Electromagnetic Calorimeters

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JLab Summer Detector/Computer Lectures https://userweb.jlab.org/~gen/talks/calor_ lect_3.pdf

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Electromagnetic Calorimeters



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 - Calorimeters
 - Generic calorimeter
 - Light collecting calorimeters
- 3 Front-End Electronics
- Procedures

5 Summary



Introduction



Calorimeters

- Generic calorimeter
- Light collecting calorimeters

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What is a calorimeter?

Particle detection main goal: measure 3-momenta \vec{P}



"True" Calorimeters

"True" calorimeters measure the temperature change of the absorber: $\Delta T = \frac{E_0}{c \cdot M} \sim \frac{1 \cdot 10^{10} eV \cdot 1.6 \cdot 10^{-19} J/eV}{10^3 J/kg/K \cdot 1kg} \approx 10^{-12} K$ too low!

- High particle flux
 - History: W. Orthmann 1 μ W sensitivity; 1930, with L. Meitner they measured the mean energy of β from ²¹⁰Bi (6% accuracy) \Rightarrow W.Pauli's neutrino hypothesis.
 - bypothesis.
 Precise beam current measurements (SLAC-1970s, JLab-2003)
- Ultra-cold temperatures (low C), superconductivity new detectors for exotic particle search, like "dark matter" candidates.



"Pseudo" Calorimeters

"Pseudo" calorimeters detect $\mathcal{O}(E_o)$: ionization, Cherenkov light

- History: N.L. Grigorov 1954 idea, 1957 implementation in cosmic ray studies (Pamir, 3900 m). Layers of an absorber and layers of proportional counters counting the number of particles in the shower (calibration needed).
- Starting in 1960s revolution in compact electronics \Rightarrow affordable ADC (Analog-to-Digital Converters). New accelerators various types of calorimeters with $\sim 10 \rightarrow 10^5$ ADC channels.

Applications

- detecting neutrals
- good energy resolution at high energies
- fast signals for trigger
- particle identification (e^{\pm}/h)

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e^{\pm} interactions



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γ interactions

Interaction in medium

- Pair production $\gamma Z \rightarrow e^+ e^- Z (K_N)$
- Pair production $\gamma e^- \rightarrow e^+ e^- e^- (K_e)$
- Compton scattering $\gamma e^- \rightarrow \gamma e^- (\sigma_{incoherent})$
- Rayleigh scattering $(\sigma_{coherent})$
- Photonuclear absorption (σ_{nuc})
- Atomic photoeffect (σ_{p.e.})





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Scaling of Material Properties

Radiation length

 X_0 - the material thickness for a certain rate of EM:

$$\mathbf{e}^{\pm} \colon \frac{dE_{loss}}{dx} \simeq \frac{E}{X_0}$$
$$\gamma \colon \lambda_{e^+e^-} \simeq \frac{9}{7} \cdot X_0$$

Derived from EM calculations: $X_0 \simeq \frac{716 \ g \cdot cm^{-2} \cdot A}{Z(Z+1) \cdot ln(287/\sqrt{Z})}$

Critical Energy

 $E_{c}: \text{ cascade stops}$ Losses: lonization = Radiation
B.Rossi: $\frac{dE_{ioniz}}{dx}|_{E_{c}} \simeq \frac{E}{X_{0}}$ $E_{c} \simeq \frac{610(710) \text{ MeV}}{Z+1.24(0.92)} \text{ solids(gasses)}$

+ Solids • Gases

Li Be B CNO Ne

10 20

Z

8

Fe Sn

50 100

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Electromagnetic Showers

Photons and light charged particles (e^{\pm}) interact with matter:

- electrons radiate $e^{\pm} \rightarrow e^{\pm}\gamma$
- photons convert $\gamma \rightarrow e^+e^-$

A cascade develops till the energy of the particles go below a certain limit.

The charged particles of the cascade (e^{\pm}) leave detectable signals.



Electromagnetic Shower: longitudinal development

Scaling variables:

$$t = \frac{x}{X_0}$$
 $y = \frac{E}{E_c}$

Simple model

A simple example of a cascade: ×2 at $\Delta t = 1$. $E(t) = \frac{E_0}{2^t} \Rightarrow t_{max} = ln \frac{E_0}{E_c} / ln 2$ $t_{max} \propto ln(\frac{E_0}{E_c})$ Detectable signal: $L_{charged} \propto E_0 / E_c$

Simulation: EGS4, GEANT





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Electromagnetic Shower: transverse size

Molière radius: $R_M = \frac{X_0 \cdot 21 MeV}{E_c}$ $R < 2 \cdot R_M$ contains 95% of the shower

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Properties of Materials

	Density	X_0	X_0	λ_I	Molière	E _{crit}	Refr.
Material	g/cm^3	g/cm²	ст	g/cm²	R _M cm	MeV	index
W	19.3	6.5	0.35	185.	0.69	10.6	
Pb	11.3	6.4	0.56	194.	1.22	9.6	
Cu	8.96	13.	1.45	134.	1.15	26.	
AI	2.70	24.	8.9	106.	3.3	56.	
С	2.25	42.	18.8	86.	3.5	111.	
Plastic	1.0	44.	42.	82.	6.1		1.58
H ₂	0.07	61.	860.	50.	50.	360.	



