Selected Topics in Hadron Spectroscopy

Mathematical Techniques

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http://cern.ch/suchung/ http://www.phy.bnl.gov/~e852/reviews.html

Covariant Helicity-coupling Amplitudes

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A general formulation is given for constructing covariant helicity-coupling amplitudes involving two-body decays with arbitrary integer spins. The decay amplitudes are given exclusively in terms of both definite orbital angular momentum and total intrinsic spin. A systematic method is developed for calculating the energy and momentum dependence of daughter particles in the decay amplitudes, and a general formula for arbitrary integer spins is given.

If a daughter particle has spin 1 or higher, the helicity-coupling amplitudes depend in general on the Lorentz factor $\gamma = E/m$, where *m* is the mass of the daughter and *E* is its energy in the parent rest frame. A significant simplification results with the exclusive use of spin tensors and momenta defined along the helicity axis for the daughter states, which is defined to be along the original *z*-axis. This technique separates out the angular distribution contained in the *D* function from the problem of finding a proper energy and momentum dependence of the helicity-coupling amplitudes.

Two-body decay: $J \rightarrow s + \sigma$

	Parent	Daughter 1	Daughter 2
Spin	J	8	σ
Parity	η_J	η_{s}	η_{σ}
4-momentum	p = (w; 0, 0, 0)	$\boldsymbol{q}=(\boldsymbol{q}_0;0,0,\boldsymbol{q})$	$\boldsymbol{k}=(k_0;0,0,-k)$
Energy	$p_0 = w$	$q^{}_0$	k_{0}
Mass	w	m	μ
Energy/Mass	1	$\gamma_s = q_0/m$	$\gamma_\sigma=k_0/\mu$
Velocity	0	eta_s	eta_{σ}
Helicity	$\lambda - u$	λ	u
γeta	0	$\gamma_s \beta_s = q/m$	$\gamma_\sigma eta_\sigma = k/\mu$
Wave function	$\phi^*(\lambda- u)$	$\omega(\lambda)$	arepsilon(- u)

where

 $\delta = \lambda - \nu$ and p = q + k, $r = q - k = (q_0 - k_0; 0, 0, 2q)$ in JRF

Consider a state with spin(parity)= $J(\eta_J)$ decaying into two states with $s(\eta_s)$ and $\sigma(\eta_\sigma)$. The decay amplitudes (helicity formalism) are given, in the rest frame of J,

(1)
$$\mathcal{M}^{J}_{\lambda\nu}(\theta,\phi,M) = \sqrt{\frac{2J+1}{4\pi}} D^{J*}_{M\delta}(\phi,\theta,0) F^{J}_{\lambda\nu}$$

If the angles are zero, then we have

(2)
$$\mathcal{M}^J_{\lambda\nu}(0,0;\delta) = \sqrt{\frac{2J+1}{4\pi}} F^J_{\lambda\nu}$$

Equivalently, the decay amplitude (canonical formalism) can be written

(3)
$$\mathcal{M}^{J}_{\ell S}(\theta,\phi;m_1,\,m_2;\,M) = G^{J}_{\ell S}(sm_1\,\sigma m_2|Sm_s)\sum_m (\ell m\,Sm_s|JM)\,Y^{\ell}_m(\theta,\phi)$$

Again, if the angles are zero, we find

(4)
$$\mathcal{M}^{J}_{\ell S}(0,0;m_1,m_2;m_s) = \sqrt{\frac{2\ell+1}{4\pi}} G^{J}_{\ell S}(sm_1\sigma m_2|Sm_s) (\ell 0 Sm_s|Jm_s)$$

Here we must have $m_1 = \lambda$, $m_2 = -\nu$ and $m_s = \lambda - \nu = \delta$, so that

(5)
$$\mathcal{M}^{J}_{\ell S}(0,0;\lambda,-\nu;\delta) = \sqrt{\frac{2\ell+1}{4\pi}} G^{J}_{\ell S}\left(\ell 0 S\delta|J\delta\right) \left(s\lambda \sigma -\nu|S\delta\right)$$

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Comparing the two amplitudes (helicity and canonical), we see that

$$(6) \left(\mathcal{M}^{J}_{\lambda \nu} = \sum_{\ell S} \mathcal{M}^{J}_{\ell S} \right) \rightarrow \left[F^{J}_{\lambda \nu} = \sum_{\ell S} \left(\frac{2\ell + 1}{2J + 1} \right)^{1/2} \left(\ell 0 \, S \delta | J \delta \right) \left(s \lambda \, \sigma \, -\nu | S \delta \right) G^{J}_{\ell S} \right]$$

where

(7)
$$\sum_{\lambda \nu} \left| F_{\lambda \nu}^J \right|^2 = \sum_{\ell S} \left| G_{\ell S}^J \right|^2$$

Generalize to relativistic cases:

$$(\ell 0 S \delta | J \delta) G^{J}_{\ell S} \quad \rightarrow \quad g^{J}_{\ell S} A^{J}_{\ell S}(\delta)$$

(8)
$$F_{\lambda\nu}^{J} = \sum_{\ell S} g_{\ell S}^{J} \left(s\lambda \sigma -\nu | S\delta \right) A_{\ell S}^{J}(\delta)$$

and compute $A^J_{\ell S}(\delta)$ in tensor formalism

(9)
$$A^{J}_{\ell S}(\delta) = \left[p^{n}, \psi(S\delta), t^{\ell}(r), \phi^{*}(J\delta)\right]_{w} \text{ or } |J\delta\rangle \to |S\delta\rangle + |\ell 0\rangle$$

where n = 0 or 1 and $[\cdots]_w$ indicates that a Lorentz invariant amplitude is to be contructed out of the four variables indicated, using the metric $\tilde{g}_{\alpha\beta}(w)$ or the totally antisymmetric rank-4 tensor $\epsilon_{\alpha\beta\gamma\delta}$, evaluated in the *J*RF.

