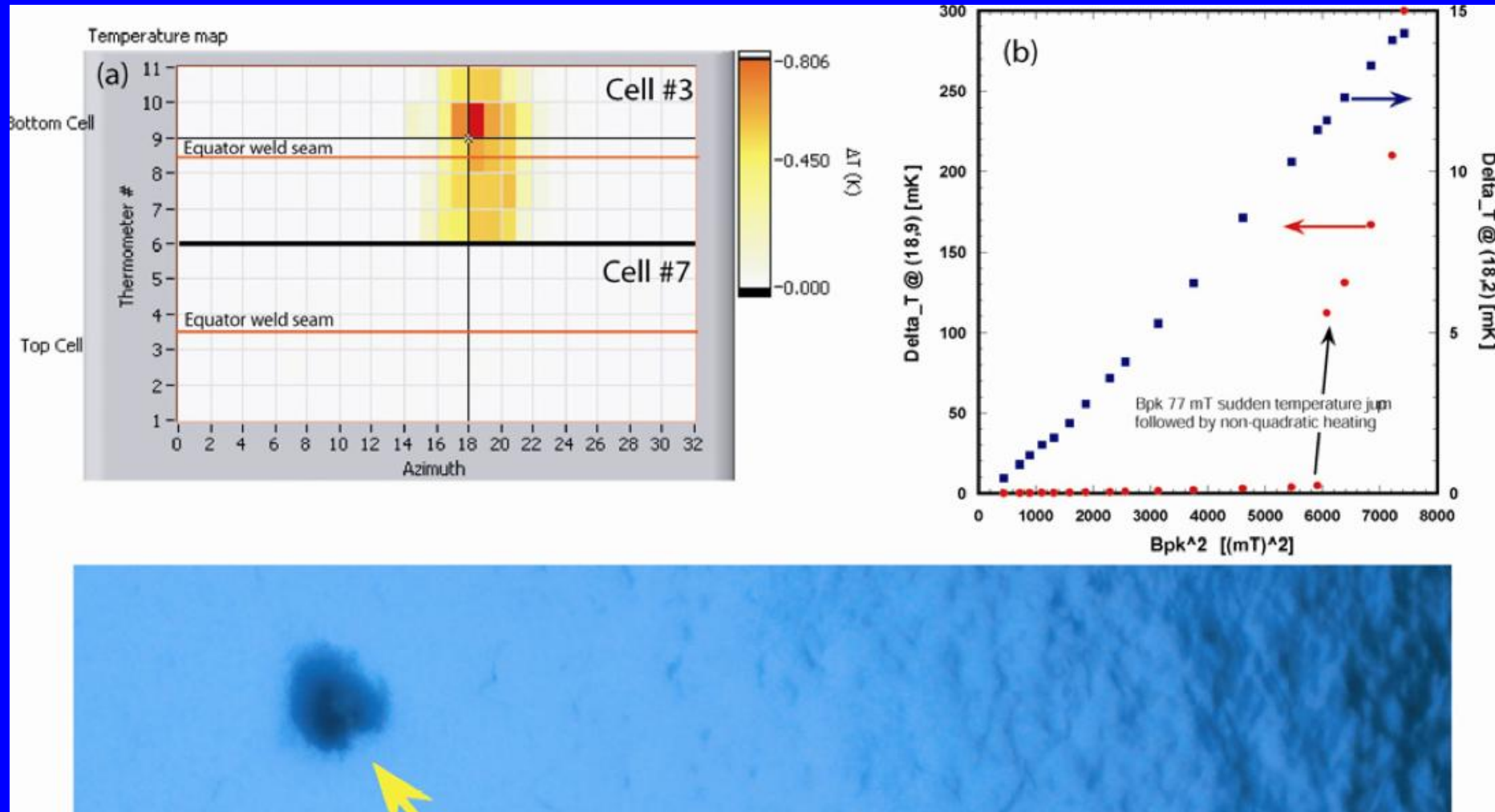


Fine grain EP

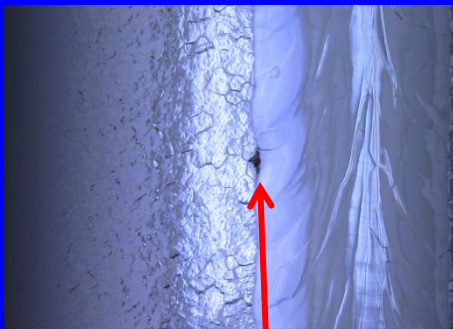
twin defects causing quench 17 MV/m. Cavity by a new vendor



Local Heating near Quench-Causing Defect



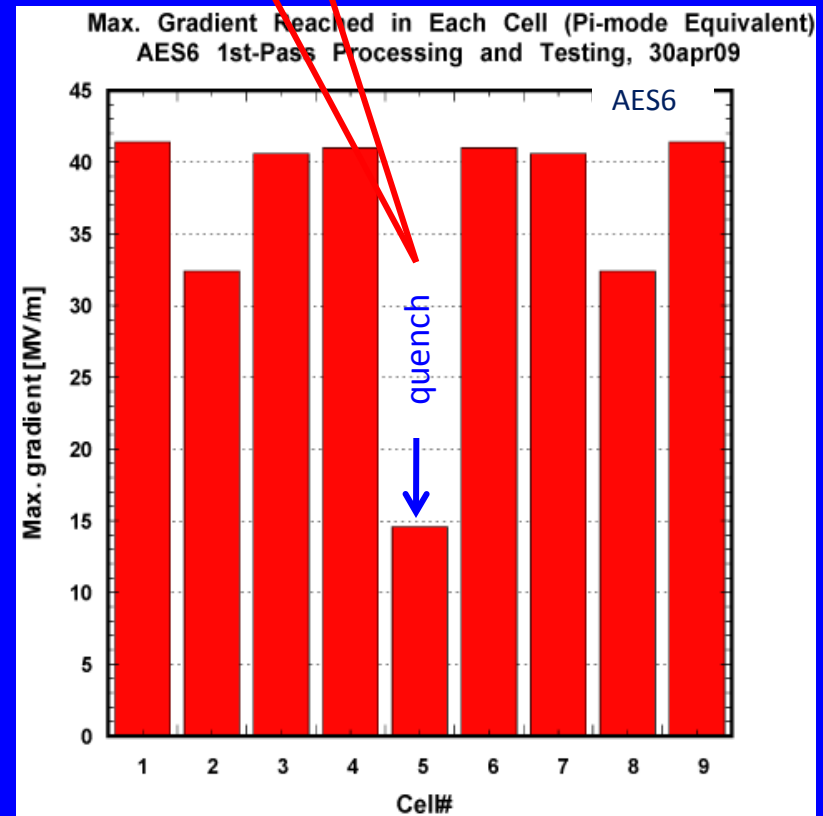
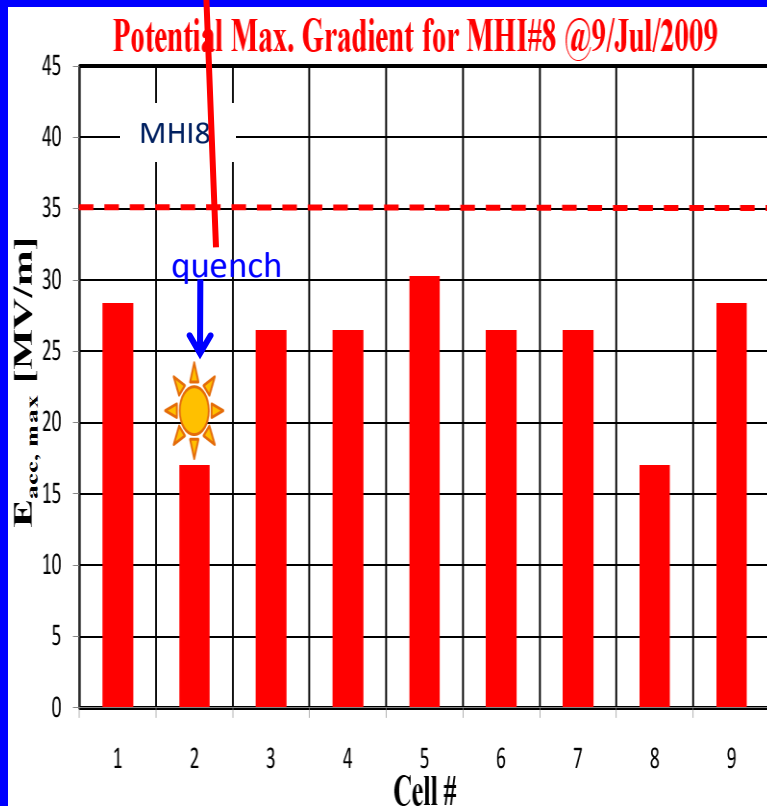
Examples with Observable Defect



