## Few-Body Hypernuclei





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- Motivation
- Numerical technique
- Light Hypernuclei
  - dependence on NN and 3N force
  - separation energies based on chiral interactions
  - CSB of four-body hypernuclei
- Conclusions & Outlook

## Hypernuclear interactions

#### Why is understanding hypernuclear interactions interesting?

- "phenomenologically"
  - hyperon contribution to the EOS, neutron stars, supernovae
  - *Λ* as probe to nuclear structure
- conceptually
  - Λ-Σ conversion process
  - experimental access to explicit chiral symmetry breaking





(SN1987a)





## Hypernuclear interactions

JÜLICH

35 YN data, no YN bound state, large uncertainties

no partial wave analysis possible

YN interaction models (Jülich 89/04, Nijmegen 89/97a-f, ESC, ...) describe all data more than perfectly, but are not phase equivalent



	1	3
SC97a	-0.7	-2.15
SC97b	-0.9	-2.11
SC97c	-1.2	-2.06
SC97d	-1.7	-1.93
SC97e	-2.1	-1.83
SC97f	-2.5	-1.73
SC89	-2.6	-1.38
Jülich '04	-2.6	-1.73

How to further constrain the YN interactions?

## Hypernuclei



- AN interactions are generally weaker than the NN interaction
  - naively: core nucleus + hyperons
  - "separation energies" are almost independent from NN(+3N) interaction
- no Pauli blocking of Λ in nuclei
  - good to study nuclear structure
  - even light hypernuclei exist in several spin states
- size of YNN interactions?
- non-trivial constraints on the YN interaction even from lightest ones  $_{3\,\mathrm{II}}(1^{+})$

$${}^{3}_{\Lambda} \mathrm{H} \left( \frac{1}{2}^{+} \right)$$

$${}^{4}_{\Lambda} \mathrm{H} \left( 0^{+} \right) - {}^{4}_{\Lambda} \mathrm{He} \left( 0^{+} \right)$$

$${}^{4}_{\Lambda} \mathrm{H} \left( 1^{+} \right) - {}^{4}_{\Lambda} \mathrm{He} \left( 1^{+} \right)$$



### Numerical technique

non-rel. Schrödinger equation  $\Psi = G_0 V \Psi$ 



decomposition in five Yakubovsky components

 $\Psi = (1+P)(\psi_{1A} + \psi_{1B} + \psi_{2A} + \psi_{2B}) + (1-P_{12})(1+P)\psi_{1C}$ solution of the Yakubovsky equations

$$\begin{split} \psi_{1A} &= G_0 t_{12} P(\psi_{1A} + \psi_{1B} + \psi_{2A}) + (1 + G_0 t_{12}) G_0 V_{123}^{(3)} \Psi \\ \psi_{1B} &= G_0 t_{12} ((1 - P_{12})(1 - P_{23}) \psi_{1B} + P \psi_{2B}) \\ \psi_{1C} &= G_0 t_{14} (\psi_{1A} + \psi_{1B} + \psi_{2A} - P_{12} \psi_{1C} + P_{12} P_{23} \psi_{1C} + P_{13} P_{23} \psi_{2B}) \\ \psi_{2A} &= G_0 t_{12} ((P_{12} - 1) P_{13} \psi_{1C} + \psi_{2B}) \\ \psi_{2B} &= G_0 t_{34} (\psi_{1A} + \psi_{1B} + \psi_{2A}) \\ \end{split}$$



improved convergence in terms of partial waves

we carefully checked convergence with respect to partial waves, stability with respect to mesh points, ...

(see Nogga et. al., PRL 88,172501 (2002))

## Known results I: independendence of NN force



Λ separation energies

$$E_{\Lambda} = E(core) - E(hypernucleus)$$

are not strongly dependent on the NN interaction

4 TL-	$0^+$		1	٨	
Λ́не	$E_B$	$E_{\Lambda}$	$E_B$	$E_{\Lambda}$	Δ
Bonn B	-8.92	1.66	-8.04	0.80	0.84
Nijm 93	-8.55	1.54	-7.69	0.72	0.79
Nijm 93 + TM	-9.32	1.56	-8.35	0.70	0.82

#### for YN interaction: SC97e

(AN, Kamada, Glöckle, 2002)

YN interaction can be discussed independently of an NN and 3N force model

# Known results II: $\Lambda$ - $\Sigma$ conversion is important



#### strong $\Lambda$ - $\Sigma$ conversion process



suppressed by isospin symmetry

- strong conversion process
- mass difference comparable to typical momenta
- no  $\pi$ -exchange  $\Lambda N$ - $\Lambda N$  interaction

 $m_K \approx 500 \text{ MeV}$ 

К

Ν

 $2m_{\pi} + m_{\Sigma} - m_{\Lambda} \approx 360 \text{ MeV}$ 

- AN is weaker than NN interaction
- Σs need to be explicitly included in any realistic calculation

effective AN interactions are

not useful to study YN forces

test: use t<sub>AN</sub> in Yakubovsky equations (here for a chiral interaction)

	w/Σ	w/o Σ	
E	1.47	1.01	
E	0.71	0.49	

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## Known results III: model dependence