The wave functions in a total intrinsic spin S is given by

$$\psi(Sm_s) = \sum_{m_a m_b} (sm_a \sigma m_b | Sm_s) \,\omega(sm_a) \,\varepsilon(\sigma m_b), \quad \psi(S - m_s) = (-)^{S - m_s} \,\psi^*(Sm_s)$$

(10)

(11)

where $\omega(sm_a)$ is a rank-*s* tensor, while $\varepsilon(\sigma m_b)$ is a rank- σ tensor. So $\psi(Sm_s)$ is a rank- $(s + \sigma)$ tensor. Note that $t^{\ell}(r)$ is a rank- ℓ tensor and $\phi(JM)$ is a rank-*J* tensor. Take the example in which both ω and ε refer to spin-1 states. Then, $\psi(Sm_s)$ is a rank-2 tensor given by

$$\begin{cases} \psi^{\alpha\beta}(22) = \omega^{\alpha}(+) \varepsilon^{\beta}(+) \\ \psi^{\alpha\beta}(21) = \frac{1}{\sqrt{2}} \Big[\omega^{\alpha}(+) \varepsilon^{\beta}(0) + \omega^{\alpha}(0) \varepsilon^{\beta}(+) \Big] \\ \psi^{\alpha\beta}(20) = \frac{1}{\sqrt{6}} \Big[\omega^{\alpha}(+) \varepsilon^{\beta}(-) + \omega^{\alpha}(-) \varepsilon^{\beta}(+) + 2\omega^{\alpha}(0) \varepsilon^{\beta}(0) \Big] \\ \psi^{\alpha\beta}(11) = \frac{1}{\sqrt{2}} \Big[\omega^{\alpha}(+) \varepsilon^{\beta}(0) - \omega^{\alpha}(0) \varepsilon^{\beta}(+) \Big] \\ \psi^{\alpha\beta}(10) = \frac{1}{\sqrt{2}} \Big[\omega^{\alpha}(+) \varepsilon^{\beta}(-) - \omega^{\alpha}(-) \varepsilon^{\beta}(+) \Big] \\ \psi^{\alpha\beta}(00) = \frac{1}{\sqrt{3}} \Big[\omega^{\alpha}(+) \varepsilon^{\beta}(-) + \omega^{\alpha}(-) \varepsilon^{\beta}(+) - \omega^{\alpha}(0) \varepsilon^{\beta}(0) \Big] \end{cases}$$

The wave functions in the *J*RF are given by

$$\begin{cases} \phi^{\alpha}(\pm) = \mp \frac{1}{\sqrt{2}} & (& 0; & 1, & \pm i, & 0 &) \\ \phi^{\alpha}(0) = & (& 0; & 0, & 0, & 1 &) \\ \chi^{\alpha}(\pm) = \mp \frac{1}{\sqrt{2}} & (& 0; & 1, & \pm i, & 0 &) \\ \chi^{\alpha}(0) = & (& 0; & 0, & 0, & 1 &) \\ \omega^{\alpha}(\pm) = \mp \frac{1}{\sqrt{2}} & (& 0; & 1, & \pm i, & 0 &) \\ \omega^{\alpha}(0) = & (& \gamma_{s}\beta_{s}; & 0, & 0, & \gamma_{s} &) \\ \varepsilon^{\alpha}(\pm) = \mp \frac{1}{\sqrt{2}} & (& 0; & 1, & \pm i, & 0 &) \\ \varepsilon^{\alpha}(0) = & (& -\gamma_{\sigma}\beta_{\sigma}; & 0, & 0, & \gamma_{\sigma} &) \end{cases}$$

(12)

Note that

(13)
$$p_{\alpha}\phi^{\alpha}(\lambda) = p_{\alpha}\chi^{\alpha}(\lambda) = q_{\alpha}\omega^{\alpha}(\lambda) = k_{\alpha}\varepsilon^{\alpha}(\lambda) = 0$$

for any λ and

14)

$$\phi(-\lambda) = (-)^{\lambda} \phi^{*}(\lambda), \quad \chi(-\lambda) = (-)^{\lambda} \chi^{*}(\lambda),$$

$$\omega(-\lambda) = (-)^{\lambda} \omega^{*}(\lambda), \quad \varepsilon(-\lambda) = (-)^{\lambda} \varepsilon^{*}(\lambda)$$

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These polarization four-vectors satisfy

15)
$$\begin{cases} p_{\alpha}\phi^{\alpha}(m) = 0\\ \phi_{\alpha}^{*}(m)\phi^{\alpha}(m') = -\delta_{mm'}\\ \sum_{m}\phi_{\alpha}(m)\phi_{\beta}^{*}(m) = \tilde{g}_{\alpha\beta}(w) \end{cases}$$

where

(16)
$$\tilde{g}_{\alpha\beta}(w) = -g_{\alpha\beta} + \frac{p_{\alpha}p_{\beta}}{w^2}$$

(17)
$$\tilde{g}_{\alpha\beta}(w) = \tilde{g}^{\alpha\beta}(w) = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

The argument w will be dropped whenever there is no ambiguity; however, one must remember that there exist in the problem two additional \tilde{g} 's, i.e. $\tilde{g}(m)$ and $\tilde{g}(\mu)$ for the states s and σ , respectively. The wave function for orbital angular momenta, in the *J*RF,

Wave Functions for Arbtrary Integer Spin:

The general spin-J wave function can be written

(19)

$$\phi^{\delta_1 \cdots \delta_J} (Jm) = [a^J(m)]^{\frac{1}{2}} \sum_{m_0} 2^{m_0/2} \sum_P \phi^{\alpha_1}(+) \cdots \phi^{\beta_1}(0) \cdots \phi^{\gamma_1}(-) \cdots$$

$$a^J(m) = \frac{(J+m)!(J-m)!}{(2J)!}, \quad \phi(J-m) = (-)^m \phi^*(Jm)$$

where the indices $\{\delta_1 \cdots \delta_J\}$ have been broken up into three distinct sets in the second summation, i.e., $\{\alpha_i\}$ with $(i = 1, m_+)$, $\{\beta_i\}$ with $(i = 1, m_0)$ and $\{\gamma_i\}$ with $(i = 1, m_-)$, where m_{\pm} stands for the numbers of $\phi(\pm)$'s and m_0 for $\phi(0)$'s. Note that