Outline





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- Generic calorimeter
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Generic Calorimeter

A matrix of separate elements:



Measured:

- A_i measured amplitudes
- $-\alpha_i$ calibration factors
 - (slow variation)
- $-\dot{x}_i|y_i$ module coordinates

 $E = \sum_{i \in k \times k} \mathcal{E}_i$

Typically k = 3, 5 $\mathcal{E}_i = \alpha_i \cdot A_i$ $x | y = f(..., x_i | y_i, E_i, ..)$ $\vec{X}_0 \Rightarrow$ direction

Important parameters

- Energy resolution $\frac{\sigma E}{E}$
- Linearity
- Coordinate resolution σx
- Timing resolution
- Stability
- Specific requirements:

radiation hardness. mag. field

Cost

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Generic Calorimeter



Important procedures

- Calibration: A_i measured
 → E_i = α_i · A_i.
 - α_i have to be measured using particles of known energies.
- Monitoring of the calibration factors α_i using detector response to a simple excitation (ex: light from a stable source).



Homogeneous and Sampling Calorimeters

Consider: EM shower in plastic scintillator Needed length $\sim 15 \cdot X_0 = 600 \ cm$ - not practical!

	Sampling calorimeters
Homogeneous calorimeters (EM)	Heavy material absorber and the
Heavy active material, no passive absorberBest energy resolutionHigher cost	Features: • Compact • Relatively cheap • Sampling fluctuations \Rightarrow impact on $\frac{\sigma E}{E}$



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Resolutions

Energy resolution

$$\frac{\sigma E}{E} = \alpha \oplus \frac{\beta}{\sqrt{E}} \oplus \frac{\gamma}{E}$$

- α constant term (calibration)
- β stochastic term (signal/shower fluctuations)

• γ - noise

Spatial resolution

$$\sigma \mathbf{X} = \alpha_1 \oplus \frac{\beta_1}{\sqrt{E}}$$



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Energy resolution

- Fluctuations of the track length (EM): $\frac{\sigma E}{E} \simeq \frac{0.005}{\sqrt{E}}$
- Statistics of the observed signal (EM): $\frac{\sigma E}{E} > \frac{0.01}{\sqrt{E}}$
- Sampling fluctuations (EM): ^{σE}/_E ≃ ^{√E}/_{√E}, where *t* is the layer thickness in X₀ (B.Rossi),

$$\sim rac{0.1 \cdot \sqrt{t}}{\sqrt{E}}$$
 for lead absorber ($t > 0.2$)

- Noise, pedestal fluctuations $\frac{\sigma E}{E} < \frac{0.01}{E}$
- Calibration drifts $\frac{\sigma E}{F} \sim 0.01$ for a large detector
- Other ...



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Spacial resolution

- Module lateral size < shower size
- Calculating the shower centroid
- EM: $\sigma x > 0.05 \cdot R_M$



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Light Collecting Homogeneous EM Calorimeters

Heavy transparent materials (small X_0) are preferable \Rightarrow compact, larger signal

- Heavy crystal scintillators: Nal, Csl, BGO, PbW etc: high light yield ⇒ good resolution, expensive
- Heavy crystal Cherenkov detectors: PbF, etc: compact, radiation hard
- Lead glass (SiO \rightarrow PbO) Cherenkov detectors: medium performance, affordable



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Light Collecting Sampling EM Calorimeters

Heavy absorber (Pb,Cu,W...) and a scintillator (plastic) or Cherenkov radiator (quartz fibers ...). Problem: how to collect the light? The most popular solutions for this moment:

- SPACAL (Pb, sc. fibers). The fibers can be bundled to the PM. Very good resolution. Difficult to manufacture.
- Sandwich with WLS fibers crossing through ("shashlik"). The fibers are bundled to the PM. Good resolution. Easy to build.



Timing resolution:

- Scintillation time
- Photodetector time Typically
- au(90%) \sim 50 ns



Light Detectors

Photomultiplier Tubes (PMT)

A vacuum vessel with a photocathode and a set of electrodes (dynodes) for electron multiplication.

