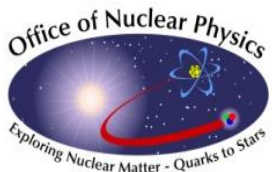


Using Weak Interaction to Probe Neutrons in Nuclei the Why and How Experimentally

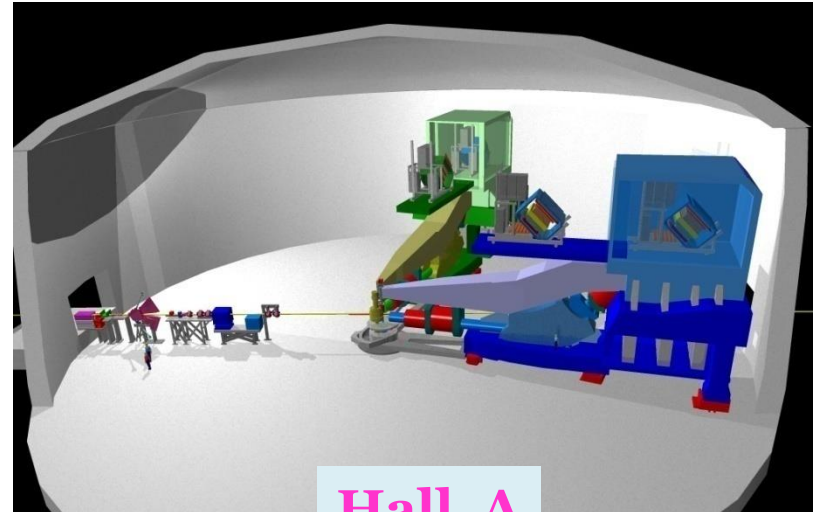
Robert Michaels
Jefferson Lab
(JLab)



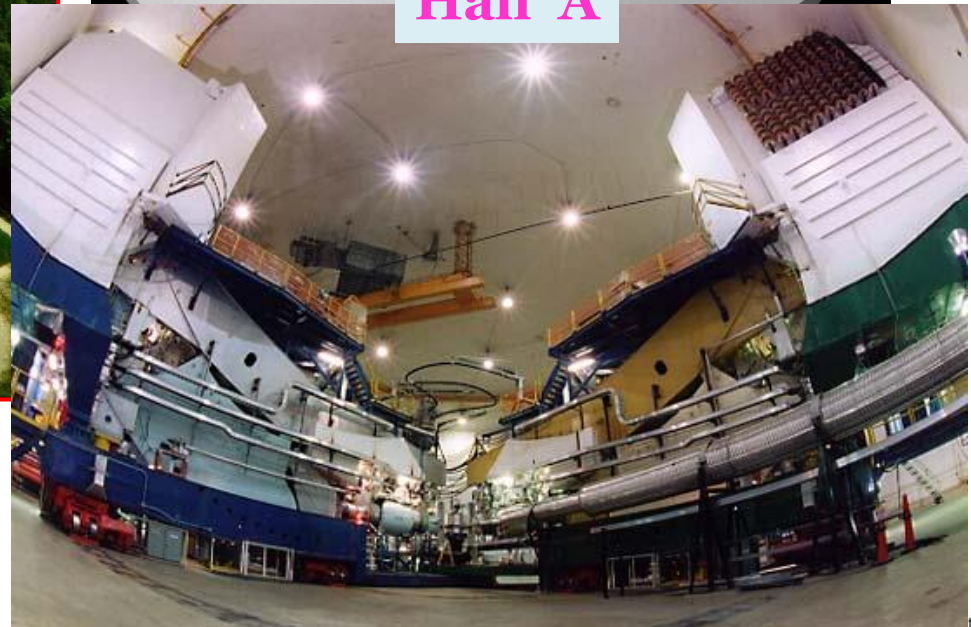
Thomas Jefferson National Accelerator Facility



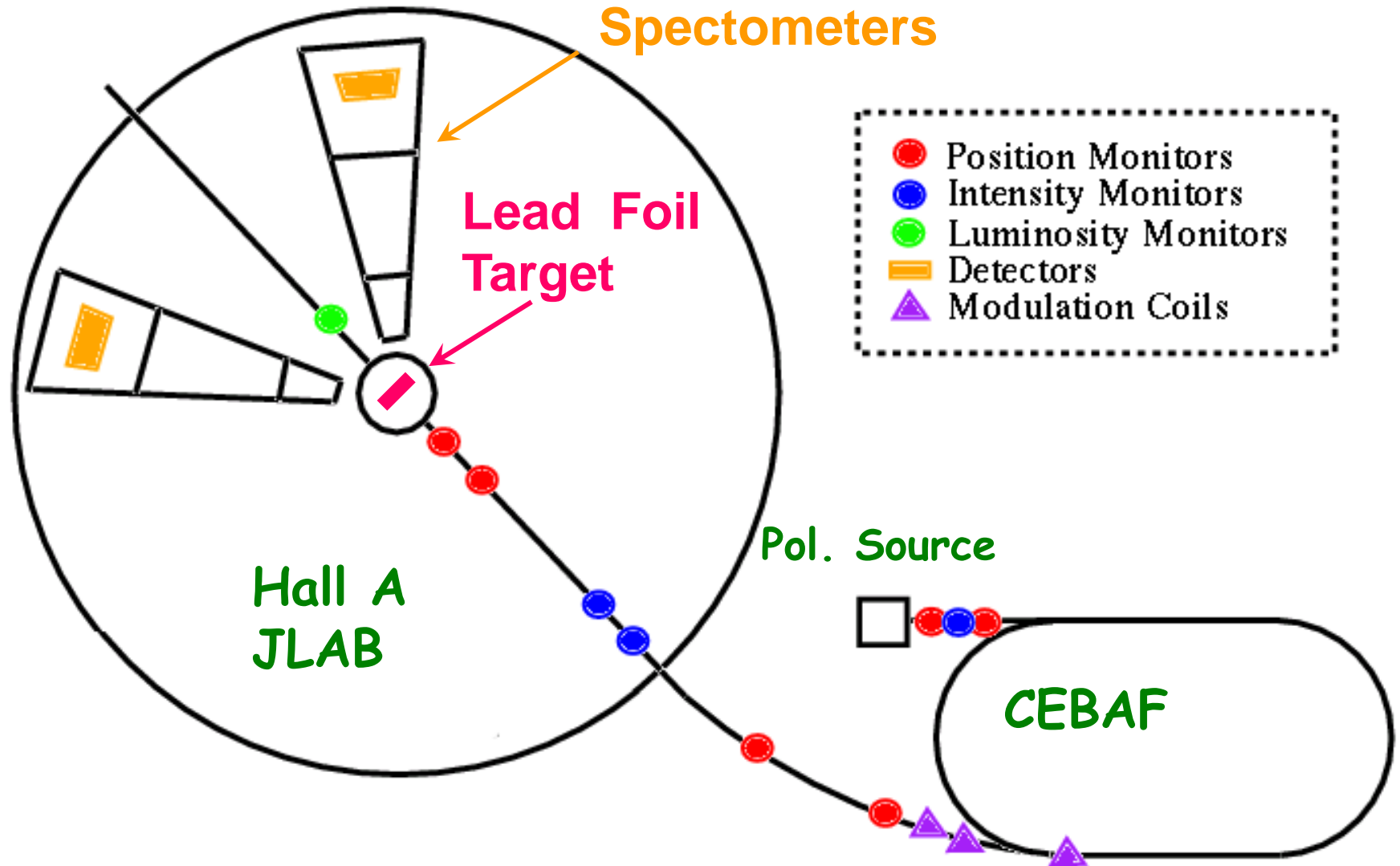
Hall A at Jefferson Lab



Hall A



PREX Setup



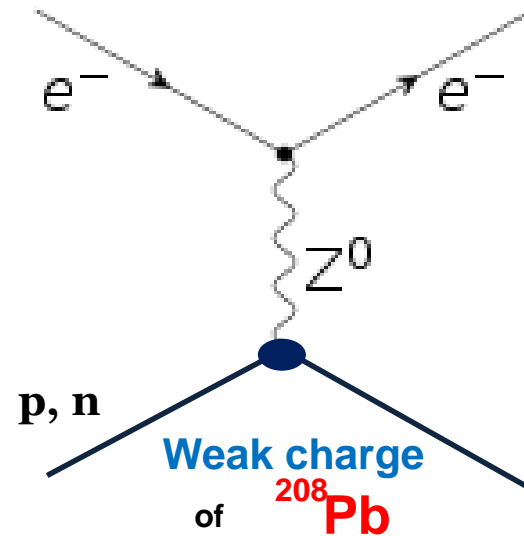
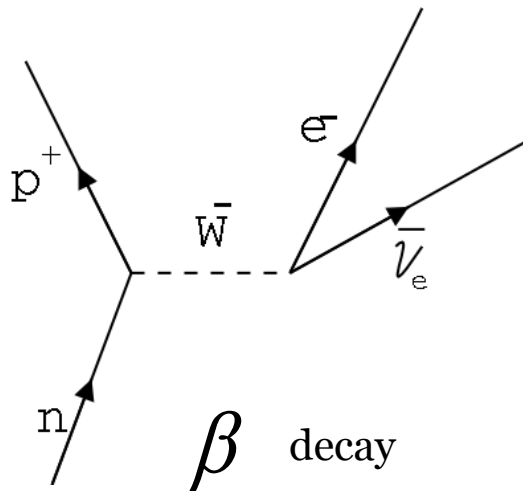
Weak Interactions

The **Glashow-Weinberg-Salam Theory** unifies the electromagnetic and weak interactions.

Left-handed fermion fields (quarks & leptons)
= doublets under SU(2)

Right-handed fields = singlets under SU(2)

Parity Violation



How to isolate the *weak interaction*

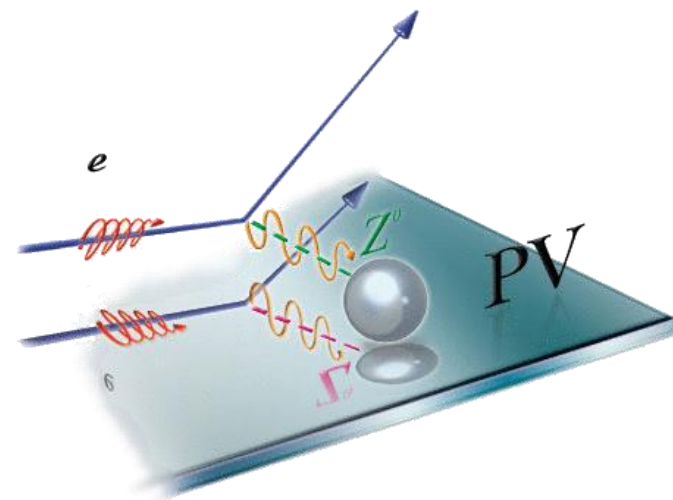


Do 2 experiments that are *mirror images* :

Weak interaction looks **different**.

