

# Beyond the Standard Model

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# Plan of the Lectures

- The Standard Model and issues beyond it
- Origin of particle masses: Higgs boson or?
- Supersymmetry
- Searches for supersymmetry: LHC & dark matter
- Extra dimensions and string theory

## Beyond the Standard Model for Montañeros

*M. Bustamante, L. Cieri and John Ellis*

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<sup>3</sup> Theory Division, Physics Department, CERN, CH 1211 Geneva 23, Switzerland

<http://arXiv.org/pdf/0911.4409.pdf>

# Summary of the Standard Model

- Particles and  $SU(3) \times SU(2) \times U(1)$  quantum numbers:

$L_L$	$\begin{pmatrix} \nu_e \\ e^- \end{pmatrix}_L, \begin{pmatrix} \nu_\mu \\ \mu^- \end{pmatrix}_L, \begin{pmatrix} \nu_\tau \\ \tau^- \end{pmatrix}_L$	$(1,2,-1)$
$E_R$	$e_R^-, \mu_R^-, \tau_R^-$	$(1,1,-2)$
$Q_L$	$\begin{pmatrix} u \\ d \end{pmatrix}_L, \begin{pmatrix} c \\ s \end{pmatrix}_L, \begin{pmatrix} t \\ b \end{pmatrix}_L$	$(3,2,+1/3)$
$U_R$	$u_R, c_R, t_R$	$(3,1,+4/3)$
$D_R$	$d_R, s_R, b_R$	$(3,1,-2/3)$

- Lagrangian:
 

$\mathcal{L} = -\frac{1}{4} F_{\mu\nu}^a F^{a\ \mu\nu}$ $+ i\bar{\psi} \not{D}\psi + h.c.$ $+ \psi_i y_{ij} \psi_j \phi + h.c.$ $+  D_\mu \phi ^2 - V(\phi)$	gauge interactions matter fermions Yukawa interactions Higgs potential
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# Gauge Interactions of the Standard Model

- Three separate gauge group factors:
  - $SU(3) \times SU(2) \times U(1)$
  - Strong  $\times$  electroweak
- Three different gauge couplings:
  - $g_3, g_2, g'$
- Mixing between the  $SU(2)$  and  $U(1)$  factors:

$$\begin{pmatrix} Z^\mu \\ A^\mu \end{pmatrix} = \begin{pmatrix} \cos(\theta_W) & \sin(\theta_W) \\ -\sin(\theta_W) & \cos(\theta_W) \end{pmatrix} \begin{pmatrix} W_3^\mu \\ B^\mu \end{pmatrix} \quad \sin^2(\theta_W) = \frac{g'^2}{g'^2 + g^2}$$

- Experimental value:  $\sin^2\theta_W = 0.23120 \pm 0.00015$

Important clue for Grand Unification



# Weak Interactions

- Interactions of lepton doublets:  $L = \begin{pmatrix} \nu_e \\ e^- \end{pmatrix}_L$
- Charged-current interactions:

$$\mathcal{L}_{cc} = \frac{-g}{\sqrt{2}} \sum_{\alpha=e,\mu,\tau} \nu_{L\alpha} \gamma_\mu l_{L\alpha} W^\mu + h.c.$$

- Neutral-current interactions:

$$\mathcal{L}_{nc} = \frac{-g}{2 \cos \theta_W} \sum_{\alpha=e,\mu,\tau} \nu_{L\alpha} \gamma_\mu l_{L\alpha} Z^\mu + h.c.$$

- Effective four-fermion interaction:

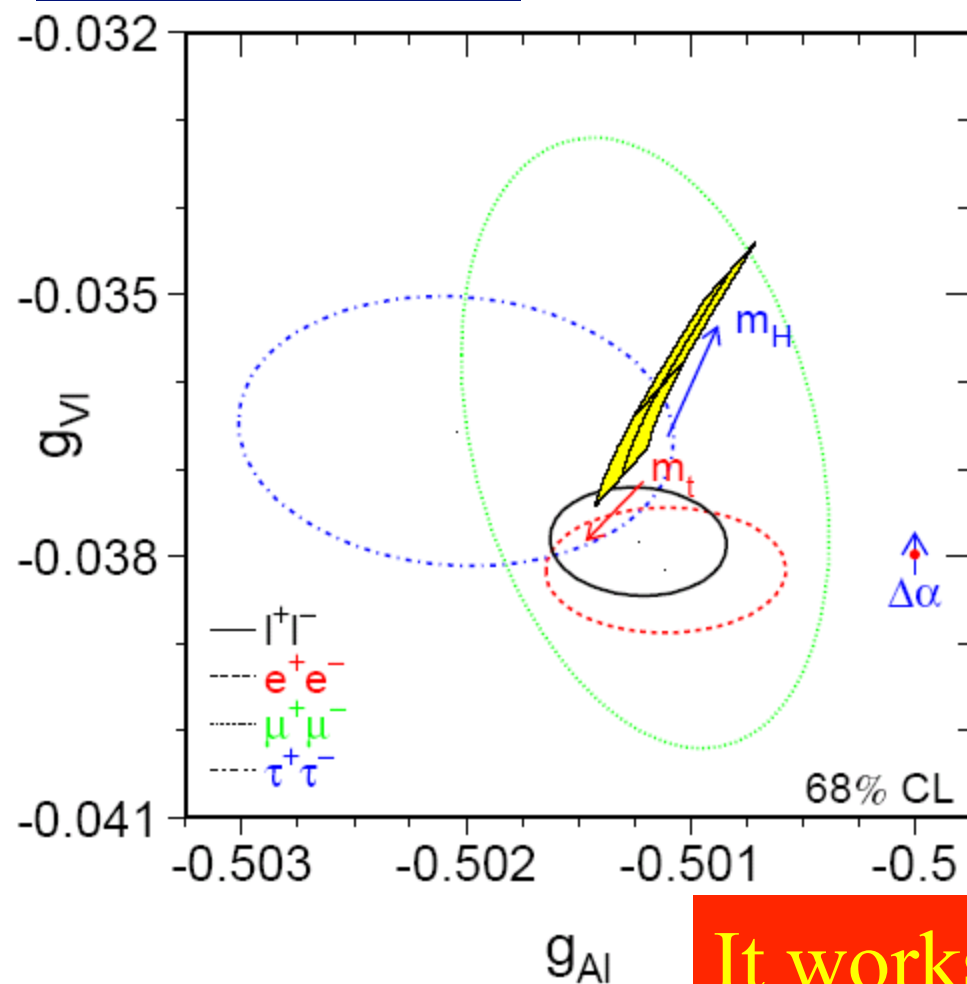
$$G_F/\sqrt{2} = g^2/8m_W^2 \quad \mathcal{L} = \frac{G_F}{\sqrt{2}} \bar{\nu}_\mu \gamma^\alpha (1 - \gamma^5) \mu \bar{e} \gamma_\alpha (1 - \gamma^5) \nu_e$$

# Status of the Standard Model

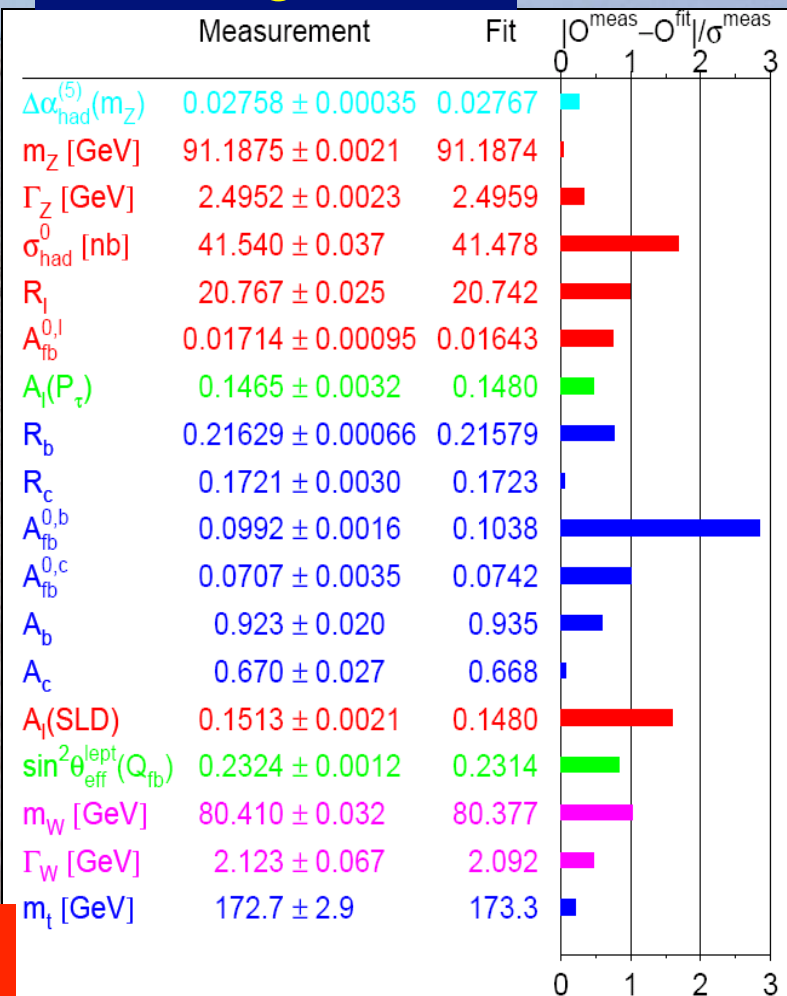
- Perfect agreement with all *confirmed* accelerator data
- Consistency with precision electroweak data (LEP et al) *only if there is a Higgs boson*
- Agreement seems to require *a relatively light Higgs boson* weighing  $< \sim 180 \text{ GeV}$
- Raises many unanswered questions:  
*mass? flavour? unification?*

# Precision Tests of the Standard Model

## Lepton couplings



## Pulls in global fit



It works!

# Parameters of the Standard Model

- Gauge sector:
  - 3 gauge couplings:  $g_3, g_2, g'$
  - 1 strong CP-violating phase
- Yukawa interactions:
  - 3 charge-lepton masses
  - 6 quark masses
  - 4 CKM angles and phase
- Higgs sector:
  - 2 parameters:  $\mu, \lambda$
- **Total: 19 parameters**

Unification?

Flavour?

Mass?



# Open Questions beyond the Standard Model

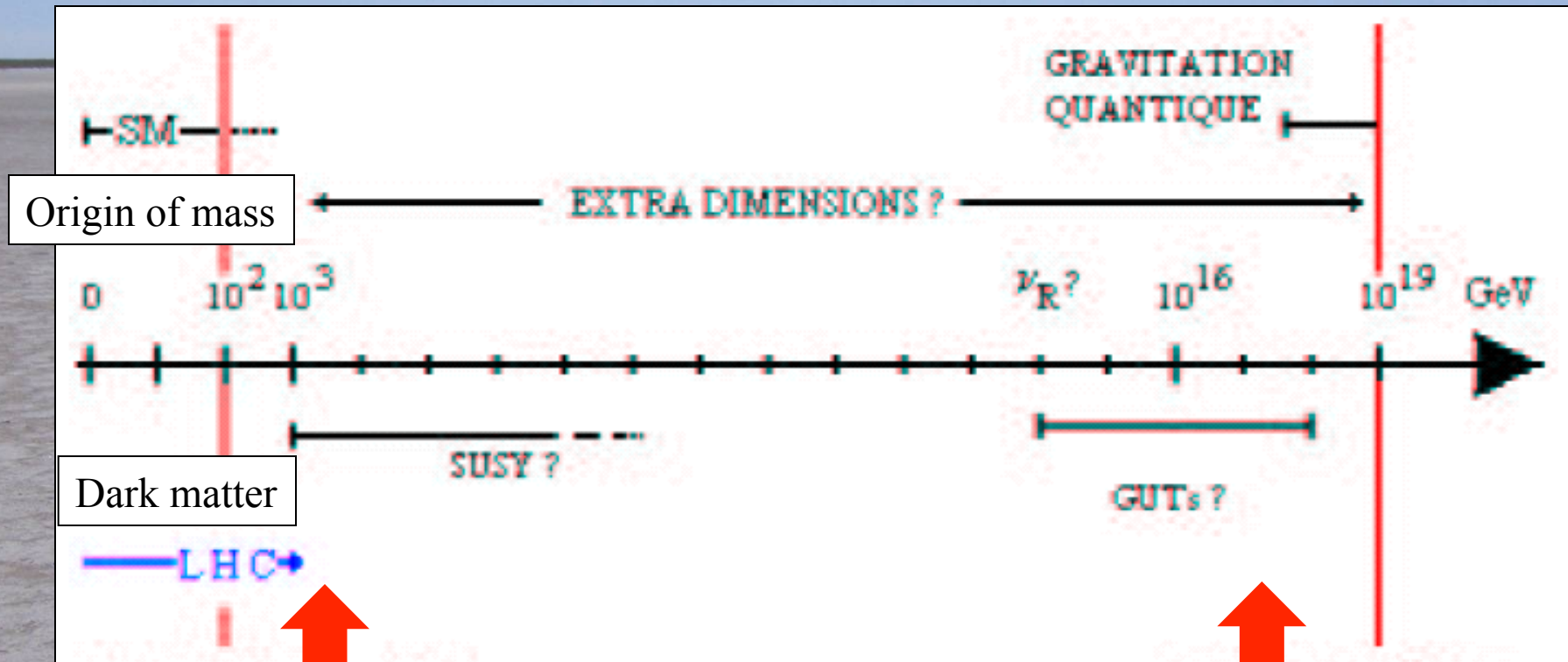
- What is the origin of particle masses?  
due to a Higgs boson? + other physics?  
solution at energy  $< 1 \text{ TeV}$  (1000 GeV)
- Why so many types of matter particles?  
matter-antimatter difference?
- Unification of the fundamental forces?  
at very high energy  $\sim 10^{16} \text{ GeV}$ ?  
probe directly via neutrino physics, indirectly via masses, couplings
- Quantum theory of gravity?  
(super)string theory: extra space-time dimensions?

Supersymmetry

Supersymmetry

Supersymmetry

# At what Energy is the New Physics?



A lot accessible  
to the LHC?

Some accessible only indirectly:  
Astrophysics & cosmology?

# Why do Things Weigh?

Newton:

Weight **proportional to** Mass

Einstein:

Energy **related to** Mass

Neither explained origin of Mass

Where do the masses  
come from?

Are masses due to Higgs boson?  
(the physicists' Holy Grail)



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# The Seminal Papers

## BROKEN SYMMETRY AND THE MASS OF GAUGE VECTOR MESONS\*

F. Englert and R. Brout

Faculté des Sciences, Université Libre de Bruxelles, Bruxelles, Belgium

(Received 26 June 1964)

## BROKEN SYMMETRIES, MASSLESS PARTICLES AND GAUGE FIELDS

P. W. HIGGS

*Tait Institute of Mathematical Physics, University of Edinburgh, Scotland*

Received 27 July 1964

VOLUME 13, NUMBER 16

PHYSICAL REVIEW LETTERS

19 OCTOBER 1964

## BROKEN SYMMETRIES AND THE MASSES OF GAUGE BOSONS

Peter W. Higgs

Tait Institute of Mathematical Physics, University of Edinburgh, Edinburgh, Scotland

(Received 31 August 1964)

## GLOBAL CONSERVATION LAWS AND MASSLESS PARTICLES\*

G. S. Guralnik,<sup>†</sup> C. R. Hagen,<sup>‡</sup> and T. W. B. Kibble

Department of Physics, Imperial College, London, England

(Received 12 October 1964)



# The Englert-Brout-Higgs Mechanism

- **Vacuum expectation value of scalar field**
- Englert & Brout: June 26<sup>th</sup> 1964
- First Higgs paper: July 27<sup>th</sup> 1964
- Pointed out loophole in argument of Gilbert if gauge theory described in Coulomb gauge
- Accepted by Physics Letters
- Second Higgs paper with explicit example sent on July 31<sup>st</sup> 1964 to Physics Letters, rejected!
- Revised version (Aug. 31<sup>st</sup> 1964) accepted by PRL
- Guralnik, Hagen & Kibble (Oct. 12<sup>th</sup> 1964)

# The Englert-Brout-Higgs Mechanism

## Englert & Brout



## Guralnik, Hagen & Kibble

We consider, as our example, a theory which was partially solved by Englert and Brout,<sup>5</sup> and bears some resemblance to the classical theory of Higgs.<sup>6</sup> Our starting point is the ordinary electrodynamics of massless spin-zero particles, characterized by the Lagrangian

$$\mathcal{L} = -\frac{1}{2}F^{\mu\nu}(\partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}) + \frac{1}{4}F^{\mu\nu}F_{\mu\nu} \\ + \varphi^{\mu}\partial_{\mu}\varphi + \frac{1}{2}\varphi^{\mu}\varphi_{\mu} + ie_0\varphi^{\mu}q\varphi A_{\mu},$$

With no loss of generality, we can take  $\eta_2 = 0$ , and find

$$(-\partial^2 + \eta_1^2)\varphi_1 = 0, \\ -\partial^2\varphi_2 = 0, \\ (-\partial^2 + \eta_1^2)A_k^T = 0,$$

where the superscript  $T$  denotes the transverse part. The two degrees of freedom of  $A_k^T$  combine with  $\varphi_1$  to form the three components of a

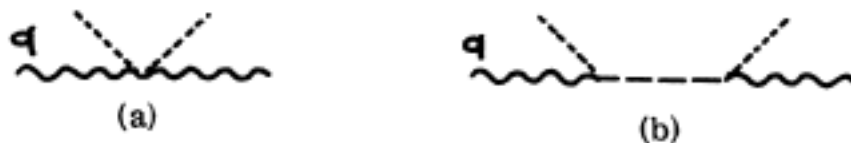


FIG. 1. Broken-symmetry diagram leading to a mass for the gauge field. Short-dashed line,  $\langle\varphi_1\rangle$ ; long-dashed line,  $\varphi_2$  propagator; wavy line,  $A_{\mu}$  propagator. (a)  $\rightarrow (2\pi)^4 ie^2 g_{\mu\nu} \langle\varphi_1\rangle^2$ , (b)  $\rightarrow -(2\pi)^4 ie^2 (q_{\mu}q_{\nu}/q^2) \times \langle\varphi_1\rangle^2$ .

# The Higgs Boson

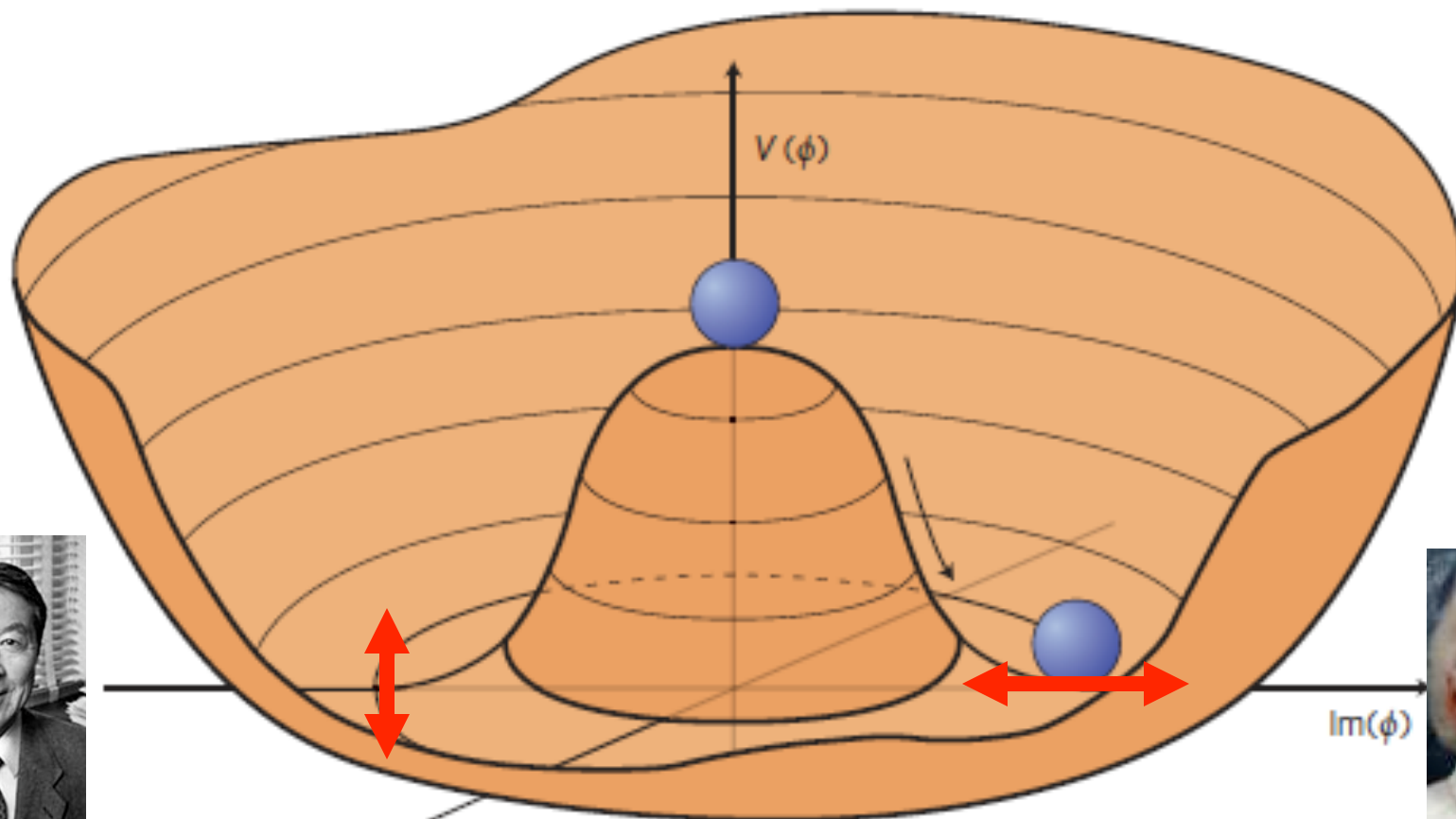
- **Higgs pointed out a massive scalar boson**

$$\{\partial^2 - 4\varphi_0^2 V''(\varphi_0^2)\}(\Delta\varphi_2) = 0, \quad (2b)$$

Equation (2b) describes waves whose quanta have  
(bare) mass  $2\varphi_0\{V''(\varphi_0^2)\}^{1/2}$ .

- “... an essential feature of [this] type of theory ... is the prediction of incomplete multiplets of vector and scalar bosons”
- Englert, Brout, Guralnik, Hagen & Kibble did not comment on its existence
- **Discussed in detail by Higgs in 1966 paper**

# Nambu **EB, GHK** and Higgs



Spontaneous symmetry breaking: massless Nambu-Goldstone boson **'eaten'** by gauge boson

**Accompanied by massive particle**



# Without Higgs ...

... there would be no atoms

- Electrons would escape at the speed of light

... weak interactions would not be weak

- Life would be impossible: there would be no nuclei, everything would be radioactive

How does the Higgs trick work?