Deep pit at boundary of under-bead of equator EBW

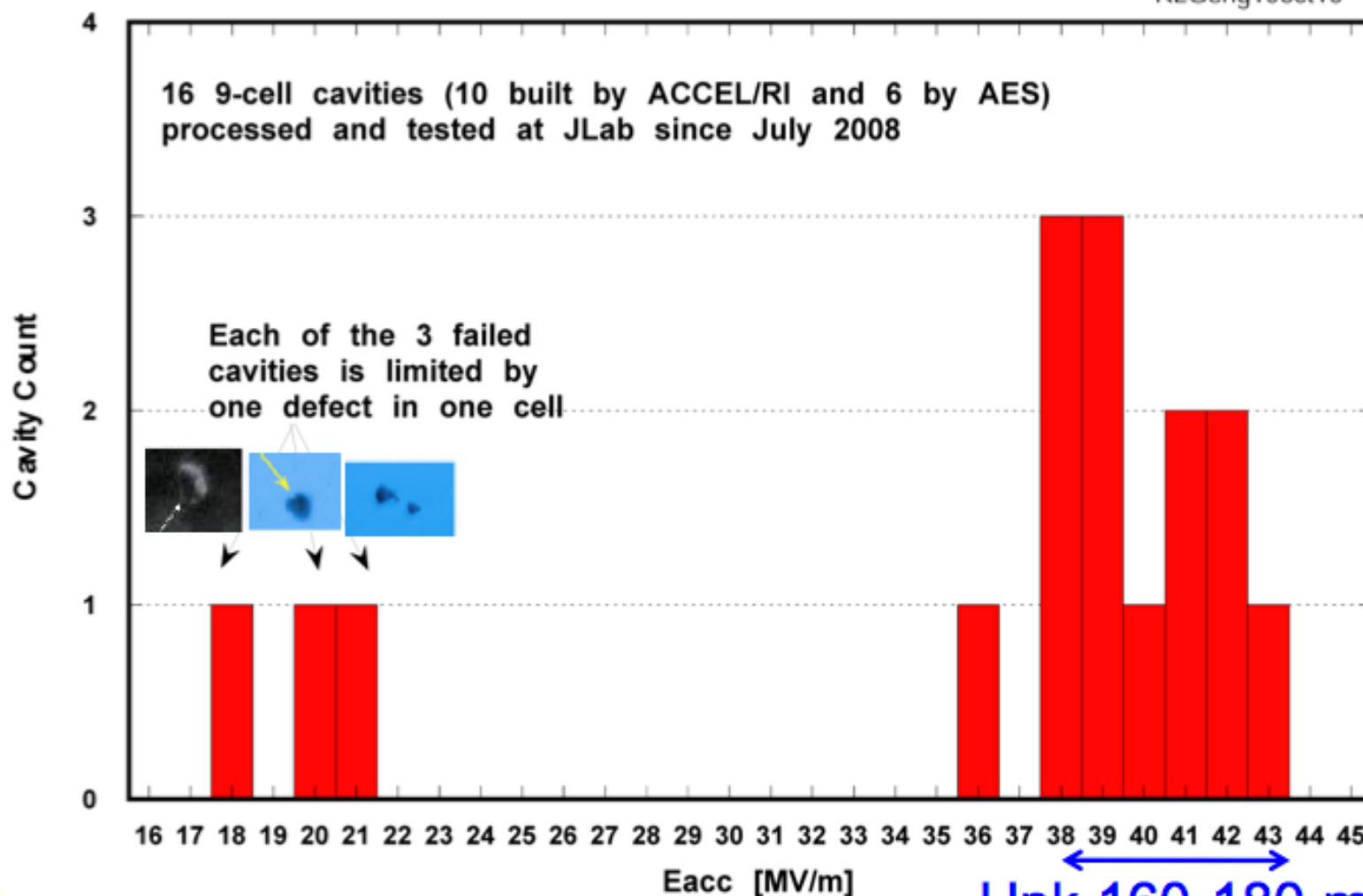


Twin defects 300-500µm dia.
8mm from equator EBW seam



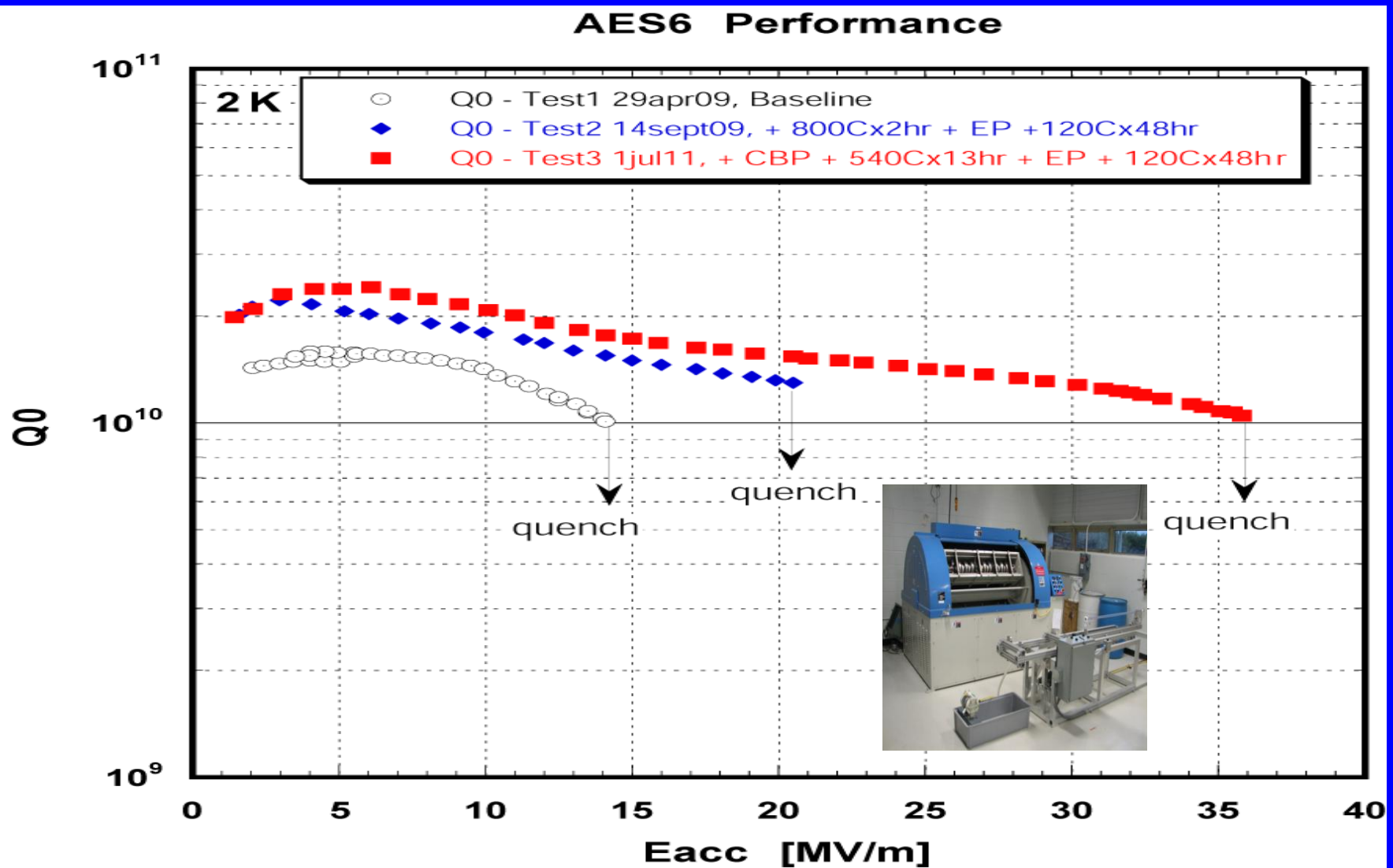
Gradient Scatter (up to 2nd-pass proc.)