EM interaction looks **same**.

The *weak interaction* changes with mirror imaging which allows to isolate it.



Positive spin



Negative spin

Incident electron

spin momentum



Target



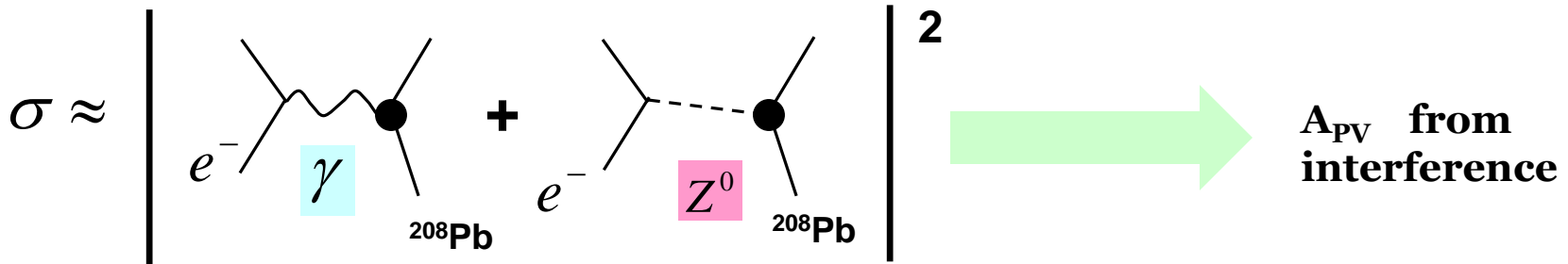
A nucleus in target



Method: Flip spin of electrons and look for difference in scattering rate.

Parity Violating Asymmetry

$$A_{PV} = \frac{\sigma_R - \sigma_L}{\sigma_R + \sigma_L} \sim 10^{-4} \times Q^2 \sim 10^{-6}$$



Applications of A_{PV} at Jefferson Lab

- **Nucleon Structure**

Strangeness $s \bar{s}$ in proton (HAPPEX, G0 expts)

- **Test of Standard Model of Electroweak** $\sin^2 \theta_W$

$e-e$ (MOLLER) or $e-q$ (PVDIS)

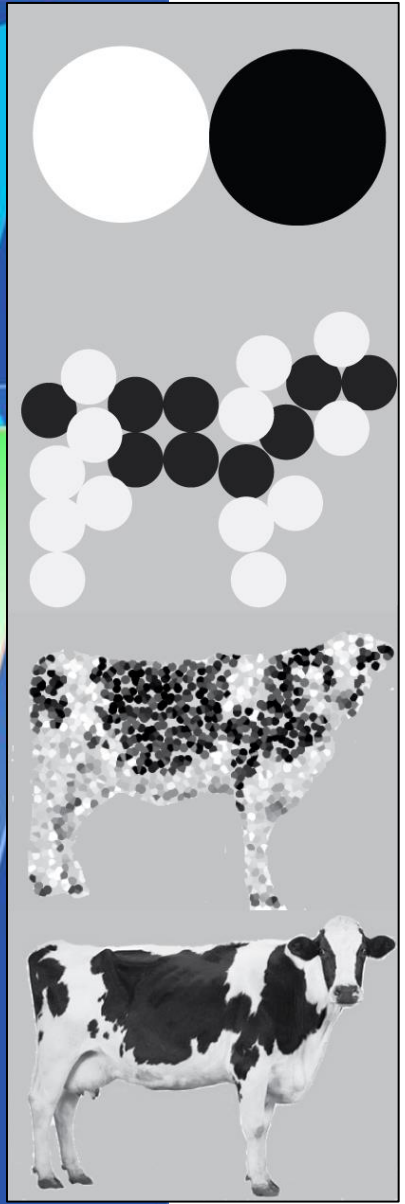
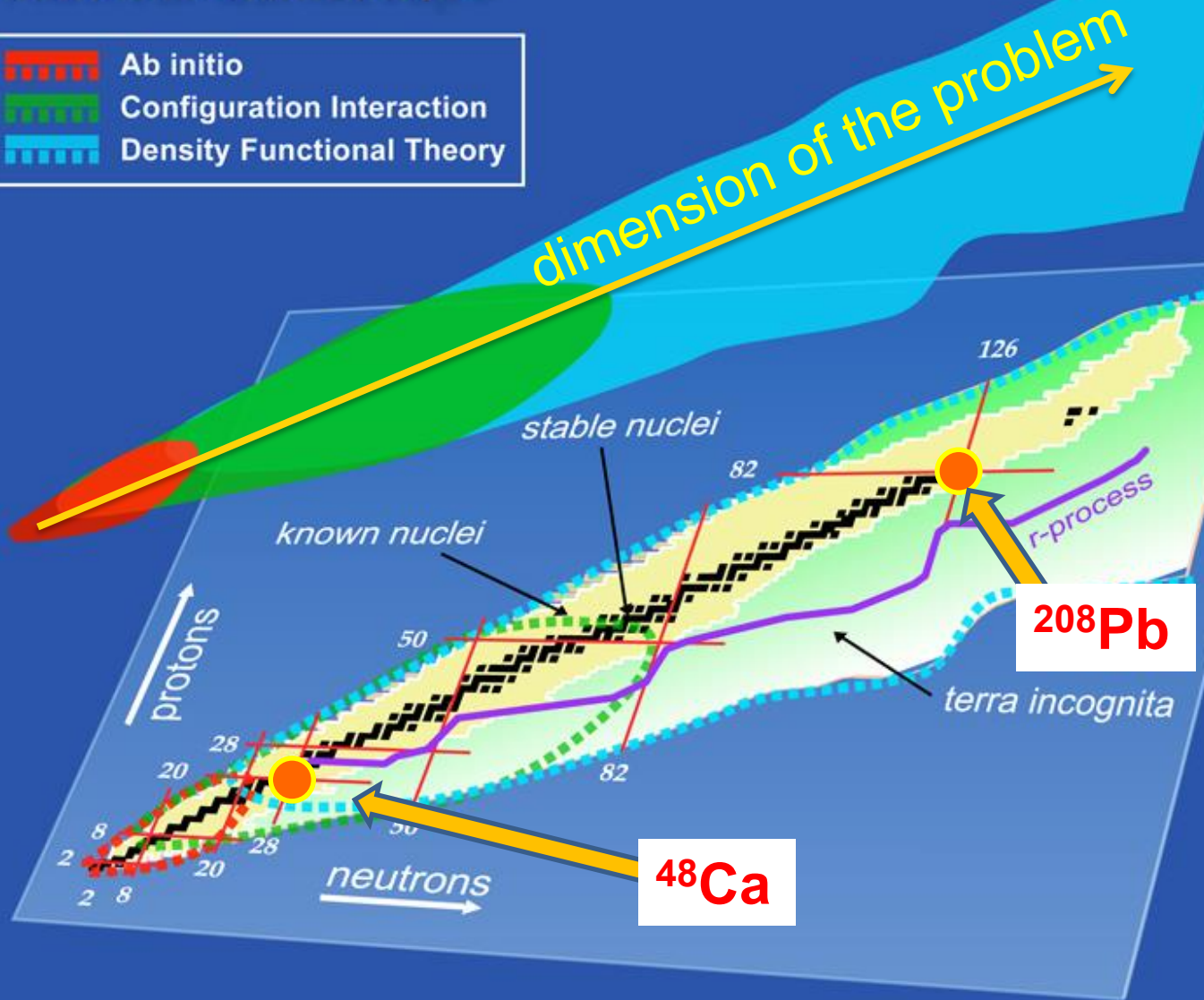
elastic $e-p$ at low Q^2 (QWEAK)

- **Nuclear Structure (neutron density) : PREX & CREX**

How to explain the nuclear landscape from the bottom up? **Theory roadmap**

Credit: Witek Nazarewicz

Nuclear Landscape



J. Phys. G 43, 044002 (2016)

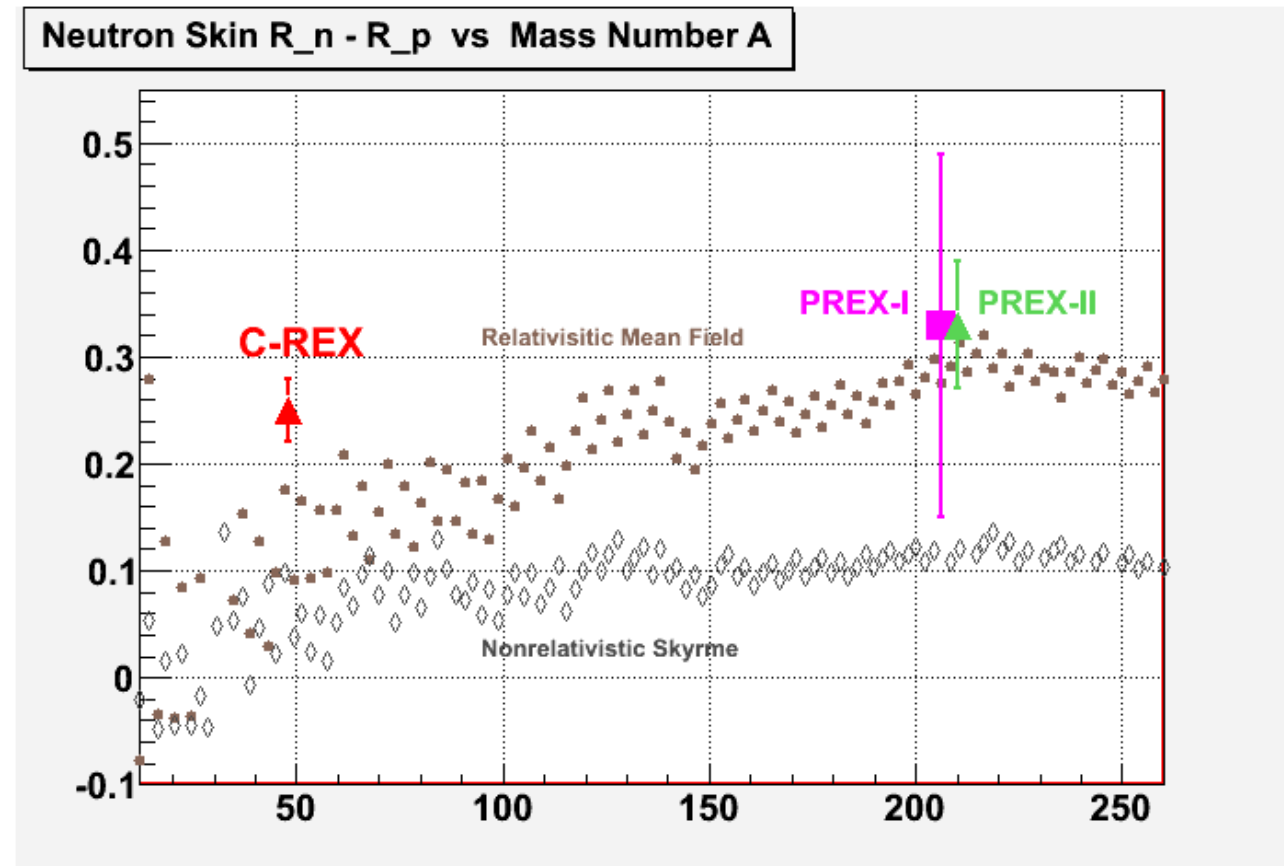
<http://iopscience.iop.org/article/10.1088/0954-3899/43/4/044002>

PREX (^{208}Pb) and CREX (^{48}Ca)

Measuring Neutron Skins $R_n - R_p$ provides new information and new constraints on nuclear structure theory

^{208}Pb more closely approximates infinite nuclear matter (and neutron stars)

The structure of ^{48}Ca can, only recently, be addressed in detailed microscopic models.