# Masses for Gauge Bosons

- Kinetic terms for SU(2) and U(1) gauge bosons:

$$\mathcal{L} = -\frac{1}{4} G_{\mu\nu}^i G^{i\mu\nu} - \frac{1}{4} F_{\mu\nu} F^{\mu\nu}$$

where  $G_{\mu\nu}^i \equiv \partial_\mu W_\nu^i - \partial_\nu W_\mu^i + ig\epsilon_{ijk} W_\mu^j W_\nu^k$   $F_{\mu\nu} \equiv \partial_\mu W_\nu - \partial_\nu W_\mu$

- Kinetic term for Higgs field:

$$\mathcal{L}_\phi = -|D_\mu \phi|^2 \quad D_\mu \equiv \partial_\mu - i g \sigma_i W_\mu^i - i g' Y B_\mu$$

- Expanding around vacuum:  $\phi = \langle 0|\phi|0 \rangle + \hat{\phi}$

$$\mathcal{L}_\phi \ni -\frac{g^2 v^2}{2} W_\mu^+ W^{\mu-} - g'^2 \frac{v^2}{2} B_\mu B^\mu + g g' v^2 B_\mu W^{\mu 3} - g^2 \frac{v^2}{2} W_\mu^3 W^{\mu 3}$$

- Boson masses:

$$m_{W^\pm} = \frac{gv}{2}$$

$$Z_\mu = \frac{gW_\mu^3 - g'B_\mu}{\sqrt{g^2 + g'^2}} : \quad m_Z = \frac{1}{2}\sqrt{g^2 + g'^2}v ; \quad A_\mu = \frac{g'W_\mu^3 + gB_\mu}{\sqrt{g^2 + g'^2}} : \quad m_A = 0$$

# Constraints on Higgs Mass

- Electroweak observables sensitive via quantum loop corrections:

$$m_W^2 \sin^2 \theta_W = m_Z^2 \cos^2 \theta_W \sin^2 \theta_W = \frac{\pi\alpha}{\sqrt{2}G_F}(1 + \Delta r)$$

- Sensitivity to top, Higgs masses:

$$\frac{3G_F}{8\pi^2\sqrt{2}}m_t^2$$

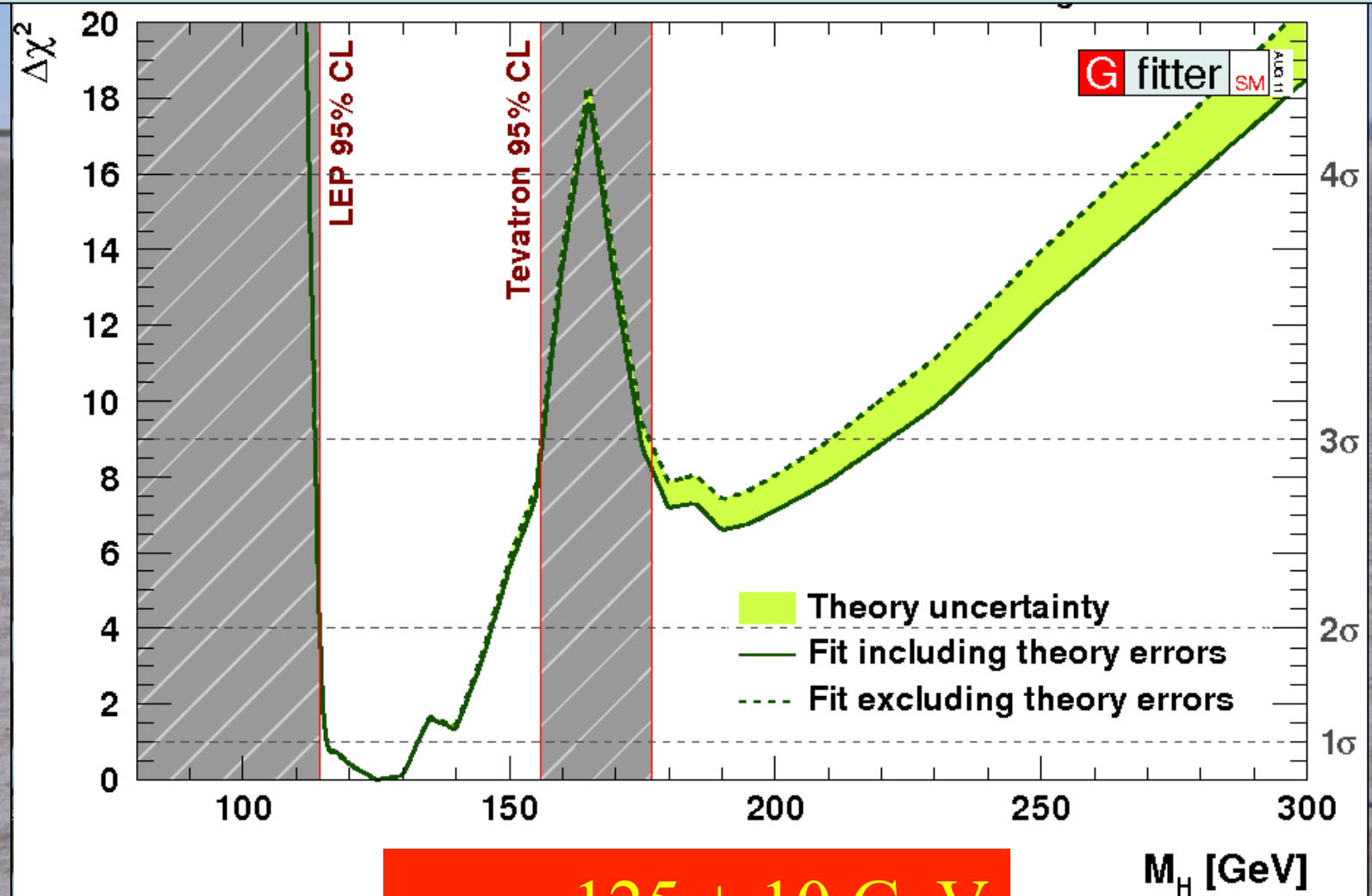
$$\frac{\sqrt{2}G_F}{16\pi^2}m_W^2\left(\frac{11}{3}\ln\frac{M_H^2}{m_Z^2} + \dots\right), M_H \gg m_W$$

- Preferred Higgs mass:  $m_H \sim 100 \pm 30 \text{ GeV}$
- Compare with lower limit from direct searches:

$$m_H > 114 \text{ GeV}$$

- No conflict!

# 2011: Combining Information from Previous Direct Searches and Indirect Data



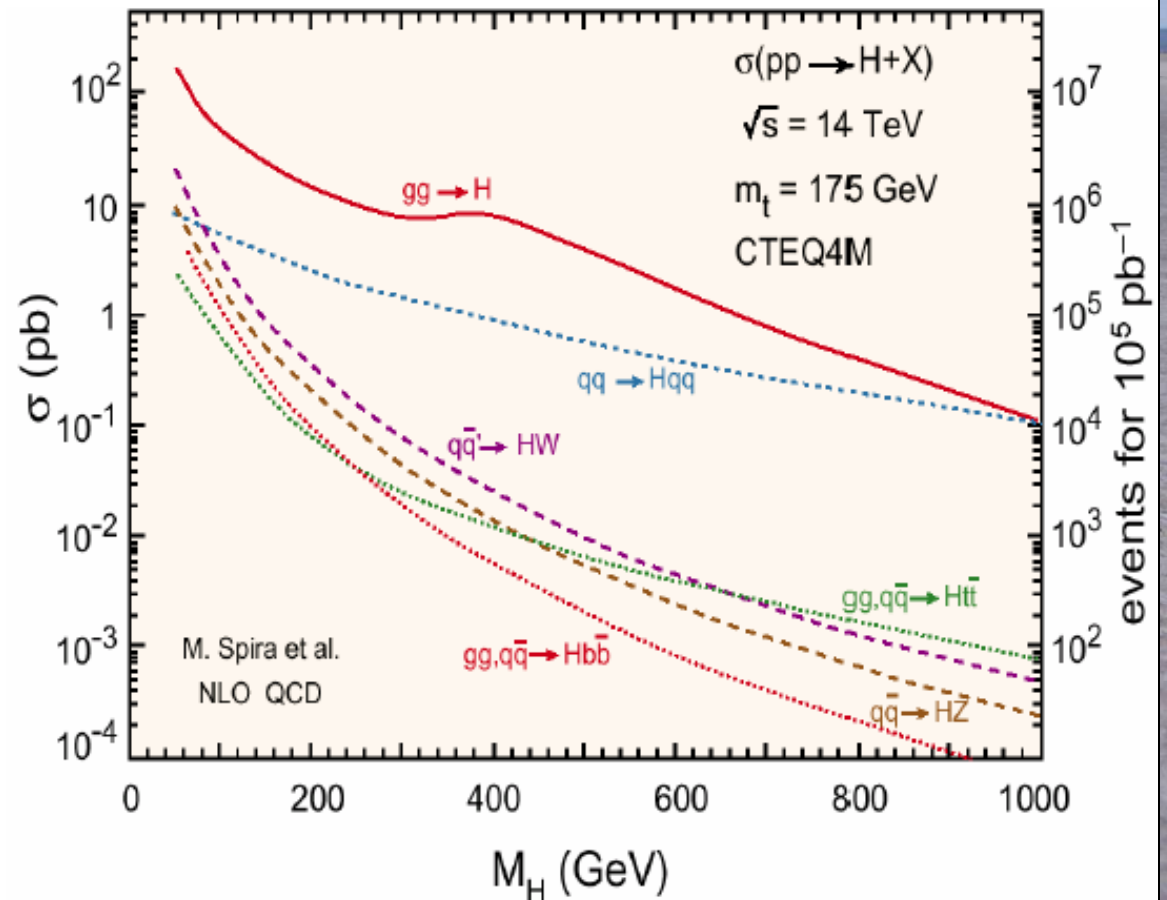
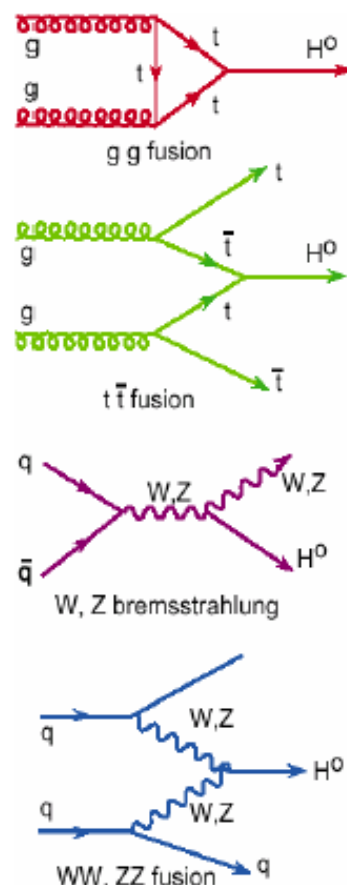
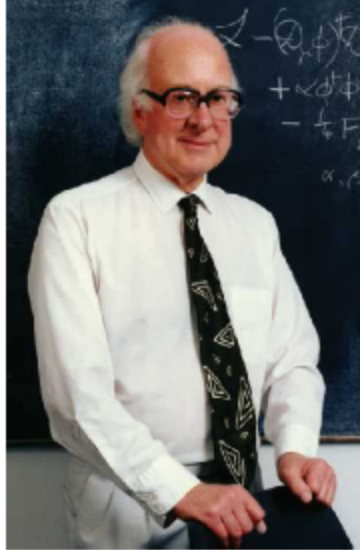
$$m_H = 125 \pm 10 \text{ GeV}$$

Gitter collaboration



A la  
recherche du  
Higgs perdu  
...

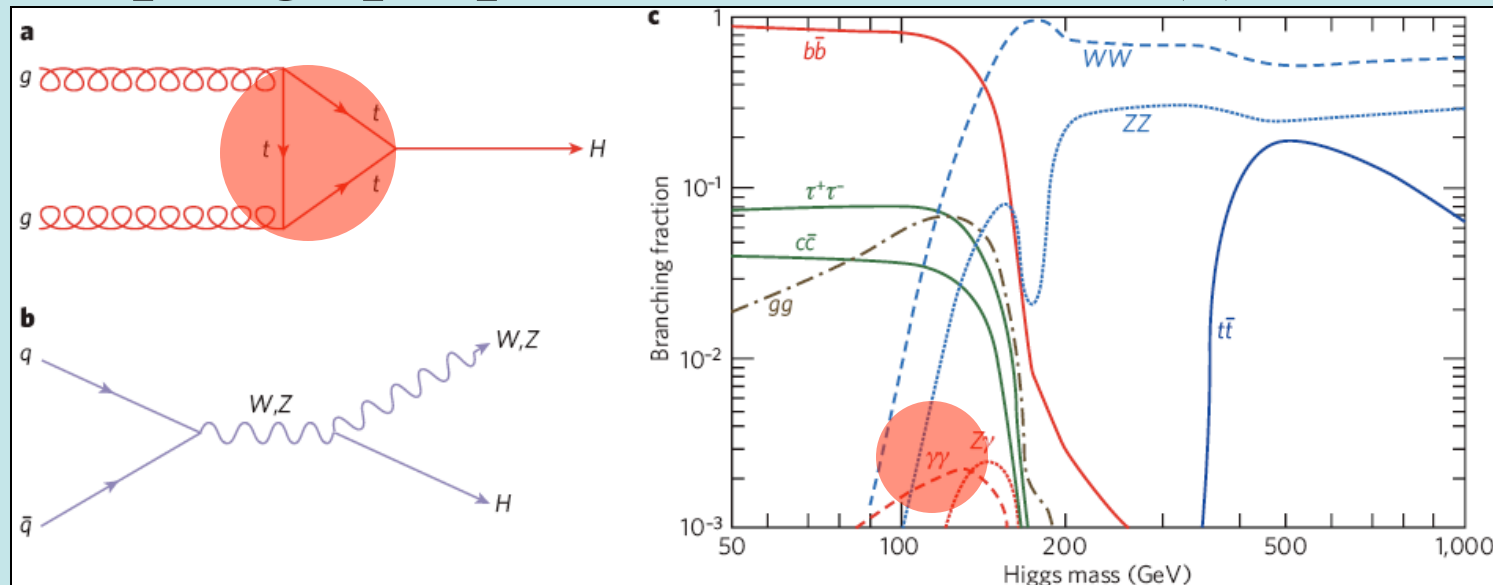
# Higgs Production at the LHC



Many production modes measurable if  $M_h \sim 125 \text{ GeV}$

# Higgs Decay Branching Ratios

- Couplings proportional to masses (?)



- Important couplings through loops:
  - gluon + gluon  $\rightarrow$  Higgs  $\rightarrow \gamma\gamma$

Many decay modes measurable if  $M_h \sim 125$  GeV

# There must be New Physics Beyond the Standard Model

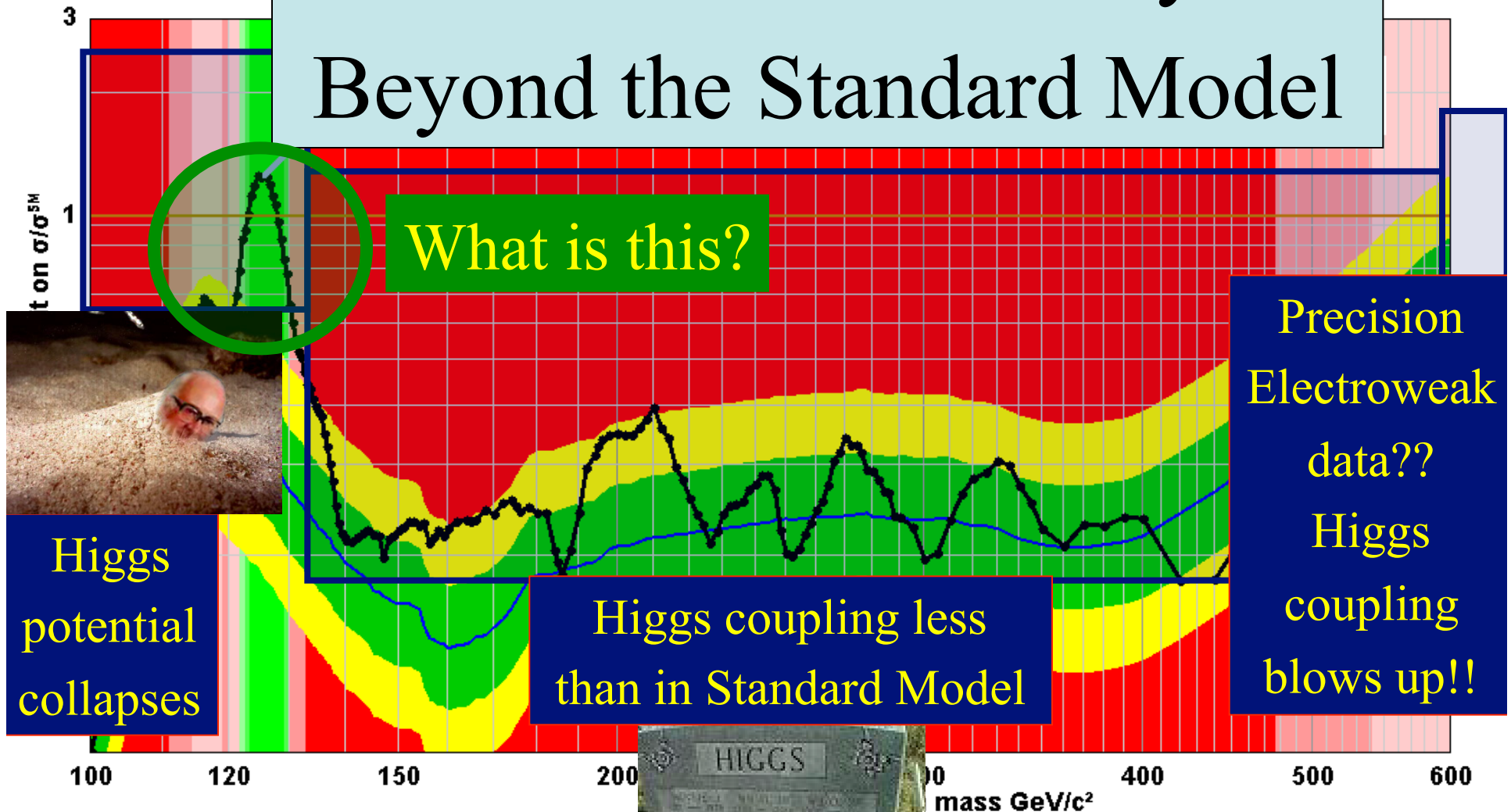
2012

What is this?

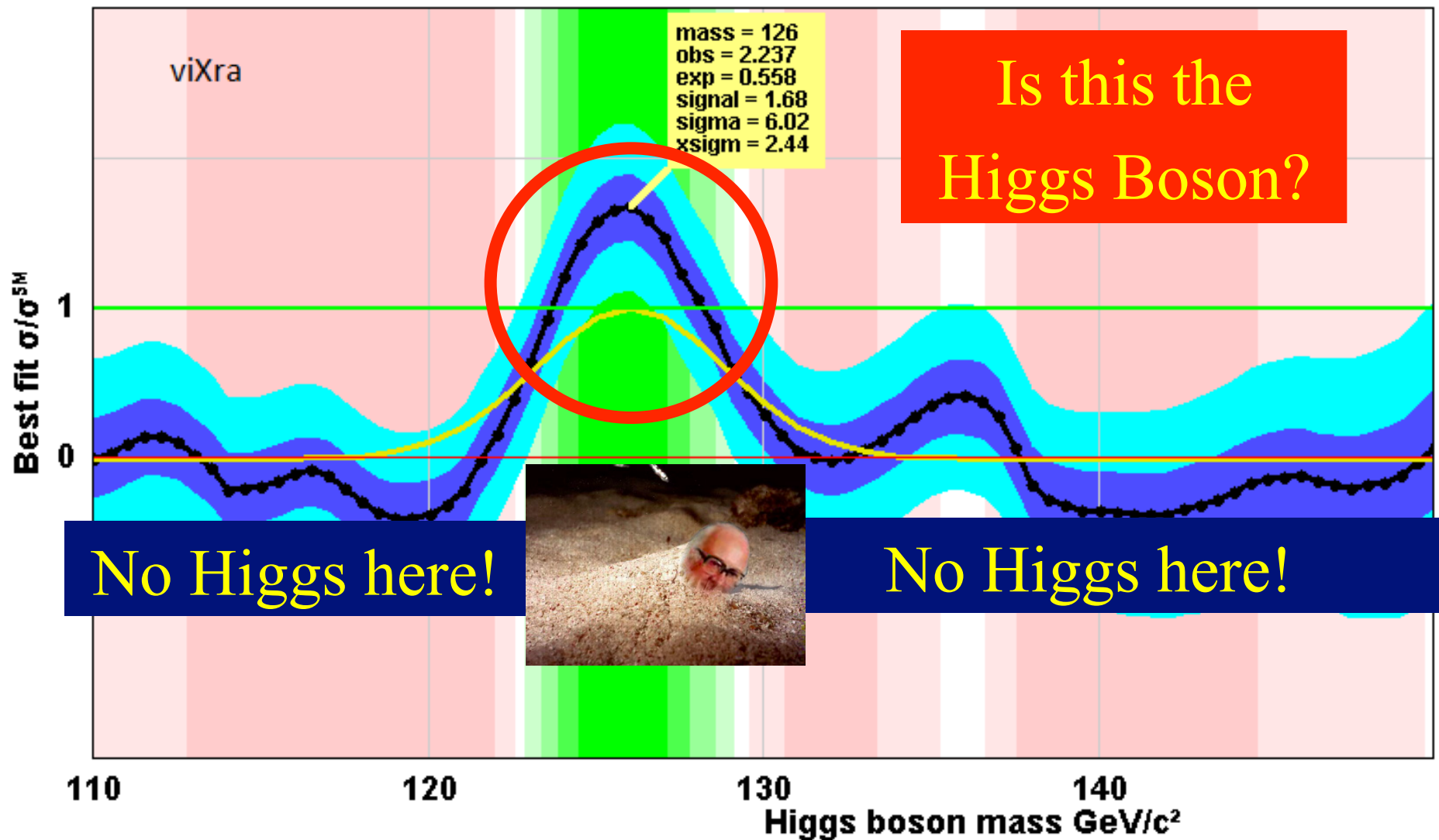
Precision  
Electroweak  
data??  
Higgs  
coupling  
blows up!!

