PV DIS: SUSY and Higher Twist



M.J. Ramsey-Musolf

Caltech Wisconsin-Madison

QuickTime[™] and a TIFF (Uncompressed) decompresson are needed to see this picture.

S. Mantry, K. Lee, G. Sacco Caltech

Outline

I. PV DIS & new physics Model independent analysis SUSY effects

II. PV DIS & Higher twist Puzzles from JLab data The twist expansion What does QCD predict? Q²-dependence, operator matrix elements

III. General remarks

I. PV DIS & New physics

Model Independent Constraints

Low energy effective PV eq interaction

$$L_{PV}^{eq} = \frac{G_{\mu}}{\sqrt{2}} \sum_{q} \left[C_{1q} \bar{e} \gamma^{\mu} \gamma_{5} e \bar{q} \gamma_{\mu} q + C_{2q} \bar{e} \gamma^{\mu} e \bar{q} \gamma_{\mu} \gamma_{5} q \right]$$

PV DIS eD asymmetry: leading twist

$$A_{PV}^{eD} = \frac{3G_{\mu}Q^2}{2\sqrt{2}\pi\alpha} \left[\frac{2C_{1u} - C_{1d} + Y(2C_{2u} - C_{2d})}{5} \right]$$

$$Y = \frac{1 - (1 - y)^2}{1 + (1 - y)^2 - \frac{y^2 R}{1 + R}}$$

$$R(x,q^2) = \frac{\sigma_L}{\sigma_R} \approx 0.2$$

Model Independent Constraints



P. Reimer, X. Zheng

Q_w and SUSY Radiative Corrections

Tree Level

$$Like Q_{W}^{p,e} \sim 1 - 4 \sin^{2}\theta_{W}$$

$$C_{2q} = -\frac{1}{2}g_{A}^{q}g_{V}^{e} = 2I_{3}^{q}(I_{3}^{e} - 2Q_{e}\hat{s}^{2})$$
Flavor-dependent
Radiative Corrections

$$C_{2q} = 2\hat{\rho}_{NC}I_{3}^{q}(I_{3}^{e} - 2Q_{e}\hat{\kappa}\hat{s}^{2}) - \frac{1}{2}\hat{\lambda}_{2}^{q}$$

Constrained by Z-pole precision observables

Scale-dependent effective weak mixing

Flavor-independent

SUSY Radiative Corrections



SUSY RPV Effects

$$\begin{split} \Delta C_{1u}^{\text{RPV}} &= -[C_{1u} - \frac{4}{3}\lambda_x] \Delta_{12k}(\tilde{e}_R^k) - \Delta'_{11k}(\tilde{d}_R^k), \\ \Delta C_{1d}^{\text{RPV}} &= -[C_{1d} + \frac{2}{3}\lambda_x] \Delta_{12k}(\tilde{e}_R^k) + \Delta'_{1k1}(\tilde{q}_L^k), \\ \Delta C_{2u}^{\text{RPV}} &= -[C_{2u} - 2\lambda_x] \Delta_{12k}(\tilde{e}_R^k) - \Delta'_{11k}(\tilde{d}_R^k), \\ \Delta C_{2d}^{\text{RPV}} &= -[C_{2d} + 2\lambda_x] \Delta_{12k}(\tilde{e}_R^k) - \Delta'_{1k1}(\tilde{q}_L^k), \end{split}$$



$$\lambda_x = \frac{\hat{s}^2 (1 - \hat{s}^2)}{1 - 2\hat{s}^2} \frac{1}{1 - \Delta \hat{r}^{\text{SM}}} \approx 0.35$$

Comparing A_d^{DIS} and $Q_w^{p,e}$



Probing SUSY with PV eN Interactions

$$\frac{\delta Q_W^e}{Q_W^e} \approx -30 \Delta_{12k}(\tilde{e}_R^k) \approx -45 \left(\frac{100 \, GeV}{m_{\tilde{e}_R^k}}\right)^2 |\lambda_{12k}|^2$$
$$\lambda_{12k} \sim 0.3 \text{ for } m_{\text{SUSY}} \sim 1 \text{ TeV \& } \delta Q_W^e / Q_W^e \sim 5\%$$