Theory from P. Ring et al. Nucl. Phys. A 624 (1997) 349

PREX
Physics
Output

Atomic
Parity
Violation

Measured Asymmetry

Correct for Coulomb
Distortions

Weak Density at one Q^2

Small Corrections for
 G_E^n G_E^s MEC

Neutron Density at one Q^2

Assume Surface Thickness
Good to 25% (MFT)

R_n

Mean Field
& Other
Models

Neutron
Stars

Slide adapted from
C. Horowitz

Fundamental Nuclear Physics :

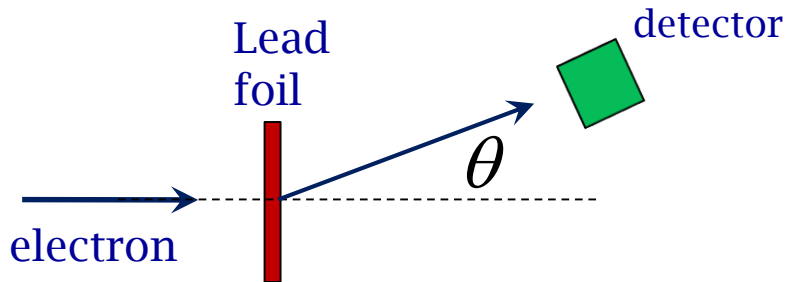
What is the size of a nucleus ?

Neutrons are thought to determine the size of heavy nuclei like ^{208}Pb .

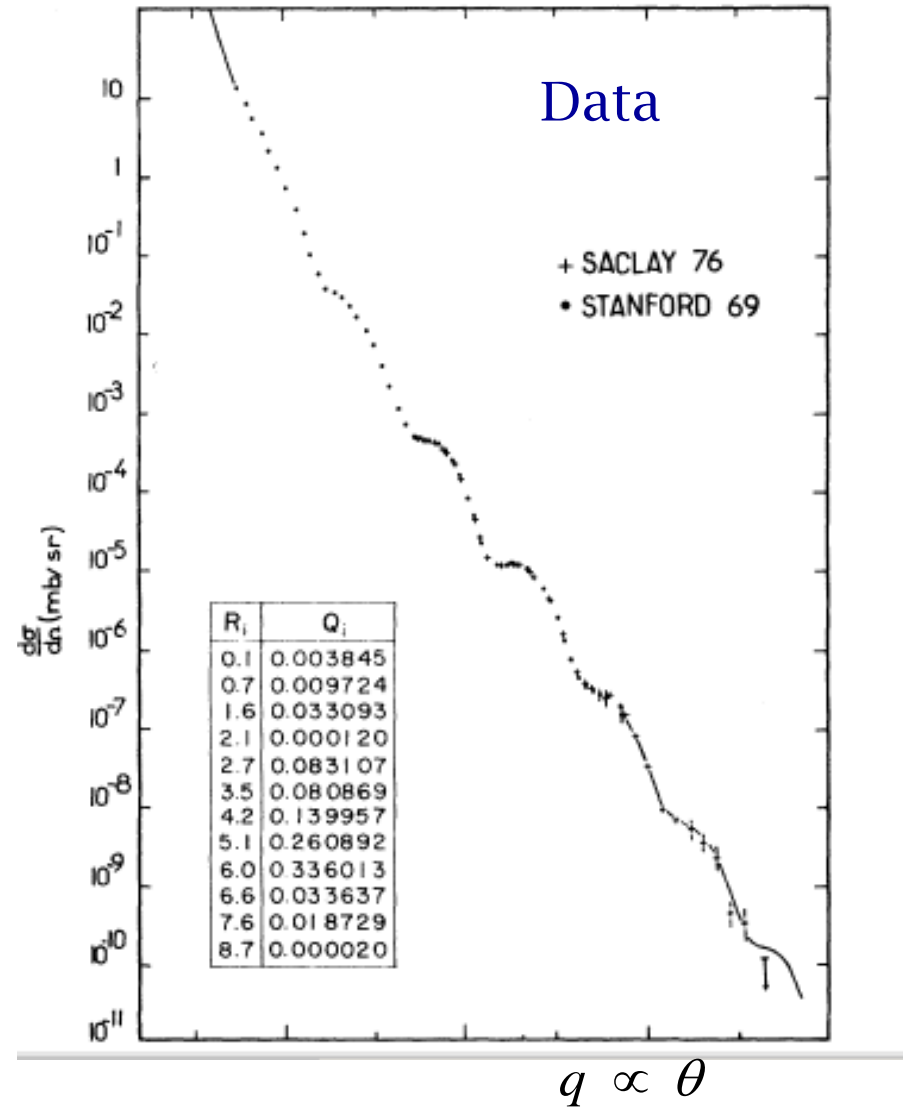
Can theory predict it ?

Scattering of High-Energy (here ~500 MeV) **Electrons** from **Lead Nuclei**.

The **nuclei** are the
“*mysterious structures*”
causing a pattern in the
scattered electrons.



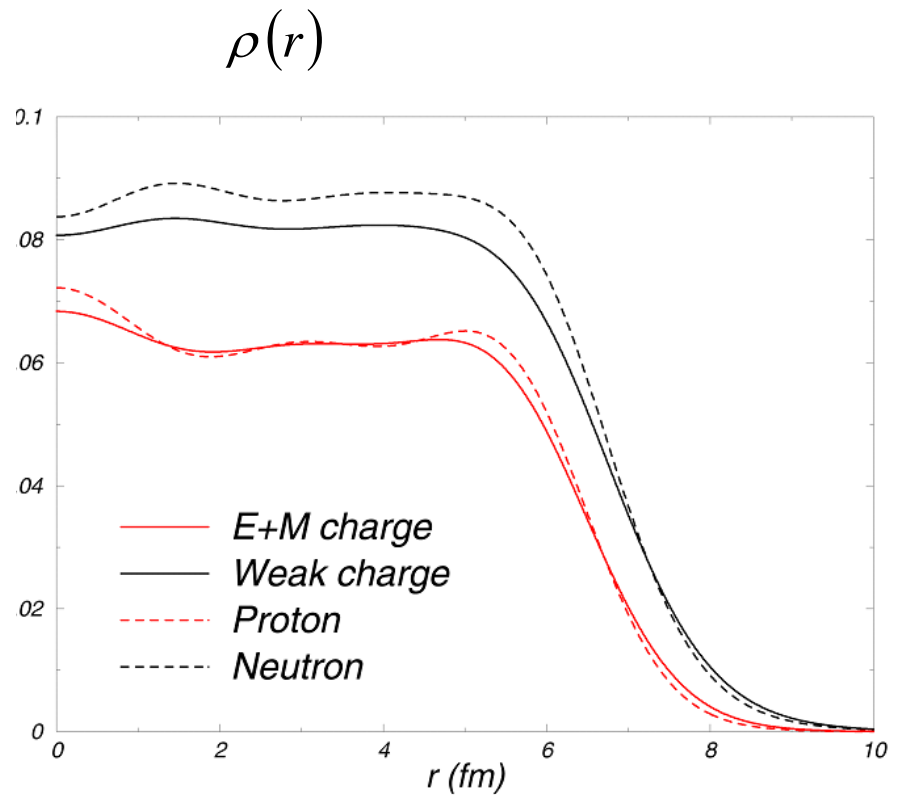
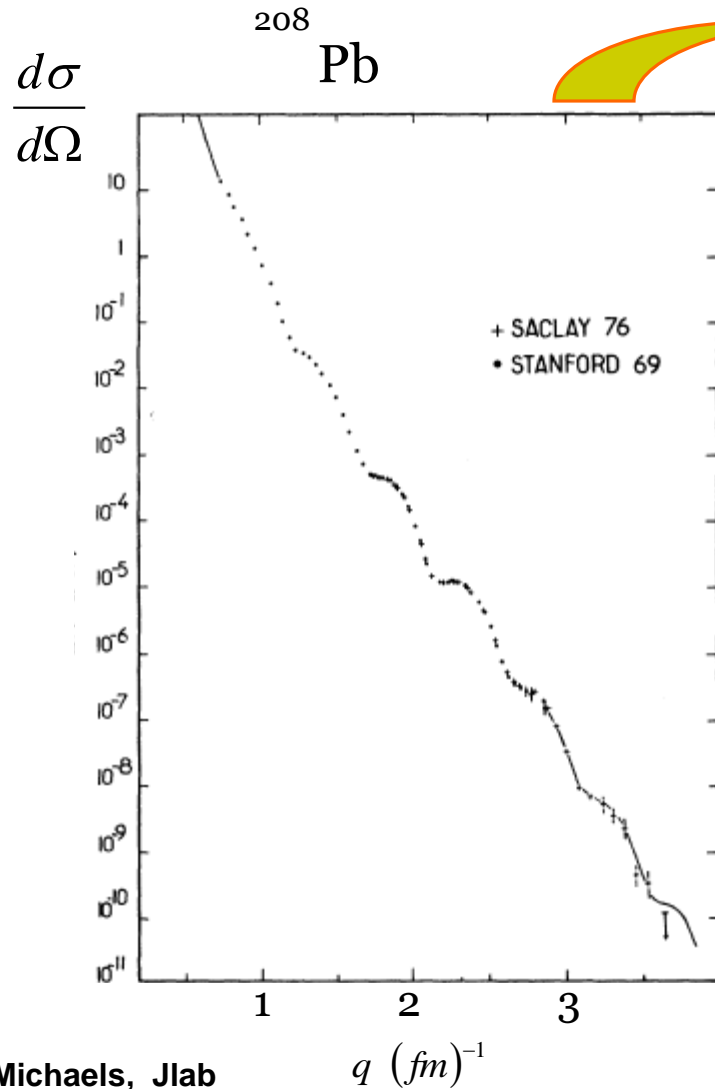
Scattering Rate vs Momentum Transfer



Reminder: Electromagnetic Scattering determines

$$\rho(r)$$

(charge distribution)