Higgs coupling less  
than in Standard Model

Higgs  
potential  
collapses



# Unofficial Combination of Higgs Search Data from July 4th





# The Particle Higgsaw Puzzle

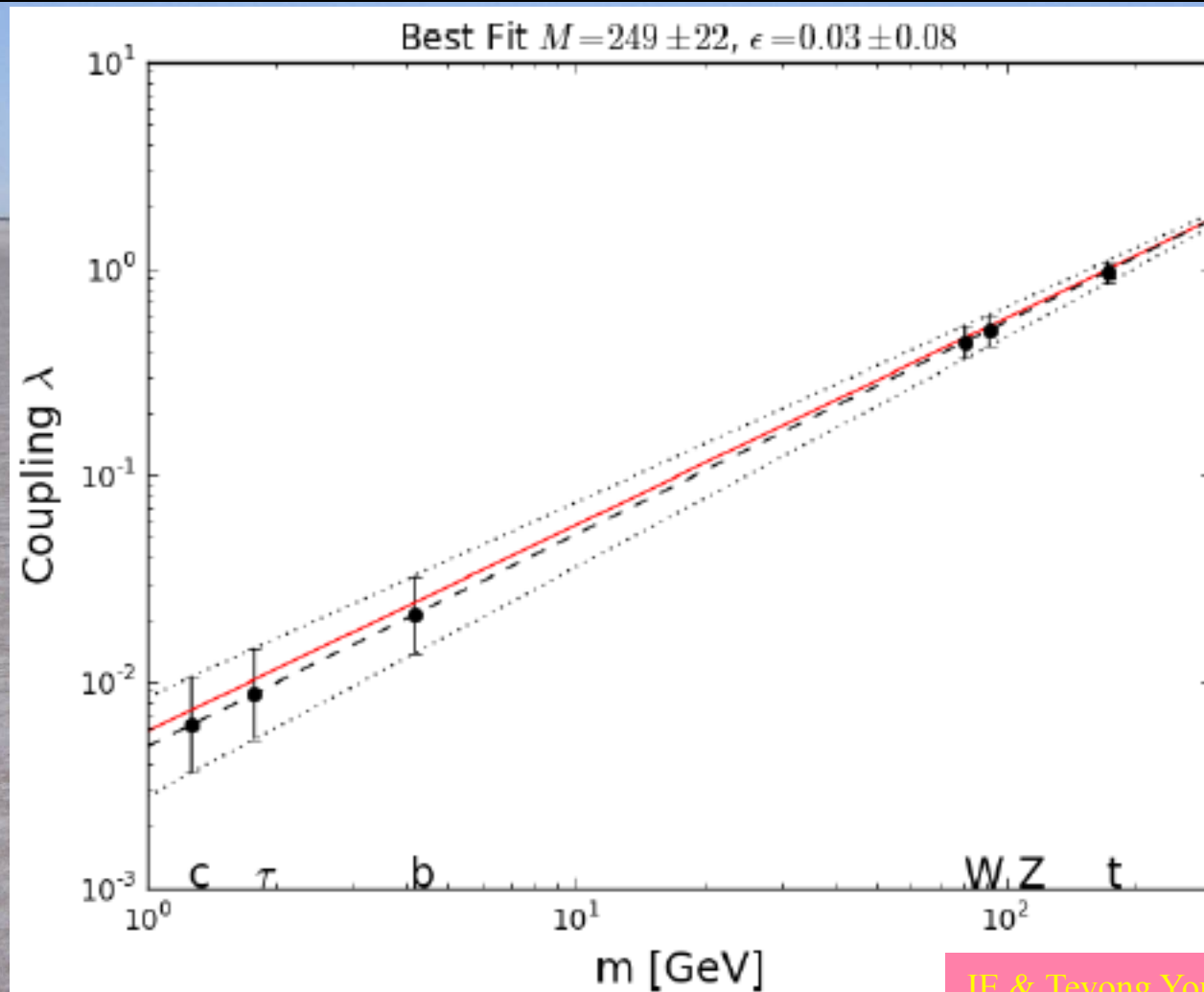
The background of the slide is a blue puzzle. In the center, one piece is missing, revealing a white surface underneath. The puzzle pieces are a light blue color with a wavy, interlocking pattern. The missing piece is located in the lower-middle section of the frame.

Is LHC finding the missing piece?  
Is it the right shape?  
Is it the right size?

# Do we already know the ‘Higgs’ has Spin Zero ?

- Decays into  $\gamma\gamma$ , so cannot have spin 1
- **Spin 0 or 2?**
- If it decays into  $\tau\tau$  or  $b\text{-bar}$ : spin 0 or 1 or orbital angular momentum
- Can diagnose spin via
  - angular distribution of  $\gamma\gamma$
  - angular correlations of leptons in  $WW$ ,  $ZZ$  decays
- Does selection of  $WW$  events mean spin 0?

# Couplings $\sim$ Mass

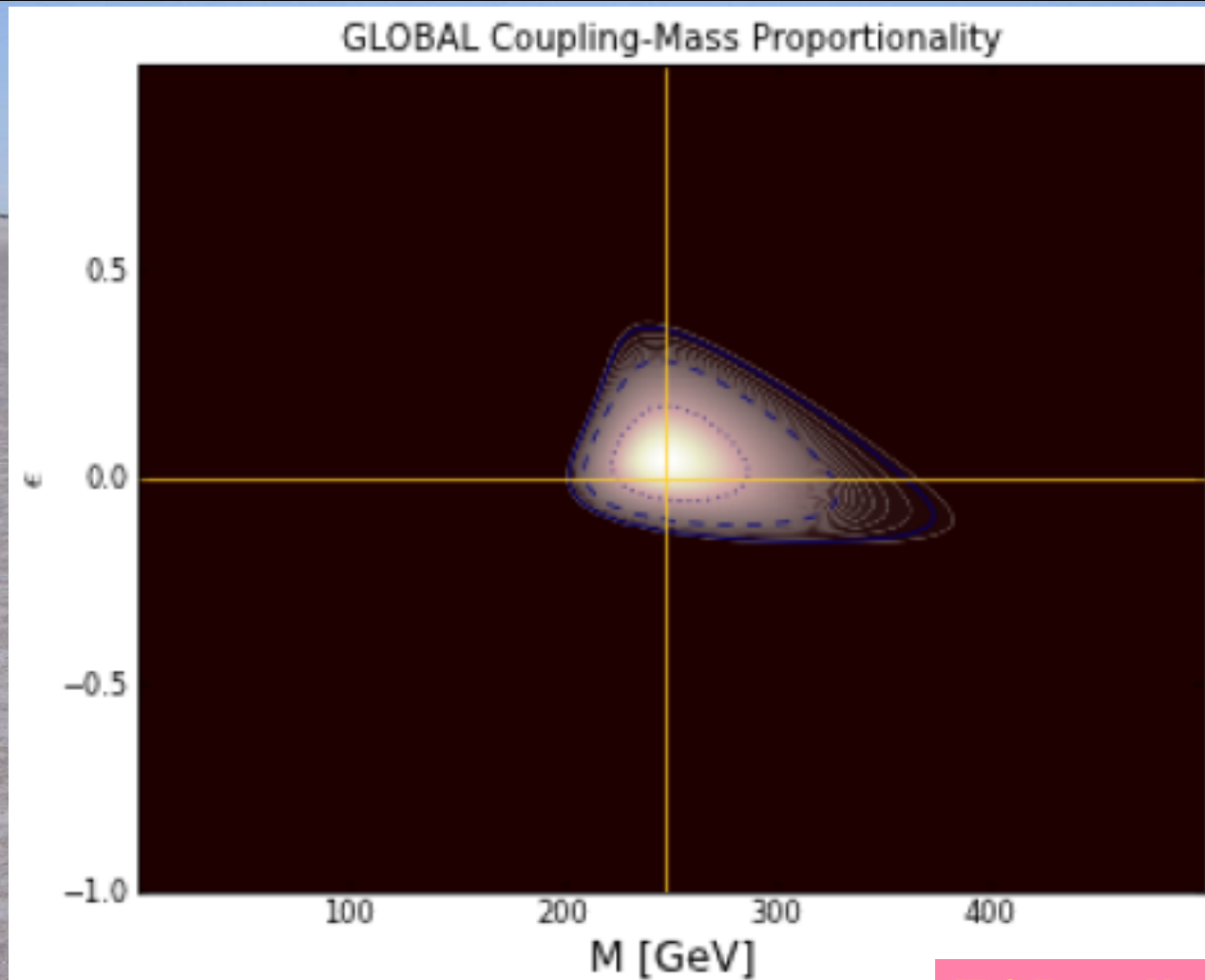


JE & Tevong You: arXiv:1207.1693

Scale as expected for Higgs Boson



# Couplings $\sim$ Mass



JE & Tevong You: arXiv:1207.1693

Scale as expected for Higgs Boson



Imagine a  
Room ...



... Open  
The Door

# What lies Beyond?





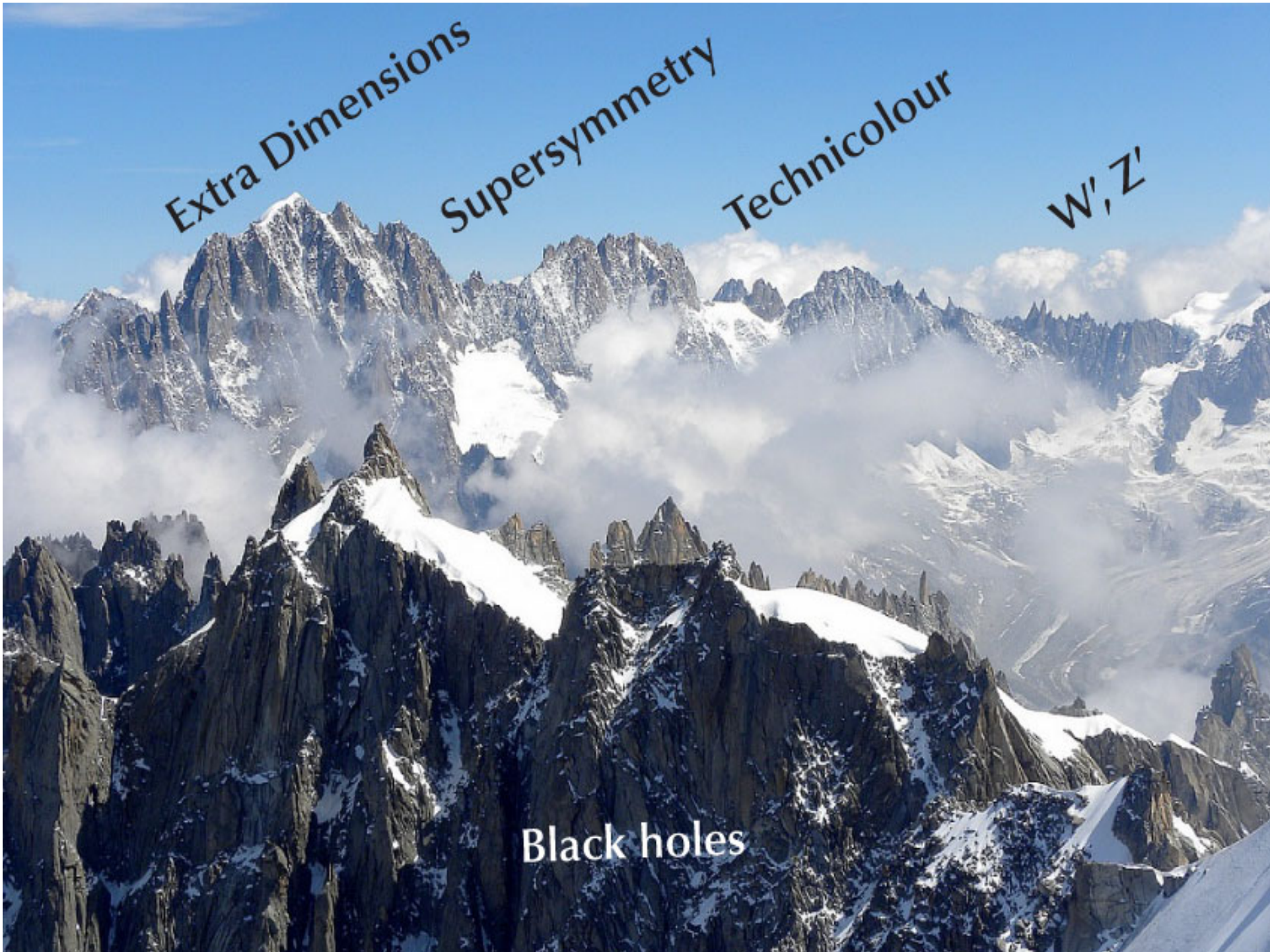
Extra Dimensions

Supersymmetry

Technicolour

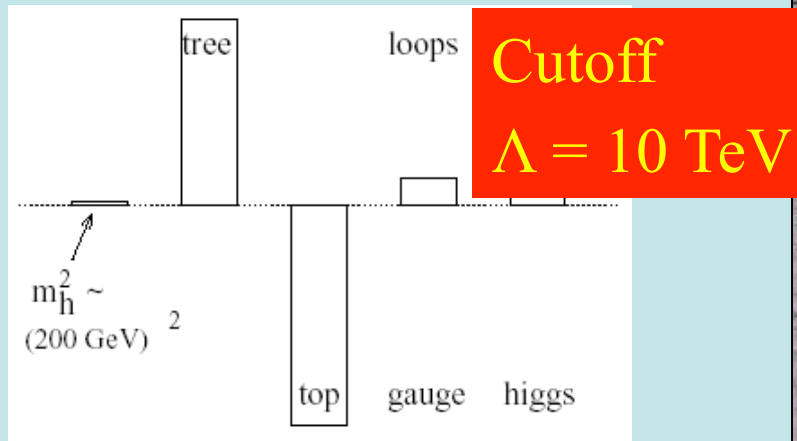
$W', Z'$

Black holes



# Elementary Higgs or Composite?

- Higgs field:  
 $\langle 0|H|0\rangle \neq 0$
- Quantum loop problems



Cut-off  $\Lambda \sim 1 \text{ TeV}$  with  
Supersymmetry?

- Fermion-antifermion condensate
- Just like QCD, BCS superconductivity
- Top-antitop condensate?  
needed  $m_t > 200 \text{ GeV}$

New technicolour force?  
-Heavy scalar resonance?  
-Inconsistent with  
precision electroweak data?



# Interpolating Models

- Combination of Higgs boson and vector  $\rho$



- Two main parameters:  $m_\rho$  and coupling  $g_\rho$
- Equivalently ratio weak/strong scale:

$$g_\rho / m_\rho$$

# General Analysis of Generalized Higgs Models

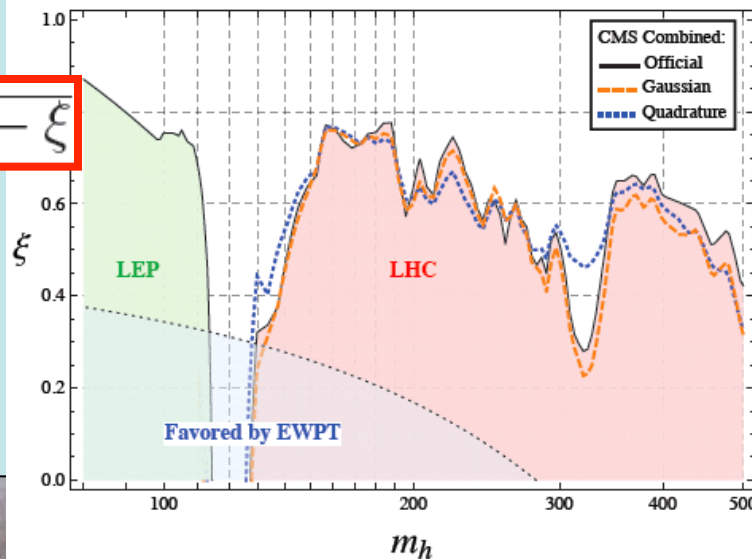
- Parameterization of effective Lagrangian:

$$\mathcal{L}^{(2)} = \frac{1}{2}(\partial_\mu h)^2 + \frac{v^2}{4} \text{Tr} (D_\mu \Sigma^\dagger D^\mu \Sigma) \left( 1 - 2a \frac{h}{v} + b \frac{h^2}{v^2} + \dots \right) \\ - \frac{v}{\sqrt{2}} \lambda_{ij}^u (\bar{u}_L^{(i)}, \bar{d}_L^{(i)}) \Sigma (u_R^{(i)}, 0)^T \left( 1 + c_u \frac{h}{v} + c_{2u} \frac{h^2}{v^2} + \dots \right)$$

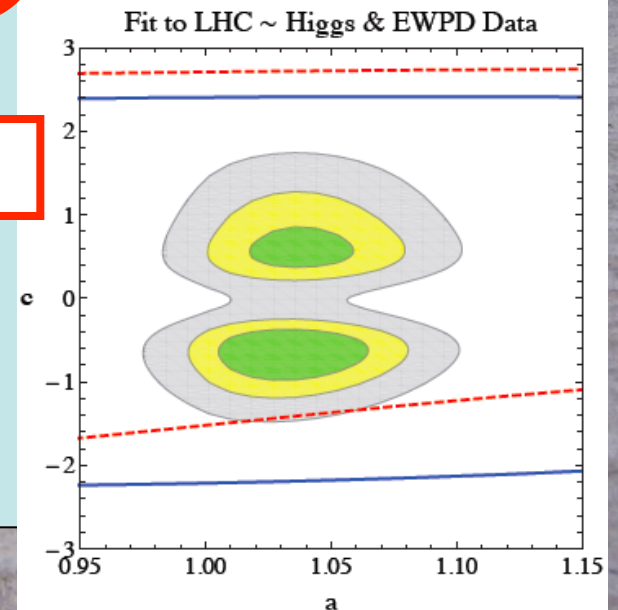
Universal Rescaling: 95% CL Exclusions

- Fits

$$a = c = \sqrt{1 - \xi}$$



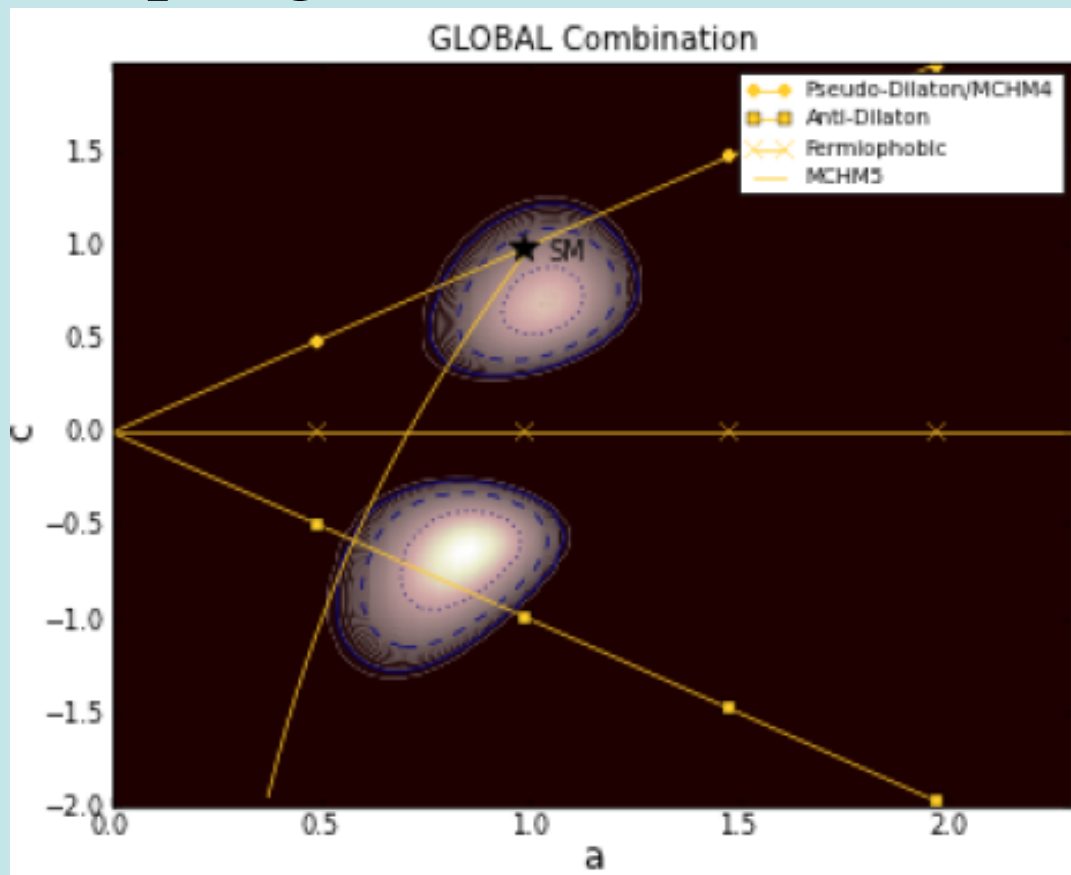
$$a \neq c$$



# Combination of LHC & Tevatron

- Rescale couplings: to bosons by  $a$ , to fermions by  $c$

Global  
analysis



- Standard Model:  $a = c = 1$

JE & Tevong You: arXiv:1207.1693

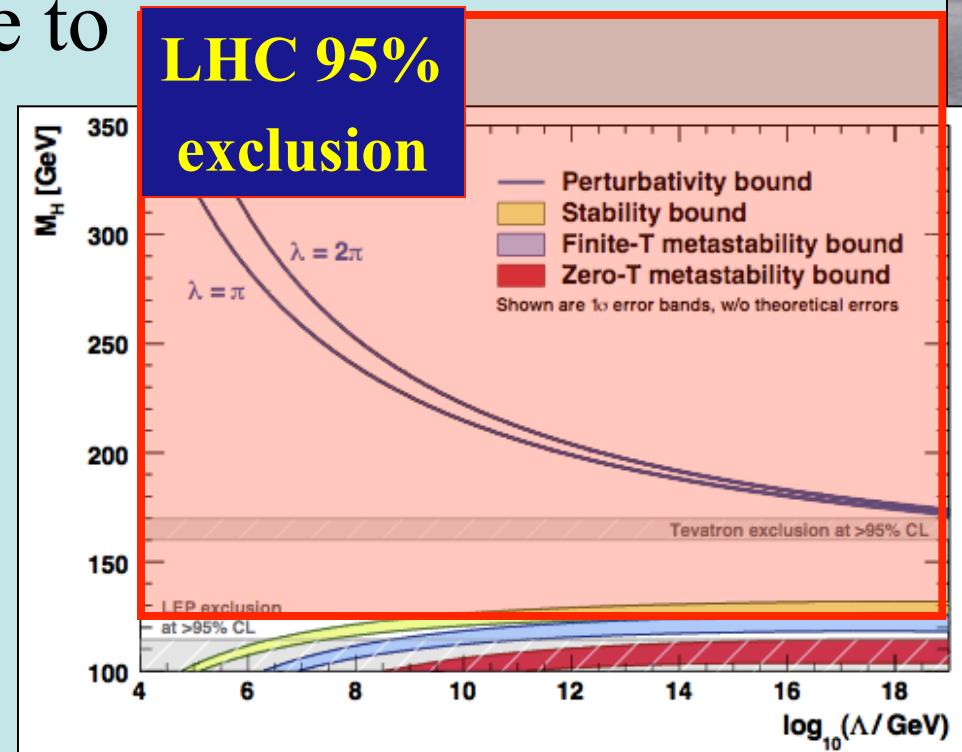
# The Stakes in the Higgs Search

- How is gauge symmetry broken?
- Is there any elementary scalar field?
- **Likely portal to new physics**
- Would have caused phase transition in the Universe when it was about  $10^{-12}$  seconds old
- May have generated then the matter in the Universe: **electroweak baryogenesis**
- A related **inflaton** might have expanded the Universe when it was about  $10^{-35}$  seconds old
- Contributes to today's **dark energy:  $10^{60}$  too much!**



# Theoretical Constraints on Higgs Mass

- Large  $M_h \rightarrow$  large self-coupling  $\rightarrow$  blow up at low-energy scale  $\Lambda$  due to renormalization
- Small: renormalization due to  $t$  quark drives quartic coupling  $< 0$  at some scale  $\Lambda \rightarrow$  vacuum unstable
- Vacuum could be stabilized by **Supersymmetry**



A photograph of a whale breaching the ocean surface, creating a splash. In the background, there are large, hazy mountains under a clear sky. A small bird is visible in flight above the whale.