RLGeng19oct10



AES6 reached 36 MV/m at Q0 1E10

“Mirror finish” CBP at FNAL, then USC, HT, EP, bake, test at JLab



Today July 15, 2011 at ILC ART Director Visit of JLab

94%

Yield at 35 MV/m achieved at
Average gradient 39 MV/m

JLAB + FNAL

Gradient Scatter (up to 2nd-pass)

16 recent data from cavities built by ACCEL/RI and AES

16 9-cell cavities (10 built by ACCEL/RI and 6 by AES)
processed and tested at JLab since July 2008

Jefferson Lab

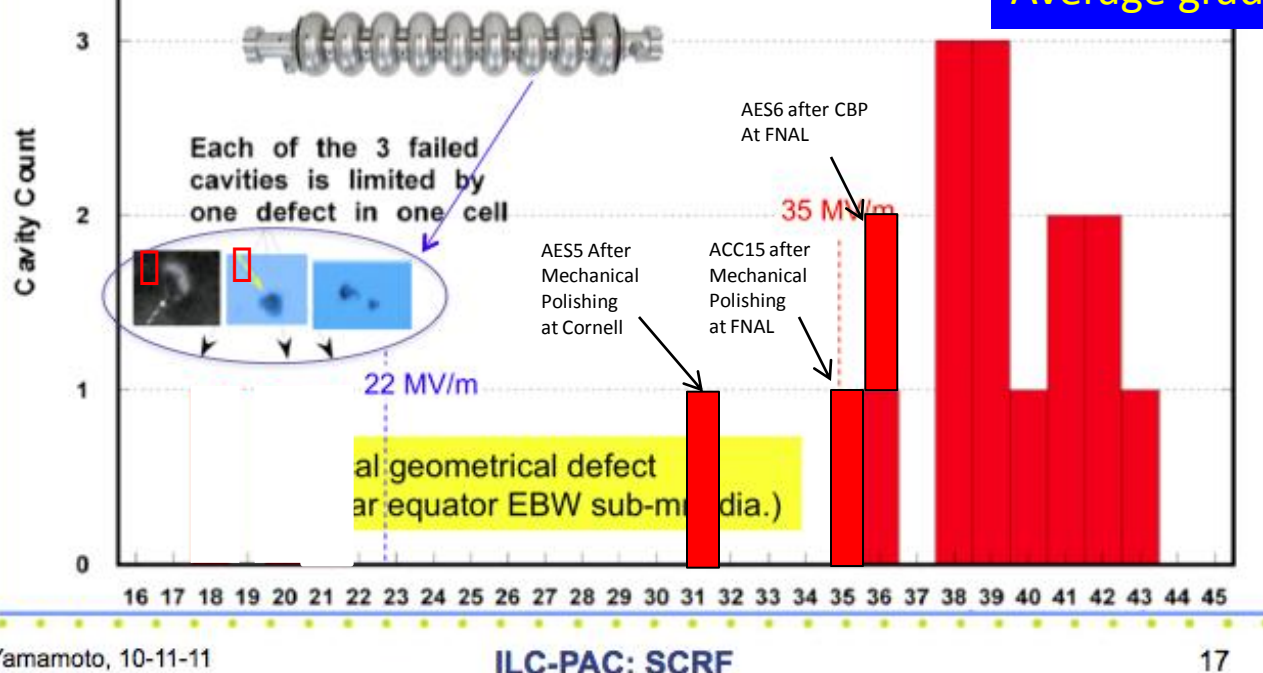
Fermilab



Cornell University

100% yield at ≥ 31 MV/m

Average gradient 39 MV/m



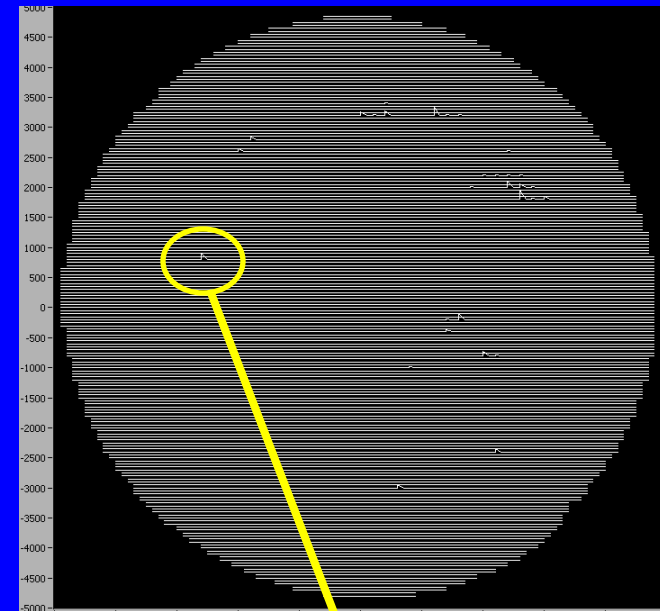
Understanding FE Behaviors w/ Samples

Surface studies of Nb samples EP'ed together w/ 9-cell cavities

JLab Scanning Field Emission SEM



- Scan Nb surface with biased tip – DC field upto 140 MV/m
- Field emission sites and I-V curve registered
- Sample transferred to SEM chamber under vacuum
- Nature of field emitter determined