(20)
$$J = m_+ + m_0 + m_-, \quad m = m_+ - m_-$$
 and $2m_\pm = J \pm m - m_0$

 m_0 ranges from 0(1), 2(3),..., to J - m =even(odd). The second sum in $\phi(Jm)$ represents a summation on the permutations

$$\{(+)(+)\cdots(0)(0)\cdots(-)(-)\cdots\}$$

It is seen readily that the number of terms in the summation is given by

(21)
$$b^{j}(m,m_{0}) = \frac{j!}{m_{+}!m_{0}!m_{-}!}$$

For spin-2 wave functions, we have

(22)
$$\begin{cases} \phi_{\alpha\beta}(2,+2) = \phi_{\alpha}(+) \phi_{\beta}(+) \\ \phi_{\alpha\beta}(2,+1) = \frac{1}{\sqrt{2}} \left[\phi_{\alpha}(+) \phi_{\beta}(0) + \phi_{\alpha}(0) \phi_{\beta}(+) \right] \\ \phi_{\alpha\beta}(2,0) = \frac{1}{\sqrt{6}} \left\{ \left[\phi_{\alpha}(+) \phi_{\beta}(-) + \phi_{\alpha}(-) \phi_{\beta}(+) \right] + 2\phi_{\alpha}(0) \phi_{\beta}(0) \right\} \end{cases}$$

and for spin 3,

$$\begin{split} \phi_{\alpha\beta\gamma}(3,+3) &= \phi_{\alpha}(+) \phi_{\beta}(+) \phi_{\gamma}(+) \\ \phi_{\alpha\beta\gamma}(3,+2) &= \frac{1}{\sqrt{3}} \left[\phi_{\alpha}(+) \phi_{\beta}(+) \phi_{\gamma}(0) + \phi_{\alpha}(+) \phi_{\beta}(0) \phi_{\gamma}(+) + \phi_{\alpha}(0) \phi_{\beta}(+) \phi_{\gamma}(+) \right] \\ \phi_{\alpha\beta\gamma}(3,+1) &= \frac{1}{\sqrt{15}} \left[\phi_{\alpha}(+) \phi_{\beta}(+) \phi_{\gamma}(-) + \phi_{\alpha}(+) \phi_{\beta}(-) \phi_{\gamma}(+) + \phi_{\alpha}(-) \phi_{\beta}(+) \phi_{\gamma}(+) \right] \\ &\quad + \frac{2}{\sqrt{15}} \left[\phi_{\alpha}(+) \phi_{\beta}(0) \phi_{\gamma}(0) + \phi_{\alpha}(0) \phi_{\beta}(+) \phi_{\gamma}(0) + \phi_{\alpha}(0) \phi_{\beta}(0) \phi_{\gamma}(+) \right] \\ \phi_{\alpha\beta\gamma}(3,0) &= \frac{1}{\sqrt{10}} \left[\phi_{\alpha}(0) \phi_{\beta}(+) \phi_{\gamma}(-) + \phi_{\alpha}(0) \phi_{\beta}(-) \phi_{\gamma}(+) + \phi_{\alpha}(-) \phi_{\beta}(+) \phi_{\gamma}(0) \right. \\ &\quad + \phi_{\alpha}(+) \phi_{\beta}(-) \phi_{\gamma}(0) + \phi_{\alpha}(+) \phi_{\beta}(0) \phi_{\gamma}(-) + \phi_{\alpha}(-) \phi_{\beta}(0) \phi_{\gamma}(+) \right] + \frac{2}{\sqrt{10}} \phi_{\alpha}(0) \phi_{\beta}(0) \phi_{\gamma}(0) \end{split}$$

Orbital Angular Momenta in Tensor Representation:

The orbital angular momenta in tensor representation can be defined through the projection operator

(23)

$$P^{\ell} = \sum_{m} \tau^{\ell}(m) \tau^{\ell *}(m)$$

In the JRF where $\vec{r} = (0, 0, r)$, we have

(24)
$$[\tau^{\ell} * (m) \otimes rrr \cdots] = c_{\ell} r^{\ell} \delta_{m 0}, \quad c_{\ell} = \ell! \left[\frac{2^{\ell}}{(2\ell)!} \right]^{1/2}$$

So we see that

(25) $[P^{\ell} \otimes rrr \cdots] = \chi(0) r^{\ell}$

where we have introduced a new wave function

(26)
$$\chi(\ell 0) = c_{\ell} \tau^{\ell}(0)$$

Here $\tau^{\ell}(m)$ is the normalized wave function for spin ℓ defined previously.

(27)
$$\chi_{ijk\cdots}(\ell 0) = \frac{(\ell!)^2}{(2\ell)!} \sum_{m_0} 2^{(\ell+m_0)/2} \sum_{P} \left[\chi(+)\cdots\chi(0)\cdots\chi(-) \right]_{ijk\cdots}$$

and $m_0 = 0(1), 2(3), 4(5), \ldots$ for ℓ =even (odd). It is clear that $\chi(\ell 0)$ is now devoid of the coefficients with $\sqrt{(\cdot \cdot \cdot)}$. Note the identity

(28)
$$t^{\ell}_{ijk\cdots}(r) = \chi_{ijk\cdots}(\ell 0) r^{\ell}$$

Indeed, we find, with $\chi(0) = (0, 0, 1), \quad \chi(\pm) = \mp \frac{1}{\sqrt{2}} (1, \pm i, 0),$

$$\chi_{i}(10) = \chi_{i}(0)$$

$$\chi_{ij}(20) = \frac{1}{3} \left[\chi_{i}(+) \chi_{j}(-) + \chi_{i}(-) \chi_{j}(+) + 2 \chi_{i}(0) \chi_{j}(0) \right]$$
(29)
$$\chi_{ijk}(30) = \frac{1}{5} \left[\chi_{i}(0) \chi_{j}(+) \chi_{k}(-) + \chi_{i}(0) \chi_{j}(-) \chi_{k}(+) + \chi_{i}(+) \chi_{j}(0) \chi_{j}(-) + \chi_{i}(-) \chi_{j}(0) \chi_{k}(+) + \chi_{i}(+) \chi_{j}(-) \chi_{j}(0) + \chi_{i}(-) \chi_{j}(+) \chi_{k}(0) + 2 \chi_{i}(0) \chi_{j}(0) \chi_{k}(0) \right]$$