# Elements of Statistics

*Fundamental Concepts*

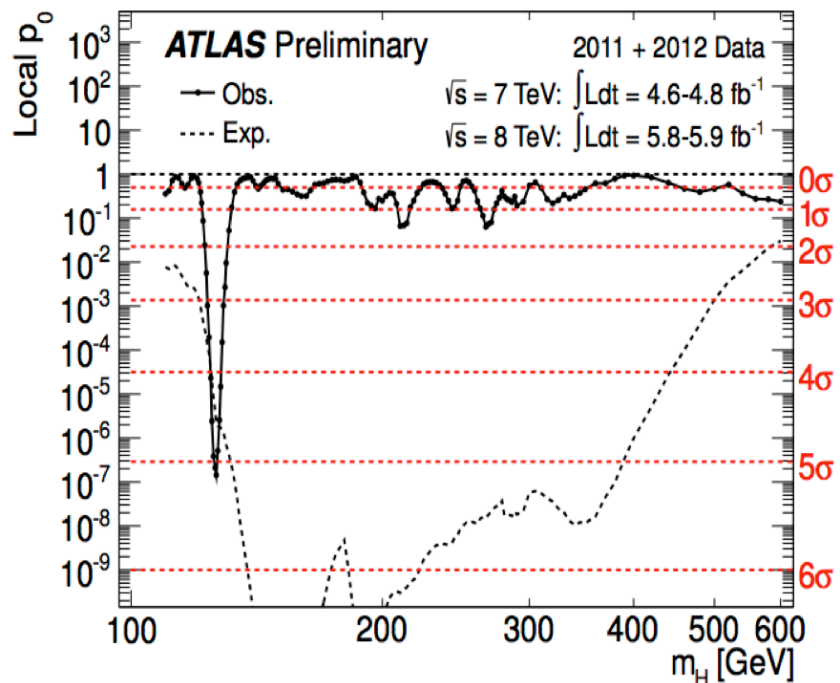
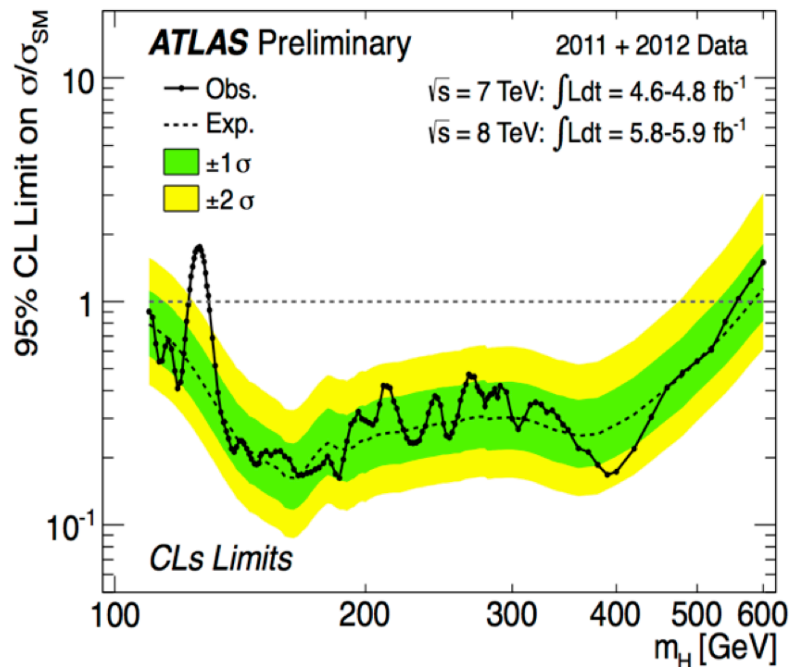
# Goal

Describe the fundamental concepts of statistics for HEP

Explore these concepts with Root-Macros for hands-on experience

*Using the random number generator ... seeing some sampling theory ....*

*Finally be able to understand the following plot !*



Appreciate there is a lot more for you/us to learn about statistical techniques

In particular concerning the treatment of systematics

Apply these results to Discovery and Exclusion in ATLAS

*So be patient and take some time to understand the techniques step by step...*

# Disclaimer :

What this lecture is not going to be about...

- It will not be a lecture on the fundamental theory of statistics
- Multivariate techniques
- Bayesian confidence intervals
- Goodness of fit theory
- In depth discussion of systematics and their treatment
- Bayesian vs. Frequentist diatribe



# Why are Statistics so Important in Particle Physics ?

Because we need to give quantitative statements about processes that have some inherent randomness...

... May this randomness be of measurement nature or quantum ...

How did it all start ?

*Liber de ludo aleae*

To study games of chance !



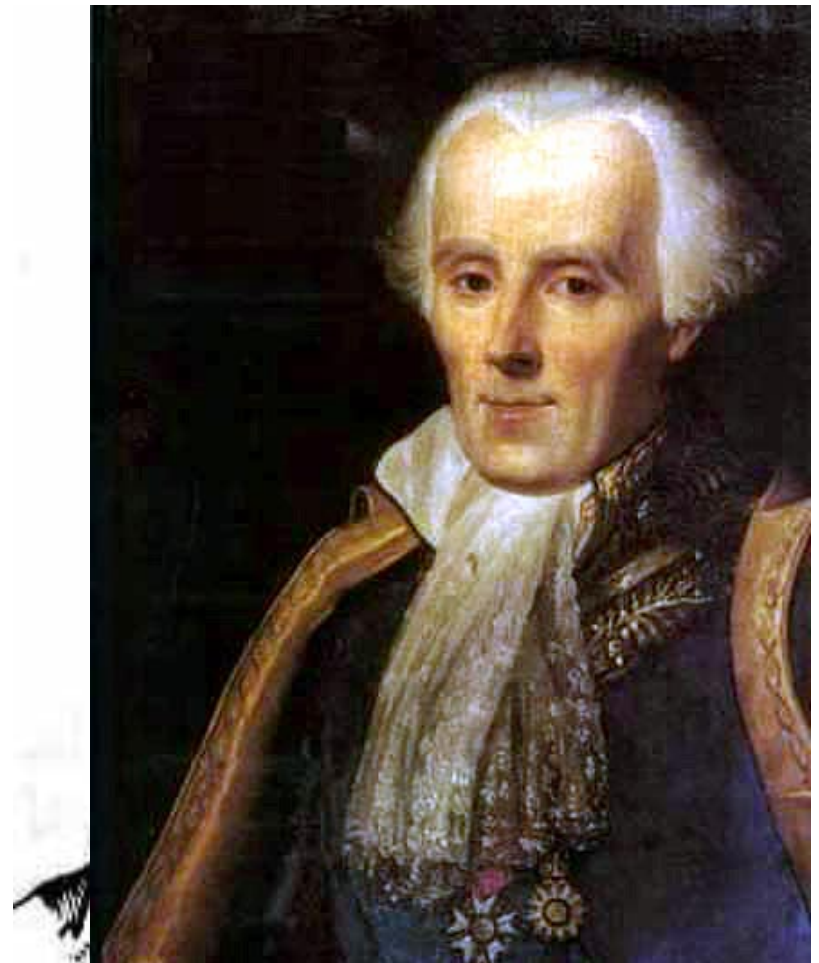
G. Cardano (1501-1576)

And many others to follow (Pascal, Fermat, etc.. )

“La theorie des probabilités n’est, au fond, que le bon sens réduit en calcul”

*“The theory of probabilities is at ultimately nothing more than common sense reduced to calculation”*

P. S. Laplace (1749-1824)



From the very innocuous seeming assumption ....

*“There is a random process characterised by a constant average event rate,  $\mu$ .”*

... many significant and fundamental results follow – perhaps the prime example of the dramatic yield of results from an assumption in all physics.

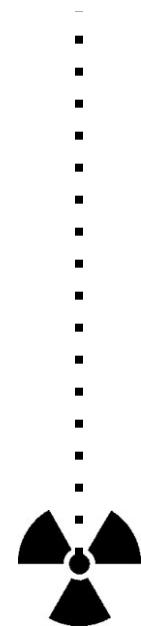
The random deviate represented by the waiting time between such events may be shown to be drawn from the exponential probability density distribution (PDF):

$$p_E(t; \mu) = \mu e^{-\mu t}$$

The random deviate represented by the number of such events within a time bin  $T$  is drawn from the Binomial Distribution, well approximated by the Poisson Distribution:

$$p_P(n; \bar{n}) = \frac{\bar{n}^n}{n!} e^{-\bar{n}}$$

Where the expectation value  $\bar{n} = T\mu$



# What is a Statistical Error ?

Imagine I have a billion white ○ and blue ● golf balls

I decide to throw one million of them into a well and decide an admixture of 15 out of one hundred blue ones...

I then know PRECISELY the probability that if you pick one at RANDOM, it will be blue...

$$p = 15\%$$

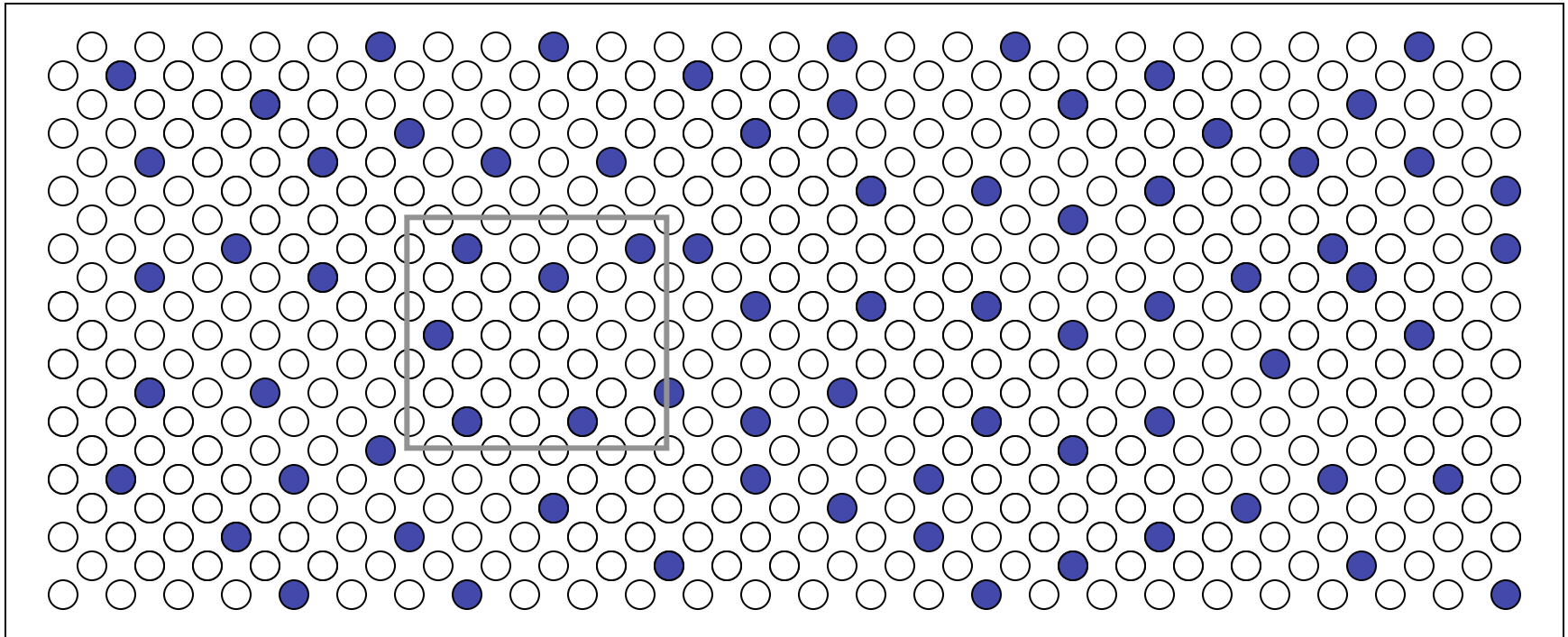
You of course don't know this number and you want to measure it...

All you have is a bucket...

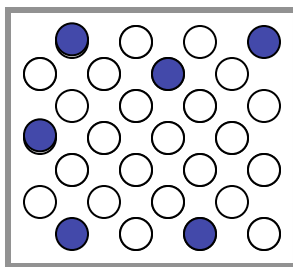
Which contains exactly 300 balls



This is approximately how the well looks like inside...



You throw the bucket and pull out the following outcome



$$n = 300$$

$$k = 36$$

Aha! You have a measurement!

The probability is...

$$P = 12\%$$

... But how precise is it ?

Remember you are not supposed to know the true value!



The difference between a measurement and the true value is the Statistical Error

Precise definition of statistical error

In this case it would be 3% absolute (20% relative), but since you don't know the true value you don't know at all what your statistical error really is !

Of course had you thrown your bucket on a different spot, you would have probably had a different measurement and the statistical error would be different...

What you want to know is your measurement error, or what the average statistical variation of your measurement is...

This can be done provided that you know the law of probability governing the possible outcomes of your experiment ... (and the true value of  $p$ , but assume that 12% is close enough)

You want to know what the probability for an outcome of  $k$  golf balls to be blue is.

For one specific outcome the probability is:

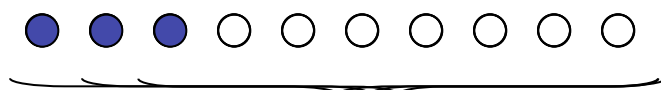
$$P = p^k \times (1 - p)^{n-k}$$

What are all possible combinations of outcomes of  $k$  blue balls out of  $n$ ?

What are all possible combination of outcomes of  $k$  blue balls out of  $n$ ?

{ For the first blue ball there are  $n$  choices, once this choice is made the second ball has  $n-1$  choices,... the  $k^{\text{th}}$  ball has  $(n-k)$  choices.

In a simple case...  $n=10$  and  $k=3$  this can be seen as:



The first ball has 10 choices  
The second ball has 9 choices  
The third ball has 8 choices

So the number of combinations is :  $n \times (n-1) \times (n-2)$

In the general case :  $n \times (n-1) \times (n-2) \times (n-3) \dots \times (n-k+1) = \frac{n!}{(n-k)!}$

Because we do not care about the order in which we have picked the balls



... avoid the double counting!

1	2	3
1	3	2
2	1	3
3	1	2
2	3	1
3	2	1

Each configuration is counted 6 times

This number corresponds in fact to the number of combinations of  $k$  blue balls out of  $k$  balls and therefore :

$$k \times (k - 1) \times (k - 2) \times (k - 3) \dots \times 1 = k!$$

Aka the number of re-arrangements of the  $k$  blue balls.

In order to account for each combination only once you just need to divide by the number of re-arrangements of the  $k$  blue balls.

So the number of combinations of  $k$  elements among  $n$  is given by :

$$C_n^k = \frac{n!}{k!(n - k)!}$$

The probability to pick  $k$  blue balls among  $n$ , given a probability  $P$  that the a ball is blue is thus :

$$P = C_n^k \times p^k \times (1 - p)^{n-k}$$

This is an absolutely fundamental formula in probability and statistics!  
It is the so called Binomial Probability!

# The Binomial Probability

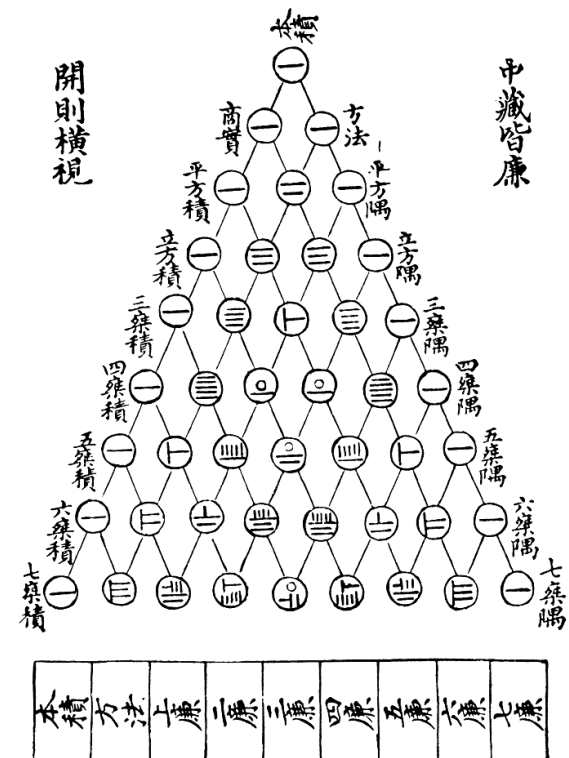
Binomial coefficients were known since more than a thousand years...

... they were also the foundation of modern probability theory!



B. Pascal (1623-1662)

古法七葉圖



The Pascal Triangle (~1000 AD)



# So what is the precision of your measurement ?

A good measure of the precision (*not the accuracy*) is the *Root Mean Square Deviation* (square root of the variance) of possible outcomes of the measurement.

You will compute it yourself. To do so you need two steps...  
(see next slide for the full derivation)

Step 1 : Compute the mean value of the binomial probability

$$\mu = nP$$

Step 2 : Compute the variance of the binomial probability

$$\text{Variance} = nP(1 - P)$$

So now you know the variance of your distribution for a given probability  $P$ ...

In your case :  $P = 12\%$  Assuming  $P$  is close enough to the true value, the precision is :

$$RMSD = \sqrt{nP(1 - P)} = 5.6$$

The relative precision ~15% is rather poor and the accuracy questionable! (Remember, your statistical error is  $45 - 36 = 9$ , although you are not supposed to know it !)

## Step 1 : Compute mean value

### Mean of the Binomial Probability

Let us denote by  $P_k^n(p)$  the binomial probability of an outcome  $k$  among  $n$  with single event probability  $p$ , and  $\mu$  its mean value, then

$$\mu = \frac{\sum_{k=0}^n k P_k^n(p)}{\sum_{k=0}^n P_k^n(p)} = \sum_{k=0}^n k C_n^k p^k (1-p)^{n-k}$$

Note that it looks like a derivative!

Nice trick: Start from the derivative...

$$\frac{\partial}{\partial p} \left[ \sum_{k=0}^n C_n^k p^k (1-p)^{n-k} \right] = 0 = \frac{\partial [1]}{\partial p}$$

$$\begin{aligned} \Rightarrow \sum_{k=0}^n k C_n^k p^{k-1} (1-p)^{n-k} - \sum_{k=0}^n (n-k) C_n^k p^k (1-p)^{n-k-1} &= 0 \\ 0 = \frac{\mu}{p} - n \sum_{k=0}^n C_n^k p^k (1-p)^{n-k-1} + \sum_{k=0}^n k C_n^k p^{k-1} (1-p)^{n-k} \end{aligned}$$

$\frac{1}{1-p}$        $\mu/(1-p)$

thus

$$\frac{\mu}{p} + \frac{\mu}{1-p} = \frac{n}{1-p} \Rightarrow (1-p)\mu + p\mu = np$$

so

$$\boxed{\mu = np} \quad \text{Aha!}$$

## Step 2 : Compute variance

### Variance of the Binomial Probability

Let us start from the previous formula for the mean value of the binomial probability.

$$\text{Variance} = \sum_{k=0}^n k^2 C_n^k p^k (1-p)^{n-k} - \underbrace{\left( \sum_{k=0}^n k C_n^k p^k (1-p)^{n-k} \right)^2}_{\mu^2}$$

given that  $\sum_{k=0}^n k C_n^k p^k (1-p)^{n-k} = \mu = np$

then  $\frac{\partial}{\partial p} \left[ \sum_{k=0}^n k C_n^k p^k (1-p)^{n-k} \right] = n$  thus:

$$\sum_{k=0}^n k^2 C_n^k p^{k-1} (1-p)^{n-k} - \sum_{k=0}^n k(n-k) C_n^k p^k (1-p)^{n-k-1} = n$$

$$\frac{1}{p} \sum_{k=0}^n k^2 C_n^k p^k (1-p)^{n-k} - \frac{n\mu}{1-p} + \frac{1}{1-p} \sum_{k=0}^n k^2 C_n^k p^k (1-p)^{n-k} = n$$

$$\Rightarrow (1-p+p) \sum_{k=0}^n k^2 C_n^k p^k (1-p)^{n-k} = np(1-p) + \underbrace{np\mu}_{\mu^2}$$

when  $\mu = np$  so:

$$\sum_{k=0}^n k^2 C_n^k p^k (1-p)^{n-k} - \mu^2 = np(1-p)$$

Variance!