Z^0 of weak interaction : sees the neutrons

T.W. Donnelly, J. Dubach, I. Sick
Nucl. Phys. A 503, 589, 1989

C. J. Horowitz, S. J. Pollock,
P. A. Souder, R. Michaels
Phys. Rev. C 63, 025501, 2001

^{208}Pb

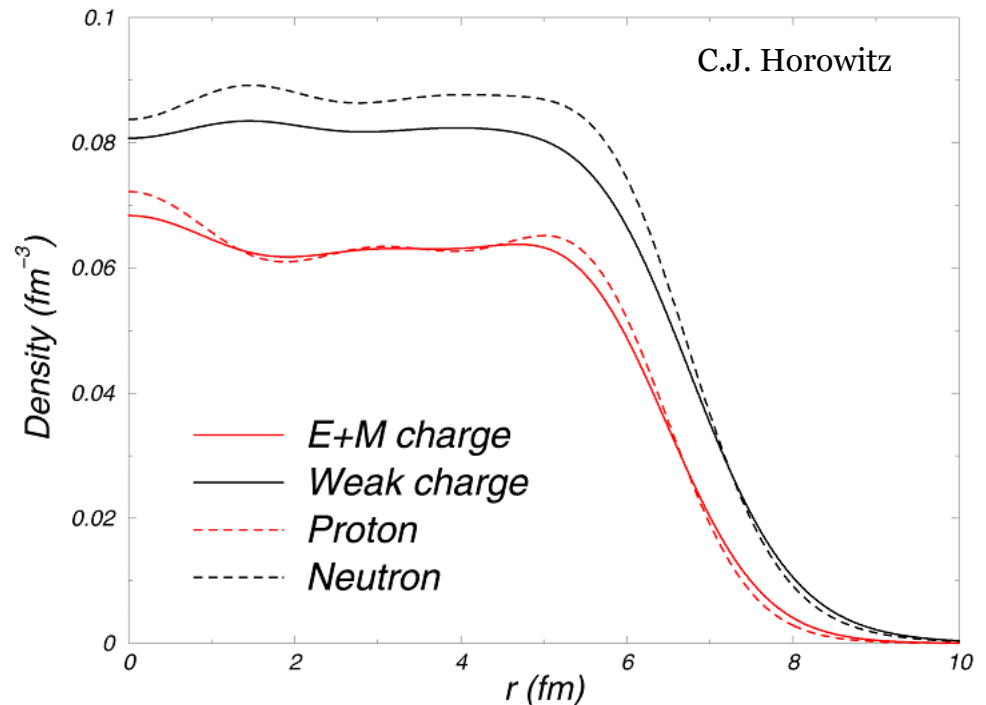
	proton	neutron
Electric charge	1	0
Weak charge	0.08	1

Neutron form factor

$$F_N(Q^2) = \frac{1}{4\pi} \int d^3r j_0(qr) \rho_N(r)$$

Parity
Violating
Asymmetry

$$A = \frac{G_F Q^2}{2\pi\alpha\sqrt{2}} \left[\underbrace{1 - 4\sin^2\theta_W}_{\approx 0} - \frac{F_N(Q^2)}{F_P(Q^2)} \right]$$

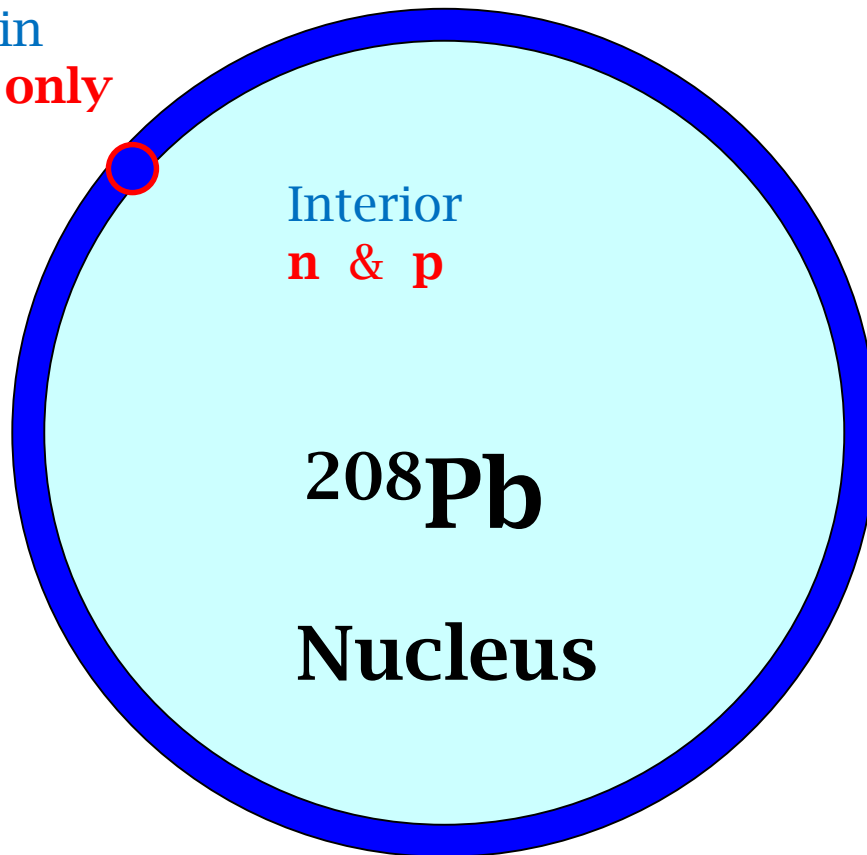


How to Measure Neutron Distributions, Symmetry Energy

- Proton-Nucleus Elastic
 - Pion, alpha, d Scattering
 - Pion Photoproduction
 - Heavy ion collisions
 - Rare Isotopes (dripline)
- Involve strong probes, the interpretation is clouded by understanding of the reaction mechanism
- Magnetic scattering → Most spins couple to zero.
 - **PREX / C-REX** (weak interaction)
 - Theory → MFT fit mostly by data *other than* neutron densities

PREX : Neutron Skin

Skin
n only



Was expected since
more neutrons (n)
than protons (p)

Observed by PREX-I
(95 % confidence)

Parity Violating
electron scattering is
“cleaner” than other
probes
(e.g. proton scattering)

Fundamental
check of nuclear
theory

Using Parity Violation

Electron - Nucleus Potential $\hat{V}(r) = V(r) + \gamma_5 A(r)$

electromagnetic

$$V(r) = \int d^3 r' Z \rho(r') / |\vec{r} - \vec{r}'|$$

^{208}Pb is spin 0

$$\frac{d\sigma}{d\Omega} = \frac{d\sigma}{d\Omega_{Mott}} |F_P(Q^2)|^2$$

Proton form factor

$$F_P(Q^2) = \frac{1}{4\pi} \int d^3 r j_0(qr) \rho_P(r)$$

Parity Violating Asymmetry

$$A = \frac{\left(\frac{d\sigma}{d\Omega}\right)_R - \left(\frac{d\sigma}{d\Omega}\right)_L}{\left(\frac{d\sigma}{d\Omega}\right)_R + \left(\frac{d\sigma}{d\Omega}\right)_L} = \frac{G_F Q^2}{2\pi\alpha\sqrt{2}} \left[\underbrace{1 - 4\sin^2\theta_W}_{\approx 0} - \frac{F_N(Q^2)}{F_P(Q^2)} \right]$$

axial

$$A(r) = \frac{G_F}{2\sqrt{2}} \left[(1 - 4\sin^2\theta_W) Z \rho_P(r) - N \rho_N(r) \right]$$

\Rightarrow $A(r)$ is small, best observed by parity violation

\Rightarrow $1 - 4\sin^2\theta_W \ll 1$ neutron weak charge \gg proton weak charge

Neutron form factor

$$F_N(Q^2) = \frac{1}{4\pi} \int d^3 r j_0(qr) \rho_N(r)$$

From low to medium to high density neutrons

Low density : nuclei

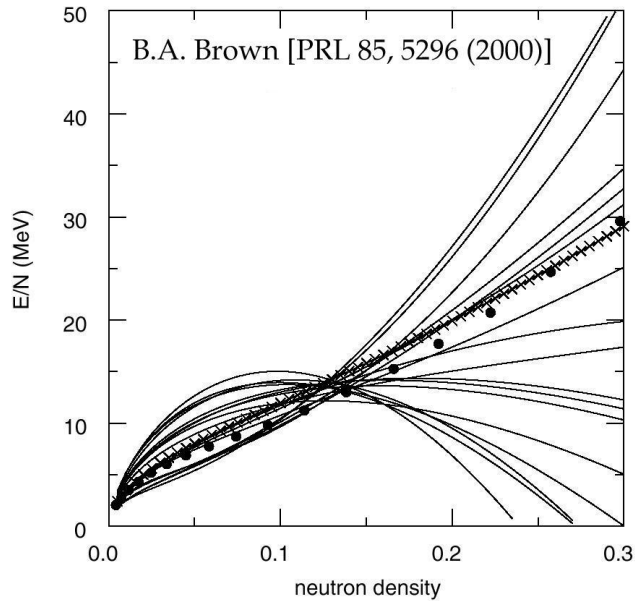
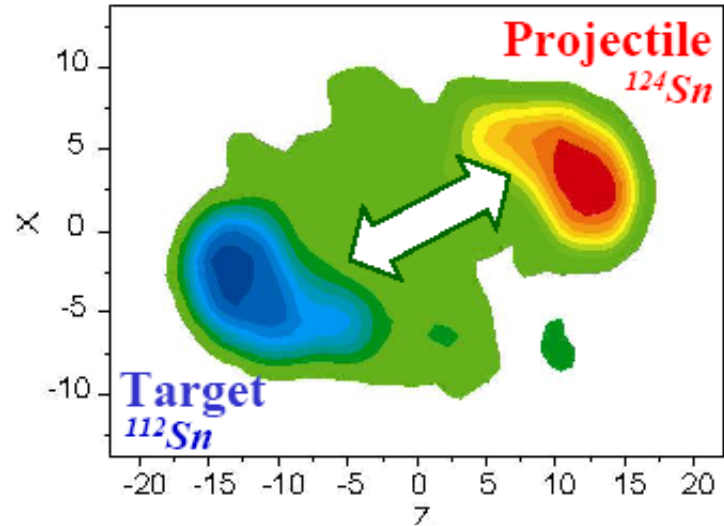


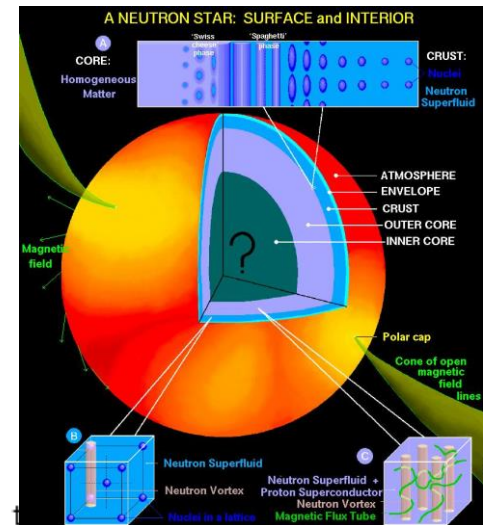
FIG. 2. The neutron EOS for 18 Skyrme parameter sets. The filled circles are the Friedman-Pandharipande (FP) variational calculations and the crosses are SkX. The neutron density is in units of $\text{neutron}/\text{fm}^3$.