For completeness, we work out the wave function for ℓ =3 and 5 as well

$$\chi_{ijkl}(40) = \frac{2}{35} \left\{ \sum_{P}^{6} \left[\chi_{i}(+) \chi_{j}(-) \chi_{k}(+) \chi_{l}(-) \right] \right. \\ \left. + 2 \sum_{P}^{12} \left[\chi_{i}(0) \chi_{j}(0) \chi_{k}(+) \chi_{l}(-) \right] + 4 \chi_{i}(0) \chi_{j}(0) \chi_{k}(0) \chi_{l}(0) \right\} \right. \\ \left. \chi_{ijklm}(50) = \frac{2}{63} \left\{ \sum_{P}^{30} \left[\chi_{i}(0) \chi_{j}(+) \chi_{k}(-) \chi_{l}(+) \chi_{m}(-) \right] \right. \\ \left. + 2 \sum_{P}^{20} \left[\chi_{i}(0) \chi_{j}(0) \chi_{k}(0) \chi_{l}(+) \chi_{m}(-) \right] + 4 \chi_{i}(0) \chi_{j}(0) \chi_{k}(0) \chi_{l}(0) \chi_{m}(0) \right\} \right.$$

It can be shown that the use of $\chi_{ijk...}(\ell 0)$ tensors for the orbital angular momenta leads to a relatively simpler algebra than that of $t_{ijk...}^{\ell}(r)$.

Examples:

i) $1^- \to 0^- + 0^-$

The Lorentz invariant amplitude is, with $\ell = 1$ and S = 0,

(30)
$$A_{10}^{(1)} = r^{\alpha} \phi_{\alpha}^{*}(0)$$
$$(JRF) \rightarrow = [r \cdot \phi^{*}(0)] = [\chi(0) \cdot \phi^{*}(0)] r = r, \quad F_{0}^{J} = g_{1} r, \quad J = 1$$

ii)
$$\underline{2^+ \to 0^- + 0^-}$$

The Lorentz invariant amplitude is, with $\ell = 2$ and S = 0,

(31)
$$A_{20}^{(2)} = r^{\alpha} r^{\beta} \phi_{\alpha\beta}^{*}(2,0) = [rr:\phi^{*}(2,0)] = [\chi(20):\phi^{*}(2,0)] r^{2}$$

$$\phi_{\alpha\beta}(2,m) = \sum_{m_1m_2} (1m_11m_2|2m) \phi_{\alpha}(m_1) \phi_{\beta}(m_2), \quad \phi_{\alpha\beta}(2,-m) = (-)^m \phi_{\alpha\beta}^*(2,m)$$

and we find

(32)
$$A_{20}^{(2)} = \sqrt{\frac{2}{3}} r^2, \quad F_0^J = \sqrt{\frac{2}{3}} g_2 r^2, \quad J = 2$$

iii)
$$3^- \to 0^- + 0^-$$

The Lorentz invariant amplitude is, with $\ell = 3$ and S = 0,

(33)
$$A_{30}^{(3)} = r^{\alpha} r^{\beta} r^{\gamma} \phi_{\alpha\beta\gamma}^*(3,0) = [rrr \therefore \phi^*(3,0)] = [\chi(30) \therefore \phi^*(3,0)] r^3$$

where, with $\phi_{\alpha\beta\gamma}(3,-m)=(-)^m\,\phi^*_{\alpha\beta\gamma}(3,m)$, and so

(34)
$$A_{30}^{(3)} = \sqrt{\frac{2}{5}} r^3, \quad F_0^J = \sqrt{\frac{2}{5}} g_3 r^3, \quad J = 3$$

iv)
$$1^- \to 1^- + 0^-$$

The Lorentz invariant amplitude is, with $\ell = 1$ and S = 1,

$$A_{11}^{(1)}(\lambda) = \varepsilon^{\alpha\beta\gamma\delta} p_{\alpha} \omega_{\beta}(\lambda) r_{\gamma} \phi_{\delta}^{*}(\lambda) = [p \,\omega(\lambda) \, r \,\phi^{*}(\lambda)] = [p \,\omega(\lambda) \,\chi(0) \,\phi^{*}(\lambda)] r$$

$$(35) \qquad (JRF) \to = w \left(\vec{\omega}(\lambda) \times \vec{r} \cdot \vec{\phi}^{*}(\lambda)\right) = w[\vec{\omega}(\lambda) \, \vec{r} \, \vec{\phi}^{*}(\lambda)]$$

$$= \pm (i \, w) \, r, \qquad \lambda = \pm 1$$

$$= 0, \qquad \lambda = 0$$

and

(36)
$$F_{\pm}^{J} = \pm (iw) g_{1} r, \quad F_{0}^{J} = 0, \qquad J = 1$$

(37)
$$(\text{parity}) \to F_{\lambda}^{J} = -F_{-\lambda}^{J}$$

v)
$$\underline{2^+ \to 1^- + 0^-}$$

The Lorentz invariant amplitude is, with $\ell = 2$ and S = 1,

(38)
$$A_{21}^{(2)}(\lambda) = \varepsilon^{\alpha\beta\gamma\delta} p_{\alpha} \omega_{\beta}(\lambda) r_{\gamma} r^{\upsilon} \phi_{\upsilon\delta}^{*}(2,\lambda) = [p, \omega(\lambda), \chi(20) \cdot \phi^{*}(\lambda)] r^{2}$$

(39) $r \xi_{\delta}(2,\lambda) = r^{\upsilon} \phi_{\upsilon\delta}(2,\lambda) \xrightarrow{\mathbf{JRF}} r \xi_i(2,\lambda) = r_k \phi_{ki}(2,\lambda)$

and we find

40)
$$\begin{cases} \xi_{\delta}(2,+2) = 0\\ \xi_{\delta}(2,+1) = \frac{1}{\sqrt{2}} \phi_{\delta}(+)\\ \xi_{\delta}(2,0) = \sqrt{\frac{2}{3}} \phi_{\delta}(0) \end{cases}$$

so that, in JRF,

(41)

$$A_{21}^{(2)}(\lambda) = w \left(\vec{\omega}(\lambda) \times \vec{r} \cdot \vec{\xi^*}(2,\lambda) \right) r = w [\vec{\omega}(\lambda) \vec{r} \vec{\xi^*}(2,\lambda)] r$$

$$= \pm \frac{1}{\sqrt{2}} (i w) r^2, \qquad \lambda = \pm 1$$

$$= 0, \qquad \lambda = 0$$

The general decay amplitude is

(42)
$$F_{\pm}^{J} = \pm \frac{1}{\sqrt{2}} (i w) g_{2} r^{2}, \quad F_{0}^{J} = 0, \quad J = 2$$

vi) $1^+ \to 1^- + 0^-$

The helicity-coupling amplitudes are, with $\ell = 0$ or 2 and S = 1,

(43)
$$\begin{cases} F_{\pm}^{J} = \sqrt{\frac{1}{3}}G_{0}^{J} + \sqrt{\frac{1}{6}}G_{2}^{J}, & J = 1\\ F_{0}^{J} = \sqrt{\frac{1}{3}}G_{0}^{J} - \sqrt{\frac{2}{3}}G_{2}^{J}, & J = 1 \end{cases}$$