- Very high gain $\sim 10^5-10^7$
- Very low electronic noise
- Size: diameter 2-40 cm
- Slow drift of the gain
- Sensitive to the magnetic field
- Relatively low QE~20%
- Radiation hard

Avalanche Photodiods (APD)

A silicon diod in avalanche mode and an electronic amplifier

- Gain $\sim 50 300$
- • High electronic noise
- Siže: 1 × 2 cm²
- Very sensitve to the bias voltage
- Not sensitive to the magnetic field
- High QE~75% at 430 nm
- Temperature sensitive -2%/K
- Radiation hardness may be a problem



Detector technology: Silicon Photomultiplier (SiPM)



New Popular Photosensor: planned to be used at many projects instead of PMTs

- Matrix of pixels: APDs in the Geiger mode
- PDE (QE \times packing factor) \sim 20%
- Gain $\sim 10^6$
- Immune to magnetic field
- Timing resolution <100 ps



- Noise (temperature dependent)
- Small size, now <12×12 mm²
- Limited range of the gain
- Non-linearity
- Radiation hardness

Extensive study of Rad. hardness

- Affected by neutron radiation
- Noise increase \propto eff. fluence
- No other serious effects
- Self-annealing: a factor of 0.5
- Self-annealing better at higher temperature



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Crystals in big experiments





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EM calorimeters with optical readout

	Density	X ₀	R _M	λ_I	Refr.	τ	Peak	Light	Np.e. GeV	rad	$\frac{\sigma E}{E}$	
Material	g/cm ³	ст	ст	ст	index	ns	λnm	yield				
Crystals												
Nal(TI)**	3.67	2.59	4.5	41.4	1.85	250	410	1.00	10 ⁶	10 ²	1.5%/E ^{1/4}	
Csl *	4.53	1.85	3.8	36.5	1.80	30	420	0.05	104	104	$2.0\%/E^{1/2}$	
CsI(TI)*	4.53	1.85	3.8	36.5	1.80	1200	550	0.40	10 ⁶	10 ³	1.5%/E ^{1/2}	
BGO	7.13	1.12	2.4	22.0	2.20	300	480	0.15	10 ⁵	10 ³	$2.\%/E^{1/2}$	
PbWO ₄	8.28	0.89	2.2	22.4	2.30	5/39%	420	0.013	104	10 ⁶	$2.0\%/E^{1/2}$	
						15/60%	440					
						100/01%						
LSO	7.40	1.14	2.3		1.81	40	440	0.7	10 ⁶	10 ⁶	$1.5\%/E^{1/2}$	
PbF ₂	7.77	0.93	2.2		1.82	Cher	Cher	0.001	10 ³	10 ⁶	$3.5\%/E^{1/2}$	
					Lea	d glass						
TF1	3.86	2.74	4.7		1.647	Cher	Cher	0.001	10 ³	10 ³	$5.0\%/E^{1/2}$	
SF-5	4.08	2.54	4.3	21.4	1.673	Cher	Cher	0.001	10 ³	10 ³	$5.0\%/E^{1/2}$	
SF57	5.51	1.54	2.6		1.89	Cher	Cher	0.001	10 ³	10 ³	$5.0\%/E^{1/2}$	
				Sa	mpling: I	ead/scintilla	tor					
SPACAL	5.0	1.6				5	425	0.3	2 · 10 ⁴	10 ⁶	$6.0\%/E^{1/2}$	
Shashlyk	5.0	1.6				5	425	0.3	10 ³	10 ⁶	10.%/ <i>E</i> ^{1/2}	
Shashlyk(K)	2.8	3.5	6.0			5	425	0.3	4 · 10 ⁵	10 ⁵	$3.5\%/E^{1/2}$	
* - hygros	scopic											

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Summary

Crystal Ball (SLAC, DESY)



- \sim 600 Nal crystals
- γ detection
- Charmonia spectra

 \Rightarrow QCD tune!