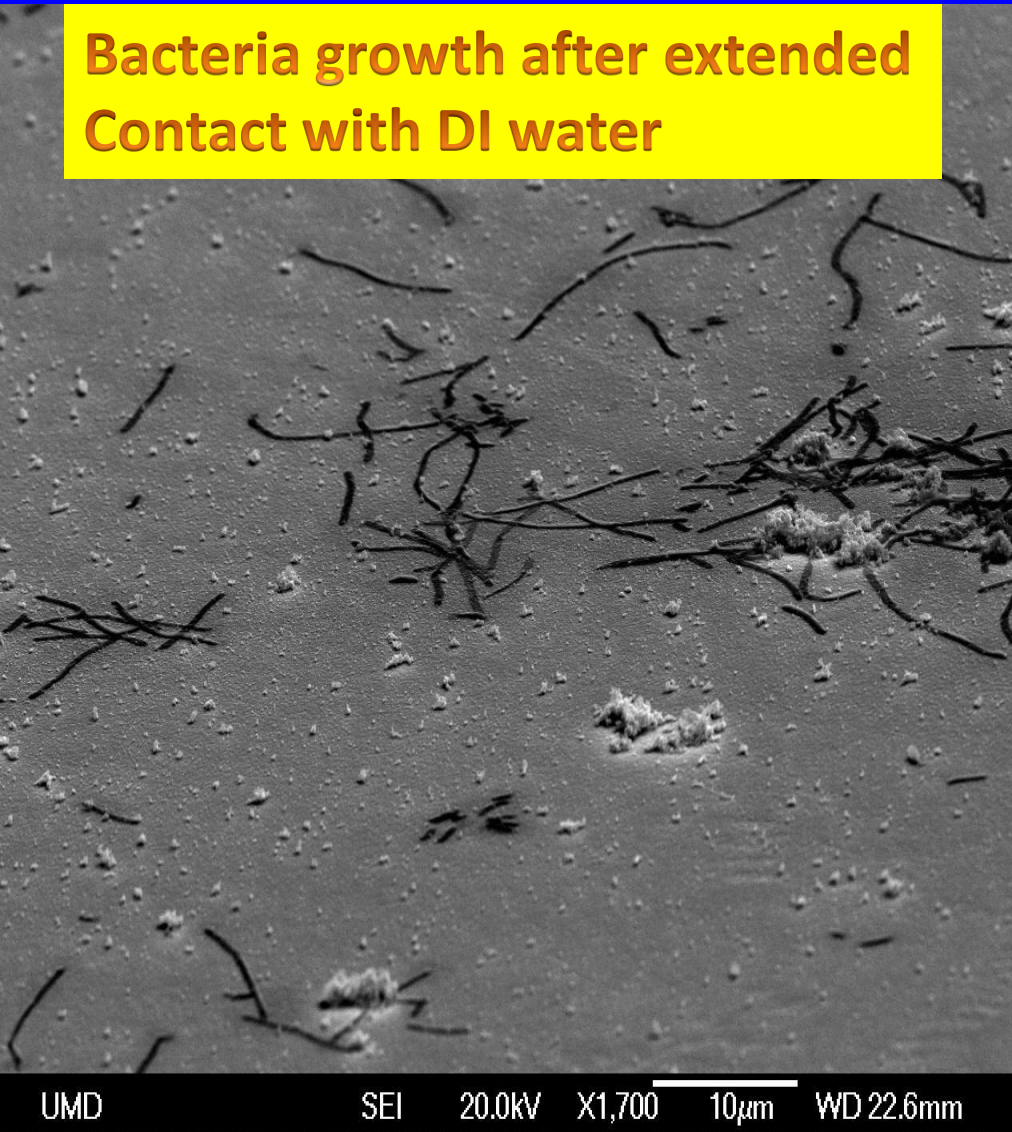


Nb-O

Dominant field emitters
on as EP'ed Nb surface

Understanding FE Behaviors w/ Samples

**Bacteria growth after extended
Contact with DI water**



**KEK Sponge cleaning
Collaboration with JLab
US-Japan Cooperation Fund**

Before

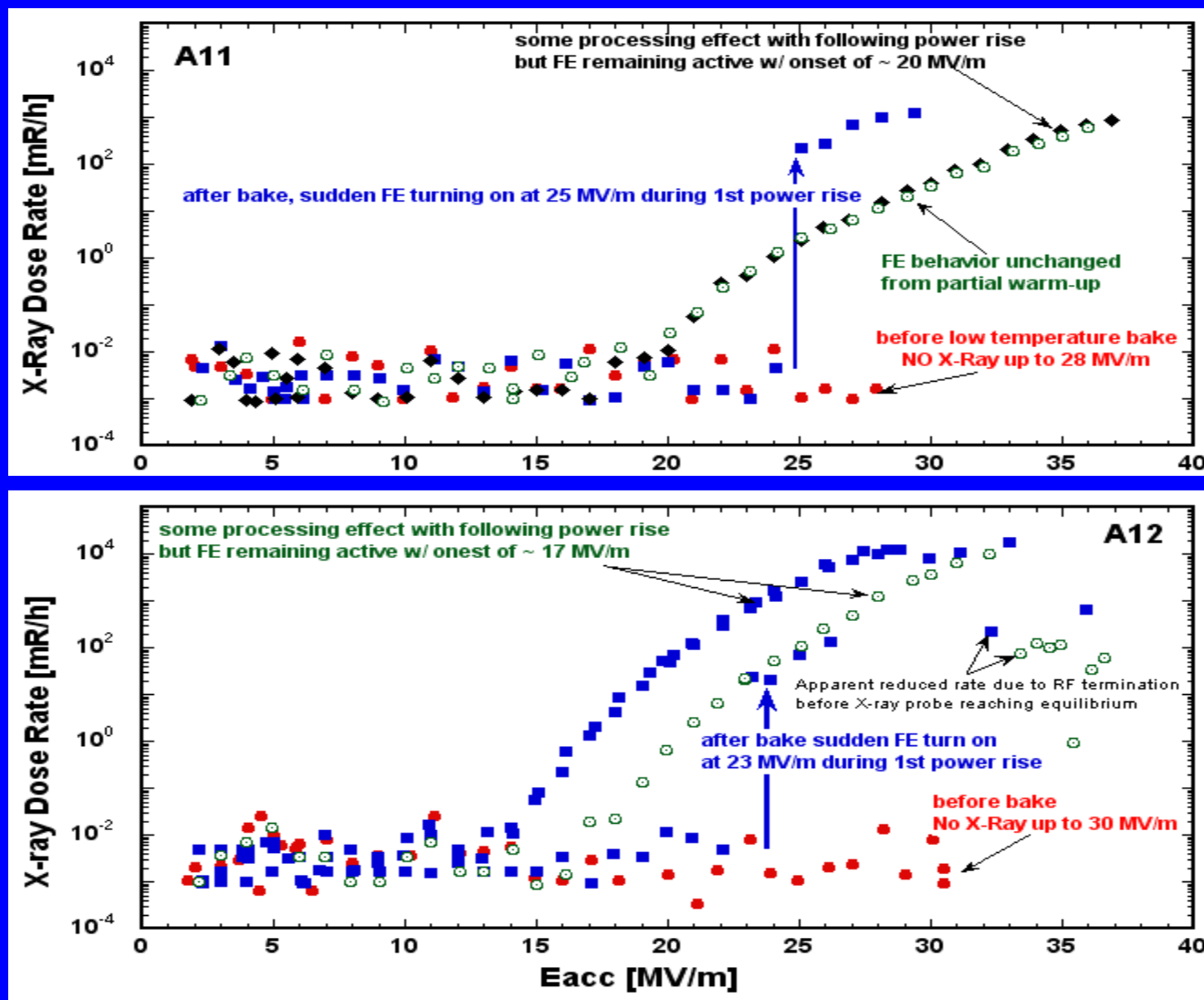
250× 28.0 kV 100μm AMRAY #0000

After

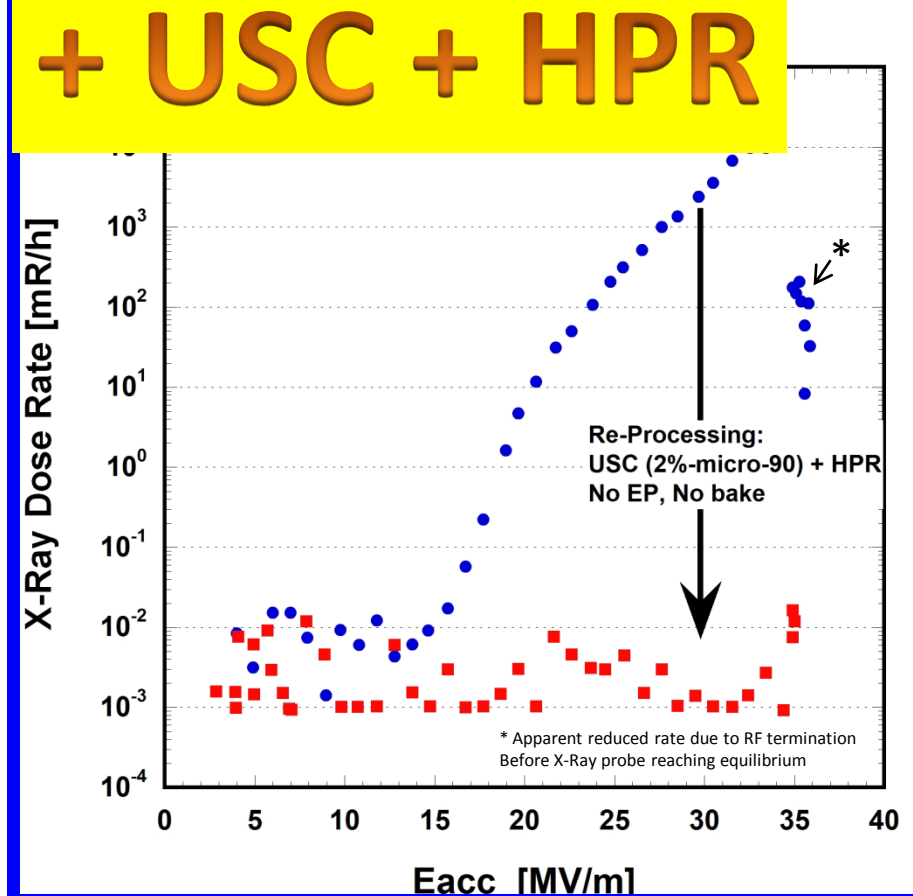
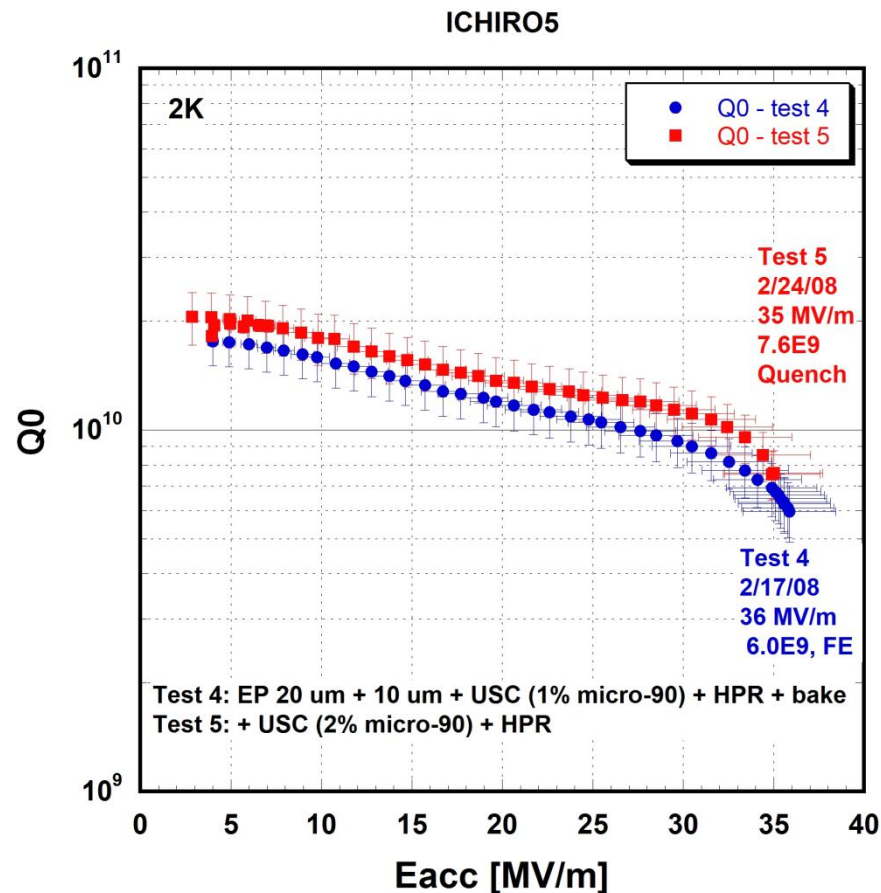
250× 28.0 kV 100μm AMRAY #0000

