



# Spin Sum Rules and 3D Nucleon Structure (2/6)

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# Outline

### Lecture 1

- Introduction
- Tour in phase space
- Galileo *vs* Lorentz

### Lecture 2

- Photon point of view
- Galileo vs Lorentz : round 2
- Nucleon 1D picture

# Forms of dynamics

**Space-time foliation Light-front components**  $a^{\pm} = \frac{1}{\sqrt{2}}(a^0 \pm a^3)$ ct ct « Space » = 3D hypersurface Z V « Time » = hypersurface label

#### **Instant-form** dynamics

 $x^0$ Time  $\vec{x}$ Space  $p^0$ Energy Momentum  $\vec{p}$ 

**Light-front form dynamics** 

 $x^+$  $\vec{x}_{\perp}, x^{-}$  $p^{-}$  $\vec{p}_{\perp}, p^+$ 

#### [Dirac (1949)]

### **Ordinary point of view**



Photon point of view

$$x^+ = \frac{1}{\sqrt{2}}(t+z)$$



**Initial frame** 



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**Boosted frame** 



(Quasi) infinite-momentum frame



## Light-front operators

$$\begin{split} K^{1}_{\perp} &= M^{+1} & J^{1}_{\perp} &= M^{-1} \\ &= \frac{1}{\sqrt{2}}(K^{1} + J^{2}) &= \frac{1}{\sqrt{2}}(K^{1} - J^{2}) \\ K^{2}_{\perp} &= M^{+2} & J^{2}_{\perp} &= M^{-2} \\ &= \frac{1}{\sqrt{2}}(K^{2} - J^{1}) &= \frac{1}{\sqrt{2}}(K^{2} + J^{1}) \end{split}$$

Transverse space-time symmetry is Galilean

$$\begin{split} [J^3, J^i_{\perp}] &= i\epsilon^{3ij}J^j_{\perp} \\ [J^3, K^i_{\perp}] &= i\epsilon^{3ij}K^j_{\perp} \\ [K^i_{\perp}, K^j_{\perp}] &= 0 \end{split}$$

$$\begin{split} [K_{\perp}^{i}, P^{-}] &= -iP_{\perp}^{i} \\ [K_{\perp}^{i}, P_{\perp}^{j}] &= -i\delta_{\perp}^{ij}P^{+} \\ [K_{\perp}^{i}, P^{+}] &= [J^{3}, P^{+}] = 0 \end{split}$$

Transverse position operator can be defined !

Longitudinal momentum plays the role of mass in the transverse plane  $P^+ \sim M$ 

Transverse boost  $p'^+ = p^+$   $\vec{p}'_\perp = \vec{p}_\perp + p^+ \vec{v}_\perp$ 

#### [Kogut, Soper (1970)]

### Quasi-probabilistic interpretation

What about the further issues with Special Relativity ?



Conclusion : quasi-probabilistic interpretation is possible !



### Non-relativistic phase space

Localized state in momentum space
$$|\Psi\rangle = |\vec{P}\rangle$$
Commomentum  
momentum $\vec{P} = \sum_{i} \vec{k}_{i}$ in position space $|\Psi\rangle = |\vec{R}\rangle$  $\Box M = |\vec{R}\rangle$  $\Box M = \frac{\sum_{i} m_{i} \vec{r}_{i}}{\sum_{i} m_{i}}$  $= \int \frac{\mathrm{d}^{3}P}{(2\pi)^{3}} e^{-i\vec{P}\cdot\vec{R}} |\vec{P}\rangle$  $\vec{R} = \frac{\sum_{i} m_{i} \vec{r}_{i}}{\sum_{i} m_{i}}$ 

**Equal-time Wigner operator** 

$$\widehat{W}(\vec{k},\vec{r},t) = \int \mathrm{d}k^0 \,\widehat{W}(k,r)$$

Intrinsic phase-space/Wigner distribution

$$\rho_W(\vec{k}, \vec{r}, \not\prec \vec{0}, \vec{0}) = \int \frac{\mathrm{d}^3 \Delta}{(2\pi)^3} \left\langle \frac{\vec{\Delta}}{2} | \widehat{W}(\vec{k}, \vec{r}, t) | - \frac{\vec{\Delta}}{2} \right\rangle$$

Identified with intrinsic variables

Time translation  $\hat{O}(t) = e^{i \hat{H} t} \hat{O}(0) e^{-i \hat{H} t}$ 

[Ji (2003)] [Belitsky, Ji, Yuan (2004)]

### Relativistic phase space

Localized state in momentum space
$$|\Psi\rangle = |P^+, \vec{P}_{\perp}\rangle$$
"CoM " momentum"in position space $|\Psi\rangle = |P^+, \vec{R}_{\perp}\rangle$ "CoM "  $\vec{R}_{\perp} = \sum_{i} \vec{k}_{i\perp}^{+}$  $= \int \frac{d^2 P_{\perp}}{(2\pi)^2} e^{-i\vec{P}_{\perp} \cdot \vec{R}_{\perp}} |P^+, \vec{P}_{\perp}\rangle$ [Soper (1977)]

[Burkardt (2000)] [Burkardt (2003)]

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#### **Equal light-front time Wigner operator**

$$\widehat{W}(k^+, \vec{k}_\perp, r^-, \vec{r}_\perp, r^+) = \int dk^- \, \widehat{W}(k, r)$$
[Ji (2003)]  
[Belitsky, Ji, Yuan (2004)]