The Lorentz invariant amplitudes are

(44)

$$A_{01}^{(1)}(\lambda) = [\omega(\lambda) \cdot \phi^{*}(\lambda)], \qquad A_{21}^{(1)}(\lambda) = [\omega \cdot \chi(20) \cdot \phi^{*}(\lambda)] r^{2}$$

$$= \omega^{\alpha}(\lambda) \phi_{\alpha}^{*}(\lambda), \qquad = \omega^{\alpha}(\lambda) t_{\alpha\beta}^{(2)}(\tilde{r}) \phi^{*\beta}(\lambda)$$

$$(JRF) \rightarrow = \omega_{i}(\lambda) \phi_{i}^{*}(\lambda), \qquad (JRF) \rightarrow = \omega_{i}(\lambda) t_{ij}^{(2)}(r) \phi_{j}^{*}(\lambda)$$

where, in the JRF,

(45)
$$t^{(2)}_{\alpha\beta}(\tilde{r}) \rightarrow t^{(2)}_{ij}(r) = r_i r_j - \frac{1}{3}r^2 \delta_{ij}$$

so that, again in the JRF,

(46)
$$A_{21}^{(1)}(\lambda) = [\vec{r} \cdot \vec{\omega}(\lambda)][\vec{r} \cdot \vec{\phi^*}(\lambda)] - \frac{1}{3}r^2[\vec{\omega}(\lambda) \cdot \vec{\phi^*}(\lambda)]$$

We see that

(47)
$$\begin{aligned} A_{01}^{(1)}(\pm) &= 1, \\ A_{01}^{(1)}(\pm) &= -\frac{1}{3}r^2, \end{aligned} \qquad \begin{aligned} A_{01}^{(1)}(0) &= \gamma_s \\ A_{21}^{(1)}(\pm) &= -\frac{1}{3}r^2, \end{aligned} \qquad \begin{aligned} A_{21}^{(1)}(0) &= \frac{2}{3}\gamma_s r^2 \end{aligned}$$

(48)
$$\begin{cases} F_{\pm}^{J} = g_{0} - \frac{1}{3}g_{2}r^{2}, & J = 1\\ F_{0}^{J} = \left(g_{0} + \frac{2}{3}g_{2}r^{2}\right)\gamma_{s}, & J = 1 \end{cases}$$

Only in the limit $\gamma_s = q_0/m \rightarrow 1$, these expressions reduce to the previous ones, with the replacement

(49)
$$G_0^J = \sqrt{3} g_0, \qquad G_2^J = -\sqrt{\frac{2}{3}} g_2 r^2$$

Now explore the meaning of the Lorentz factor $\gamma_s = q_0/m$. We first examine the Taylor series expansion of γ_s

$$\gamma_s = \frac{q_0}{m} = \left[1 + \left(\frac{q}{m}\right)^2\right]^{1/2}$$

(50)
$$\simeq 1 + \frac{1}{2} \left(\frac{q}{m}\right)^2 - \frac{1}{2 \cdot 4} \left(\frac{q}{m}\right)^4 + \frac{1 \cdot 3}{2 \cdot 4 \cdot 6} \left(\frac{q}{m}\right)^6$$

$$-\frac{1\cdot 3\cdot 5}{2\cdot 4\cdot 6\cdot 8} \left(\frac{q}{m}\right)^8 + \frac{1\cdot 3\cdot 5\cdot 7}{2\cdot 4\cdot 6\cdot 8\cdot 10} \left(\frac{q}{m}\right)^{10} - \cdots$$

where q = r/2 in the JRF. So the orbital angular momentum is no longer quantized.

The decay amplitude is

(51)
$$\mathcal{M}^{J}_{\lambda}(\theta,\phi,M) \propto D^{J*}_{M\lambda}(\phi,\theta,0)F^{J}_{\lambda}$$

Assume that $\omega(\lambda)$ is a 'stable' particle. Then λ becomes an external variable which is unobserved; there should be no interference between different helicities, and the angular distribution must include a sum over the helicities. Assume further that the only non-zero density-matrix element is ρ_{00} .

Then, the angular distribution in the *J*RF becomes, after integrating over ϕ ,

(52)
$$I(\theta) \propto \sum_{\lambda} \left[d_{0\lambda}^{J}(\theta) \right]^{2} |F_{\lambda}^{J}|^{2}, \qquad J = 1$$

This leads to a distribution

(53)
$$I(\theta) \propto |F_0^J|^2 \cos^2(\theta) + |F_+^J|^2 \sin^2(\theta)$$

so that, assuming g_0 and g_2 are relatively real,

(54)
$$I(\theta) \propto g_0^2 [(\gamma_s^2 - 1)\cos^2(\theta) + 1] + \frac{1}{9}g_2^2 r^4 [(4\gamma_s^2 - 1)\cos^2(\theta) + 1] \\ \pm \frac{2}{3}g_0g_2 r^2 [(2\gamma_s^2 + 1)\cos^2(\theta) - 1]$$

Note

(55)
$$r = 2q$$
 and $\gamma_s^2 = 1 + \left(\frac{q}{m}\right)^2$

in the JRF.

vii) $2^- \to 2^+ + 0^-$

This problem involves three orbital angular momenta, $\ell = 0$, 2 and 4. Correspondingly, the invariant amplitudes are

