Electrons for Neutrinos: How electron scattering data can improve oscillation experiments

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Collaboration

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Neutrino Interaction

- $\begin{array}{c} CC \ QE \\ \hline v_l \\ \hline v_l \\ \hline v_l + n \rightarrow l^- + p \\ \hline \overline{v}_l + p \rightarrow l^+ + n \end{array}$
- Weak eigenstates ≠ mass eigenstates
- Neutrino mixing PNMS (Pontecorvo-Maki-Nakagawa-Sakata) matrix

$$egin{bmatrix}
u_e \
u_\mu \
u_ au \end{bmatrix} = egin{bmatrix} U_{e1} & U_{e2} & U_{e3} \ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \ U_{ au 1} & U_{ au 2} & U_{ au 3} \end{bmatrix} egin{bmatrix}
u_1 \
u_2 \
u_2 \
u_3 \end{bmatrix}$$

Oscillations

Weak interaction

Neutrino Oscillations



Incoming Energy Reconstruction



Cherenkov detectors (T2K,..)

- Assuming Quasielastic (QE) interaction
- Using solely the final state lepton

$$E_{QE} = \frac{2M\epsilon + 2ME_l - m_l^2}{2(M - E_l + |k_l|\cos\theta)}$$



Tracking detectors (DUNE,..)

- Detect outgoing hadron and lepton
- Need good hadronic resolution

$$E_{\rm cal} = E_l + E_p^{\rm kin} + \epsilon$$

for (e,e'p)

$\epsilon\,$ is the nucleon separation energy $^{\sim}\,20~MeV$

F. Hauenstein, SESAPS

Osciallations need E_v reconstruction



6

Oscillations need E_v reconstruction

$$P(v_{\mu} \to v_{x}) = \sin^{2}(2\theta) \times \sin^{2}\left(\frac{\Delta m^{2}L}{4E_{\nu}}\right)$$



Oscillations need E_v reconstruction

$$P(v_{\mu} \to v_{x}) = \sin^{2}(2\theta) \times \sin^{2}\left(\frac{\Delta m^{2}L}{4 E_{v} real}\right)$$



Energy reconstruction

Wide neutrino
 beam distribution



 Reconstruction requires knowledge of the nuclear interactions -> event generators



Systematic Effects by Nuclear Models



Events generated with GiBUU, reconstruct with Genie

Events generated with GiBUU, reconstruct with GiBUU

→ Imperfect event generators
→ systematic errors!

P. Coloma et. al, Phys. Rev. D 89, 073015 (2014)

F. Hauenstein, ODU, SESAPS

Use electron scattering to improve models

- Known incident energy
- High intensity
- Similar interaction with nuclei
 - Single boson exchange
 - CC Weak current [vector plus axial] e⁻–N QE

•
$$j_{\mu}^{\pm} = \overline{u} \frac{-ig_W}{2\sqrt{2}} (\gamma^{\mu} - \gamma^{\mu}\gamma^5) u$$

- EM current [vector]
 - $j^{em}_{\mu} = \bar{u} \gamma^{\mu} u$
- Same nuclear physics





Nuclear Physics





Nuclear Physics



Plus pion production ...



- Large amount of electron scattering data available
- Known beam energy
- Analyze electron data as if it is "Neutrino data"
 - Select specific interaction (e,e') or (e,e'p)
 - Scale electron data using Mott cross section
 - ➤ Test energy reconstruction
 - Compare with event generators
 - > As a start focus on quasi-elastic (QE) scattering

CLAS6 at Jefferson Lab

- Large (~ 2π) acceptance
- Open electron trigger
- Charged hadron threshold
 - $-P_{p} > 300 \text{ MeV/c}$
 - $-P_{\pi^{+/-}} > 150 \text{ MeV/c}$



CLAS6 coverage









CLAS6 E2a Data

Targets:

- ³He
- ⁴He
- ¹²C
- ⁵⁶Fe

Energies:

- 1.1 GeV
- 2.2 GeV
- 4.4 GeV



Reconstructing the initial energy

- Choose 0π events to enhance the Quasielastic (QE) sample
 - Subtract "undetected pions"
- Reconstruct the incident lepton energy:

$$-E_{QE} = \frac{2M_N\epsilon + 2M_NE_l - m_l^2}{2(M_N - E_l + k_l\cos\theta_l)}$$

- ϵ : nucleon separation energy, M_N nucleon mass
- $\{m_l, E_l, k_l, \theta_l\}$ scattered lepton mass, energy, momentum and angle
- broadened by nucleon fermi motion

$$-E_{cal} = E_e + T_p + \epsilon$$
 [for (e,e'p)]

Reconstructing the Incident Energy



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Reconstruction Incoming Energy worse for higher masses



Reconstruction Incoming Energy worse for higher beam energies ⁵⁶Fe



²²

Energy reconstruction dependence on P^{\perp}



Event Generators

- Bad reconstruction is OK if the generator describes reality
 - ightarrow Can be checked by comparing with data
- Several generators used in neutrino experiments
 - Neut (T2K)
 - NuWro (MicroBoone, Minerva)
 - GiBUU (KM3Net, ...)
 - GENIE (MicroBoone, Minerva, DUNE,...)



Data-Genie Comparisons



Data-Genie Comparisons - QE peak

C(e,e'p) 2.26 GeV, 0.8 < x < 1.2





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Peaks in same location

Summary Data-Genie: E_{beam} Reconstruction

Fe	e⁻ Data	ν GENIE
2.26 GeV	26%	62%
4.46 GeV	14%	62%

Fraction of Fe(e, e'p) events with E_{Cal} within 5% of E_{beam}

Possible effect on DUNE Oscillation



- (Chris Marshall, LBNL)
- Proof of principle to show potential impact
- Threw events with vA Genie
 - Reconstructed with vA Neut or eA data
- Compared E_{rec} for eA to E_{rec} for vA
- Used 2.26 GeV eA E_{rec} for all incident energies



CLAS12

- ~ 4π acceptance
- forward detector (8 40°)
 - Toroidal magnetic field
- Hermetic central detector (40 – 135°)
 - 5 T solenoidal field
- Backward Angle Neutron Detector
- x10 larger luminosity than CLAS 6



- 45 beam days approved with an A rating for
 - 1.1, 2.2, 4.4, and 6.6 GeV beam energies
 - d, He, C, O, Ar, and Sn targets



- Electron scattering can contribute dramatically to neutrino experiments
 - Similar physics
 - Lots of data available
 - Lots more to come
- Worse energy reconstruction for higher beam momenta and larger p_T
- Disagreement between data and GENIE
- Impact on neutrino oscillation parameters



Backup

Data-Genie Comparisons - E_{beam} Reconstruction

$$E_{cal} = E_e + T_p + \epsilon$$

Background Subtraction

Non-QE interactions lead to multi hadron final states.

Gaps in CLAS acceptance will make them look like (e,e'p) events.

Data Driven Correction:

- 1. Use measured (e,e'p π) events,
- 2. Rotate π around q to determine its acceptance,
- 3. Subtract (e,e'p π) contributions

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- 1. Use measured (e,e'p π) events,
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- 4. Do the same for 2p, 3p, 2p+ π etc.

Background Subtraction

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We're Also Improving Genie

- 1. Corrected expression for Mott cross section in QE
- 2. MEC/2p2h
 - 1. Added boost back to lab frame
 - 2. Corrected mass for cluster of particles
 - 3. Corrected Form Factors
- 3. Resonance
 - 1. Replaced old calculation with GSL Minimizer (now gives correct peak location)
 - 2. Switched to Berger-Seghal model
 - 3. Used corrected coupling constant for EM interactions
- 4. Nucleon momentum distributions
 - 1. Switched to Local Fermi Gas Model

Beginning work on NuWro and GiBUU.

Consulting with the relevant experts on each code.