Medium density : Heavy Ion Collisions



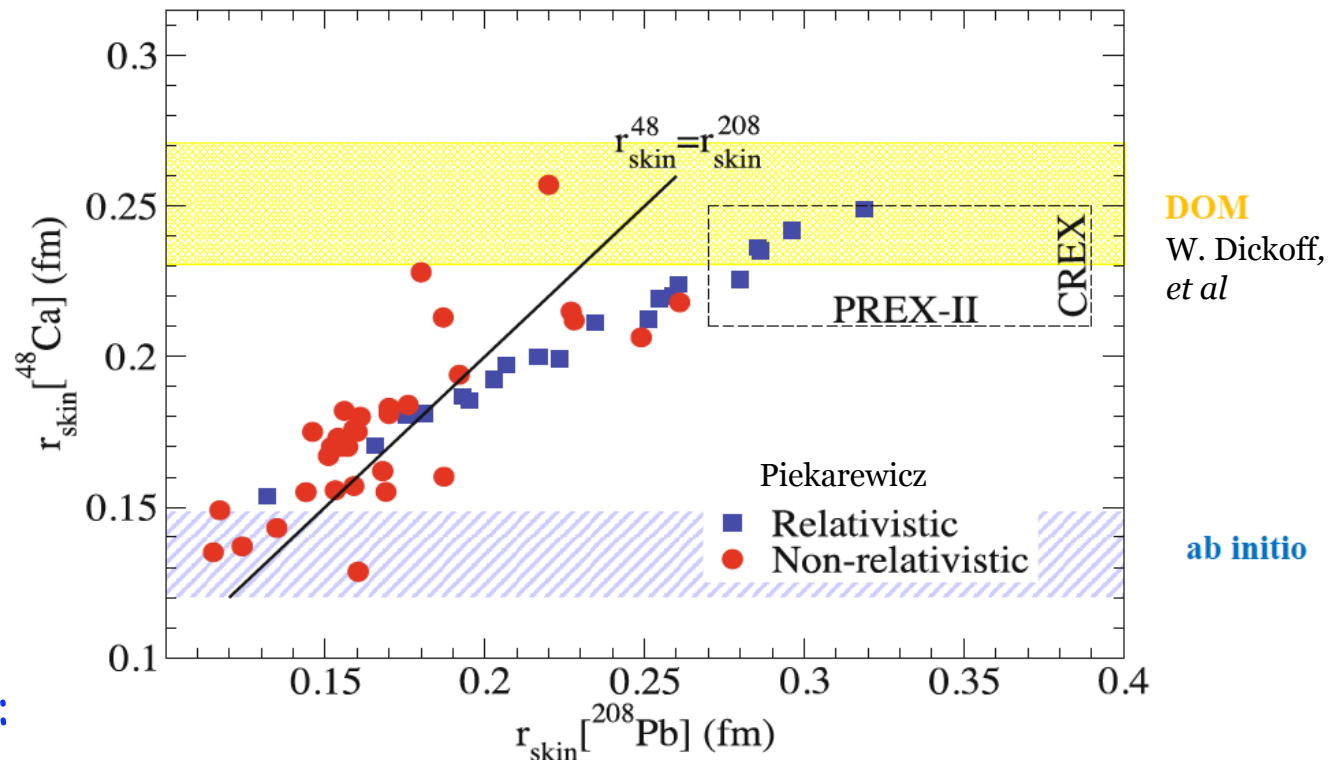
Danielewicz, Lacey, and Lynch,
Science 298 (2002) 1592.

Highest density : Neutron Stars



“Ab Initio” (exact microscopic) calculations of R_{skin} for ^{48}Ca have recently been published. G. Hagen et al., Nature Physics 12, 186 (2016).

Can be compared to Density Functional Theory (the red and blue points) and Dispersive Optical Model (DOM).



• “Ab initio”:

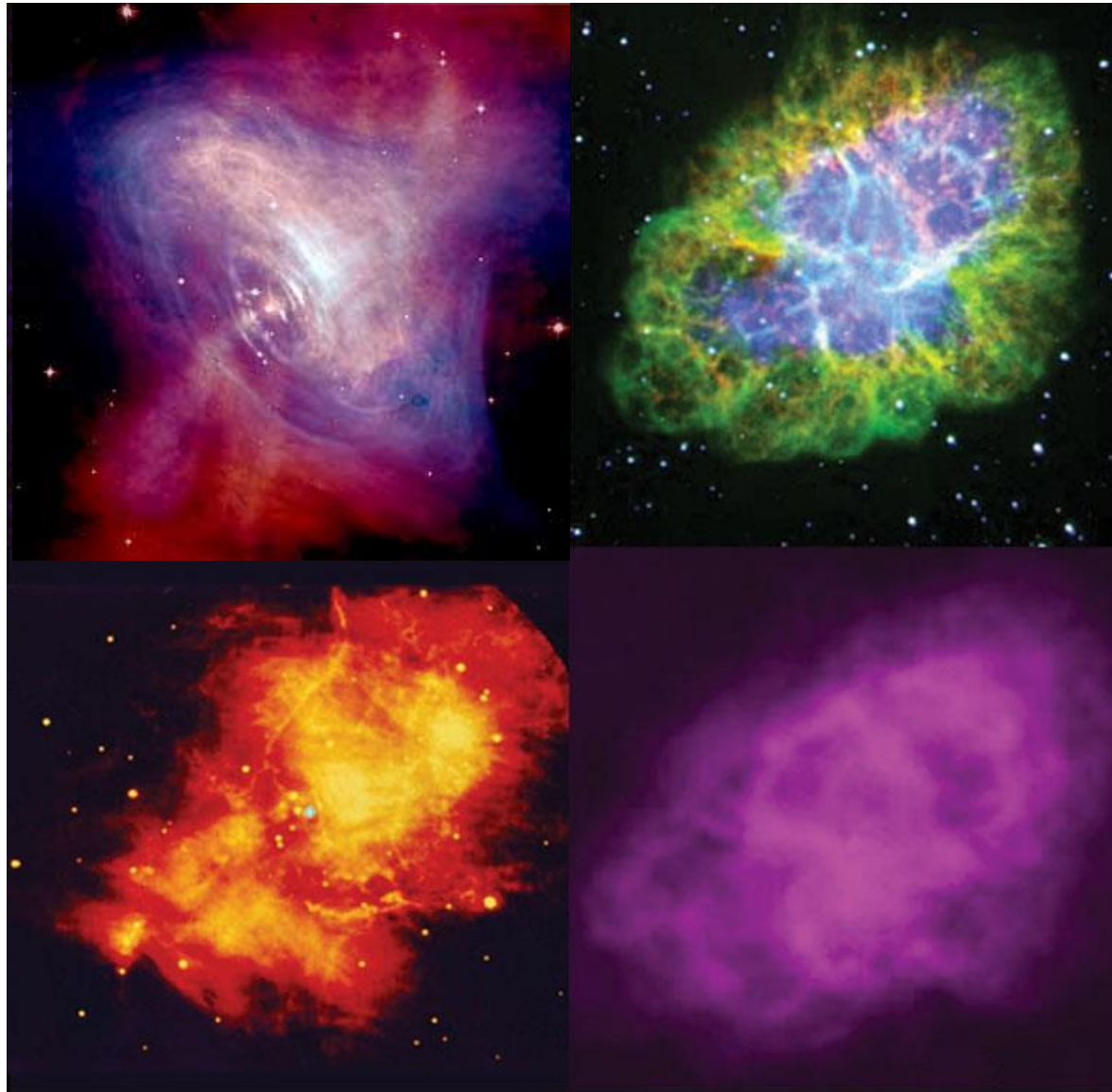
G. Hagen et al., Nature Phys. 12, 186 (2016)

Application :

Neutron Stars

What is the nature of extremely dense matter ?

Do collapsed stars form “exotic” phases of matter ? (strange stars, quark stars)



Crab Nebula (X-ray, visible, radio, infrared)

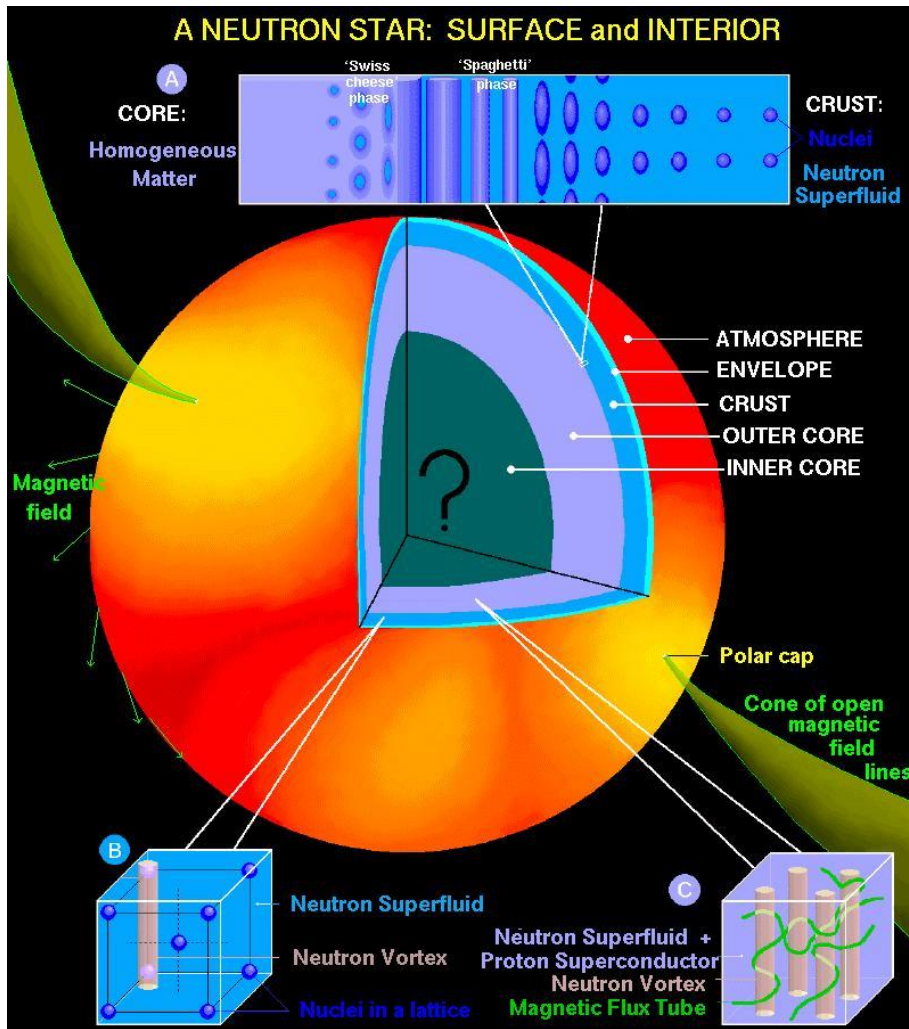


Fig from: Dany Page.