 $\partial v \beta \beta$ sensitivity $\lambda'_{111} \le 2 \times 10^{-4} \left(\frac{m_{\tilde{q}}}{100 \, \text{GeV}}\right)^2 \left(\frac{m_{\tilde{g}}}{100 \, \text{GeV}}\right)^{1/2}$ λ_{111} ~ 0.06 for m_{SUSY} ~ 1 TeV LFV Probes of RPV: $\mu ! e \gamma$ $|\lambda_{131}\lambda_{231}| \leq 2.3 imes 10^{-4} \left(rac{m_{ ilde{\ell}}}{100 \; {
m GeV}}
ight)^2$ $|\lambda'_{111}\lambda'_{211}| \le 7.6 \times 10^{-5} \left(\frac{m_{\tilde{q}}}{100 \text{ GeV}}\right)^2$ $\lambda_{k31} \sim 0.15$ for $m_{SUSY} \sim 1$ TeV

$$egin{aligned} &|\lambda_{131}\lambda_{231}| \leq 1.1 imes 10^{-5} \left(rac{m_{ ilde{\ell}}}{100 \; {
m GeV}}
ight)^2 \ &|\lambda_{111}'\lambda_{211}'| \leq 6.0 imes 10^{-7} \left(rac{m_{ ilde{q}}}{100 \; {
m GeV}}
ight)^2 \end{aligned}$$

 $\lambda_{k31} \sim 0.03$ for $m_{SUSY} \sim 1$ TeV

Lepton Flavor & Number Violation



II. PV DIS & Higher Twist

Probing the strange sea with PV

First look beyond the guark model Not surprising: $m_{s} / \Lambda_{\gamma} \sim 0.15$ Great success for the JLab mission

-0.1

-0.15



 $G_{\rm M}^{\rm s} = 0.28 + / - 0.20$ $G_{\rm F}^{\rm s}$ = -0.006 +/- 0.016 ~3% +/- 2.3% of proton magnetic moment ~20% +/- 15% of isoscalar magnetic moment $\sim 0.2 + / - 0.5\%$ of Electric distribution

Consistent with s-quark contributions to $m_P \& J_P$ but smaller than early theoretical expectations

Courtesy of Kent Pashke (U Mass)



Data from JLab E94-110 (nucl-ex/0410027, submitted to PRL) Courtesy C Keppel



Twist Expansion

Light cone expansion: $x^2 \sim 0$

 $T\left\{J^a_\mu(x)J^b_\nu(0)\right\}$





Jaffe '97

Twist Expansion

Light cone expansion: $x^2 \sim 0$

$$T\left\{J^{a}_{\mu}(x)J^{b}_{\nu}(0)\right\}\sim\Gamma_{\mu\nu}\sum_{n,k}x^{\mu_{1}}\dots x^{\mu_{n}}C^{(n)}_{k}(x^{2})\hat{O}^{(n)}_{k,\ \mu_{1}\dots\mu_{n}}(0)$$

Momentum sum rules

$$M_{j}^{(n)}(Q^{2}) \equiv \int_{0}^{1} dx_{B} x_{B}^{n-j} F_{j}(x_{B}, Q^{2}) \propto \sum_{k} \tilde{C}_{1,k}^{(n)}(Q^{2}/\mu^{2}, g) A_{k}^{(n)}(\mu)$$

$$\langle P | \hat{O}_{k, \ \mu_{1}...\mu_{n}}^{(n)}(0) | P \rangle = A_{k}^{(n)} (P_{\mu_{1}}...P_{\mu_{n}} + \cdots)$$

$$Twist = d(n) - j$$

$$\tilde{C}_{j,k}^{(n)}(Q^{2}) \sim \ln \frac{Q^{2}}{\mu^{2}} \qquad twist 2$$

$$\tilde{C}_{j,k}^{(n)}(Q^{2}) \sim \frac{1}{Q^{2}} \ln \frac{Q^{2}}{\mu^{2}} \qquad twist 4$$

Twist Expansion: Q² Evolution

Twist two



Singlet moments: 2 x 2 mixing only

 $\hat{O}_{NS}^{(n)} = \frac{(i)^{n-1}}{n!} \Big\{ \bar{q}(x) \frac{\lambda^a}{2} \gamma \Big\}$ $\hat{O}_{Sq}^{(n)} = \frac{(i)^{n-1}}{n!} \Big\{ \bar{q}(x) \gamma_{\mu_1}$