	$^3_{\Lambda}{ m H}$ in MeV	$0^+$ $^4_{ m in N}$	ie /ev <sup>1+</sup>	$^1a_{\Lambda p}$ in	$^{3}a_{\Lambda p}$ fm	$P_{\Sigma}$
SC97d	-	1.3	0.8	-1.7	-1.9	1.5 %
SC97e	0.02	1.5	0.7	-2.1	-1.8	1.6 %
SC97f	0.08	1.7	0.5	-2.5	-1.7	1.8 %
SC89	0.15	2.1	0.02	-2.6	-1.4	4.1 %
Jülich 04	0.13	1.9	2.3	-2.6	-1.7	0.9 %
Expt	0.13	2.4	1.2	?	?	-

mostly from (AN, Kamada, Glöckle, 2002)

- none of these interaction models predicts the hypernuclei correctly
- no strict relation of the scattering lengths to any separation energy

With this in mind:

- qualitative study of predications based on LO and NLO interactions w/o SU(3) breaking
- first attempt to estimate N2LO/3BF contribution by variation of  $\lambda$
- qualitative study of **CSB** of  ${}^4_{\Lambda}{
  m H} {}^4_{\Lambda}{
  m He}$

## Chiral NN & YN interactions



#### reminder:



(from Epelbaum, 2008)

additional constraints required (only 35 data, but 26 parameters at NLO)

- SU(3) broken by physical  $m_{\pi}, m_{\kappa}, m_{\eta}$
- no SU(3) breaking in F<sub>π</sub>, F<sub>K</sub>, F<sub>η</sub>
- "minimize" P-waves and <sup>1</sup>P<sub>1</sub>-<sup>3</sup>P<sub>1</sub> mixing only 13 parameters determined by data
- realizations for  $\lambda = 450 \dots 700 \text{ MeV}$



- one possible realization at NLO more constraints required
- (J. Haidenbauer et al., 2013 & previous talk)

## Chiral interactions at LO, NLO



	1	3
LO	-1.9	-1.2
NLO	-2.9	-1.51.7
Jülich '04	-2.6	-1.7

(Polinder et al., NPA 779, 244 (2006), Haidenbauer et al., NPA 915, 24 (2013) see Johann Haidenbauer's talk)

- hypertriton binding energy provides constraint on spin dependence of the YN interaction
- better description of the energy dependence in NLO
- significantly increased scattering lengths in NLO compared to LO



## How important are 3B forces?





(from Epelbaum, 2008)

- we explicitly include the  $\Sigma$ ! (otherwise the 3BF should be LO)
- the missing 3BF are either **short-ranged** or induced by **decouplet** baryons ( $\Sigma^*$ ,  $\Delta$ )

Important tool to estimate 3BF in absence of explicit calculations: cutoff variations allow one to get lower bounds on their contribution

## Hypertriton separation energies





- singlet scattering length for one cutoff chosen so that hypertriton binding energy is OK
- cutoff variation
  - is **lower bound** for magnitude of higher order contributions
  - correlation with  $\chi^2$  of YN interaction ?
- long range 3BFs need to be explicitly estimated

## Separation energies for ${}^{4}_{\Lambda}H$





- LO/NLO results: LO uncertainty in 0<sup>+</sup> is underestimated by cutoff variation
- NLO results in line with model results, implies underbinding
- long range 3BFs need to be explicitly estimated
- but: for this version of NLO, results are inconsistent with experiment
  - note: this NLO does not allow for SU(3) breaking in contact part of YN
  - ad-hoc p-waves

## Separation energies of ${}^{4}_{\Lambda}$ H





- LO/NLO cutoff dependence does not indicate 3BF contribution long range 3BF needs to be studied
- results cutoff dependence for small Λ?
   related to non-optimal description of data?
- LO/NLO: splitting stabilizes
- but: NLO results are inconsistent with experiment

## CSB at NLO & for model interactions



#### **Contributions to the difference** $E_{\Lambda} \begin{pmatrix} 4 \\ \Lambda \end{pmatrix} - E_{\Lambda} \begin{pmatrix} 4 \\ \Lambda \end{pmatrix} + E_{\Lambda} \begin{pmatrix} 4 \\$



• NN force contribution due to small deviation of Coulomb

- YN force contribution:
  - SC89 CSB is strong
  - NLO CSB is zero, only Coulomb acts (Σ component)
- kinetic energy contribution is driven by Σ component

# CSB and $\Sigma$ probability $\Sigma$ probabilities in $~^4_{\Lambda} H$



- spin/isospin structure of hypernuclei drives
   Σ components
- kinetic energy contribution is given linearly by differences of Σ components





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## **Conclusions & Outlook**



- YN interactions are interesting and not well understood
  - $\Lambda$ - $\Sigma$  conversion, explicit chiral symmetry breaking
  - well known: YN models fail
  - NLO of chiral interactions: still freedom to adjust YN forces
  - but: further estimates of three-baryon interactions (in progress)
- hypernuclei are an essential source of information on YN
  - *it is not trivial to describe the simplest systems consistently*
  - experiments for very light hypernuclei are important! The data needs to be accurate (better data for the hypertriton?) We need to be sure that these data are reliable.
- CSB for four-body hypernuclei is a puzzle
  - obviously related to  $\Lambda$ - $\Sigma$  conversion Can we engineer chiral interactions with different conversion strength?
  - experiments for **very light** hypernuclei are important! Is today's data reliable?
- extension of complete calculations to larger systems (access more data) (see also Roland Wirth's talk)