J.M. Lattimer & M. Prakash, Science 304 (2004) 536.

Inputs:

- Eq. of state (EOS)

$$P(\rho)$$

PREX helps here

- Hydrostatics (Gen. Rel.)
- Astrophysics Observations

Luminosity L

Temp. T

Mass M from pulsar timing

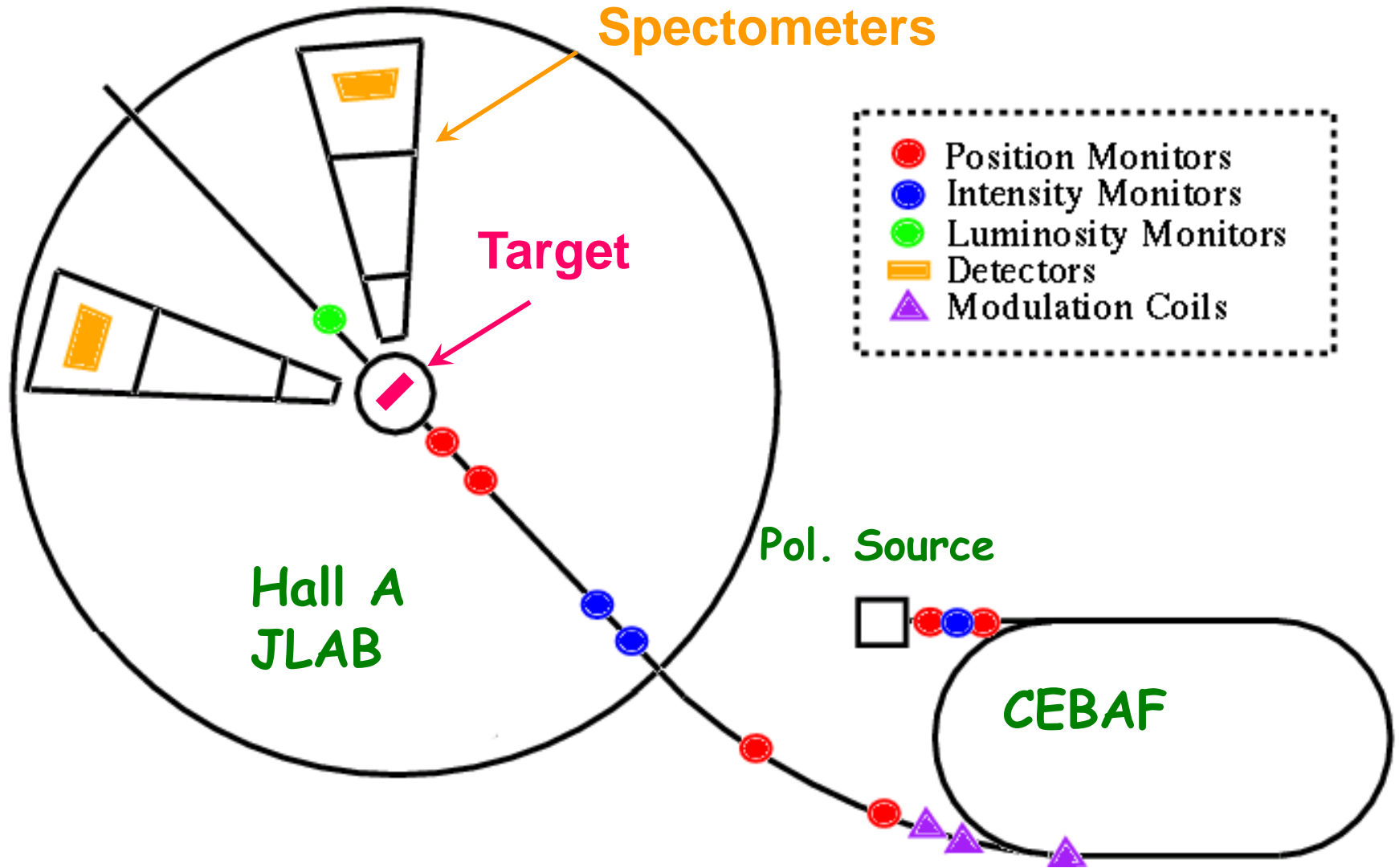
$$L = 4\pi\sigma_B R^2 T^4$$

(with corrections ...)

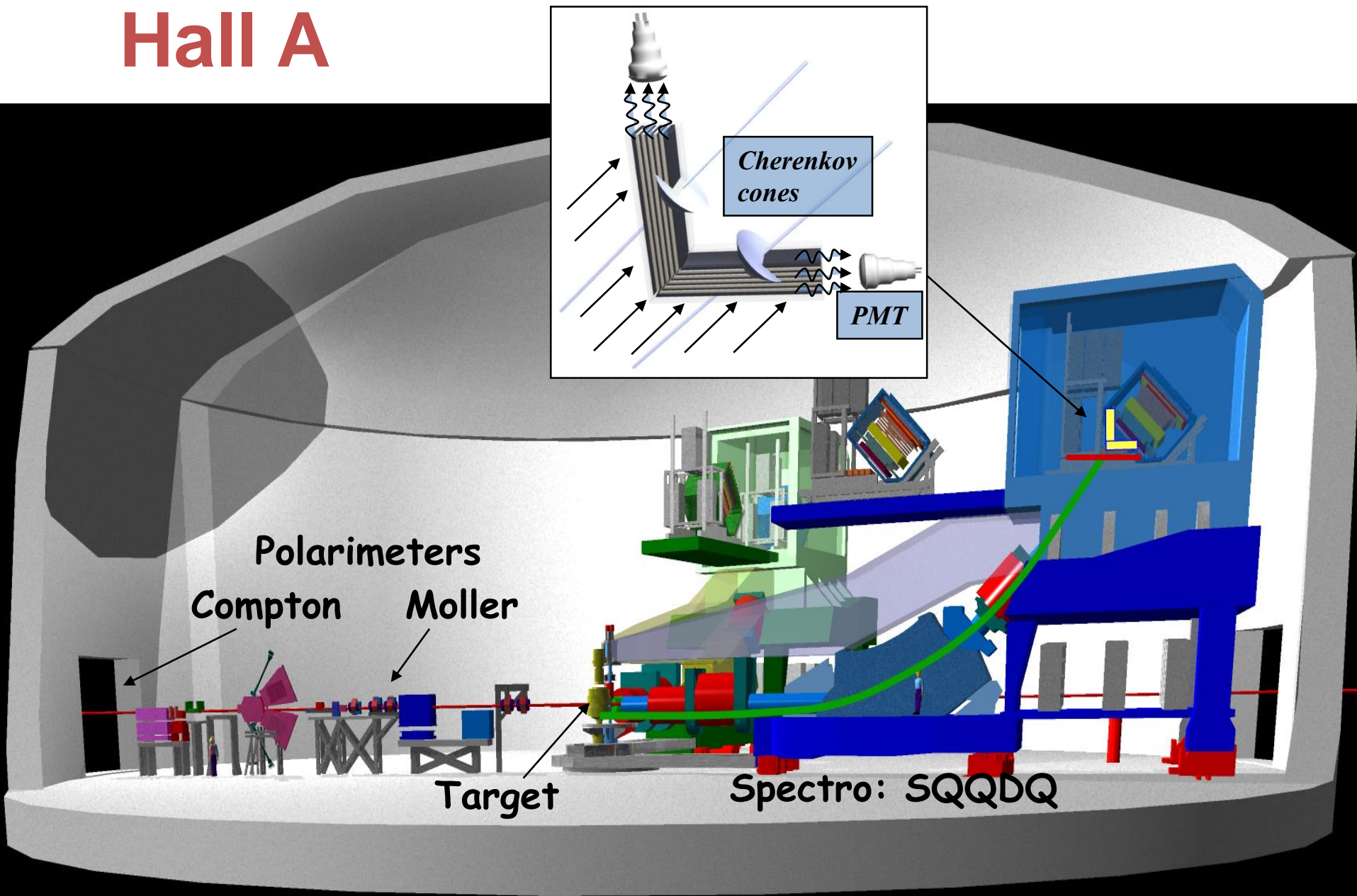
Mass - Radius relationship

Experiment Setup

Parity: "The entire lab is the experiment"

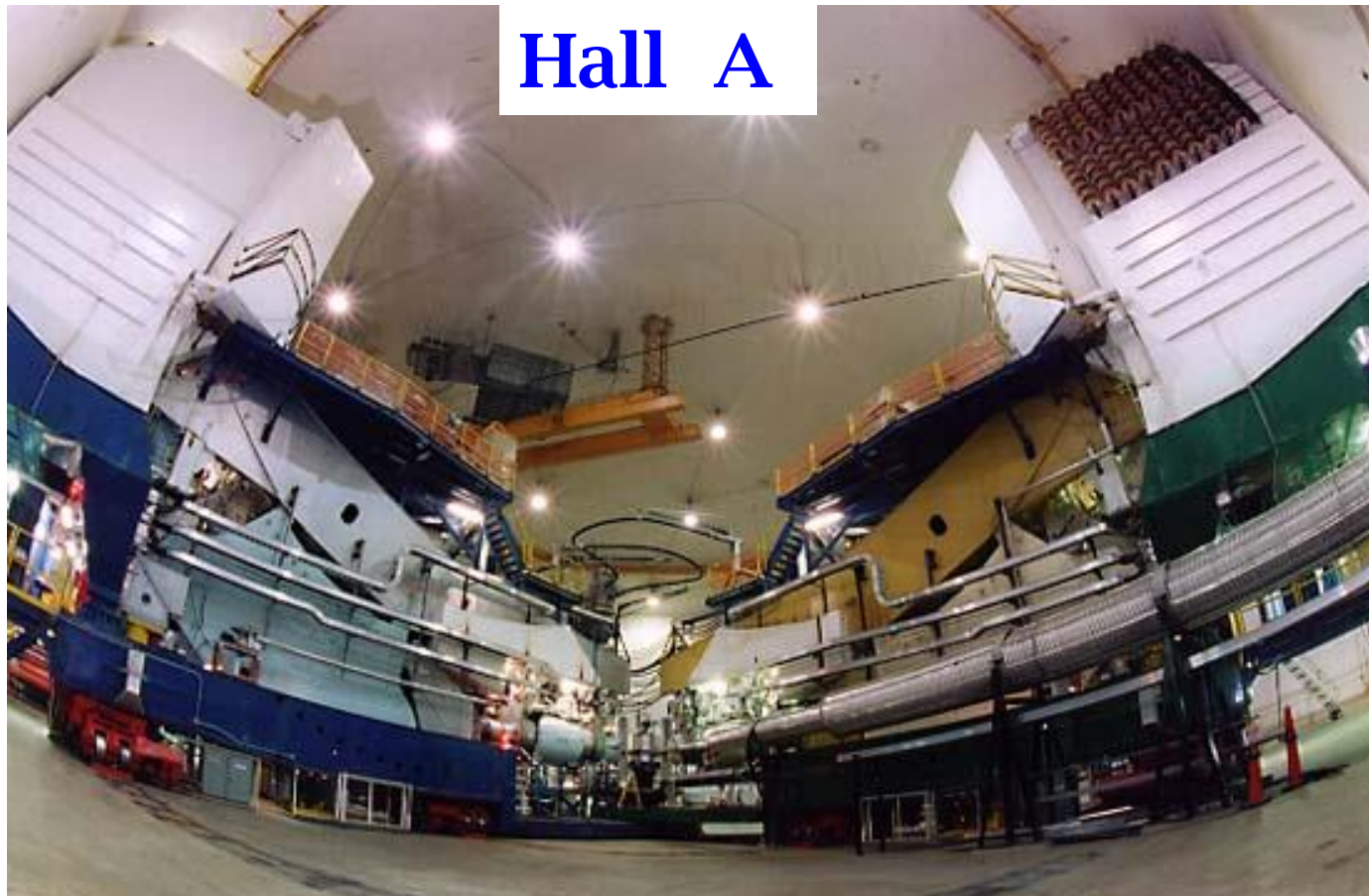


Hall A



Spectrometers at Jefferson Lab

These machines are “microscopes” for looking at quarks.