(56)
$$\begin{cases} A_{02}^{(2)}(\lambda) = [\omega(\lambda) : \phi^*(\lambda)] \\ A_{22}^{(2)}(\lambda) = [\cdot\omega(\lambda) \cdot t^{(2)}(r) \cdot \phi^*(\lambda) \cdot] = [\cdot\omega(\lambda) \cdot \chi(20) \cdot \phi^*(\lambda) \cdot] r^2 \\ A_{42}^{(2)}(\lambda) = [\omega(\lambda) : t^{(4)}(r) : \phi^*(\lambda)] = [\omega(\lambda) : \chi(40) : \phi^*(\lambda)] r^4 \end{cases}$$

where

(57)
$$t_{ijkl}^{(4)}(r) = r_i r_j r_k r_l - \frac{1}{7} r^2 \sum_{P}^{6} (\delta_{ij} r_k r_l) + \frac{1}{35} r^4 \sum_{P}^{3} (\delta_{ij} \delta_{kl}) = \chi_{ijkl}(40) r^4$$

and so

(58)
$$\begin{cases} F_2^J = g_0 - \frac{1}{3} g_2 r^2 + \frac{2}{35} g_4 r^4 \\ F_1^J = \left(g_0 + \frac{1}{6} g_2 r^2 - \frac{8}{35} g_4 r^4\right) \gamma_s \\ F_0^J = g_0 \left(\frac{2}{3} \gamma_s^2 + \frac{1}{3}\right) + \frac{1}{3} g_2 r^2 \left(\frac{4}{3} \gamma_s^2 - \frac{1}{3}\right) + \frac{12}{35} g_4 r^4 \left(\frac{2}{3} \gamma_s^2 + \frac{1}{3}\right) \end{cases}$$

viii) $3^+ \to 2^+ + 0^-$

This problem involves $\ell = 1$, 3 or 5. So we need rank-3 and rank-5 tensors corresponding to $\ell = 3$ and 5. The relevant orbital-momentum tensors are

(59)

$$t_{ijk}^{(3)}(r) = r_{i} r_{j} r_{k} - \frac{1}{5} r^{2} (\delta_{ij} r_{k} + \delta_{jk} r_{i} + \delta_{ki} r_{j})$$

$$= \chi_{ijk}(30) r^{3}$$

$$t_{ijklm}^{(5)}(r) = r_{i} r_{j} r_{k} r_{l} r_{m} - \frac{1}{9} r^{2} \sum_{P}^{10} (\delta_{ij} r_{k} r_{l} r_{m}) + \frac{1}{63} r^{4} \sum_{P}^{15} (\delta_{ij} \delta_{kl} r_{m})$$

$$= \chi_{ijklm}(50) r^{5}$$

The decay amplitudes can be written

$$\begin{aligned} A_{12}^{(3)}(\lambda) &\propto \left[\omega(\lambda) : \phi^*(\lambda) \cdot \chi(0)\right] r \\ A_{32}^{(3)}(\lambda) &\propto \left[\cdot \omega(\lambda) \cdot \chi(30) : \phi^*(\lambda) \cdot\right] r^3 \\ A_{52}^{(3)}(\lambda) &\propto \left[\omega(\lambda) : \chi(50) \therefore \phi^*(\lambda)\right] r^5 \end{aligned}$$

(60)

It can be shown that (the required algebra tedious and difficult)

$$\begin{split} F_2^J &= \frac{1}{\sqrt{3}} \left(g_1 - \frac{2}{5} g_3 r^2 + \frac{2}{21} g_5 r^4 \right) r \\ F_1^J &= \sqrt{\frac{2}{15}} \gamma_s \left(2g_1 + \frac{1}{5} g_3 r^2 - \frac{10}{21} g_5 r^4 \right) r \\ F_0^J &= \sqrt{\frac{3}{5}} \left\{ g_1 \left(\frac{2}{3} \gamma_s^2 + \frac{1}{3} \right) + \frac{4}{15} g_3 \left(\frac{3}{2} \gamma_s^2 - \frac{1}{2} \right) r^2 + \frac{20}{63} g_5 \left(\frac{2}{3} \gamma_s^2 + \frac{1}{3} \right) r^4 \right\} r \end{split}$$

We need to find a different techinique which results in a relatively simpler algebra. Or, better still, find a general formula which does not require writing down invariant amplitudes.

S. U. Chung and Jan Friedrich, Phys. Rev. D78, 074027 (2008)

ix) $0^+ \to 1^- + 1^-$

The allowed orbital angular momenta are $\ell = 0$ and 2. Since J = 0, we must have $S = \ell = 0$ or 2 and $\lambda = \nu = 0$ or 1. So we find in the JRF

(61)
$$\begin{cases} A_{00}^{(0)}(0) = \tilde{g}_{\alpha\beta}(w) \,\psi^{\alpha\beta}(00) = -\frac{1}{\sqrt{3}} (2 + \gamma_s \gamma_\sigma) \\ A_{22}^{(0)}(0) = [\chi(20) : \psi(20)]_w \, r^2 \\ = \frac{1}{3} \sqrt{\frac{2}{3}} \, (1 + 2\gamma_s \gamma_\sigma) \end{cases}$$

The helicity-coupling amplitudes are

(62)
$$\begin{cases} F_{\pm\pm}^{(0)} = -g_{00}^{(0)} \left(\frac{1}{3}\gamma_s\gamma_\sigma + \frac{2}{3}\right) + \frac{1}{3}g_{22}^{(0)} \left(\frac{2}{3}\gamma_s\gamma_\sigma + \frac{1}{3}\right)r^2 \\ F_{00}^{(0)} = g_{00}^{(0)} \left(\frac{1}{3}\gamma_s\gamma_\sigma + \frac{2}{3}\right) + \frac{2}{3}g_{22}^{(0)} \left(\frac{2}{3}\gamma_s\gamma_\sigma + \frac{1}{3}\right)r^2 \end{cases}$$