Understanding FE Behaviors w/ Real Cavities

Observation of Baking Induced Field Emission in EP'ed Cavity

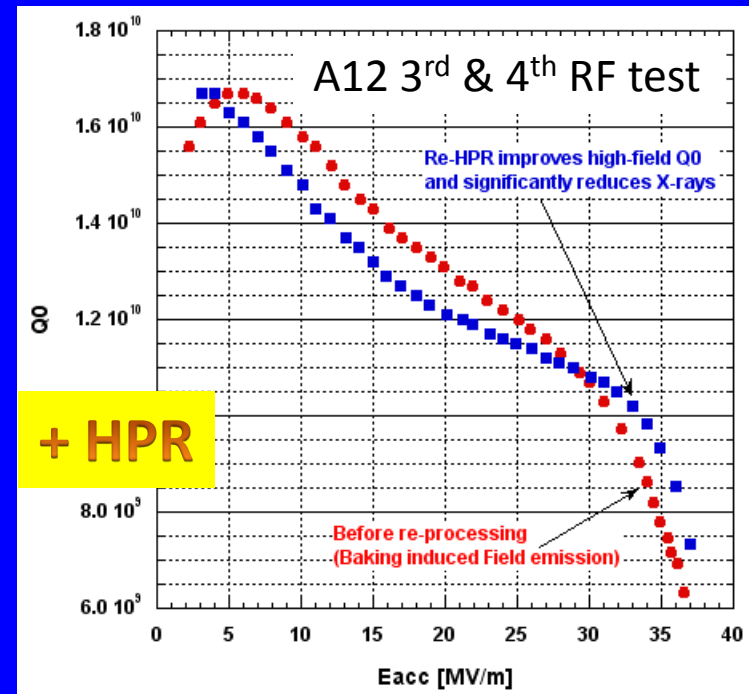
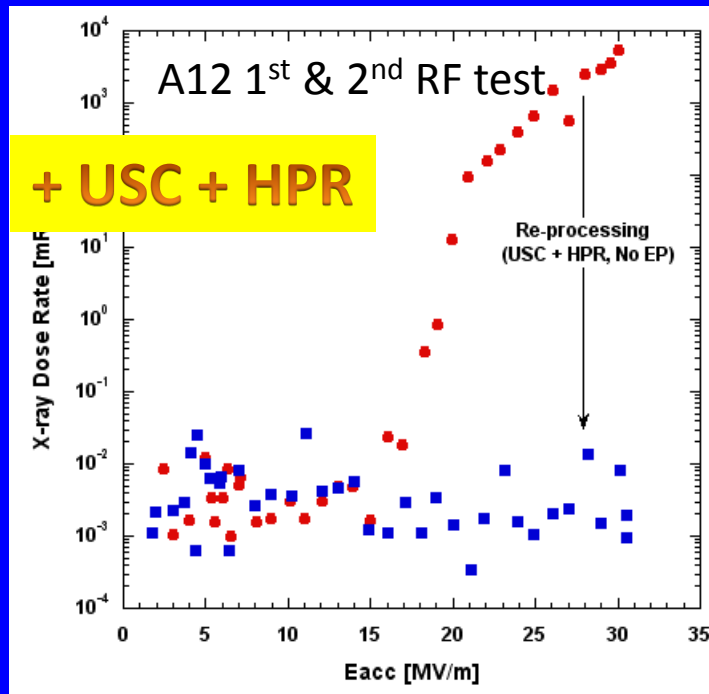


First Example of Reducing/Eliminating Field Emission by Re-cleaning



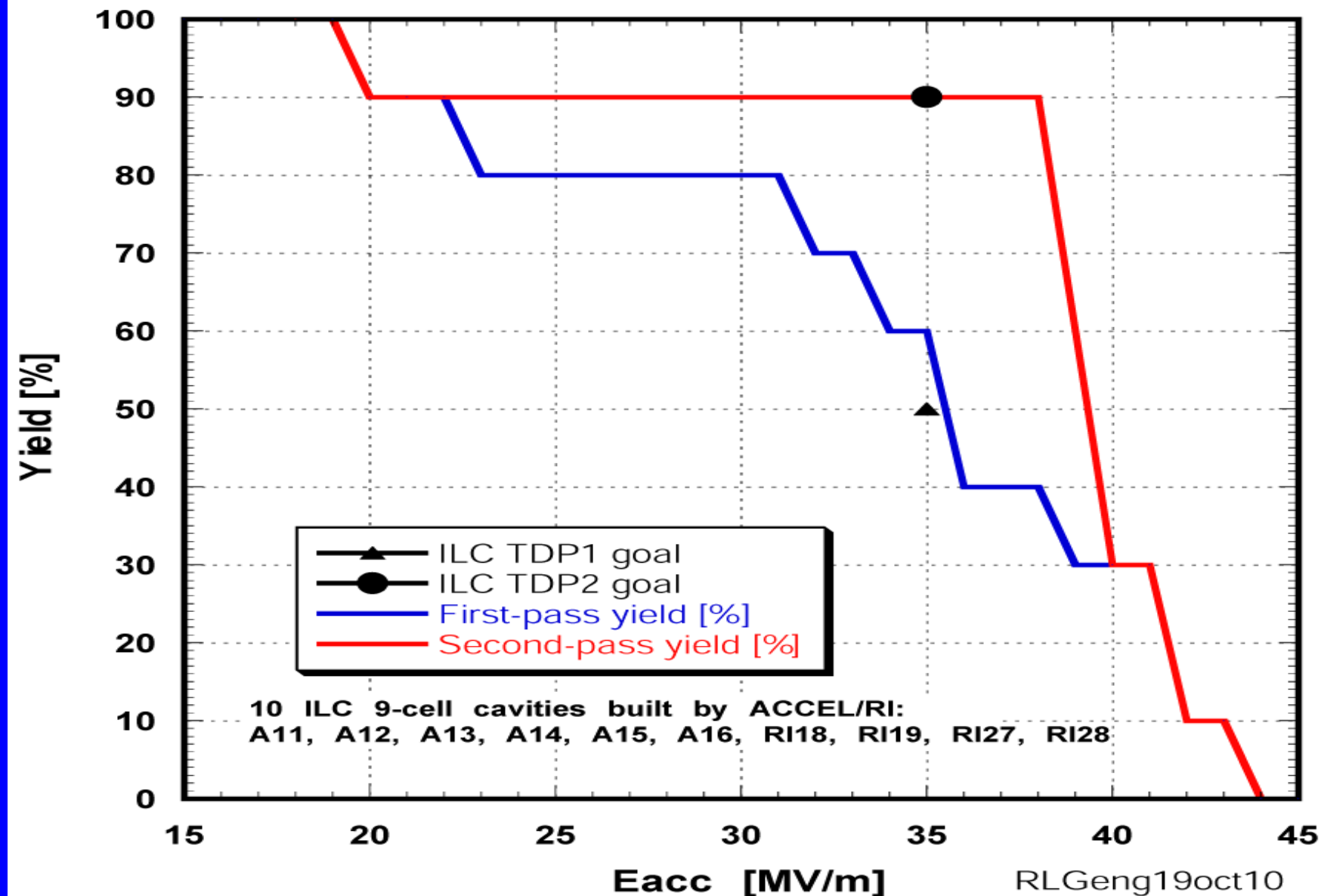
More details of multiple processing and testing results can be found in JLab report at ILC SCRF meeting, April 21-25, 2008, FNAL