Thus

$$\boxed{\text{Variance} = np(1-p)}$$

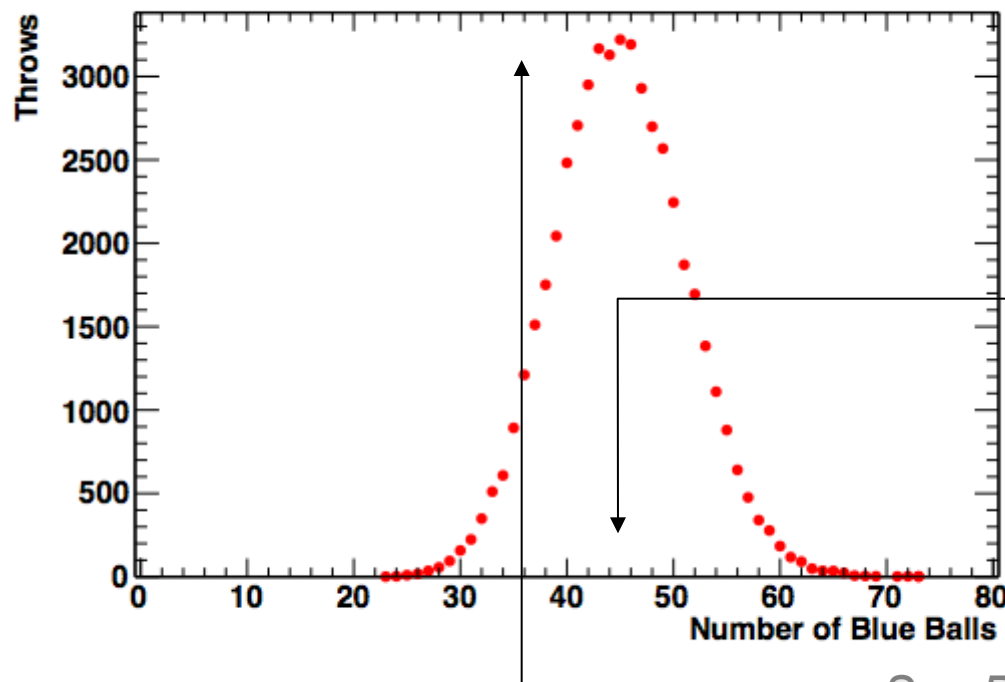
But wait...

Now you are curious to see what happens if you repeat your measurement!

You have noticed that the average binomial probability is the expected value!

Intuitively you will therefore try to repeat and average your measurements...

You will do it 50,000 times and meticulously plot the number of counts. This is what you get :



$$N_{\text{throws}} = 50000$$

The average number of blue balls in 50,000 throws :

$$\langle \text{Number}_{\text{Blue}} \rangle = 44.98$$

$$\langle P \rangle = 14.99\%$$

Your initial measurement (36) !

*See Binomial.C*

Now you decide that your measurement is the average, what is its precision ?

## What is the variance of the average ?

Let's start from one straightforward and general property of the Variance for two random variables  $X$  and  $Y$  :

$$\begin{aligned} \text{Var}(aX + bY) &= \langle (aX + bY - \langle aX + bY \rangle)^2 \rangle = \langle [a(X - \langle X \rangle) + b(Y - \langle Y \rangle)]^2 \rangle \\ &= a^2 \text{Var}(X) + b^2 \text{Var}(Y) + 2ab \text{Cov}(X, Y) \end{aligned}$$

Where the covariance is :  $\text{Cov}(X, Y) = \langle (X - \langle X \rangle)(Y - \langle Y \rangle) \rangle$

This formula generalizes to... 
$$\text{Var}\left(\sum_{i=0}^n a_i X_i\right) = \sum_{i=0}^n a_i^2 \text{Var}(X_i) + \sum_{0 \leq i < j \leq n} a_i a_j \text{Cov}(X_i, X_j)$$

Therefore assuming that each of the bucket throws measurement  $N_{Blue}^k$  is independent from the previous one, the mean value being a simple sum of the measurements divided by the number of throws :

$$\langle \text{Number}_{Blue} \rangle = \frac{1}{N_{Throws}} \sum_{k=1}^{N_{Throws}} N_{Blue}^k$$

The ensemble variance then is :

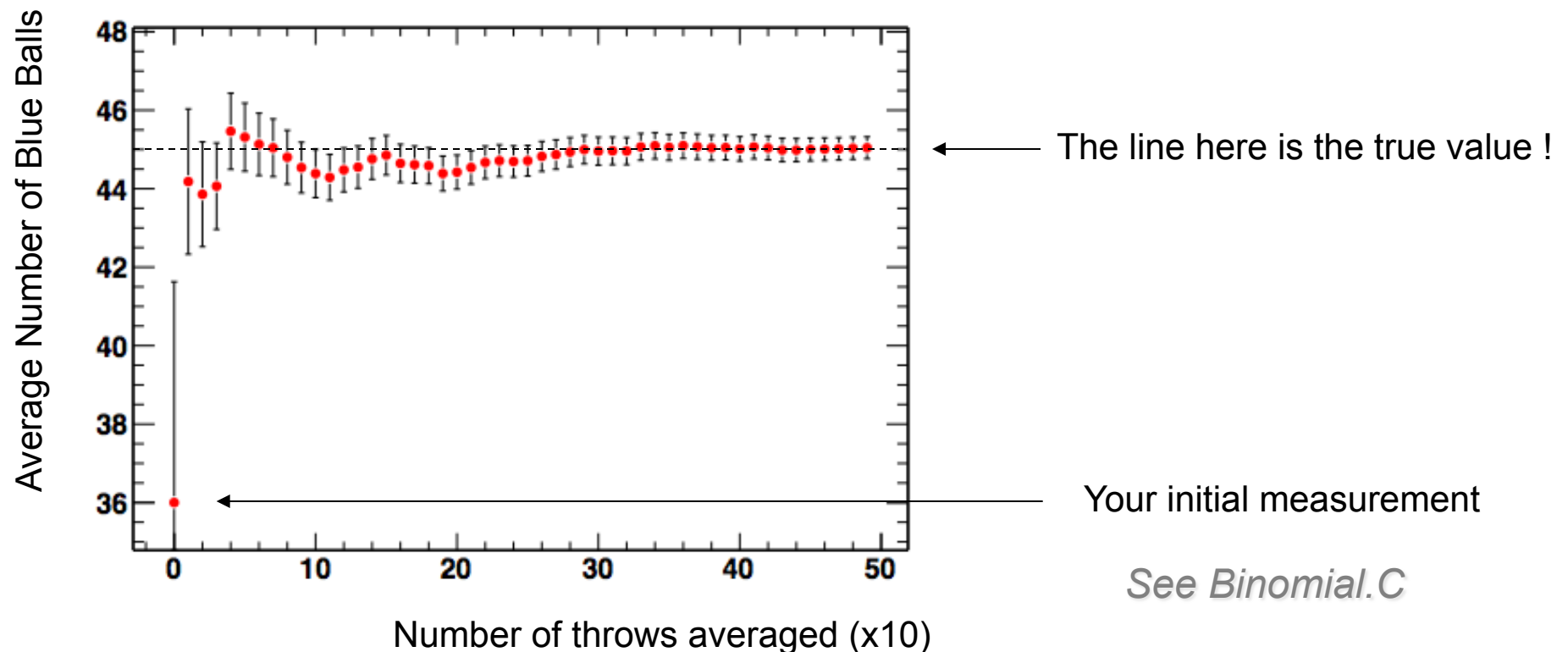
$$\hat{\sigma}^2 = \text{Var}\left(\frac{1}{N_{Throws}} \sum_{k=1}^{N_{Throws}} X_i\right) = \frac{1}{N_{Throws}^2} \sum_{k=1}^{N_{Throws}} \text{Var}(X_i) = \frac{1}{N_{Throws}^2} N_{Throws} \text{Var}(X_i) = \frac{nP(1-P)}{N_{Throws}}$$



The precision being given by the *Root Mean Square Deviation* :

$$RMSD = \sqrt{\frac{nP(1-P)}{N_{Throws}}} = \frac{RMSD_{Individual}}{\sqrt{N_{Throws}}} = 0.01\%$$

Very interesting behavior : Although you do not know the true value  $p$ , you see that the average is converging towards it with increasing precision!



This is an illustration of the **LAW of LARGE NUMBERS** ! Extremely important, intuitive but not trivial to demonstrate...

What is the meaning of our first measurement  $N_{\text{blue}} = 36$  ?

Now that we know (after 50,000 throws) to a high precision that the probability of a blue ball is very close to 15%.

The frequency of an outcome as low as 12% is ~10% (not so unlikely!)

What difference would it make if you had known true value ?

Frequency at which the measurement is within the precision as estimated from the truth :

$$|P_{\text{meas}} - p| \leq \sqrt{np(1-p)} \Rightarrow 70\% \text{ (of the cases the measurement is within the true statistical RMSD)}$$

Frequency at which the true value is within the precision as estimated from the measurement :

$$|P_{\text{meas}} - p| \leq \sqrt{nP_{\text{Meas}}(1-P_{\text{Meas}})} \Rightarrow 67\% \text{ (of the cases the true value is within the measured error)}$$

*See Coverage.C*

The true value coverage is similar in the two cases, keep these values in mind...

Here all results are derived from a simulation in terms of frequencies...

Computing Binomial probabilities with large numbers of  $N$  can be quite difficult !

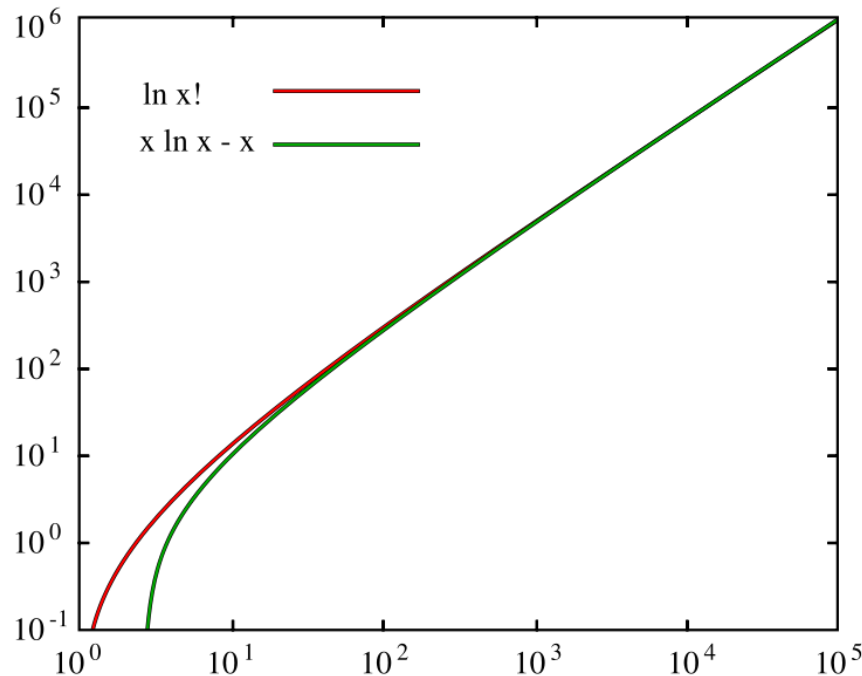
# The Gaussian or Normal Probability

Is there a way to simplify the computation ? Not so trivial to compute 300! directly...

A very nice approximation of the Binomial Probability can be achieved using Stirling's Formula !

$$n! \approx \sqrt{2\pi n} \left(\frac{n}{e}\right)^n$$

$$\ln n! \approx n \ln n - n + \frac{1}{2} \ln 2\pi n$$



Formula is valid for large values of  $n$ ...

$$C_n^k p^k (1-p)^{n-k} \approx \frac{1}{\sqrt{2\pi\sigma^2}} e^{-\frac{(k-\langle k \rangle)^2}{2\sigma^2}}$$

$$\sigma = \sqrt{np(1-p)}$$

(See derivation in the next slide)

## Binomial convergence towards Normal

### Derivation of the Normal Probability

... Again start from the binomial probability, but this time use the Stirling formula:  $n! \sim \sqrt{2\pi n} \left(\frac{n}{e}\right)^n$

$$\left[ \begin{array}{l} \ln(k!) \sim k \ln k - k + \frac{1}{2} \ln(2\pi k) \\ \text{Very useful!} \end{array} \right.$$

$$\text{then } \ln C_n^k = \ln n! - \ln k! - \ln(n-k)!$$

$$\begin{aligned} \text{thus } \ln [C_n^k p^k (1-p)^{n-k}] &= n \ln n - n + \frac{1}{2} \ln(2\pi n) \\ &- [k \ln k - k + \frac{1}{2} \ln(2\pi k)] - [(n-k) \ln(n-k) - (n-k) + \frac{1}{2} \ln(2\pi(n-k))] \\ &+ k \ln p + (n-k) \ln(1-p) \end{aligned}$$

Again using a derivative trick

$$\begin{aligned} \frac{\partial}{\partial k} [\ln(C_n^k p^k (1-p)^{n-k})] &= -\ln k + \ln(n-k) + \ln p - \ln(1-p) \\ &= \underbrace{-\frac{1}{2k} + \frac{1}{2(n-k)}}_{O\left(\frac{1}{n}\right)} \ln\left(\frac{p(n-k)}{k(1-p)}\right) \end{aligned}$$

We see that the maximum value will occur at  $p(n-k) = k(1-p)$  or  $k = np$  Which is also the average value -

Then looking at the second derivative:

$$\frac{\partial^2}{\partial k^2} [\ln(C_n^k p^k (1-p)^{n-k})] = -\left(\frac{1}{n-k} + \frac{1}{k}\right) + O\left(\frac{1}{k^2}\right)$$

then take at maximum value  $k = np$ :

$$\frac{\partial^2}{\partial k^2} [\ln(C_n^k p^k (1-p)^{n-k})] \sim \frac{-1}{np(1-p)} = -\frac{1}{\sigma^2} \quad \text{where } \sigma = \sqrt{np(1-p)}$$

Therefore with a second order Taylor series expansion:

$$\begin{aligned} \ln(C_n^k p^k (1-p)^{n-k}) &\sim \underbrace{\ln A}_{\text{constant}} + \frac{(k-np)^2}{2} \cdot \left(\frac{-1}{\sigma^2}\right) \\ \text{therefore } C_n^k p^k (1-p)^{n-k} &\sim A e^{-\frac{(k-np)^2}{2\sigma^2}} \end{aligned}$$

introducing the notation  $\tilde{k} = np$  corresponding to the maximum of the probability - then

$$C_n^k p^k (1-p)^{n-k} \sim A e^{-\frac{(\tilde{k}-k)^2}{2\sigma^2}}$$

then using the normalization of the probability

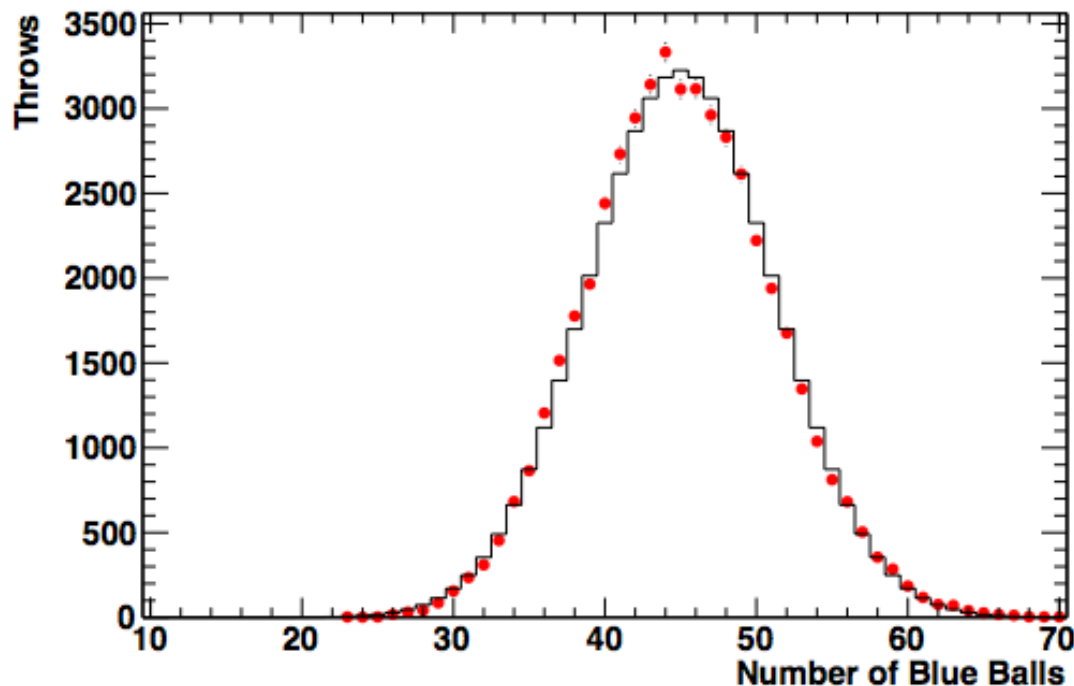
$$\begin{aligned} \int_{-\infty}^{\infty} A e^{-\frac{(\tilde{k}-k)^2}{2\sigma^2}} dk &= 1 \quad \text{and the Gauss integral} \\ \int_{-\infty}^{+\infty} e^{-\frac{z^2}{2\sigma^2}} dz &= \sqrt{2\pi\sigma^2} \quad \text{so: } C_n^k p^k (1-p)^{n-k} = \frac{e^{-\frac{(\tilde{k}-k)^2}{2\sigma^2}}}{\sqrt{2\pi\sigma^2}} \end{aligned}$$



# Validity of the Normal Convergence (Approximation)

Does the approximation apply to our bucket experiment ( $n=300$  and  $p=15\%$ ) ?

*See NormalConvergence.C*



C. F. Gauss (1777-1855)

Not bad (although not perfect) !

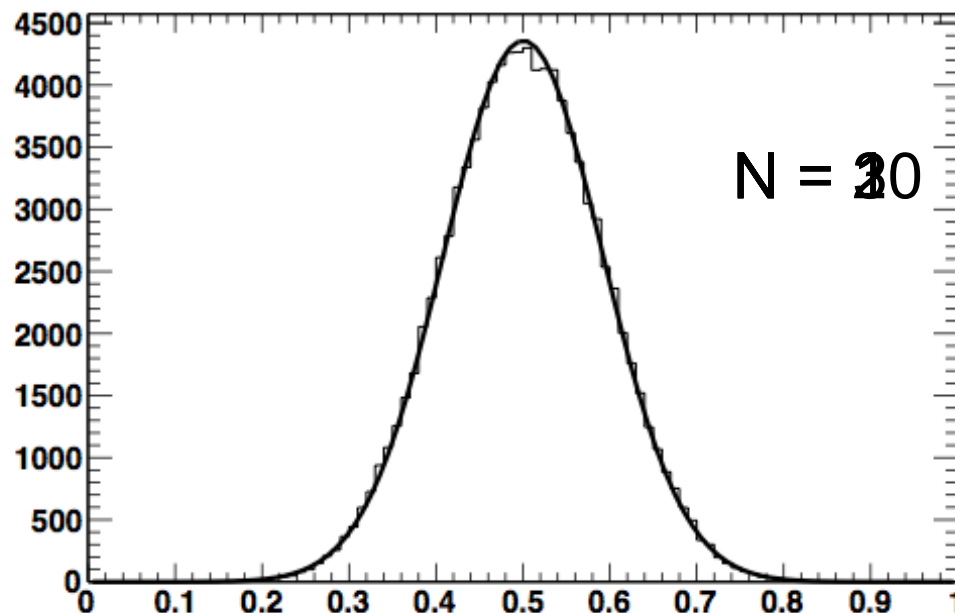
In practice you can use the normal law when approximately  $n > 30$  and  $np > 5$

# What is so “Normal” About the Gaussian?

## The Central Limit Theorem... ... at Work !

When averaging various independent random variables (*and identically distributed*) the distribution of the average converges towards a Gaussian distribution

See CLT.C



$$\text{RMS} = \frac{[0,1]}{\sqrt{12}} \times \frac{1}{\sqrt{30}}$$

At  $N=10$  an excellent agreement with a gaussian distribution is observed

The CLT is one of the main reasons for the great success of the Gaussian law...

On the one hand the CLT is very powerful to describe all those phenomena that result from the superposition of various other phenomena... but on the other hand it is just a limit...

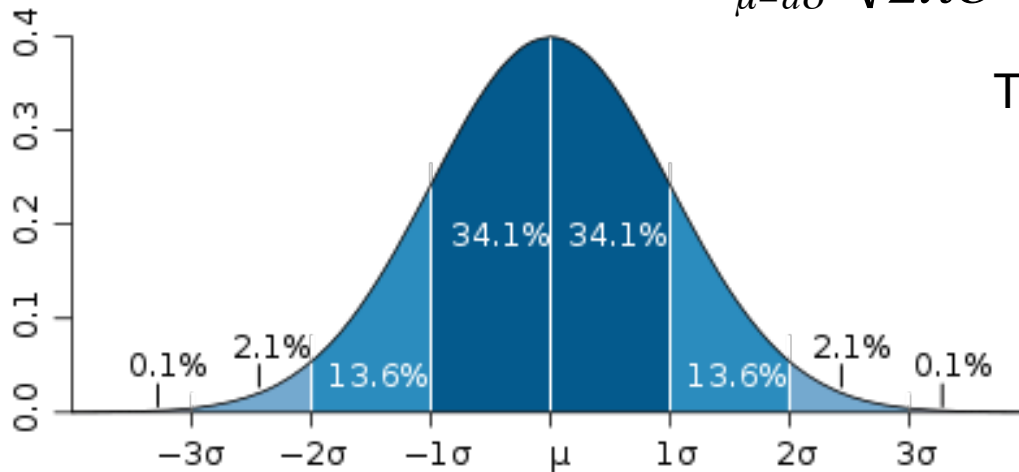
# The Notion of Standard Error

Starting from the gaussian PDF :

$$G_{PDF}(x, \mu, \sigma) = \frac{1}{\sqrt{2\pi\sigma^2}} e^{-\frac{(x-\mu)^2}{2\sigma^2}}$$

Let's give a first definition of a central confidence interval as the deviation from the central value...

$$P(a\sigma) = \int_{\mu-a\sigma}^{\mu+a\sigma} \frac{1}{\sqrt{2\pi\sigma^2}} e^{-\frac{(x-\mu)^2}{2\sigma^2}} dx$$



Then for :

- $a = 1$  :  $P(a\sigma) = 68.3\%$
- $a = 2$  :  $P(a\sigma) = 95.4\%$
- $a = 3$  :  $P(a\sigma) = 99.7\%$

*See NormalCoverage.C*

If you knew the true value of the “error” ( $\sigma$ ) then you could say that the in the gaussian limit that the true value has 68.3% probability to be within the  $1\sigma$ , but in many practical examples (such as the well) the true value of the error is not known...

# How does the Bucket Experiment Relate to Particle Physics?

The bucket experiment is the measurement of an abundance (blue balls)...

This is precisely what we call in particle physics cross sections...

... except that the bucket contains all collisions collected in an experiment so...

- We try to fill it as much as possible ( $N$  is very large and not constant!)
- The processes we are looking for are very rare ( $p$  is very small)

The very large  $N$  makes it difficult to compute the binomial probability...