In the nonrelativistic limit, we have $\gamma_s = \gamma_\sigma = 1$, so that

$$\begin{cases} F_{\pm\pm}^{(0)} = -g_{00}^{(0)} + \frac{1}{3} g_{22}^{(0)} r^2 \\ F_{00}^{(0)} = g_{00}^{(0)} + \frac{2}{3} g_{22}^{(0)} r^2 \end{cases} \begin{cases} F_{\pm\pm}^J = \sqrt{\frac{1}{3}} G_{00}^J + \sqrt{\frac{1}{6}} G_{22}^J \\ F_{00}^J = -\sqrt{\frac{1}{3}} G_{00}^J + \sqrt{\frac{2}{3}} G_{22}^J \end{cases}$$

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x) $1^+ \rightarrow 1^+ + 1^-$ [also for $1^- \rightarrow 1^- + 1^-$ or $\pi_1(2000) \rightarrow \omega \rho$]

The allowed orbital angular momenta are $\ell = 1$ and 3. Since J = 1, we can have S = 0, 1 or 2. The invariant amplitudes are, with $\delta = \lambda - \nu$,

$$\begin{cases} A_{1S}^{(1)}(\delta) = [\chi(0) \cdot \psi(S\delta) \cdot \phi^*(\delta)]_w r, & S = 0, 1, 2 \\ B_{1S}^{(1)}(\delta) = [\phi^*(\delta) \cdot \psi(S\delta) \cdot \chi(0)]_w r, & S = 0, 1, 2 \\ A_{3S}^{(1)}(\delta) = [\psi(S\delta) : \chi(30) \cdot \phi^*(\delta)]_w r^3, & S = 2 \end{cases}$$

So we see that

(63)

$$(64) \begin{cases} F_{++}^{(1)} = -\frac{1}{3}g_{10}^{(1)}\gamma_{s}\gamma_{\sigma}r + \frac{1}{3}(g_{12}^{(1)} + f_{12}^{(1)})\gamma_{s}\gamma_{\sigma}r + \frac{1}{5}g_{32}^{(1)}\left(\frac{1}{3} + \frac{2}{3}\gamma_{s}\gamma_{\sigma}\right)r^{3} \\ F_{0+}^{(1)} = \frac{1}{2}(g_{11}^{(1)}\gamma_{s} - f_{11}^{(1)}\gamma_{\sigma})r + \frac{1}{2}(g_{12}^{(1)}\gamma_{s} + f_{12}^{(1)}\gamma_{\sigma})r \\ - \frac{1}{5}g_{32}^{(1)}\left(\frac{1}{2}\gamma_{s} + \frac{1}{2}\gamma_{\sigma}\right)r^{3} \\ F_{+0}^{(1)} = -\frac{1}{2}(g_{11}^{(1)}\gamma_{s} - f_{11}^{(1)}\gamma_{\sigma})r + \frac{1}{2}(g_{12}^{(1)}\gamma_{s} + f_{12}^{(1)}\gamma_{\sigma})r \\ - \frac{1}{5}g_{32}^{(1)}\left(\frac{1}{2}\gamma_{s} + \frac{1}{2}\gamma_{\sigma}\right)r^{3} \\ F_{00}^{(1)} = \frac{1}{3}g_{10}^{(1)}\gamma_{s}\gamma_{\sigma}r + \frac{2}{3}(g_{12}^{(1)} + f_{12}^{(1)})\gamma_{s}\gamma_{\sigma}r + \frac{2}{5}g_{32}^{(1)}\left(\frac{1}{3} + \frac{2}{3}\gamma_{s}\gamma_{\sigma}\right)r^{3} \end{cases}$$

x) $1^+ \rightarrow 1^+ + 1^-$ (continued)

So the amplitudes depend on six parameters — $g_{10}^{(1)}$, $g_{11}^{(1)}$, $g_{12}^{(1)}$, $g_{32}^{(1)}$, $f_{11}^{(1)}$ and $f_{12}^{(1)}$. A potential seventh parameter, $f_{10}^{(1)}$, occurs with the same functional dependence on γ 's as $g_{10}^{(1)}$, and hence it can be obsorbed into $g_{10}^{(1)}$.

The helicity-coupling amplitudes are, in the previous formulation,

$$\begin{cases} \sqrt{2}F_{++}^{(1)} = \sqrt{\frac{2}{3}}g_{10}^{(1)}r - \sqrt{\frac{2}{15}}g_{12}^{(1)}r + \frac{1}{\sqrt{5}}g_{32}^{(1)}r^{3} \\ \sqrt{2}F_{0+}^{(1)} = \left[\frac{1}{\sqrt{2}}g_{11}^{(1)}r - \sqrt{\frac{3}{10}}g_{12}^{(1)}r - \frac{1}{\sqrt{5}}g_{32}^{(1)}r^{3}\right]\gamma_{s} \\ \sqrt{2}F_{+0}^{(1)} = \left[-\frac{1}{\sqrt{2}}g_{11}^{(1)}r - \sqrt{\frac{3}{10}}g_{12}^{(1)}r - \frac{1}{\sqrt{5}}g_{32}^{(1)}r^{3}\right]\gamma_{\sigma} \\ F_{00}^{(1)} = \left[-\frac{1}{\sqrt{3}}g_{10}^{(1)}r - \frac{2}{\sqrt{15}}g_{12}^{(1)}r + \sqrt{\frac{2}{5}}g_{32}^{(1)}r^{3}\right]\gamma_{s}\gamma_{\sigma} \end{cases}$$

(65)

Consider now the decay $1^- \rightarrow 1^- + 1^-$. Carrying out the Bose symmetrization requirement, we find

$$\textbf{(66)} \quad \begin{cases} F_{++}^{(1)} = 0 \\ F_{0+}^{(1)} = \frac{1}{2}(g_{11}^{(1)} - f_{11}^{(1)}) \left(\frac{1}{2}\gamma_s + \frac{1}{2}\gamma_\sigma\right) r + \frac{1}{4}(g_{12}^{(1)} - f_{12}^{(1)}) \left(\gamma_s - \gamma_\sigma\right) r \\ F_{+0}^{(1)} = -\frac{1}{2}(g_{11}^{(1)} - f_{11}^{(1)}) \left(\frac{1}{2}\gamma_s + \frac{1}{2}\gamma_\sigma\right) r + \frac{1}{4}(g_{12}^{(1)} - f_{12}^{(1)}) \left(\gamma_s - \gamma_\sigma\right) r \\ F_{00}^{(1)} = 0 \end{cases}$$

The amplitudes depend on two parameters $(g_{11}^{(1)} - f_{11}^{(1)})$ and $(g_{12}^{(1)} - f_{12}^{(1)})$. The terms corresponding to the second parameter, with $\ell + S = \text{odd}$, vanish in the nonrelativistic limit $\gamma_s = \gamma_\sigma = 1$.

xi) $\underline{1^- \rightarrow 1^+ + 1^-}$

The allowed orbital angular momenta are $\ell = 0$ and 2. Since J = 1, we must have S = 1 or 2.