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KTeV (FNAL)



• 3256 CsI crystals • $\pi^{\circ} \rightarrow \gamma \gamma$ detection • $\sigma E/E \approx 2.0\% \sqrt{E} + 0.5\%$

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BaBar (SLAC)





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PANDA at GSI - under construction



- \sim 15000 PbWO₄ crystals
- New APD $1 \times 1 \text{ cm}^2$



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SpaCal (CERN, Frascatti, JLab)

scintillating fibers / lead matrix



- Fibers/lead 50% / 50% in volume
- X_o = 1.2 cm
- 5 g/cm³

- CERN original R&D
- KLOE (DAFNE) 5000 PMTs
- KLOE $\sigma E/E \approx 5.7\%/E^{1/2}$
- KLOE $\sigma \tau \approx 50/E^{1/2} + 50 \text{ ps}$



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SpaCal: Barrel Calorimeter in Hall D





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SpaCal: Barrel Calorimeter in Hall D

Built at Regina

- Lead swaging (grooves)
- Glue lead and fibers layer by layer
- Out and polish







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Barrel Calorimeter - Construction of Modules



- all 48 modules built
- very regular matrix

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a ruler at the opposite side

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Shashlyk: Experiment KOPIO





• $\sigma E/E \approx 2.0 \oplus 3.0\%/E^{1/2}$ • $\sigma \tau \approx 70/E^{1/2} \oplus 14/E$ ps

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Front-End Electronics

Requirements

- Resolution $\sim 10^{-3}$
- Dynamic range > 10²: needed to measure the shower profile and the coordinates
- Differential linearity <1%
- Digitization speed (>1 MHz)
- Readout speed (>100 kHz)
- Cost

Existing generic solutions

- Charge integrating ADC
- Flash ADC
- Combinations (pipeline ADC)



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Charge Integrating ADC



- Many products on the market
- Precise: 12-15 bits
- Gate must come in time \Rightarrow long

(>300-500 ns) delay for each channel is needed (cables)

- Slow conversion time > 10 μ s \Rightarrow not suitable for trigger logic
- Problems at very high rate:

pileup, deadtime

Pedestal



Integrating ADC



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Flash ADC



- Cost ×10 of the QDC (250 MHz, 12 bits)
- Huge memory buffers needed
- Resolution n bits $\Rightarrow 2^n$ comparators
- Pipeline readout no dead time
- No delay cables needed
- Pileup can be partially resolved
- Timing resolution without extra discr.& TDCs
- FPGA computing trigger logic
- Became the mainstream



In	tr	n	а		\sim	11	\sim	n
		U	u	u	0	u	U	

Procedures

Calibration

- The detector has to be calibrated at least once.
 - Test beam
 - Better: in-situ, using an appropriate process:
 - e^+e^- collider: Bhabha scattering $e^+e^- \rightarrow e^+e^-$, $e^+e^- \rightarrow e^+e^-\gamma$
 - LHC: Z→e⁺e^{-/} (1 Hz at low luminocity)
 - h+h $\rightarrow \pi^0$ +X, $\pi^0 \rightarrow \gamma \gamma$
 - RCS (JLab): e[−]p→e[−]p

Procedure: for event n:

$$\mathcal{E}^{(n)} = \sum_{i \in k \times k} \alpha_i \cdot \mathbf{A}_i^{(n)}$$

$$\chi^2 = \sum_n (E^{(n)} - \sum_{i \in k \times k} \alpha_i \cdot A_i^{(n)}) / \sigma_n$$

- System of linear equations
- \Rightarrow *N* × *N* matrix nearly diagonal
- Easy to solve



Monitoring

Instabilities:

- All avalanche-type devices tend to drift (PMT, gas amplification ...)
- Optical components may lose transparency
- Temperature dependence
- Many other sources of instability ...

Calibration is typically done once per many days of running \Rightarrow signal monitoring in between is needed.



Procedures

Summary

Light collecting devices



- Stable pulsed light source:
 - Xe flash lamp: 1% stability, >100 ns pulse
 - Laser: 2-5% stability, ≪1 ns pulse
 - LED: 1-3% stability in thermostate, >30 ns pulse
- Usually the light source has to be monitored
- Light distribution
- Material transparency: not easy to monitor (λ-dependence)
- Scintillation yield no monitoring this way



Summary

Calorimeters are used for:

- Detecting neutrals
- Energy and coordinate measurements
- Trigger
- Separation of hadrons against e^{\pm} , γ and muons

The calorimeters are of increasing importance with higher energies. They became the most important/expensive/large detectors in the current big projects (LHC etc).



Procedures

Summary (continued)

There are various techniques to build calorimeters for different resolution, price, radiation hardness and other requirements.

The typical energy resolutions are:

- EM: from $\frac{\sigma E}{E} \sim \frac{2\%}{\sqrt{E}} \oplus 0.3\%$ for scintillating crystals to about $\frac{\sigma E}{E} \sim \frac{10\%}{\sqrt{E}} \oplus 0.8\%$ for sampling calorimeters.
- HD calorimeters: $\frac{\sigma E}{E} \sim \frac{30-50\%}{\sqrt{E}} \oplus 3\%$

The coordinate resolutions could be about 1-3 mm for EM calorimeters and 20-30 mm for HD ones.