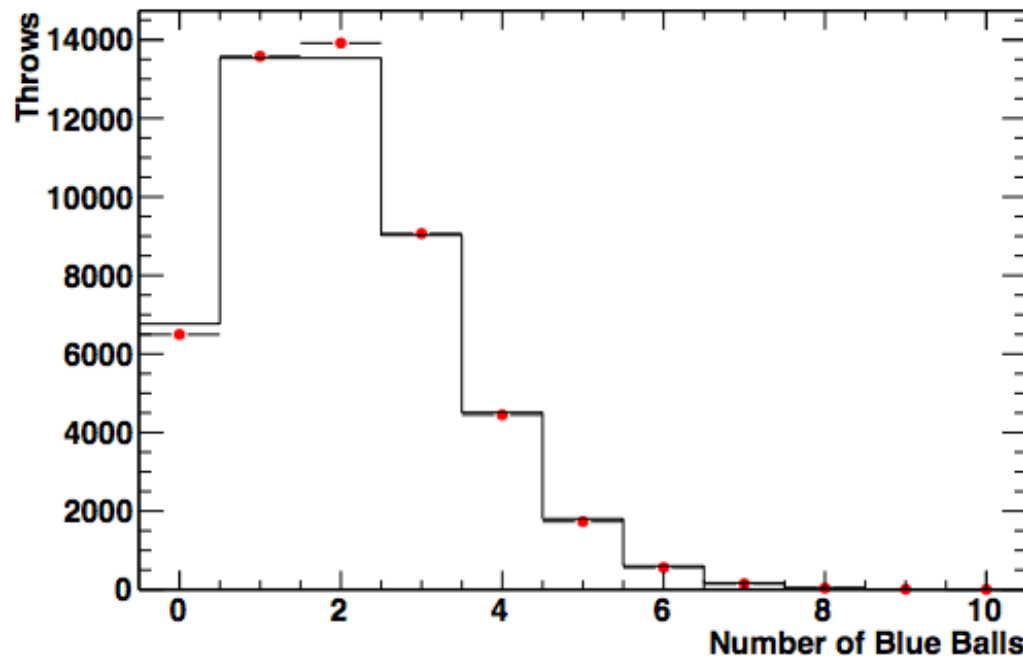
# The Poisson Probability

In the large  $n$  and small  $p$  limit and assuming that  $np = \mu$  is finite you can show (see next slide) that ...

$$C_n^k p^k (1-p)^{n-k} \approx \frac{(np)^k}{k!} e^{-(np)} = \frac{\mu^k}{k!} e^{-\mu}$$

Much simpler formulation! In practice you can use the normal law when approximately  $n > 30$  and  $np < 5$

*See PoissonConvergence.C*



**N=100 and p=20%**



S. D. Poisson (1781-1840)

Interesting to note that Poisson developed his theory trying not to solve a game of chance problem but a question of Social Science !

### Derivation of the Poisson Probability

... from the binomial probability :  $C_n^k p^k (1-p)^{n-k}$   
which can be written in the limit when

$p$  is very small and  $np$  is finite.

$$C_n^k = \frac{n!}{k! (n-k)!} = \frac{1}{k!} [n(n-1) \dots (n-k+1)] \sim \frac{n^k}{k!}$$

$$\begin{aligned} \text{and for } p \text{ small } (1-p)^{n-k} &= 1 - (n-k)p + \frac{(n-k)(n-k-1)}{2!} p^2 \\ &+ \dots \approx 1 - np + \frac{(np)^2}{2!} + \dots \\ &\approx e^{-np} \end{aligned}$$

$$\begin{aligned} \text{therefore } C_n^k p^k (1-p)^{n-k} &\approx \frac{n^k}{k!} p^k e^{-np} \\ &= \frac{(np)^k}{k!} e^{-np} \end{aligned}$$

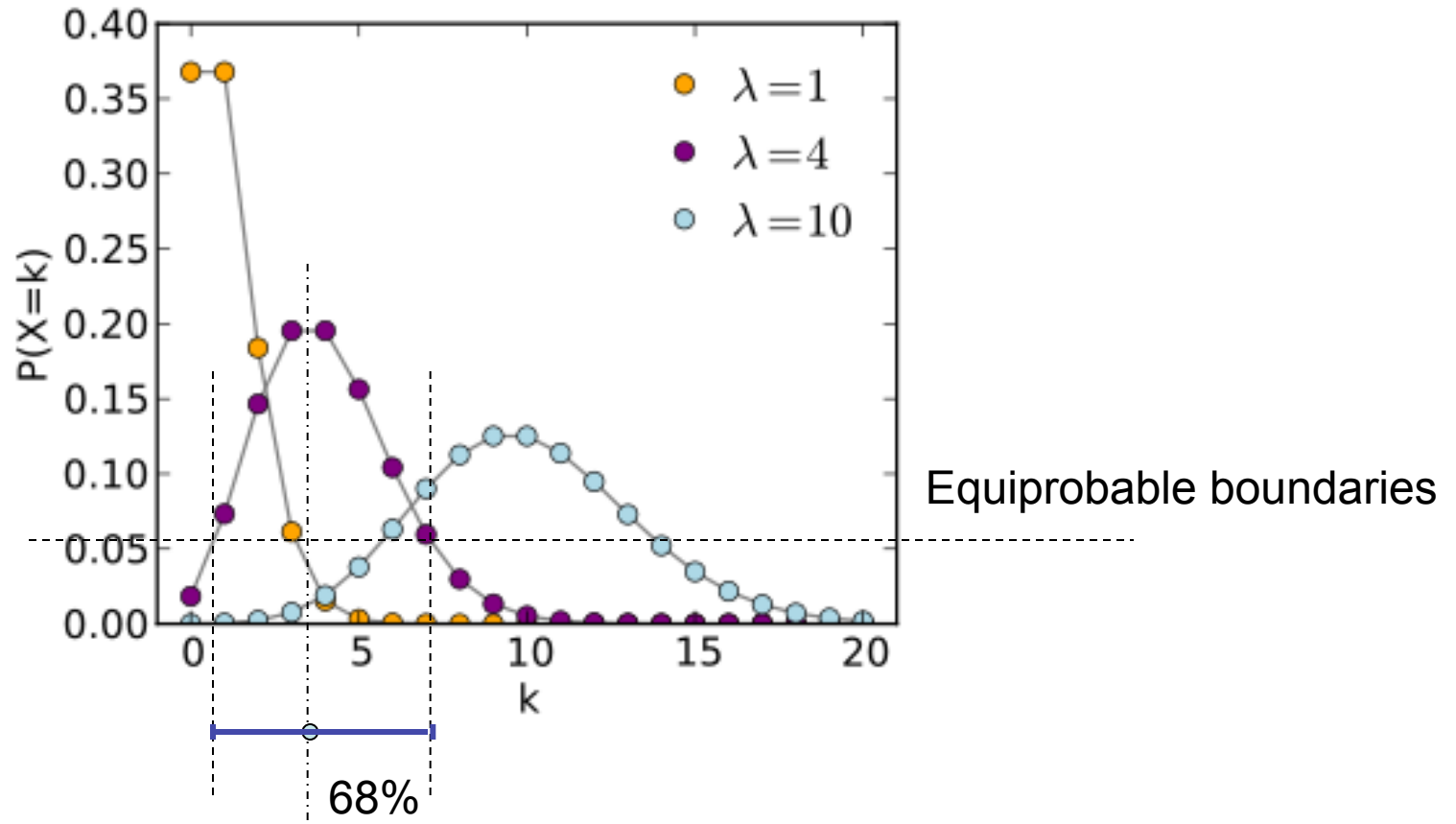
thus denoting  $\mu = np$  the poisson probability  
of an outcome  $k$  given an expectation of  $\mu$

$$\text{is } P(k, \mu) = \frac{\mu^k}{k!} e^{-\mu}$$

This approximation was essentially when ( $n > 50$   
and  $p < 0.1$ ).

# Poisson Intervals (or Errors)

Now how will you define a central confidence interval in a non symmetric case ?



The integration needs to start from the most probable value downwards...

Here is our first encounter with the necessity of an ordering !

# What have we learned ?

...and a few by-products...

1.- Repeating measurements allows to converge towards the true value of an observable more and more precisely ...

But never reach it with infinite precision !!!

Even more so accounting for systematics...  
(what if the balls do not have an homogeneous distribution ?)

2.- Binomial variance is also useful to compute the so-called binomial error, mostly used for efficiencies :

$$\sigma_{\varepsilon} = \frac{\sigma_{\mu}}{N} = \sqrt{\frac{\varepsilon(1-\varepsilon)}{N}} \quad \mu = np$$

For an efficiency you must consider n fixed !

3.- We came across a very important formula in the previous slides

$$\text{Var}\left(\sum_{i=0}^n a_i X_i\right) = \sum_{i=0}^n a_i^2 \text{Var}(X_i) + \sum_{0 \leq i < j \leq n} a_i a_j \text{Cov}(X_i, X_j)$$

That generalizes (with a simple Taylor expansion) to...

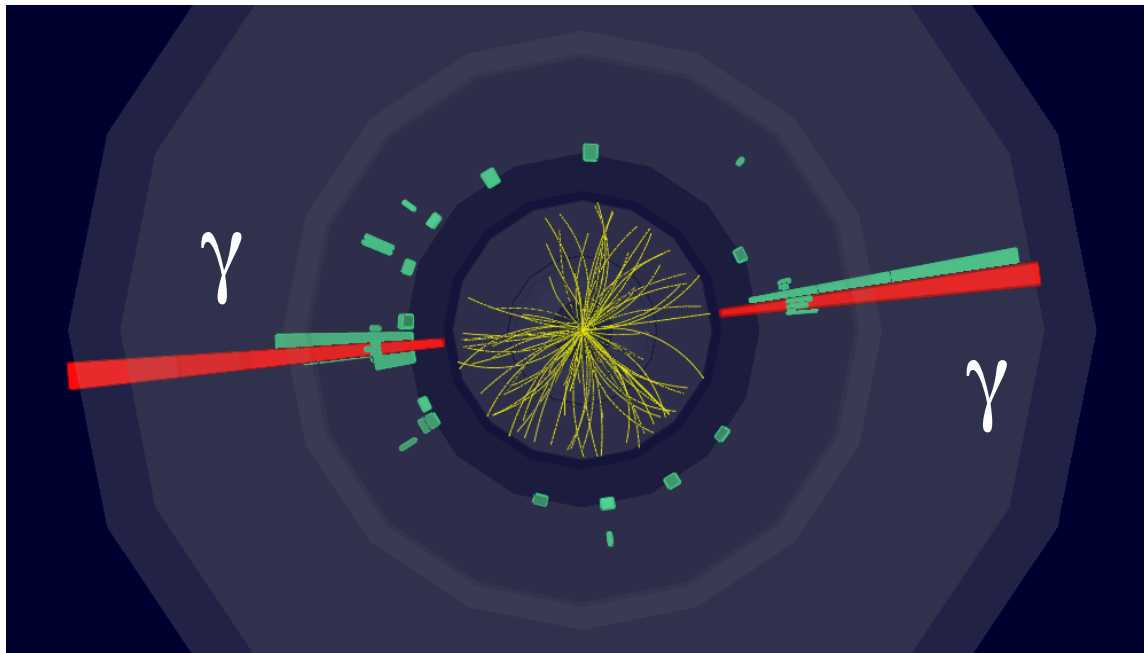
$$\text{var}(f(x_1, \dots, x_n)) = \sum_{i=0}^n \left(\frac{\partial f}{\partial x_i}\right)^2 \text{var}(x_i) + \sum_{0 \leq i < j \leq n} \frac{\partial f}{\partial x_i} \frac{\partial f}{\partial x_j} \text{cov}(x_i, x_j)$$



# Likelihood

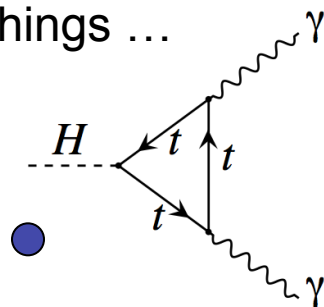
Unfortunately in High Energy physics experiments, events (balls) don't come in single colors (white or blue) ... Their properties are not as distinct !

For instance take this simple event :

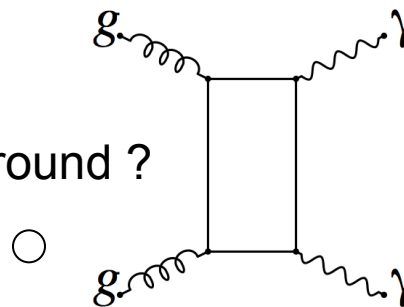


Could be many things ...

Higgs ?



Background ?



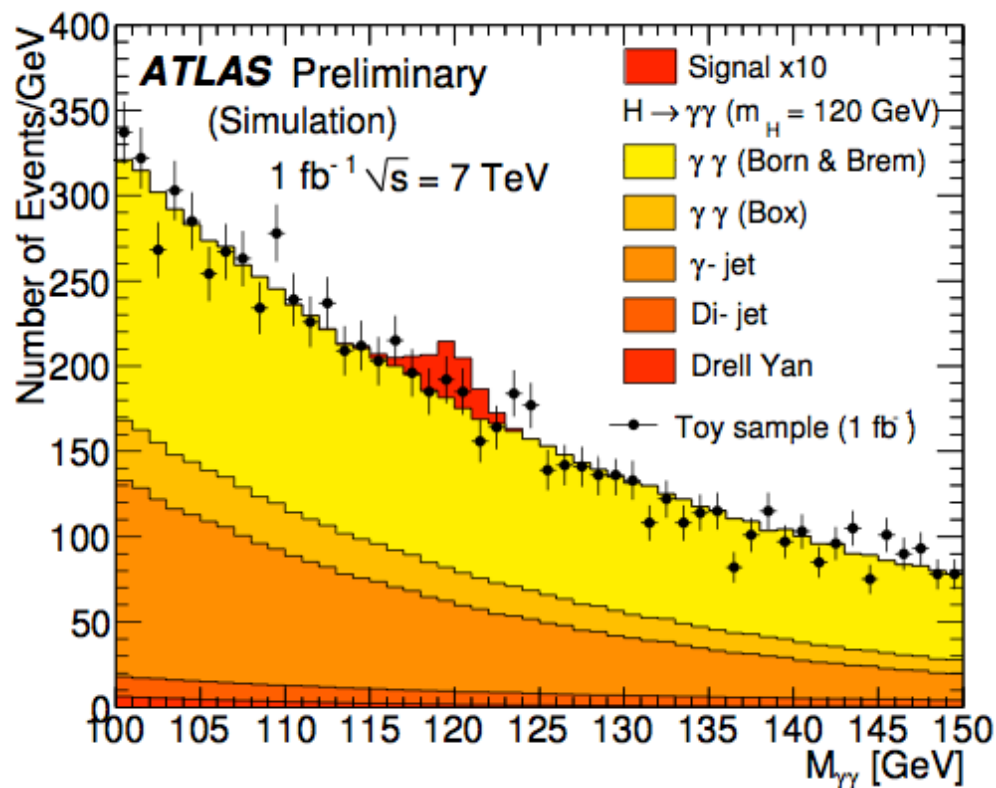
Let alone that they can be indistinguishable (quantum interference)

# How can we distinguish between the two ?

Very vast question, let's first start with how to measure their properties

(Which is also a very vast question!)

One clear distinctive feature is that the signal is a narrow mass resonance, while the background is a continuum !



To measure properties in general (*a.k.a. parameter estimation*) among the most commonly used tools is the maximum likelihood fit...

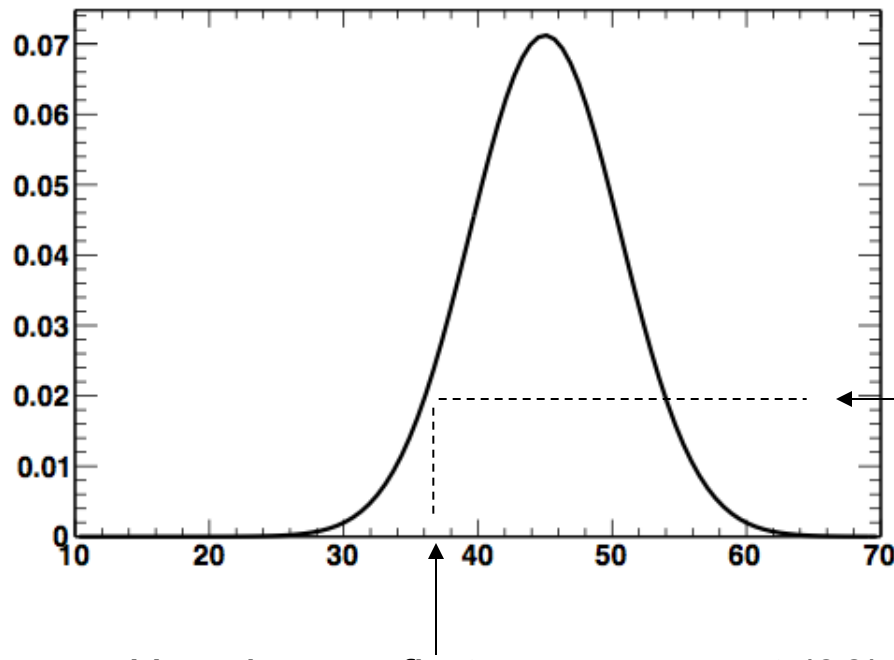
# What is a Likelihood ?

A simple way of defining a Likelihood is a Probability Density Function (PDF) which depends on a certain number of parameters...

Simplistic definition is a function with integral equal to 1...

Let's return to the well experiment but under a different angle this time...

(but this applies to any parameter estimate)



Under certain hypothesis :

- Gaussian centered at 45 ( $p=15\%$ )
- Width equal to error for 1 bucket (~6.2 blue balls)

Here is its probability !  
*or Likelihood*

Not so likely !

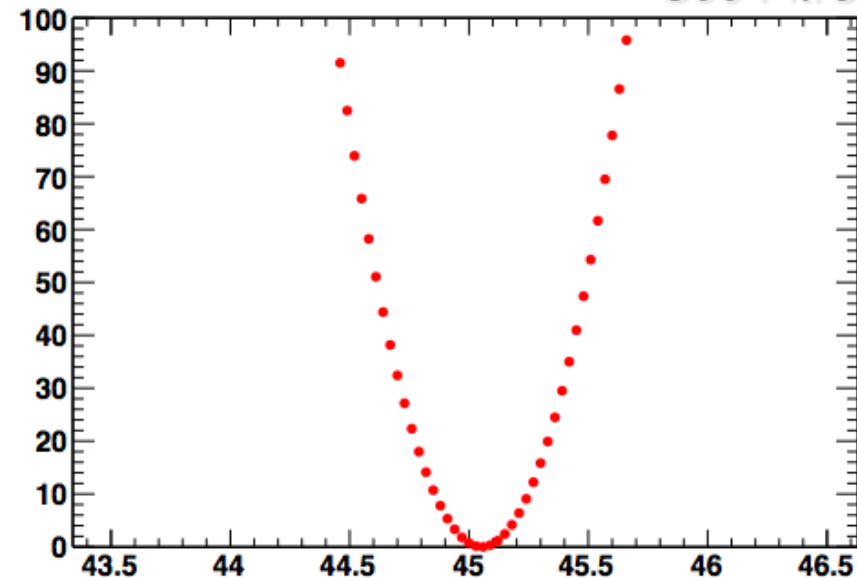
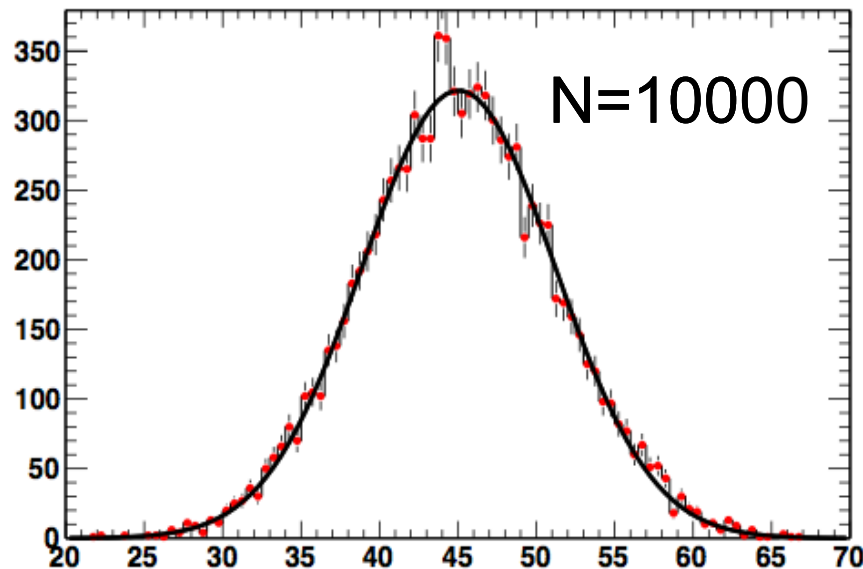
Here is your first measurement (36) !

What happens when we throw more buckets ?

$$L(\mu) = \prod_{i=1}^N f_{\mu}(n_i)$$

Then the probability of each bucket can be multiplied!

*See Fit.C*



This probability will soon be very very small ( $O(0.1)^{100}$ ...). It is easier to handle its log :

$$\ln(L(\mu)) = \sum_{i=1}^n \ln(f_{\mu}(n_i))$$

Then to estimate a parameter one just has to maximize this function of the parameter  $\mu$  (or minimize  $-2\ln L$  you will see why in a slide)...

See how the accuracy translates in the sharpness of the minimum!