(67)
$$\begin{cases} A_{01}^{(1)}(\delta) = [p, \ \psi(1\delta), \ \phi^*(\delta)]_w \\ A_{2S}^{(1)}(\delta) = [p, \ \psi(S\delta) \cdot \chi(20), \ \phi^*(\delta)]_w \ r^2, \quad S = 1, 2 \\ B_{2S}^{(1)}(\delta) = [p, \ \chi(20) \cdot \ \psi(S\delta), \ \phi^*(\delta)]_w \ r^2, \quad S = 1, 2 \\ C_{2S}^{(1)}(\delta) = [p, \ \psi(S\delta), \ \chi(20) \cdot \ \phi^*(\delta)]_w \ r^2, \quad S = 1, 2 \end{cases}$$

The helicity-coupling amplitudes are

$$\begin{cases} F_{++}^{(1)} = (iw) \left[g_{01}^{(1)} - \frac{1}{3} g_{21}^{(1)} r^2 - \frac{1}{3} f_{21}^{(1)} r^2 + \frac{2}{3} h_{21}^{(1)} r^2 \right] \\ F_{0+}^{(1)} = (iw) \left[g_{01}^{(1)} \left(\frac{1}{2} \gamma_{\sigma} + \frac{1}{2} \gamma_s \right) - \frac{1}{3} h_{21}^{(1)} \left(\frac{1}{2} \gamma_{\sigma} + \frac{1}{2} \gamma_s \right) r^2 \\ + \frac{1}{6} g_{21}^{(1)} (2\gamma_{\sigma} - \gamma_s) r^2 + \frac{1}{6} f_{21}^{(1)} (2\gamma_s - \gamma_\sigma) r^2 + \frac{1}{6} h_{22}^{(1)} (\gamma_{\sigma} - \gamma_s) r^2 \\ - \frac{1}{2} g_{22}^{(1)} \left(\frac{2}{3} \gamma_{\sigma} + \frac{1}{3} \gamma_s \right) r^2 + \frac{1}{2} f_{22}^{(1)} \left(\frac{2}{3} \gamma_s + \frac{1}{3} \gamma_{\sigma} \right) r^2 \right] \\ F_{+0}^{(1)} = (iw) \left[g_{01}^{(1)} \left(\frac{1}{2} \gamma_{\sigma} + \frac{1}{2} \gamma_s \right) - \frac{1}{3} h_{21}^{(1)} \left(\frac{1}{2} \gamma_{\sigma} + \frac{1}{2} \gamma_s \right) r^2 \\ + \frac{1}{6} g_{21}^{(1)} (2\gamma_{\sigma} - \gamma_s) r^2 + \frac{1}{6} f_{21}^{(1)} (2\gamma_s - \gamma_{\sigma}) r^2 - \frac{1}{6} h_{22}^{(1)} (\gamma_{\sigma} - \gamma_s) r^2 \\ + \frac{1}{2} g_{22}^{(1)} \left(\frac{2}{3} \gamma_{\sigma} + \frac{1}{3} \gamma_s \right) r^2 - \frac{1}{2} f_{22}^{(1)} \left(\frac{2}{3} \gamma_s + \frac{1}{3} \gamma_{\sigma} \right) r^2 \right] \end{cases}$$

There are a total of seven parameters for the amplitudes.

This example applies equally well to the case: $1^+ \rightarrow 1^- + 1^-$, where the two vector states are identical. The resulting helicity-coupling amplitudes are

$$(69) \begin{cases} F_{++}^{(1)} = 0 \\ F_{0+}^{(1)} = (iw) \left[\frac{1}{4} \left(g_{21}^{(1)} - f_{21}^{(1)} \right) \left(\gamma_{\sigma} - \gamma_{s} \right) - \frac{1}{2} \left(g_{22}^{(1)} - f_{22}^{(1)} \right) \left(\frac{1}{2} \gamma_{\sigma} + \frac{1}{2} \gamma_{s} \right) \right] r^{2} \\ F_{+0}^{(1)} = (iw) \left[\frac{1}{4} \left(g_{21}^{(1)} - f_{21}^{(1)} \right) \left(\gamma_{\sigma} - \gamma_{s} \right) + \frac{1}{2} \left(g_{22}^{(1)} - f_{22}^{(1)} \right) \left(\frac{1}{2} \gamma_{\sigma} + \frac{1}{2} \gamma_{s} \right) \right] r^{2} \end{cases}$$

and the amplitudes depend only on two parameters $(g_{21}^{(1)} - f_{21}^{(1)})$ and $(g_{22}^{(1)} - f_{22}^{(1)})$. The first term is *purely* relativistic and so it vanishes in the limit $\gamma_s = \gamma_\sigma = 1$. In addition, one may expect that the first term should remain small and insignificant, since γ_s should be nearly equal to γ_σ independent of the size of w compared to m or μ .

The helicity-coupling amplitudes are, in the nonrelativistic limit,

$$\begin{cases} \sqrt{2}F_{++}^{(1)} = \frac{1}{\sqrt{3}}G_{01}^{(1)} - \sqrt{\frac{2}{3}}G_{21}^{(1)} \\ \sqrt{2}F_{0+}^{(1)} = \frac{1}{\sqrt{3}}G_{01}^{(1)} + \frac{1}{\sqrt{6}}G_{21}^{(1)} - \frac{1}{\sqrt{2}}G_{22}^{(1)} \\ \sqrt{2}F_{+0}^{(1)} = \frac{1}{\sqrt{3}}G_{01}^{(1)} + \frac{1}{\sqrt{6}}G_{21}^{(1)} + \frac{1}{\sqrt{2}}G_{22}^{(1)} \end{cases}$$

(70)

There are three parameters for three amplitudes.

To Be Continuned...