Spectrometers at Jefferson Lab

Hall A



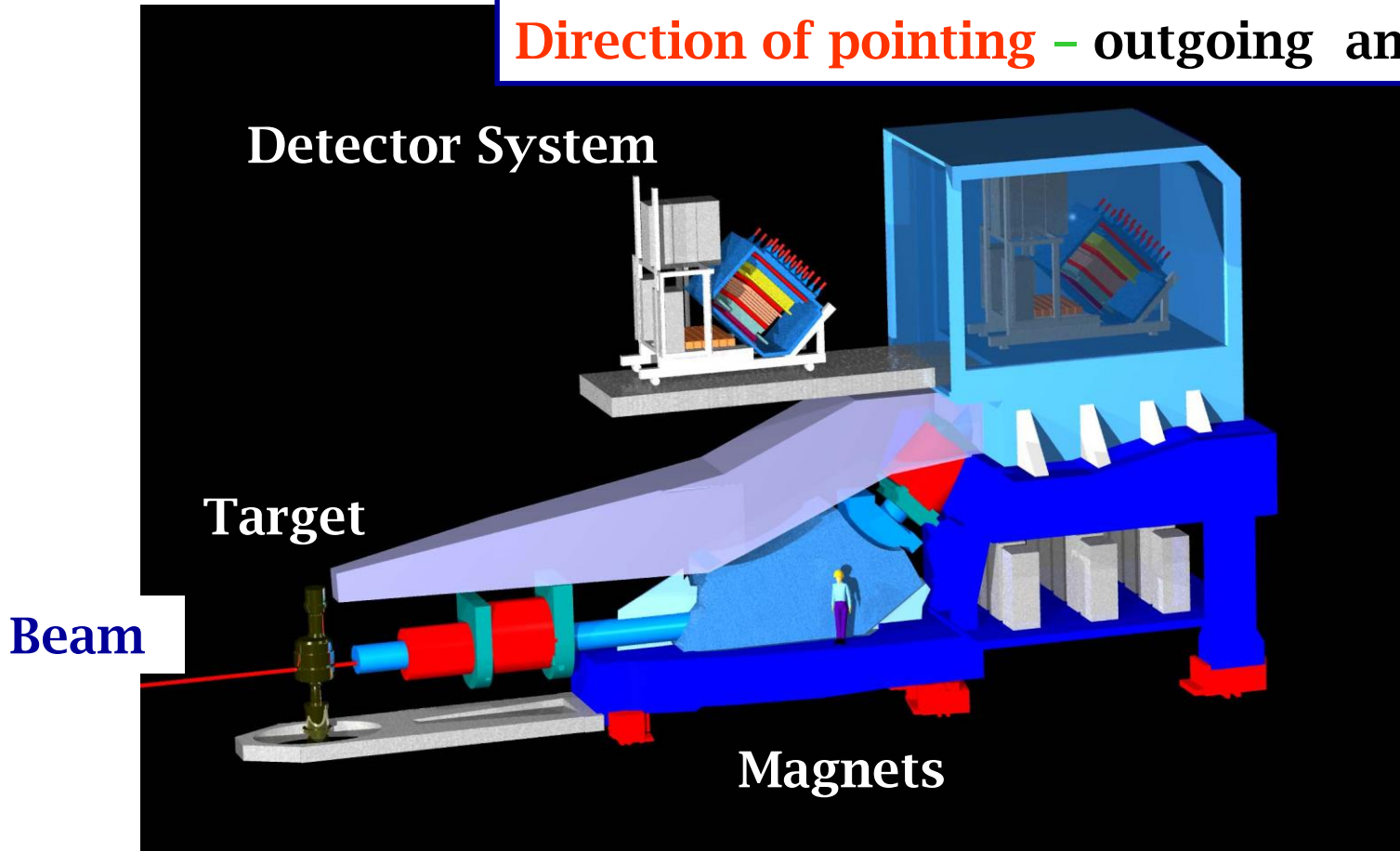
R. Michaels, Jlab
HUGS Lecture

Spectrometers Measure :

Magnets - measure momentum

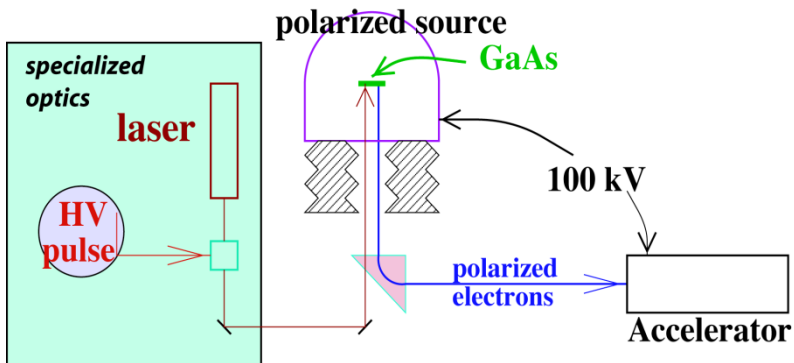
Detectors - identifies particles

Direction of pointing - outgoing angle



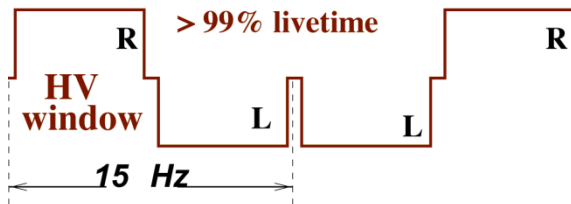
How to do a Parity Experiment

(integrating method)



rapid, random, helicity flipping

Rapid, Random Helicity Flips



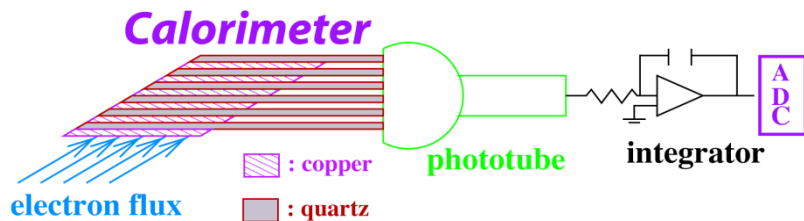
Measure flux F for each window

$$A_{\text{window pair}} = \frac{F_R - F_L}{F_R + F_L}$$

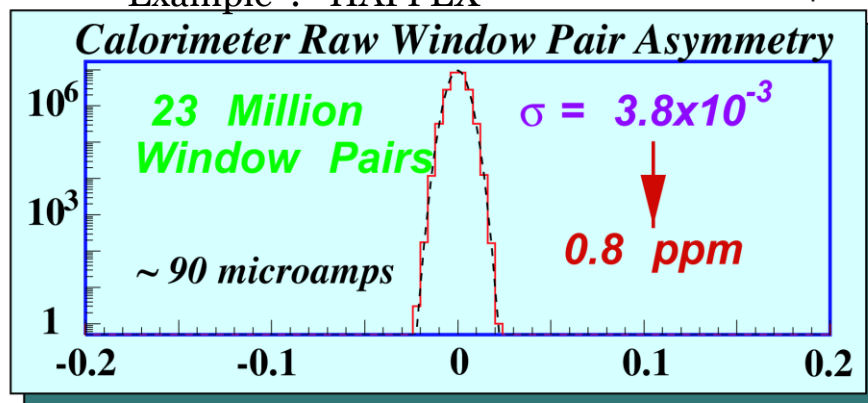
Flux Integration Technique:

HAPPEX: 2 MHz

PREX: 500 MHz



Signal Average N Windows Pairs: $A \pm \frac{\sigma(A)}{\sqrt{N_{\text{windows}}}}$
 Example : HAPPEX



No non-gaussian tails to $\pm 5\sigma$

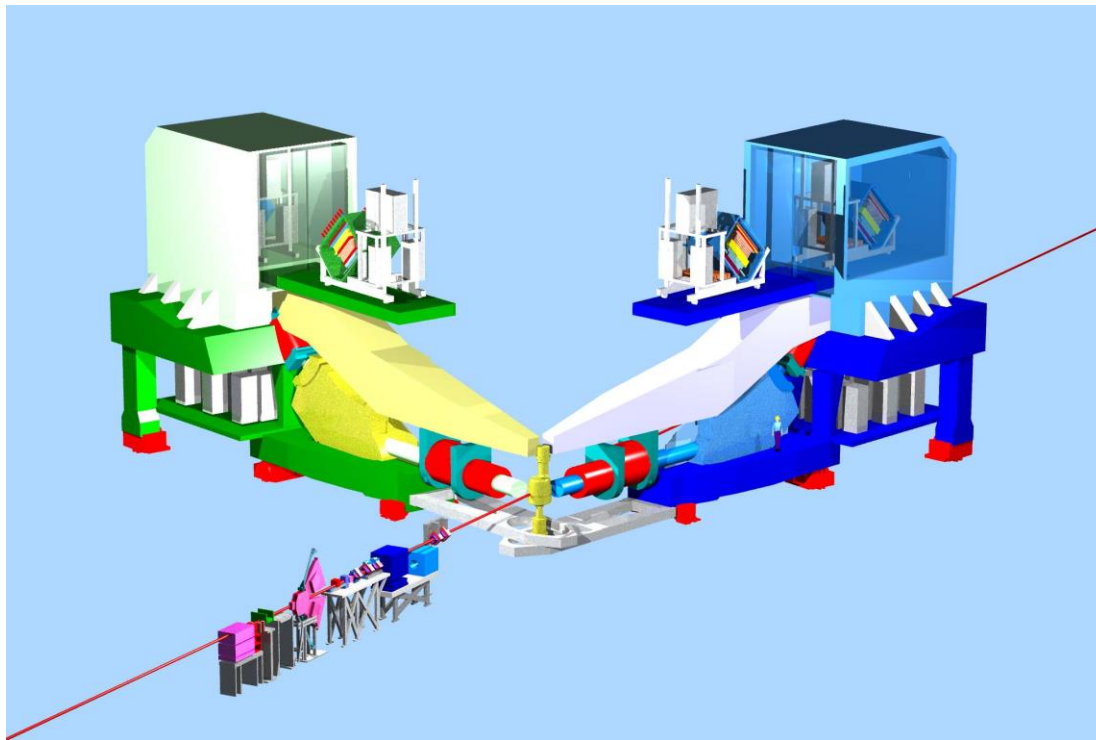
Experiment: Flip spin and count # scattered electrons in each spin state