 $\hat{O}_{Sg}^{(n)} = \frac{2(i)^{n-2}}{n!} Tr \left\{ G_{\mu_1}^{\nu} \right\}$

 $\Delta \cdot Q_n^{1(k,l)} = g \overline{\psi}_R \mathbb{A} \overline{d}^{l} \overline{d}^k \psi_R \overline{\psi}_R \mathbb{A} \overline{d}^{n-2-k-l} \psi_R ,$ $\Delta \cdot Q_n^{2(k,l)} = g \overline{\psi}_R \tau_a \mathbf{\Delta} \vec{d}^{-l} \vec{d}^k \psi_R \overline{\psi}_R \mathbf{\Delta} \vec{d}^{n-2-k-l} \tau_a \psi_R ,$ $\Delta \cdot Q_n^{3(k,l)} = g \overline{\psi}_R \mathbb{A} \overline{d}^{l} \overline{d}^k \psi_R \overline{\psi}_L \mathbb{A} \overline{d}^{n-2-k-l} \psi_L ,$ $\Delta \cdot Q_n^{4(k,l)} = g \overline{\psi}_R \tau_a \mathbb{A} \vec{d}^{-l} \vec{d}^k \psi_R \overline{\psi}_L \mathbb{A} \vec{d}^{n-2-k-l} \tau_a \psi_L ,$ $\Delta \cdot Q_n^{5(k,l)} = g \overline{\psi}_L \, \underline{A} \overline{d}^{l} \overline{d}^k \psi_L \overline{\psi}_L \, \underline{A} \overline{d}^{n-2-k-l} \psi_L \,,$ $\Delta \cdot Q_n^{6(k,l)} = g \overline{\psi}_L \tau_a \Delta \vec{d}^{l} \vec{d}^k \psi_L \overline{\psi}_L \Delta \vec{d}^{n-2-k-l} \tau_a \psi_L ,$ $\Delta \cdot Q_n^{7(k)} = \overline{\psi} d^{k*} f \gamma_5 d^{n-1-k} \psi ,$ $\Delta \cdot Q_n^{8(k)} = i \overline{\psi} d^{-k} f d^{n-1-k} \psi ,$ $\Delta \cdot Q_n^{9(k,l)} = g \overline{\psi} d^k f_a^a (d^l f_a)_a d^{n-3-k-l} \Delta \psi ,$ $\Delta \cdot Q_n^{10(k,l)} = igf_{abc} \overline{\psi} \overline{d}^k f_a^{\alpha} (\vec{d}^l f_{\alpha})_b \vec{d}^{n-3-k-l} \Delta \tau_c \psi ,$ $\Delta \cdot Q_n^{11(k,l)} = g d_{abc} \overline{\psi} \overline{d}^k f_a^{\alpha} (\overline{d}^l f_{\alpha})_b \overline{d}^{n-3-k-l} \Delta \tau_c \psi ,$ $\Delta \cdot Q_n^{12(k,l)} = ig \overline{\psi} \overline{d}^{k*} f_a^{\alpha} (\overline{d}^l f_{\alpha})_a \overline{d}^{n-3-k-l} \Delta \gamma_5 \psi ,$ $\Delta \cdot Q_n^{13(k,l)} = g f_{abc} \overline{\psi} \overline{d}^{-k} * f_a^a (\vec{d}^l f_\alpha)_b \vec{d}^{n-3-k-l} \Delta \gamma_5 \tau_c \psi ,$ $\Delta \cdot Q_n^{14(k,l)} = igd_{abc} \overline{\psi} \overline{d}^{k*} f_a^{\alpha} (\overline{d}^l f_{\alpha})_b \overline{d}^{n-3-k-l} \Delta \gamma_5 \tau_c \psi .$