#### Intrinsic relativistic phase-space/Wigner distribution

$$\rho_{W}(x,\vec{k}_{\perp},\vec{b}_{\perp}) = \frac{P^{+}}{\langle P^{+}|P^{+}\rangle} \int \mathrm{d}r^{-} \frac{\mathrm{d}^{2}\Delta_{\perp}}{(2\pi)^{2}} \langle P^{+},\frac{\vec{\Delta}_{\perp}}{2}|\widehat{W}(xP^{+},\vec{k}_{\perp},r^{-},\vec{b}_{\perp},r^{+})|P^{+},-\frac{\vec{\Delta}_{\perp}}{2}\rangle$$

$$x = \frac{k^{+}}{P^{+}} = \frac{1}{2} \int \frac{\mathrm{d}^{2}\Delta_{\perp}}{(2\pi)^{2}} \langle P^{+},\frac{\vec{\Delta}_{\perp}}{2}|\widehat{W}(xP^{+},\vec{k}_{\perp},0,\vec{b}_{\perp},0)|P^{+},-\frac{\vec{\Delta}_{\perp}}{2}\rangle$$
post invariant l

**Boost invariant !** 

$$\int \mathrm{d}k^+ f(k^+) = \int \mathrm{d}x \, P^+ f(xP^+)$$

Space-time translation

 $\widehat{O}(r) = e^{i\widehat{P}\cdot r}\,\widehat{O}(0)\,e^{-i\widehat{P}\cdot r}$ Normalization  $\langle P^+|P^+\rangle = 2P^+ 2\pi \,\delta(0)$ 

#### [C.L., Pasquini (2011)]

### Relativistic phase space

### Our intuition is instant form and not light-front form

### <u>**NB</u></u>: \rho\_W(x, \vec{k}\_{\perp}, \vec{b}\_{\perp}) is invariant under light-front boosts</u>**

It can be thought as instant form phase-space/Wigner distribution in IMF !



#### [C.L., Pasquini (2011)]

### Link with parton correlators

#### **Phase-space/Wigner distributions are Fourier transforms**

#### **General parton correlator**

$$W(x,\xi,\vec{k}_{\perp},\vec{\Delta}_{\perp}) = \frac{1}{2} \int \frac{\mathrm{d}z^{-} \,\mathrm{d}^{2} z_{\perp}}{(2\pi)^{3}} \,e^{ik\cdot z} \,\langle p' | \phi^{*}(-\frac{z}{2}) \,\phi(\frac{z}{2}) | p \rangle \big|_{z^{+}=0}$$



### Link with parton correlators

#### Adding spin to the picture



#### Leading twist

$$\gamma^{+} \sim \delta_{\lambda'\lambda}$$
$$\gamma^{+}\gamma_{5} \sim (\tau_{3})_{\lambda'\lambda}$$
$$i\sigma^{j+}\gamma_{5} \sim (\tau_{i})_{\lambda'\lambda}$$



### Link with parton correlators



Adding spin and color to the picture

$$W_{\Lambda'\Lambda}^{[\Gamma]}(x,\xi,\vec{k}_{\perp},\vec{\Delta}_{\perp};\mathcal{W}) = \frac{1}{2} \int \frac{\mathrm{d}z^{-} \,\mathrm{d}^{2}z_{\perp}}{(2\pi)^{3}} \,e^{ik\cdot z} \,\langle p',\Lambda'|\overline{\psi}(-\frac{z}{2})\,\Gamma\,\mathcal{W}_{-\frac{z}{2}\frac{z}{2}}\,\psi(\frac{z}{2})|p,\Lambda\rangle\big|_{z^{+}=0}$$

Wilson line

$$\mathcal{W}_{ba} = \mathcal{P}\left[e^{ig\int_a^b \mathrm{d}x^- A^+}\right]$$

Gauge transformation

$$\psi(x) \mapsto U(x)\psi(x)$$
$$\mathcal{W}_{yx} \mapsto U(y)\mathcal{W}_{yx}U^{-1}(x)$$

Very very very complicated object not directly measurable Let's look for simpler measurable correlators















### Parton distribution functions (PDFs)



**PDF correlator** 

$$\mathcal{F}_{\Lambda'\Lambda}^{[\Gamma]}(x) = \frac{1}{2} \int \frac{\mathrm{d}z^{-}}{2\pi} e^{ixP^{+}z^{-}} \langle P, \Lambda' | \overline{\psi}(-\frac{z}{2}) \Gamma \mathcal{W} \psi(\frac{z}{2}) | P, \Lambda \rangle \Big|_{z^{+}=z_{\perp}=0}$$
$$= \int \mathrm{d}^{2}k_{\perp} W_{\Lambda'\Lambda}^{[\Gamma]}(x, 0, \vec{k}_{\perp}, \vec{0}_{\perp})$$

#### Parametrization

$$\mathcal{F}_{\Lambda'\Lambda}^{[\Gamma]}(x) = \sum_{i} \left[ \overline{u}(P,\Lambda') \,\Gamma_i \, u(P,\Lambda) \right] \, \mathrm{PDF}_i(x)$$

Constrained by Lorentz and discrete space-time symmetries

# Parton distribution functions (PDFs)













### Summary

### Lecture 2

- Light-front coordinates are very convenient
- Relativistic phase space has 2+3 dimensions
- PDFs provide 1D pictures of the nucleon



$$PDF(x) = \int d^2k_{\perp} d^2b_{\perp} \rho_W(x, \vec{k}_{\perp}, \vec{b}_{\perp})$$