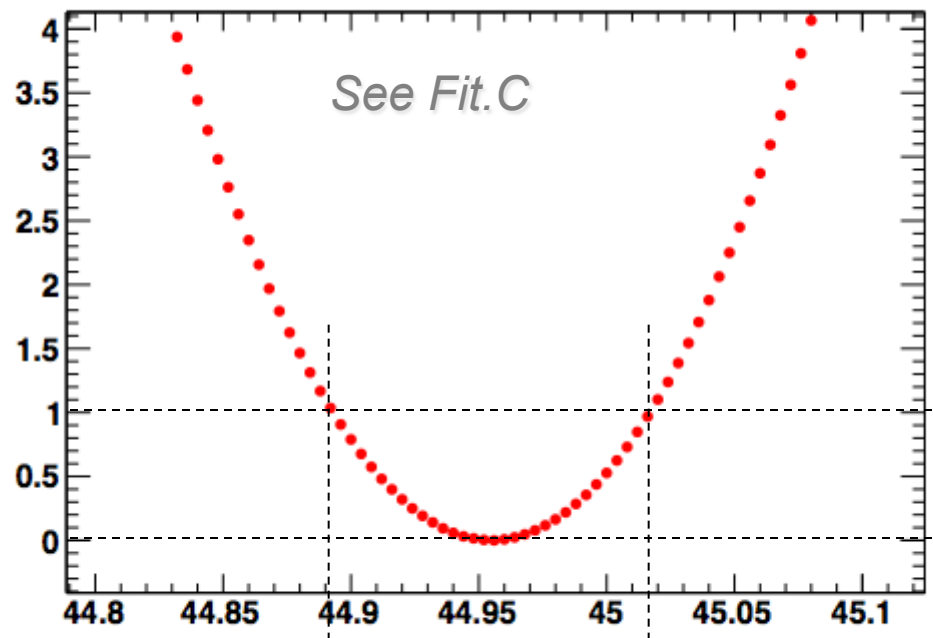


In our simple (but not unusual) case we can see that :

$$-2\ln(L(\mu)) = -2 \sum_{i=1}^n \ln(f_{\mu}(n_i)) = -2 \sum_{i=1}^n \ln\left(\frac{1}{\sqrt{2\pi}\sigma} e^{-\frac{(n_i - \mu)^2}{2\sigma^2}}\right) = \underbrace{\sum_{i=1}^n \frac{(n_i - \mu)^2}{\sigma^2}} + cste$$

This is also called  $\chi^2$

There is an exact equivalence between maximizing the Likelihood or minimizing the  $\chi^2$  (Least Squares Method) in the case of a gaussian PDF



You can also see that the error on the measured value will be given by a variation of  $-2 \ln L$  of one unit :

$$\Delta(-2\ln(L(\mu))) = 1$$

$$\bar{\mu} = 44.95 \pm \underbrace{0.06}$$

Which is precisely  $\frac{\sigma}{\sqrt{n}}$

“Proof” : Estimation of variance of model parameters (errors)

$$S = \sum_{i=1}^n \left[ \frac{y_i - f(x_i; a_j)}{\sigma_i} \right]^2 \quad Y = \begin{bmatrix} \sigma_1^2 & \text{cov}(1, 2) & \text{cov}(1, 3) & \dots \\ \cdot & \sigma_2^2 & \text{cov}(2, 3) & \dots \\ \cdot & \cdot & \sigma_3^2 & \dots \\ \cdot & \cdot & \cdot & \dots \end{bmatrix} \quad (Y^{-1})_{ij} = \frac{1}{2} \frac{\partial^2 S}{\partial a_i \partial a_j}$$

The likelihood function

The mean is estimated by a model

$a_j$  are model parameters

The errors in the parameters are estimated by the diagonal elements of the covariance matrix, which for linear least squares is given by the inverse of the double partial derivative.

$$\sigma^2 = \left| \frac{1}{2} \frac{\partial^2 S}{\partial \theta^2} \right|^{-1} \quad \text{In 1d .... simple case}$$

If we expand  $S$  in Taylor series about the minimum

$$\begin{aligned} S(\theta) &= S(\theta^*) + \frac{1}{2} \frac{\partial^2 S}{\partial \theta^2} (\theta - \theta^*)^2 \\ &= S(\theta^*) + \frac{1}{\sigma^2} (\theta - \theta^*)^2. \end{aligned}$$

At the point  $\theta = \theta^* + \sigma$ , we thus find that

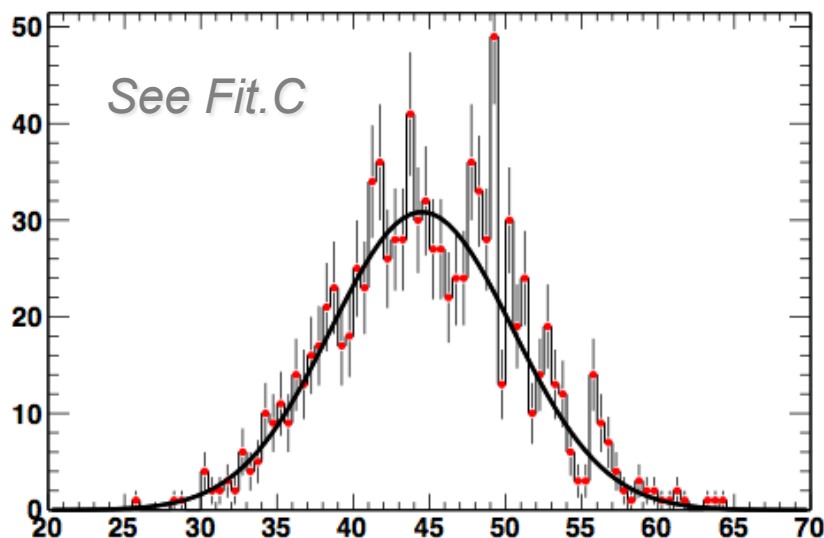
$$S(\theta^* + \sigma) = S(\theta^*) + 1.$$

A robust procedure varies the twice log likelihood function about the minimum  $\theta^*$  by 1 to find the root of the variance in the model parameter.

For higher dimensionality, the tested model parameter is stepped while the others are varied to maintain the minimum condition.

# What have we learned?

How to perform an unbinned likelihood fit :



For  $n=1000$  the fit yields

$$\bar{\mu} = 44.91 \pm 0.19$$

Using a simple binned fit (as shown here with 100 bins) in the same data yields :

$$\bar{\mu} = 44.81 \pm 0.20$$

LSM between the PDF and the bin value

This can of course be applied to any parameter estimation, as for instance the di-photon reconstructed mass !

The  $\chi^2$  value is itself a statistic (random variable).

One can repeat the measurement, (throw of the bucket, collection of the data), and one would get a different data set, and then calculate a different  $\chi^2$ .

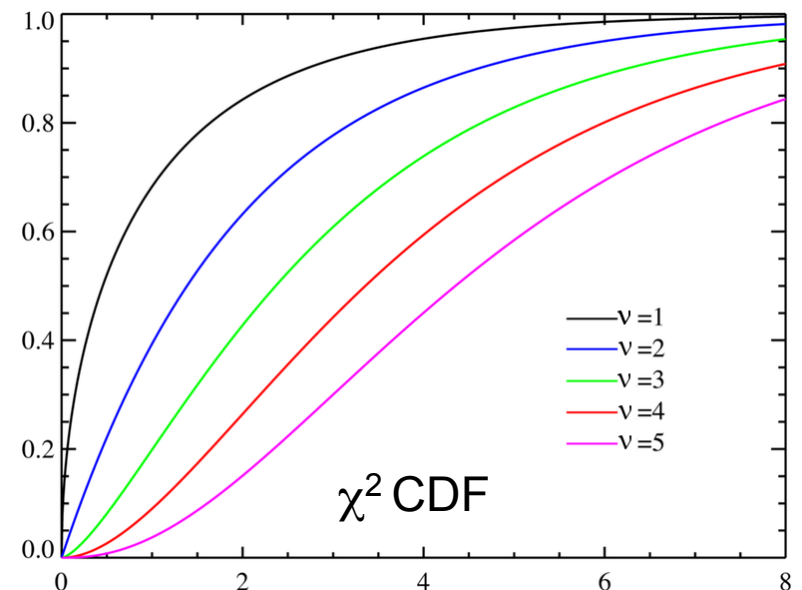
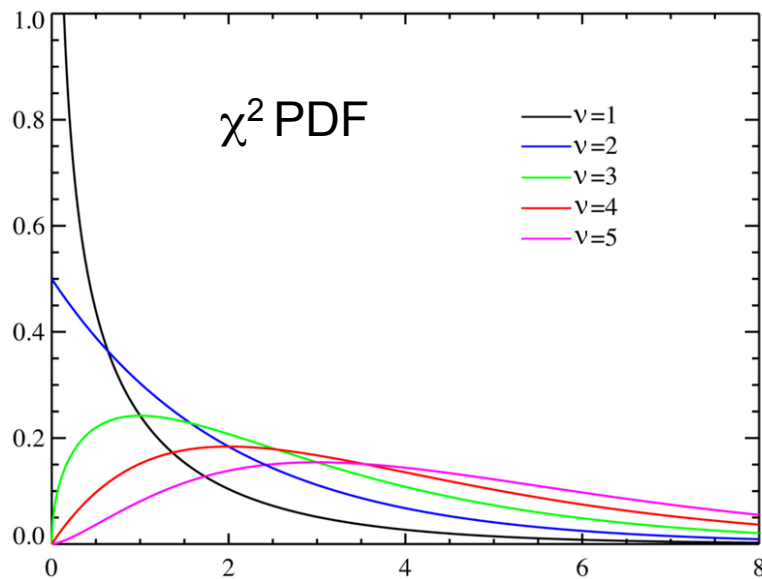
This means that the value of  $\chi^2$  belongs to a distribution.

As we could write down the  $\chi^2$  exactly when the single point distribution was gaussian, it follows that the  $\chi^2$  distribution is amenable to analysis, and can be calculated as:

$$P(u) = \frac{(u/2)^{(\nu/2)-1} e^{-u/2}}{2\Gamma(\nu/2)}$$

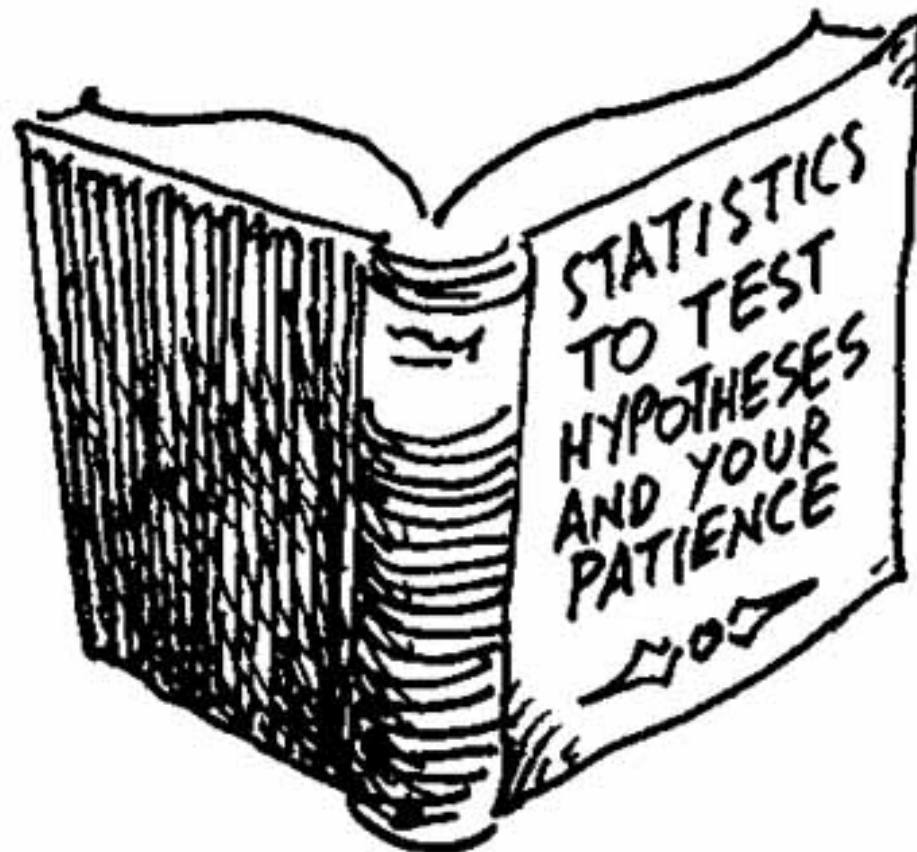
We have used  $u = \chi^2$  to avoid confusion with the exponent.

$\Gamma(\nu/2)$  represents the gamma function and  $\nu$  the degrees of freedom (see later).



# Hypothesis Testing

How to set limits or claim discovery ?



Hypothesis Testing in HEP Boils Down to One Question :  
Is there a Signal ?



# Exclusion, Observation or Discovery ?

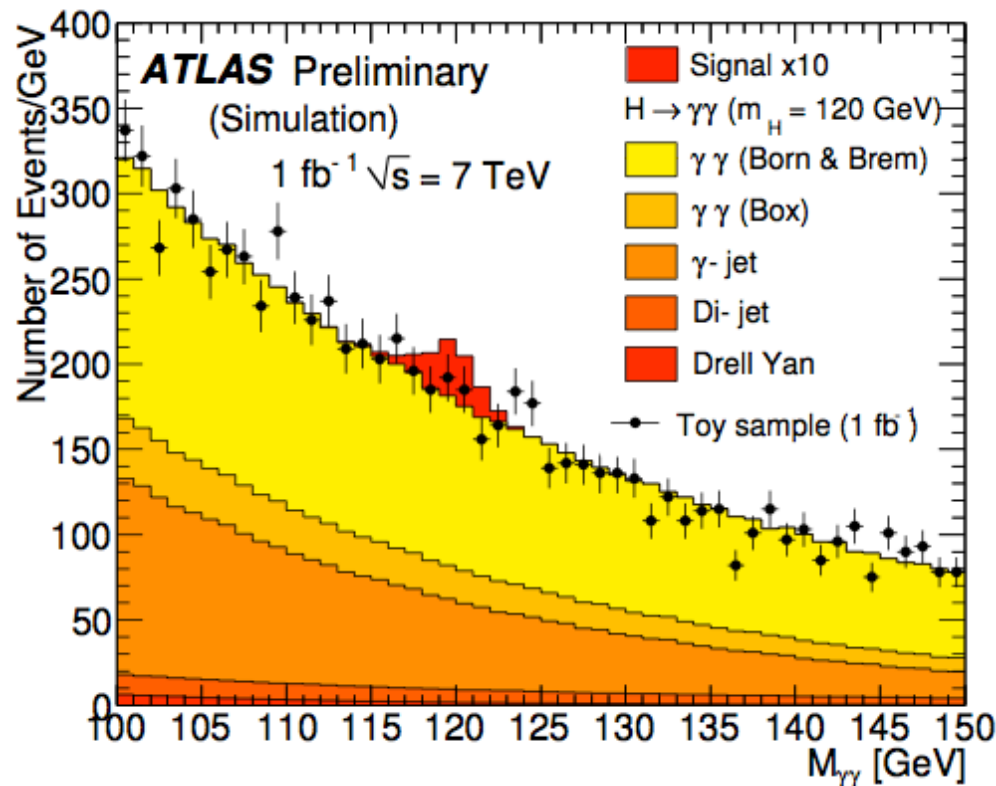
The goal here is to assess quantitatively the compatibility of our observation with two hypotheses :

No-Signal ( $H_0$ ) and presence of Signal ( $H_1$ )...

We need to be able to have estimate whether an experiment is more Signal-like or Background-Like.

Neyman construction (1933)

Let's again take the example of the  $H \rightarrow \gamma\gamma$  analysis at LHC (in ATLAS)



# The Neyman-Pearson Lemma

The underlying concept in ordering experiments is really to quantify the compatibility of the observation with the signal hypothesis ( $H_1$ ) ...

The problem of testing Hypotheses was studied in the 30's by Jerzy Neyman and Egon Pearson...

They have shown that the ratio of likelihoods of an observation under the two hypotheses is the most powerful tool (or test-statistic or order parameter) to

$$E = \frac{P(H_1 | x)}{P(H_0 | x)}$$

# The F-Test

Consider the case where the test statistic is defined as

$$F = \frac{\chi^2(H_1 | x) / v_1}{\chi^2(H_0 | x) / v_2} = \frac{\frac{1}{v_1} \sum \frac{(f(x_i; h, \hat{\theta}) - y_i)^2}{\sigma^2}}{\frac{1}{v_2} \sum \frac{(f(x_i; \hat{\theta}) - y_i)^2}{\sigma^2}}$$

With reference to the High Energy Physics example, in  $H_1$ ,  $h$  is the height of a (gaussian) peak (of assumed known width) on a smooth background characterised by a function of parameters  $\theta$ , and in  $H_0$ , there is only the smooth background.

The ratio of two  $\chi^2$  distributions will be well defined because the  $\chi^2$  is well defined.

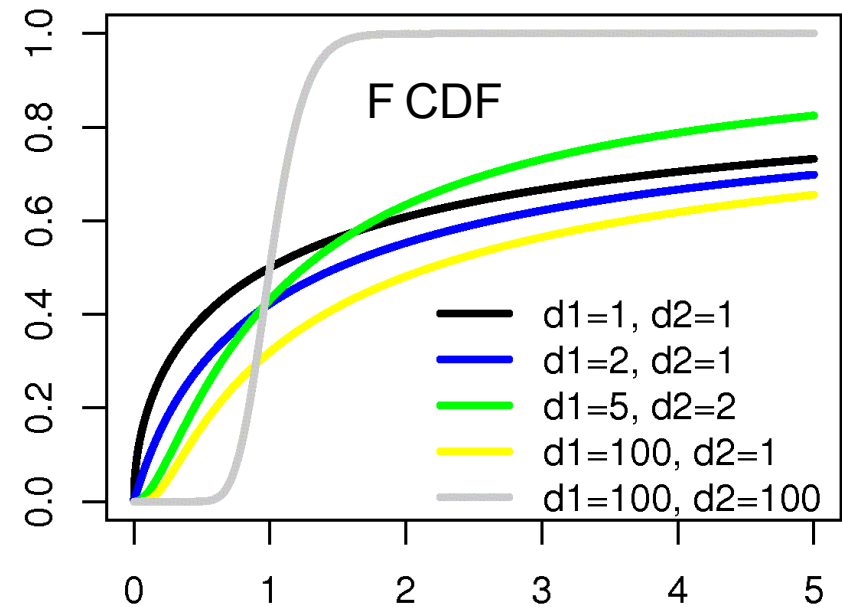
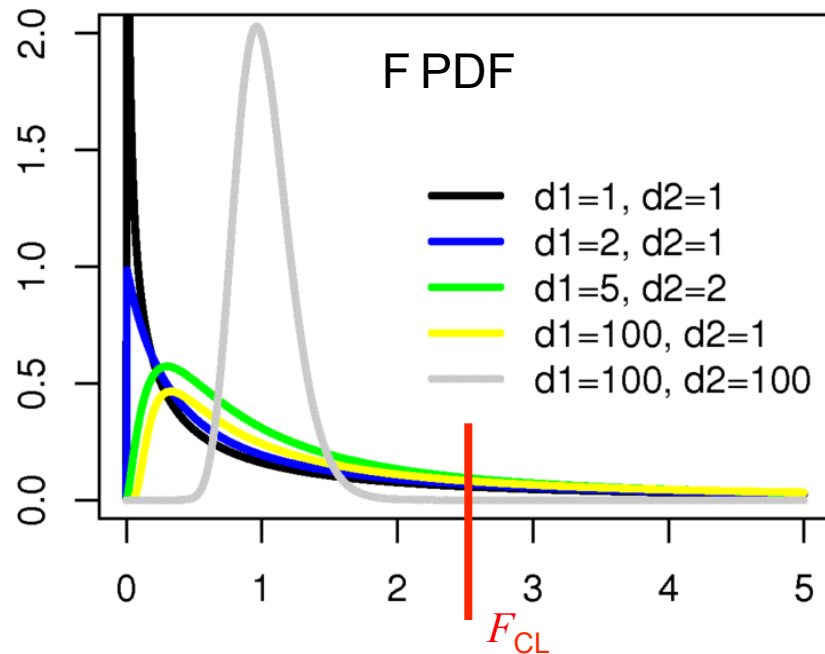
The ratio is the F statistic, which itself belongs to a distribution.

$$Q(F | v_1, v_2) = I_{\frac{v_2}{v_2 + v_1 F}} \left( \frac{v_2}{2}, \frac{v_1}{2} \right)$$

Where  $I$  is the incomplete beta function.

Note : We are asking if the two distributions (with and without the peak) are different.

$\nu_1$  and  $\nu_2$  are the degrees of freedom for  $H_1$  and  $H_0$  respectively.  
 $H_1$  is described by  $f(x, h, \theta)$  which has  $n$  data points and  $m$  free parameters.  
 Then,  $\nu_1 = n - m$ .  $H_0$  will have one more degree of freedom than  $H_1$ .

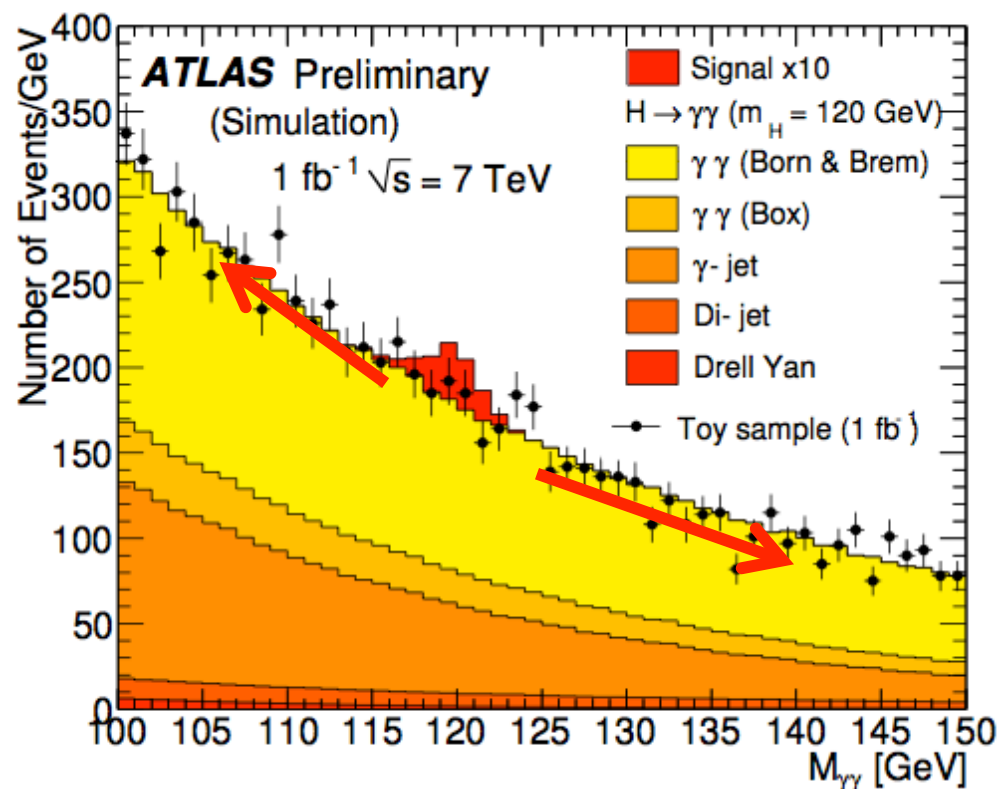


A confidence limit for the rejection (acceptance) of  $H_0$ , the null hypothesis, that there is no peak, corresponds to discovery (exclusion).

In this analysis, the confidence limit is set at CL%, and the  $F$  distribution is integrated to the the  $F$ -value of  $F_{CL}$ . Based on the cumulate  $F$  distribution to the point  $F_{CL}$ , we are CL% certain that a measured  $F$ -value larger than  $F_{CL}$  is not statistically acceptable as being consistent with  $H_0$ .