More Examples of Reducing/Eliminating Field Emission by Re-cleaning



Another example is A6: last S0 test at JLab 37 MV/m, limited by field emission. After shelf storage over a year, A6 re-cleaning (USC + HPR) and shipped under vacuum, RF test at FNAL saw an improved Q(Eacc) over the last test at JLab.

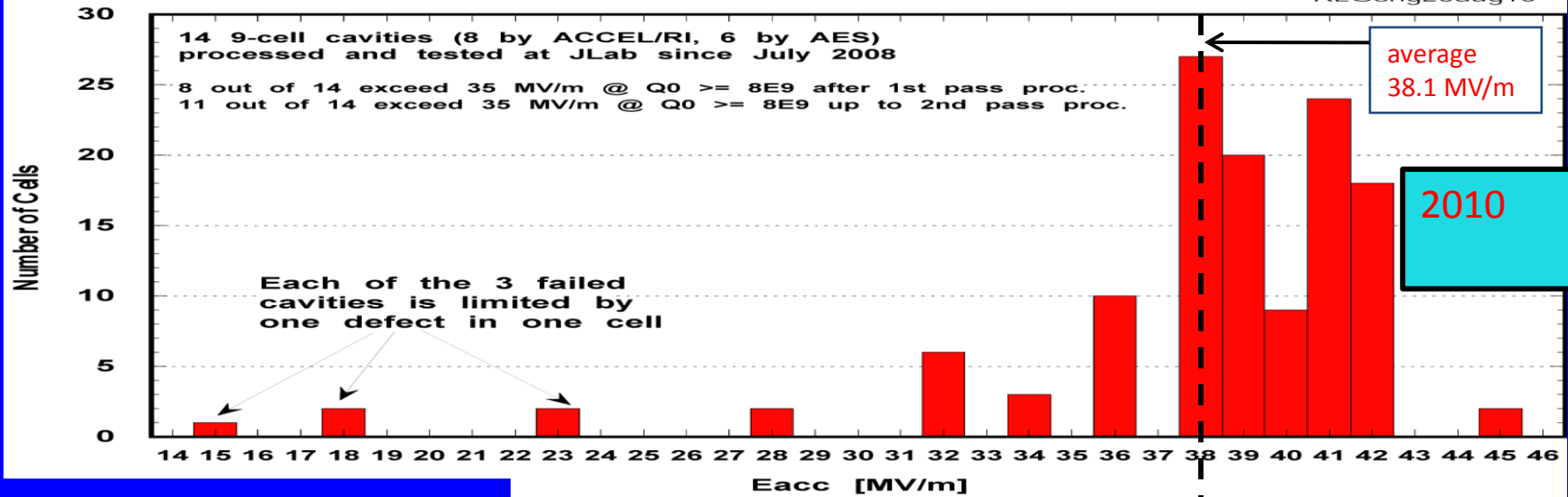
Gradient Yield of 10 ILC Cavities Built by One Vendor Processed and Tested at JLab since July 2008