Twist Expansion : Q² Evolution

Equivalence to parton model: twist two



Twist Expansion : Q² Evolution



 $\Delta \cdot O_n^{1(k,l)} = g \overline{\psi}_R \Delta \overline{d}^{-l} \overline{d}^k \psi_R \overline{\psi}_R \Delta \overline{d}^{n-2-k-l} \psi_R ,$ $\Delta \cdot Q_n^{2(k,l)} = g \overline{\psi}_R \tau_a \Delta \vec{d}^{-l} \vec{d}^k \psi_R \overline{\psi}_R \Delta \vec{d}^{n-2-k-l} \tau_a \psi_R ,$ $\Delta \cdot Q_n^{3(k,l)} = g \overline{\psi}_R \Delta \overline{d}^{-l} \overline{d}^k \psi_R \overline{\psi}_L \Delta \overline{d}^{n-2-k-l} \psi_L ,$ $\Delta \cdot Q_n^{4(k,l)} = g \overline{\psi}_R \tau_a \Delta \vec{d}^{-l} \vec{d}^k \psi_R \overline{\psi}_L \Delta \vec{d}^{n-2-k-l} \tau_a \psi_L ,$ $\Delta \cdot Q_n^{5(k,l)} = g \overline{\psi}_L \mathbf{\Delta} \overline{d}^{l} \overline{d}^k \psi_L \overline{\psi}_L \mathbf{\Delta} \overline{d}^{n-2-k-l} \psi_L ,$ $\Delta \cdot Q_n^{6(k,l)} = g \overline{\psi}_L \tau_a \mathbf{\Delta} \overline{d}^{l} \overline{d}^k \psi_L \overline{\psi}_L \mathbf{\Delta} \overline{d}^{n-2-k-l} \tau_a \psi_L ,$ $\Delta \cdot Q_n^{7(k)} = \overline{\psi} d^{k*} f \gamma_5 d^{n-1-k} \psi ,$ $\Delta \cdot Q_n^{8(k)} = i \overline{\psi} d^k f d^{n-1-k} \psi ,$ $\Delta \cdot Q_n^{9(k,l)} = g \overline{\psi} \overline{d}^k f_a^{\alpha} (\overline{d}^l f_{\alpha})_a \overline{d}^{n-3-k-l} \Delta \psi ,$ $\Delta \cdot Q_n^{10(k,l)} = igf_{abc} \overline{\psi} \overline{d}^k f_a^{\alpha} (\overline{d}^l f_{\alpha})_b \overline{d}^{n-3-k-l} \Delta \tau_c \psi ,$ $\Delta \cdot Q_n^{11(k,l)} = g d_{abc} \overline{\psi} \overline{d}^k f_a^{\alpha} (\overline{d}^l f_{\alpha})_b \overline{d}^{n-3-k-l} \breve{\Delta} \tau_c \psi ,$ $\Delta \cdot Q_n^{12(k,l)} = ig \overline{\psi} \overline{d}^{k*} f_a^{\alpha} (\overline{d}^l f_{\alpha})_a \overline{d}^{n-3-k-l} \Delta \gamma_5 \psi ,$ $\Delta \cdot Q_n^{13(k,l)} = g f_{abc} \overline{\psi} d^{-k} f_a^{\alpha} (\vec{d}^l f_{\alpha})_b \vec{d}^{n-3-k-l} \Delta \gamma_5 \tau_c \psi ,$ $\Delta \cdot O_n^{14(k,l)} = igd_{abc} \overline{\psi} \overline{d}^{k*} f_a^{\alpha} (\overline{d}^l f_{\alpha})_b \overline{d}^{n-3-k-l} \Delta \gamma_5 \tau_c \psi .$



Data from JLab E94-110 (nucl-ex/0410027, submitted to PRL) Courtesy C Keppel



Probing Higher Twist with PV

PV Deep Ineslastic eD (J Lab 12 GeV)

Early days of PVES to probe strangeness:

Extensive theoretical work to ensure results would be intepretable and meaningful (many thanks to T.W. Donnelly)

We can do the same for PV DIS

5 .5C

Theoretical Challenges pQCD evolution of twist four moments

Lattice QCD for $\tau = 4$ matrix elements

Organizing the program: what kinematics, complementarity with PC F_{1,2}, ...

preliminary

III. General remarks

- PV DIS provides a comprehensive probe of QCD beyond the parton model & possible deviations from the SM EW sector
- Looking beyond the parton model is a natural continuation of the JLab strange quark searches
- Rich set of challenges for theory & experiment: higher twist, CSB in pdfs, d/u...
- Important complement to 12 GeV Moller that would be a focused and powerful probe of EW SM & possible new physics