This analysis is didactic and illustrative, but it suffers from several drawbacks. It does not respect the “*look elsewhere*” effect, it assumes a normal distribution for the data, it cannot easily take into account the full systematics of the measurement, amongst other issues.

The “*look elsewhere*” effect considers that we do not know where the peak should be. The estimated probability of the peak must be multiplied by the number of ways that it could have been manifested (roughly the factor of the measurement interval divided by the peak width – assuming the peak width is also not free).



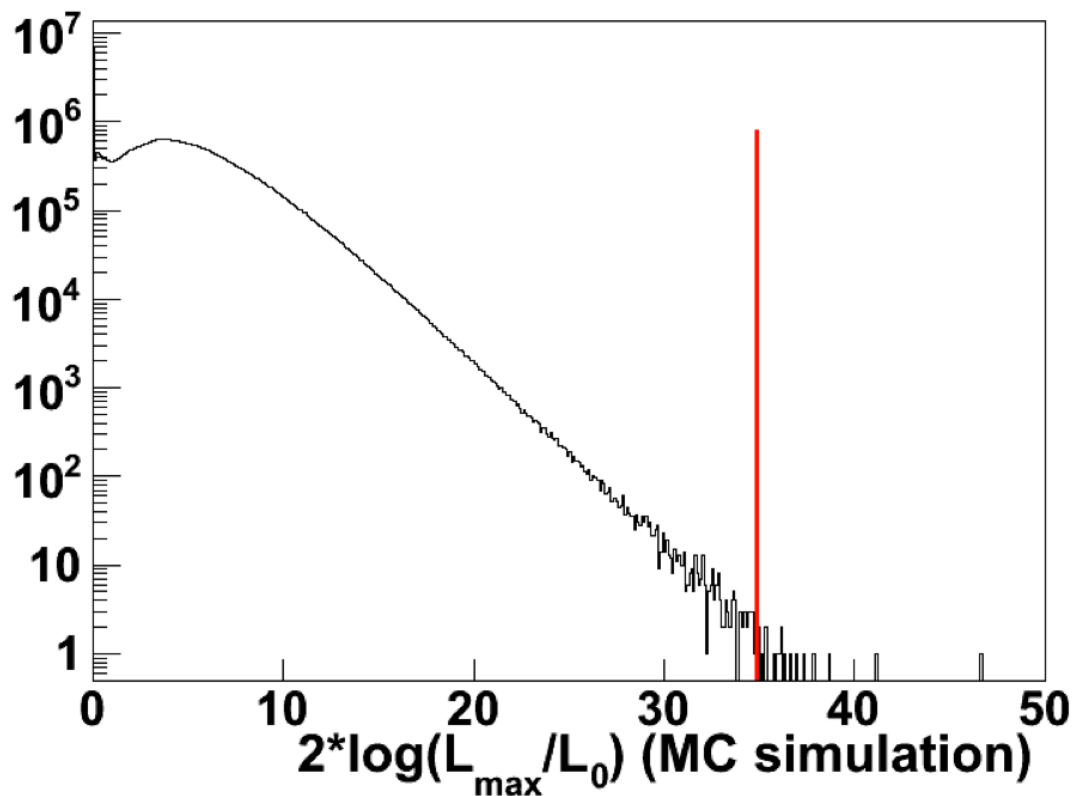
An improvement is to develop toy Monte Carlo pseudo experiments for  $H_1$  and  $H_0$ .

$$E = \frac{P(H_1 | x)}{P(H_0 | x)}$$



The “ $\chi^2$ ” statistic for  $H_1$  and  $H_0$  can be calculated using the synthetic data. The toy MC pseudo experiment can be repeated many times, billions of times, and the PDF’s of the “ $\chi^2$ ” statistic for  $H_1$  and  $H_0$  can be numerically assembled.

The same can be done for the statistic 
$$E = \frac{P(H_1 | x)}{P(H_0 | x)}$$



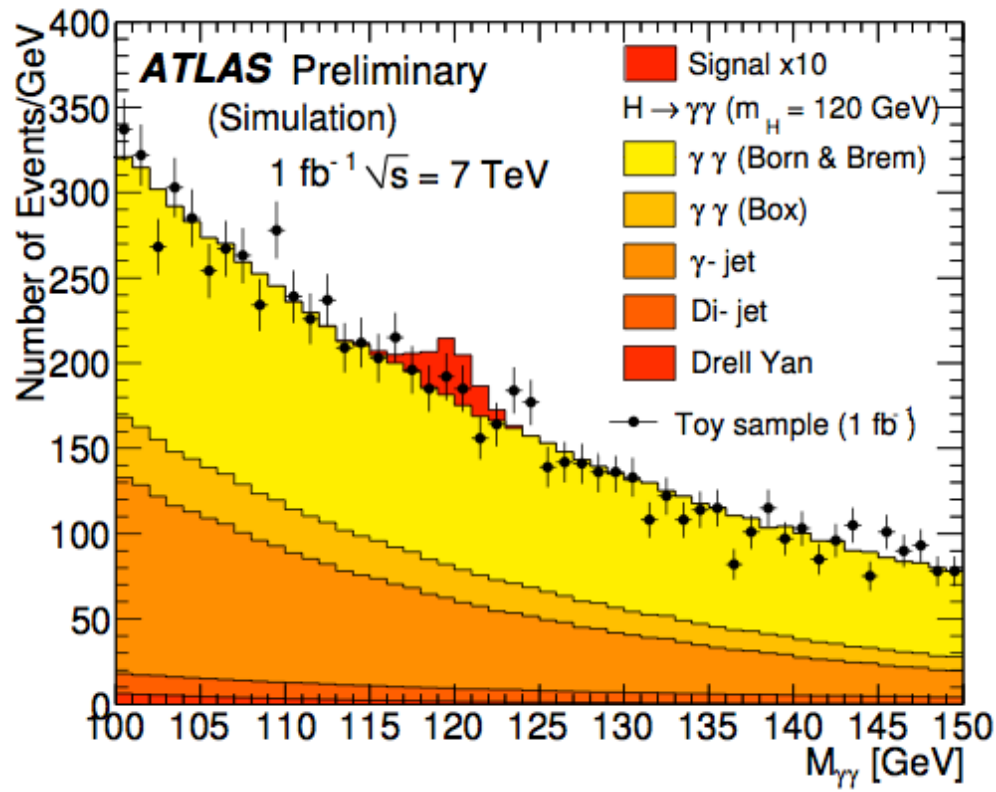
The “*look elsewhere*” effect will be accommodated if the peak position is a free parameter, and it could then range freely in the position where the statistical fluctuations allow it to be found most favorably. Other effects (width variations, systematics are conceivably able to be included in developing the PDF’s.

The process of setting a CL% and determining a p-value from the CDF can now follow based on these distributions.

# The Profile Likelihood

A very useful tool to compute limits, observation or discovery sensitivities and treat systematics is the Profile Likelihood ... based on toy MC pseudo experiments.

Let's again take the example of the  $H \rightarrow gg$  analysis at LHC (*in ATLAS*)



We have a simple model for the background :

$$b(m, \theta) = \theta_1 e^{-\theta_2 m}$$

Relies only on two parameters

Assume a very simple model for the signal :

$$s(m, \mu) = \mu s \times \text{Gauss}(m)$$

The Gaussian is centered at  $120 \text{ GeV}/c^2$  and a width of  $1.4 \text{ GeV}/c^2$

# The Profile Likelihood

The overall fit model is very simple :

$$L(\mu, \theta | data) = \prod_{i \in data} (s(m_i, \mu) + b(m_i, \theta))$$

This model relies essentially only on two types of parameters :

- The signal strength parameter ( $\mu$ ) It is essentially the signal normalization
- The nuisance parameters ( $\theta$ ) Background description in the “*side bands*”

$$\lambda(\mu) = \frac{L(\mu, \hat{\hat{\theta}}(\mu) | data)}{L(\hat{\mu}, \hat{\theta} | data)}$$

← Test of a given signal hypothesis  $\mu$   
 ← Best fit of the data

Prescription similar to the Feldman Cousins

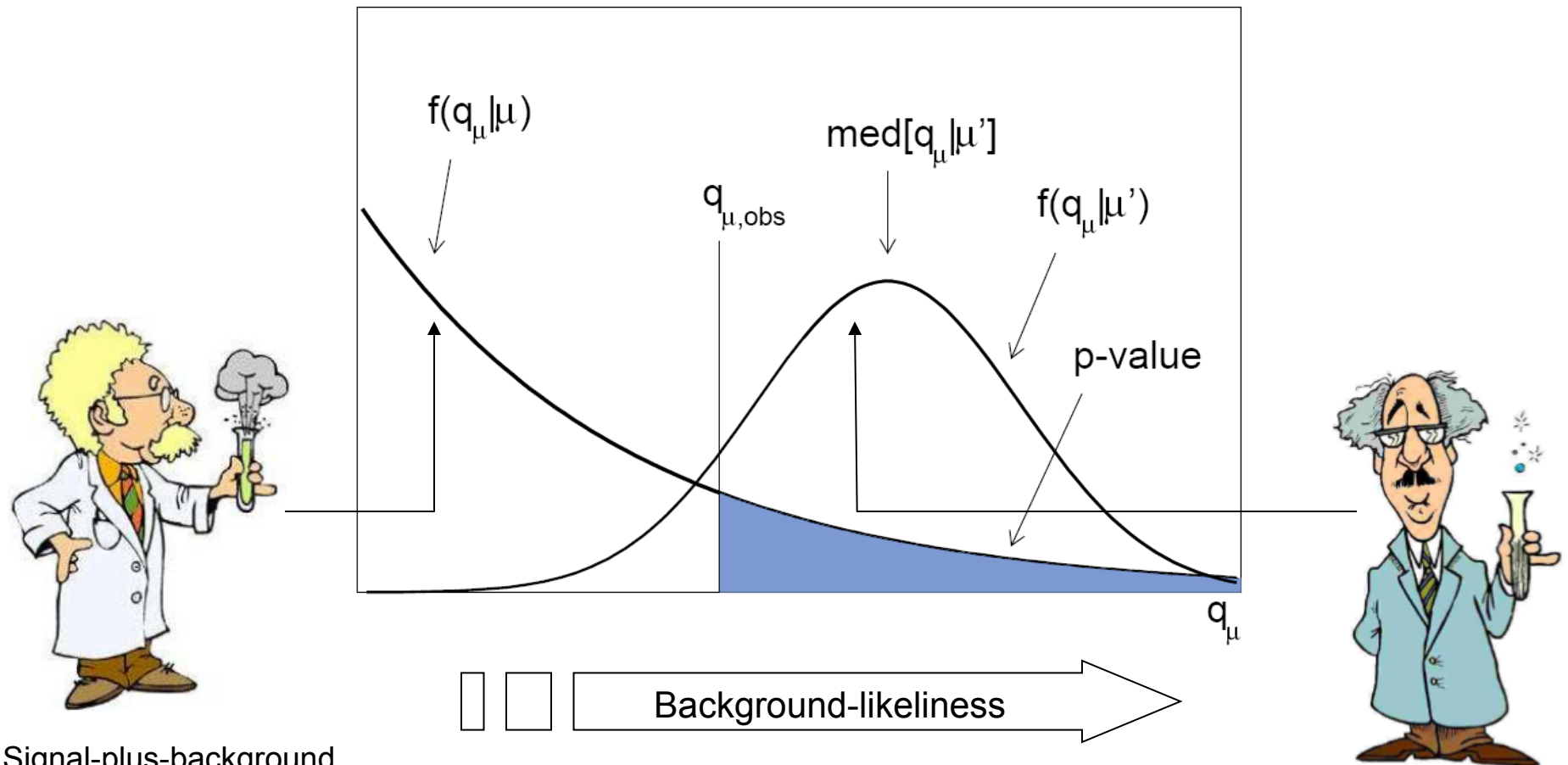
Usually work with the estimator :  $q_\mu = -2 \ln(\lambda(\mu))$  Because ...

# Wilks' Theorem

Under the  $H_\mu$  Signal hypothesis the PL is distributed as a  $\chi^2$  with 1 d.o.f. !

(v.i.z a well know analytical function)

To estimate the overall statistical behavior, toy MC full experiments are simulated and fitted !

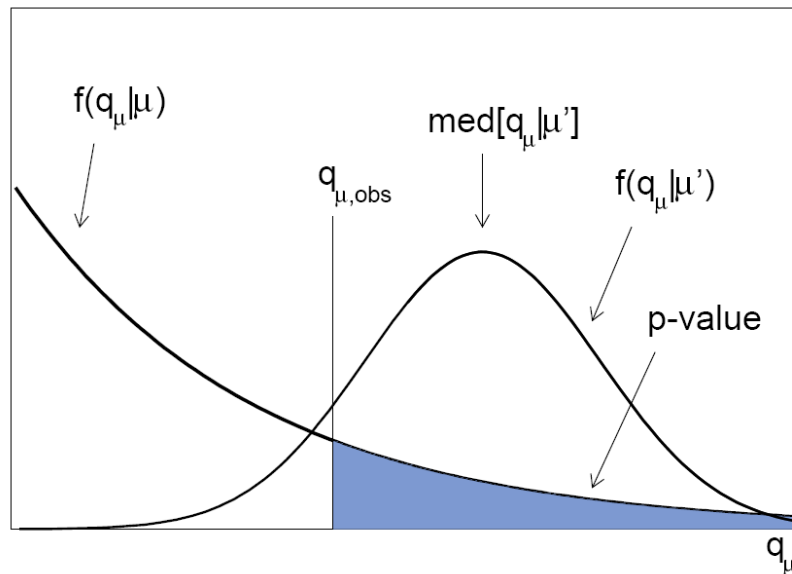


Signal-plus-background  
Toy experiments

Background only  
Toy experiments ( $\mu'=0$ )

# 95% CL Limits

The observed 95% CL upper limit on  $\mu$  is obtained by varying  $\mu$  until the p value :



$$1 - CL_{s+b} = p = \int_{q_{obs}}^{+\infty} f(q_\mu | \mu) dq_\mu = 5\%$$

Analytically simple

This means in other words that if there is a signal with strength  $\mu$ , the false exclusion probability is 5%.

The 95% CL exclusion sensitivity is obtained by varying  $\mu$  until the p value :

$$p = \int_{\underbrace{\text{med}(q_\mu | 0)}}^{+\infty} f(q_\mu | \mu) dq_\mu = 5\%$$

Background only experiments

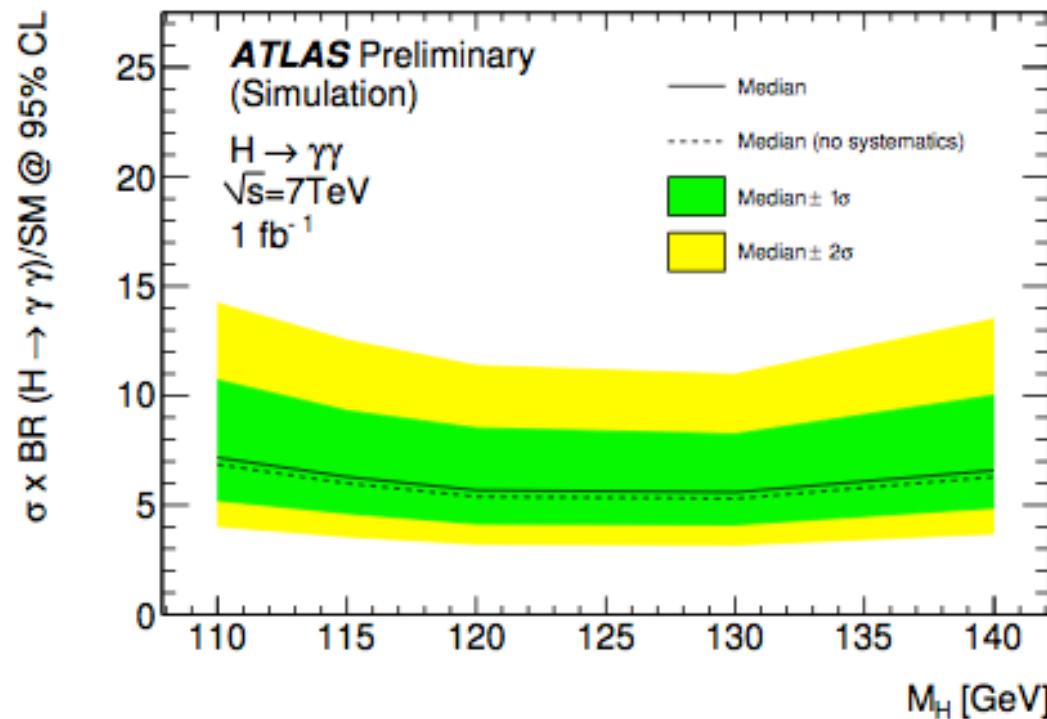


# Exclusion Results

Performing this analysis for several mass hypotheses and using  $CL_{s+b}$  the exclusion has the same problem as the simple Poisson exclusion with background...

No-Signal ( $H_0$ ) and presence of Signal ( $H_1$ )...

i.e. a signal of 0 can be excluded with a fluctuation of the background



We thus apply the (conservative) “modified frequentist” approach that requires :

$$CL_s = CL_{s+b} / CL_b = 5\% \quad \text{where} \quad CL_b = \int_{q_{obs}}^{+\infty} f(q_\mu | 0) dq_\mu$$

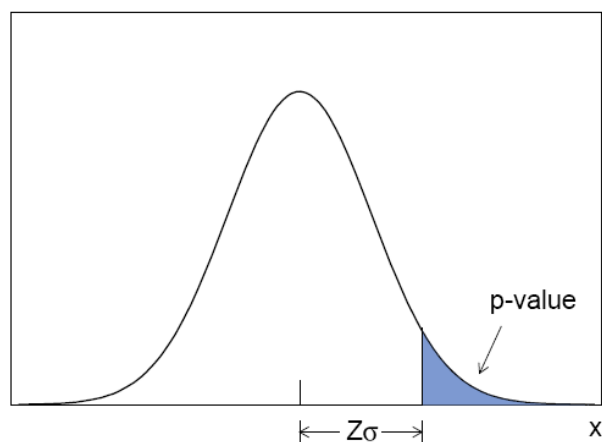
# Observation and Discovery

The method is essentially the same, only the estimator changes...we now use  $q_0$

In this case the  $f(q_0|0)$  will be distributed as a  $\chi^2$  with 1 d.o.f. (Wilks' theorem)

$$p = \int_{q_{obs}}^{+\infty} f(q_0|0) dq_0$$

- To claim an observation ( $3 \sigma$ ) : the conventional p-value required is  $1.35 \cdot 10^{-3}$
- To claim an observation ( $5 \sigma$ ) : the conventional p-value required is  $2.87 \cdot 10^{-7}$



Corresponds to the “one sided” convention

This means in other words that in absence of signal, the false discovery probability is  $p$ .

« a probability of 1 in 10 000 000 is *almost* impossible to estimate »

R. P. Feynman

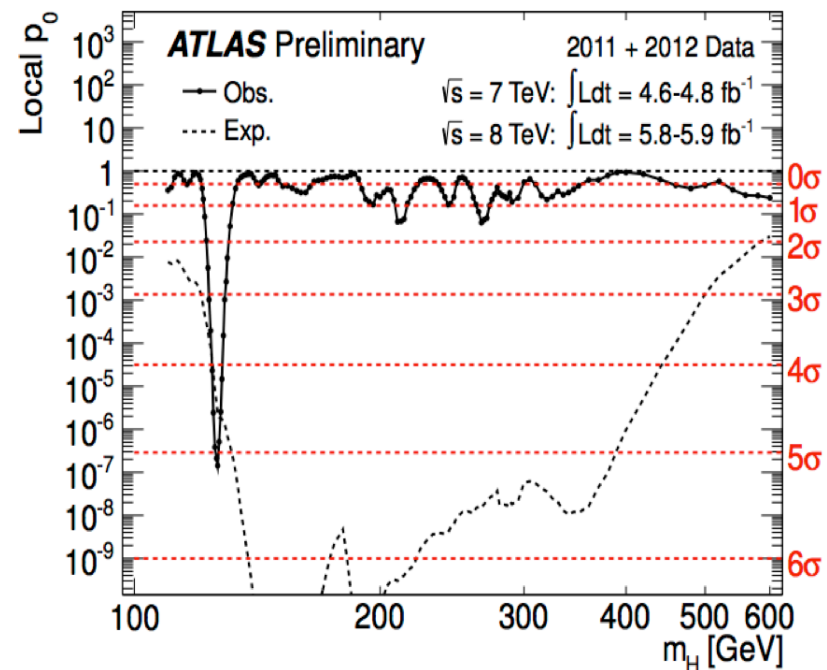
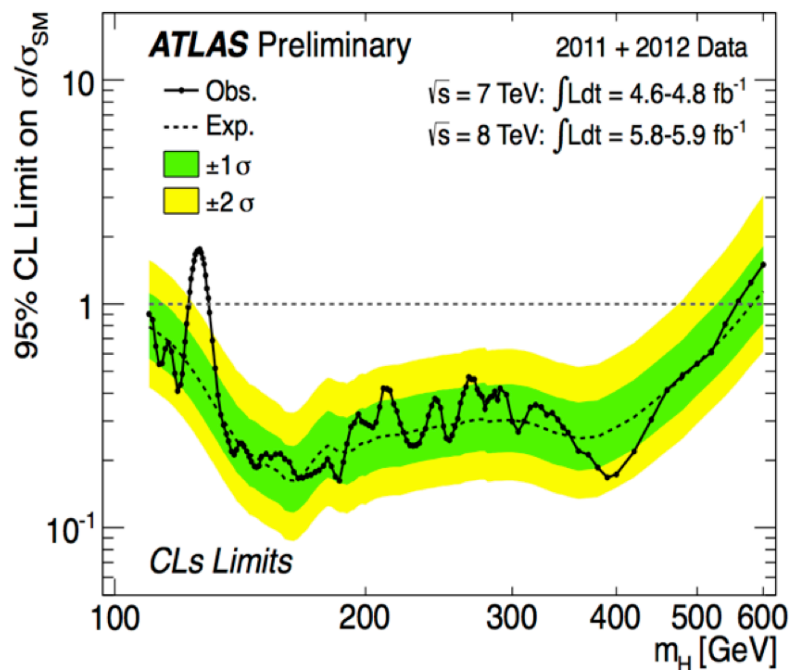
(What do you care what other people think?)

# Conclusion

We went through an overview of the fundamental concepts of statistics for HEP

If possible take some time to play with the Root-Macros for hands-on experience

*You should now be able to understand the following plot !*



There is a lot more for you/us to learn about statistical techniques

In particular concerning the treatment of systematics

So be patient and take some time to understand the techniques step by step...

... and follow Laplace's advice about statistics !