RLGeng19oct10

Gradient Reached by Each Cell

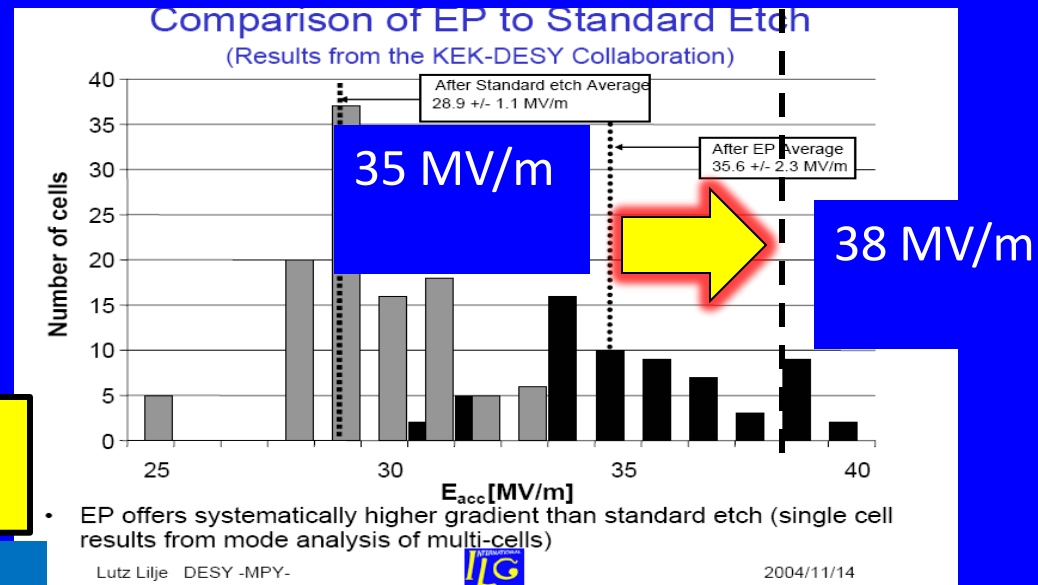
RLGeng25aug10



Gradient
State-of-the-art
then and now

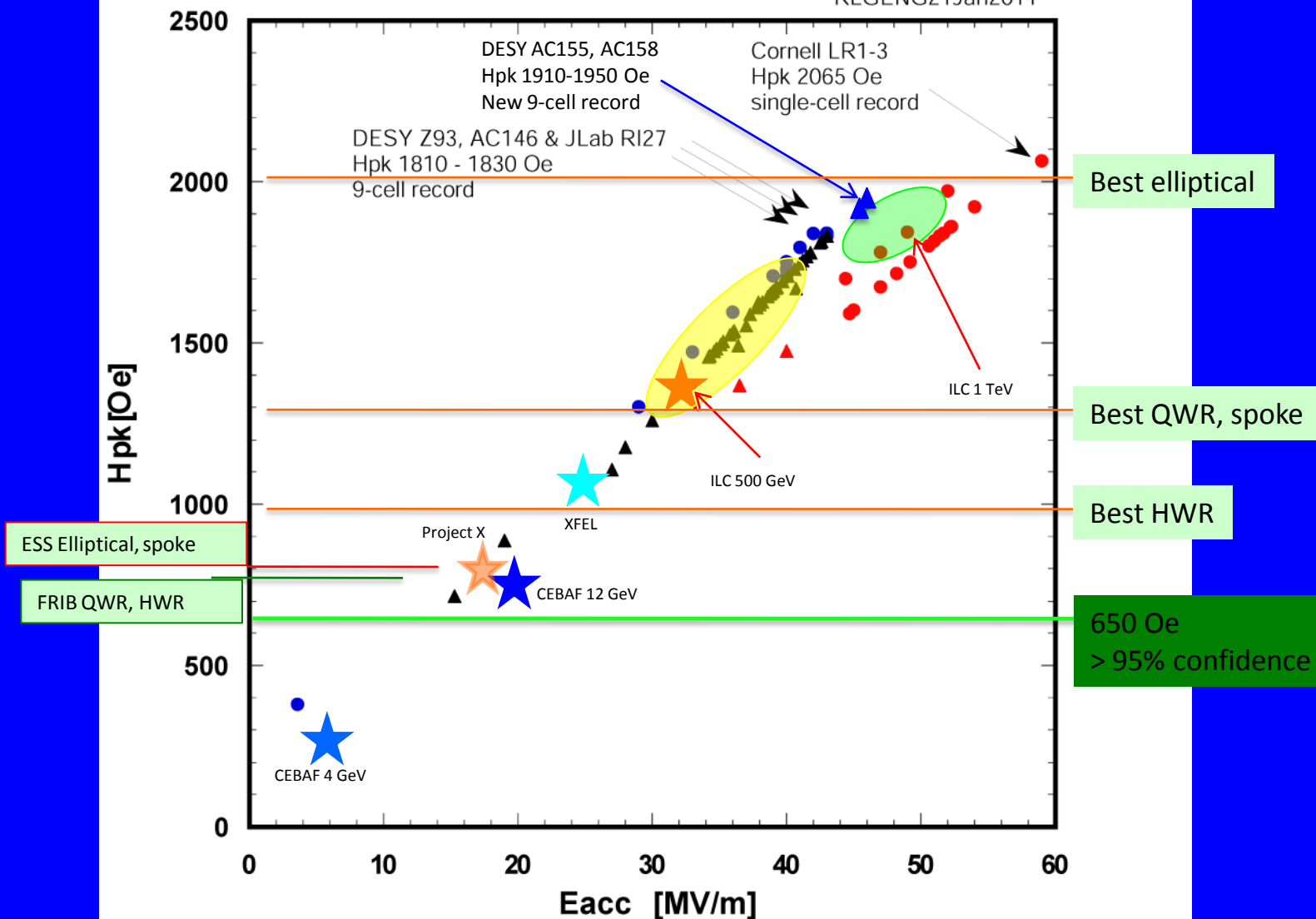
2004

2004 DESY EP 9-cell cavities
Gradient distribution in cells from
pass-band measurements (~ 8 cavities)



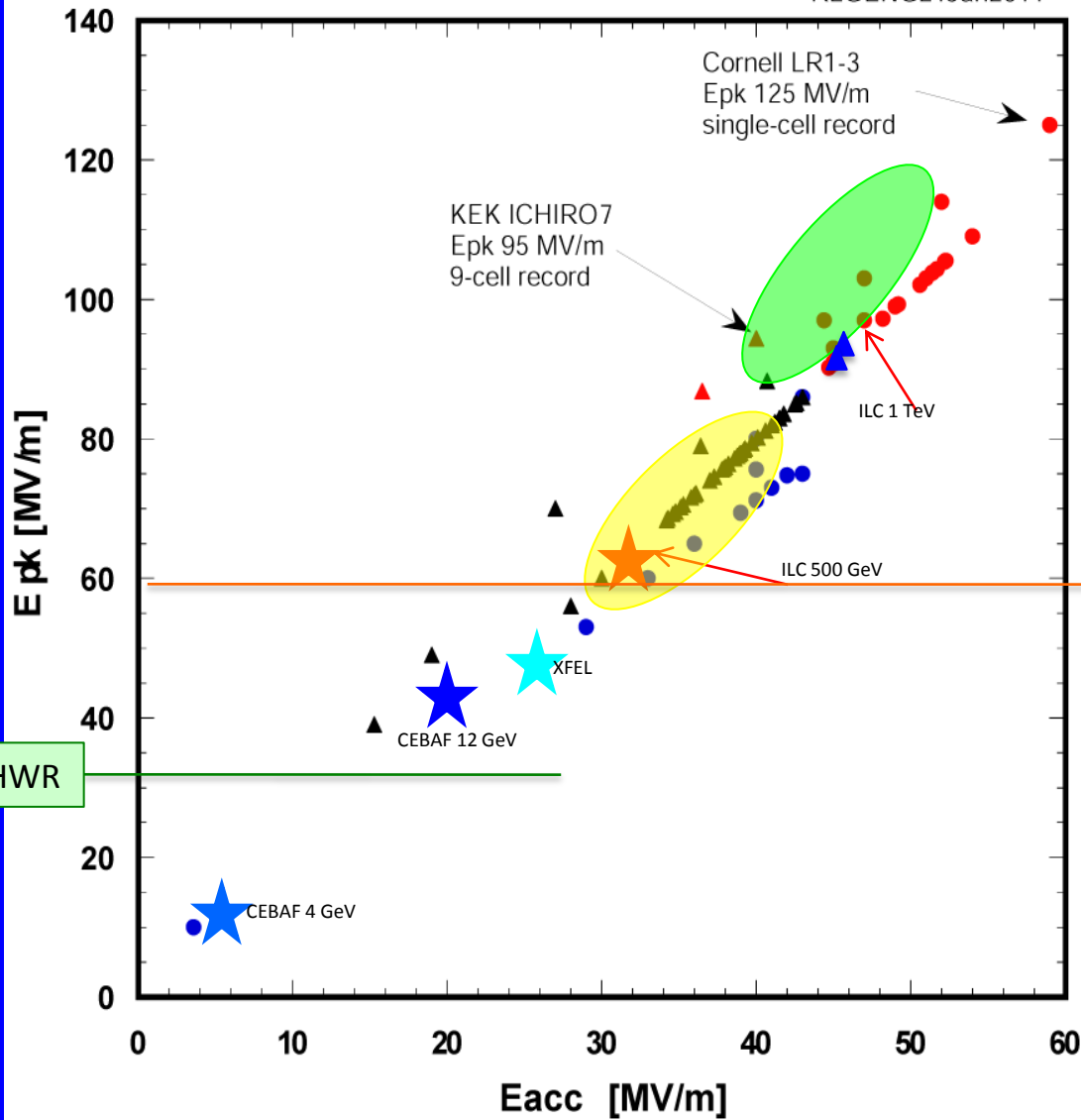
Achieved Peak Surface Magnetic Field in L-band SRF Niobium Cavities (Circle: Single-Cell Cavity; Triangle: Multi-Cell Cavity)

RLGENG21Jan2011



Achieved Peak Surface Electric Field in L-band SRF Niobium Cavities (Circle: Single-Cell Cavity; Triangle: Multi-Cell Cavity)

RLGENG21Jan2011



FRIB QWR, HWR

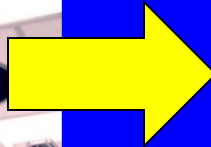
Best QWR, HWR

SRF Cavity Gradient R&D Impacts & Benefits



1991 (BCP)

Cryomodule cavity 5 MV/m
320 cavities installed in linacs

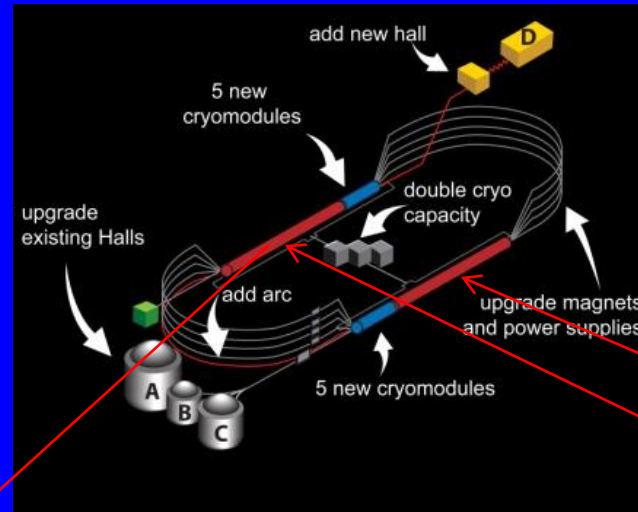


2011 (BCP+HTA+EP+HPR+LTB)

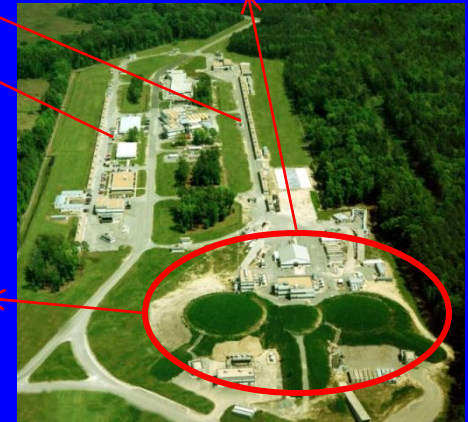
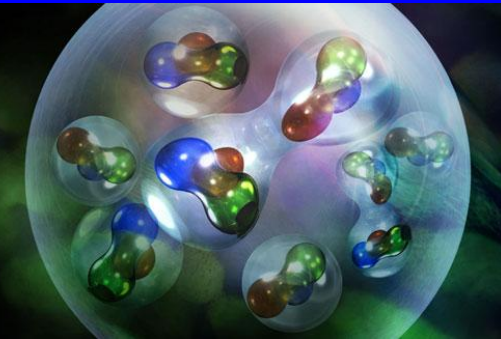
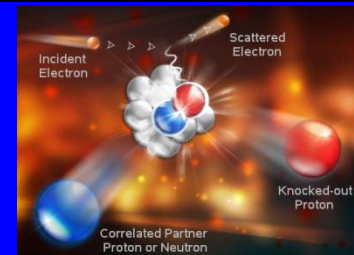
Cryomodule cavity 19.2 MV/m
80 cavities needed, many qualified to
>25 MV/m, some up to 35-43 MV/m



40 old cryomodules two linacs
in operation since 1993

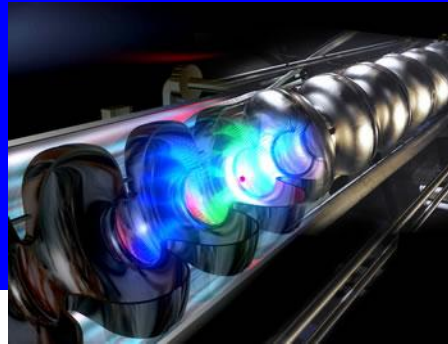
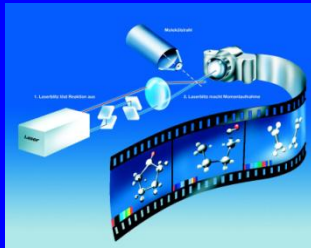
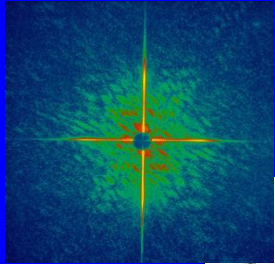


**CW
SRF Linac**

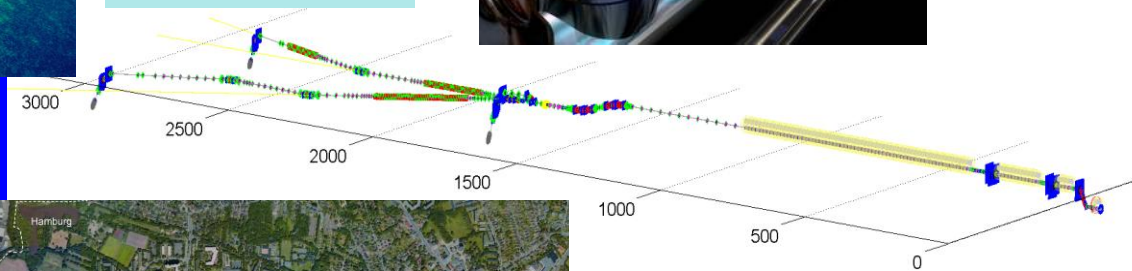


CEBAF upgrade, under construction now, will double its energy to 12 GeV. The present 6 GeV machine has 42 old cryomodules. The additional 6 GeV is achieved by adding only 10 new modules with high gradient cavities.

SRF Cavity Gradient R&D Impacts & Benefits



2011 (EP+HTA+EP/BCP+HPR+LTB)
Cryomodule cavity 24.3 MV/m
640 cavities needed
DESY qualified many cavities up to 35-43 MV/m



*Pulsed
SRF Linac*

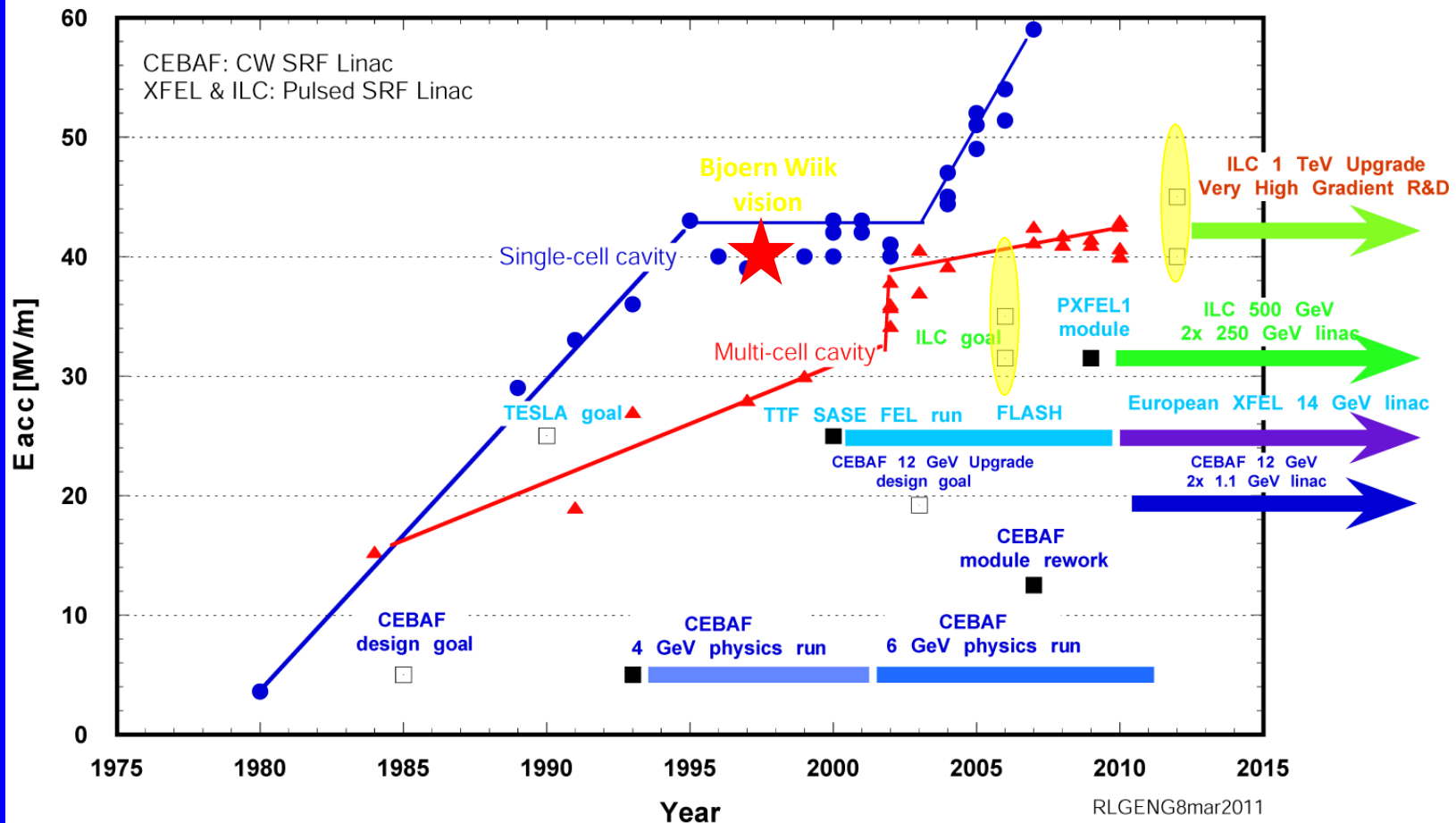


80 cryomodules needed

As a result of DESY's TTF experience and FLASH operation, European XFEL, under construction now, will reach 14 GeV with 640 high gradient cavities.

SRF Cavity Gradient Progress

L-Band SRF Niobium Cavity Gradient Envelope and Gradient R&D Impact to SRF Linacs



Steady progress in SRF cavity gradient makes SRF an enabling technology
SRF based electron linacs (CW & pulsed) have track record of successful operations

Conclusion

- High gradient SRF cavity R&D at JLab a success
 - Defended ILC design gradient choice
 - Built a technical base of high gradient expertise at JLab
 - Provided direct benefit to CEBAF upgrade project
 - Final surface processing of upgrade cavities
 - Validated the first US industrial vendor for high performance SRF cavity fabrication
- Our understanding of gradient limiting mechanisms including quench limit and field emission is much improved by instrumented cavity testing and cryogenic temperatures.
- The program provided a unique opportunity for JLab SRF workers, scientists, technicians and students, to advance SRF science and technology.