Detector 1
(count electrons)

Detector 2
(count electrons)



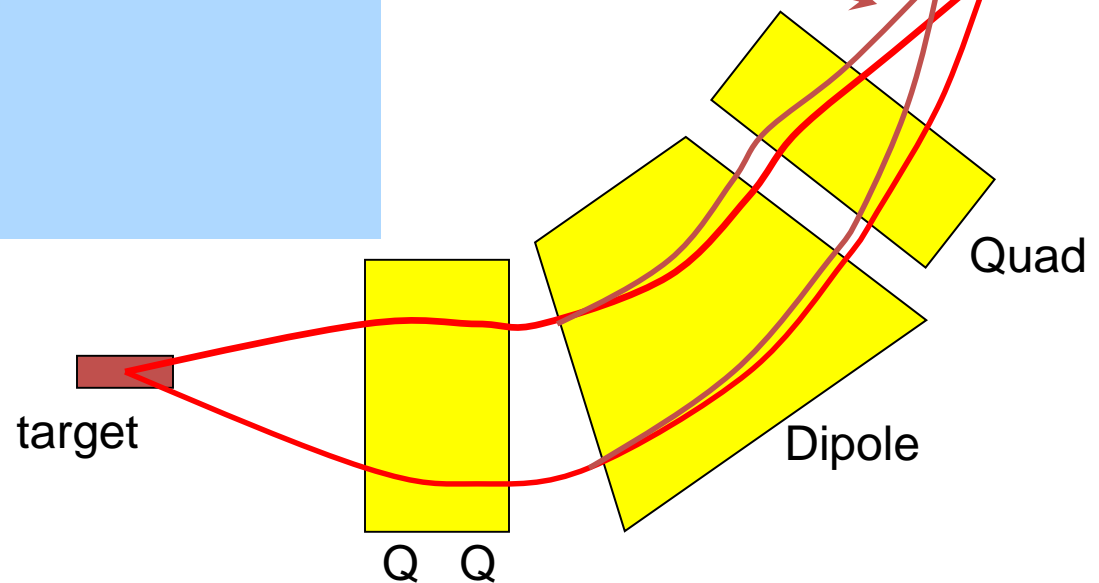
High Resolution Spectrometers



Spectrometer Concept:
Resolve Elastic

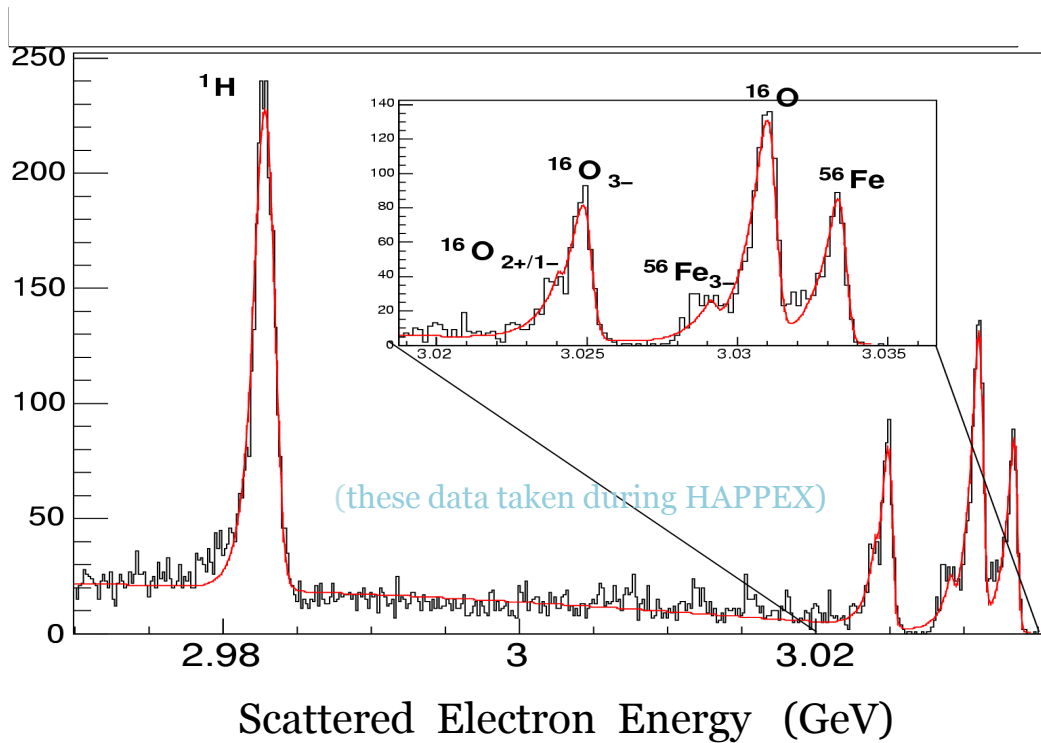
1st excited state Pb 2.6 MeV

Elastic → detector
Inelastic →



Left-Right symmetry to
control transverse
polarization systematic

Measure θ from Nuclear Recoil



δE =Energy loss
 E =Beam energy
 M_A =Nuclear mass
 θ =Scattering angle

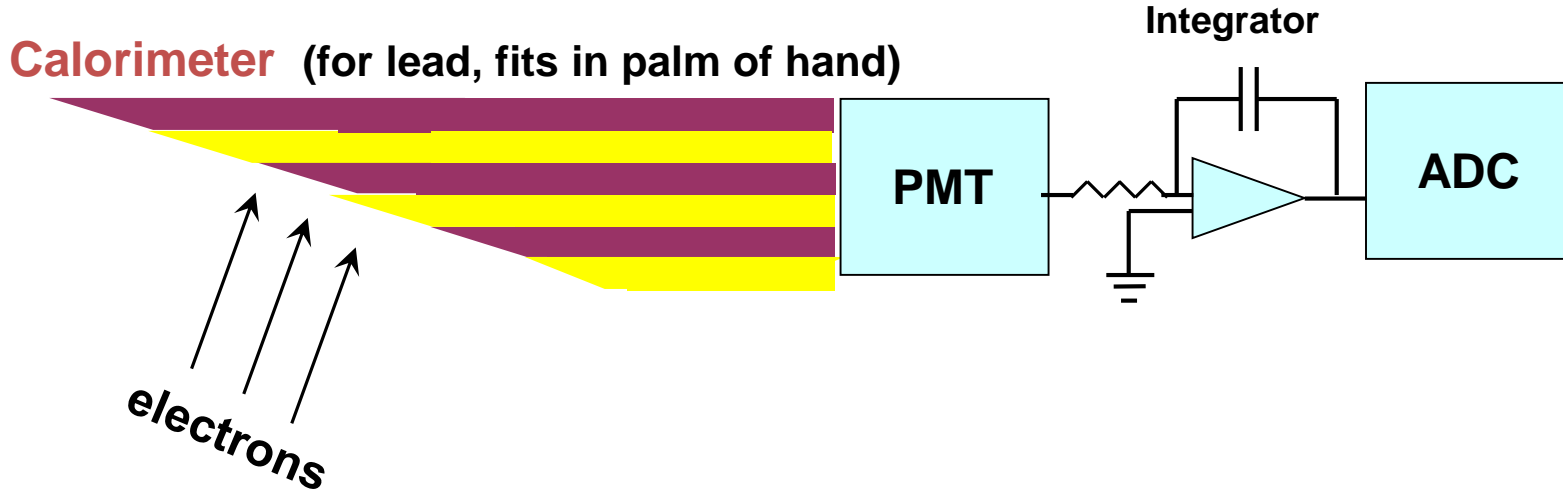
$$\frac{\delta E}{E} \approx \frac{\theta^2}{2} \frac{E}{M_A}$$

Recoil is large for H, small for nuclei

(3X better accuracy than survey)

Integrating Detection

- Integrate in 30 msec helicity period.
- **Deadtime free.**
- 18 bit ADC with $< 10^{-4}$ nonlinearity.
- But must separate backgrounds & inelastics (\rightarrow HRS).



PREX-I

Physics Result

at $Q^2 = 0.00906 \text{ GeV}^2$

$$A_{PV} = \frac{\sigma_R - \sigma_L}{\sigma_R + \sigma_L} =$$

$$0.6571 \pm 0.0604(\text{stat}) \pm 0.0130(\text{syst})$$

ppm

9.2 %

2.0 %

- **Statistics limited (9%)**
- **Followup expt, PREX-II to get to 3%**
- **Systematic error goal achieved ! (2%)**

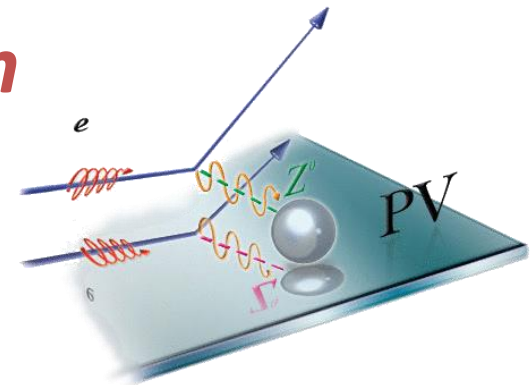
A list of the **Systematic Errors** for **PREX** with explanations to follow for the first two ...

Error Source	Absolute (ppm)	Relative (%)
Beam Asymmetries (1)	0.0072	1.1
Polarization (2)	0.0083	1.3
Detector Linearity	0.0076	1.2
BCM Linearity	0.0010	0.2
Rescattering	0.0001	0
Transverse Polarization	0.0012	0.2
Q ² (2)	0.0028	0.4
Target Thickness	0.0005	0.1
¹² C Asymmetry (1)	0.0025	0.4
Inelastic States	0	0
TOTAL	0.0140	2.1

(1) **Nonzero correction** (the rest assumed zero)

(2) **Normalization Correction** applied

Isolating the *weak interaction* is not as easy as you think !



Positive spin

Mirror Image



Negative spin

Incident electron

spin



momentum



Target



A nucleus in target

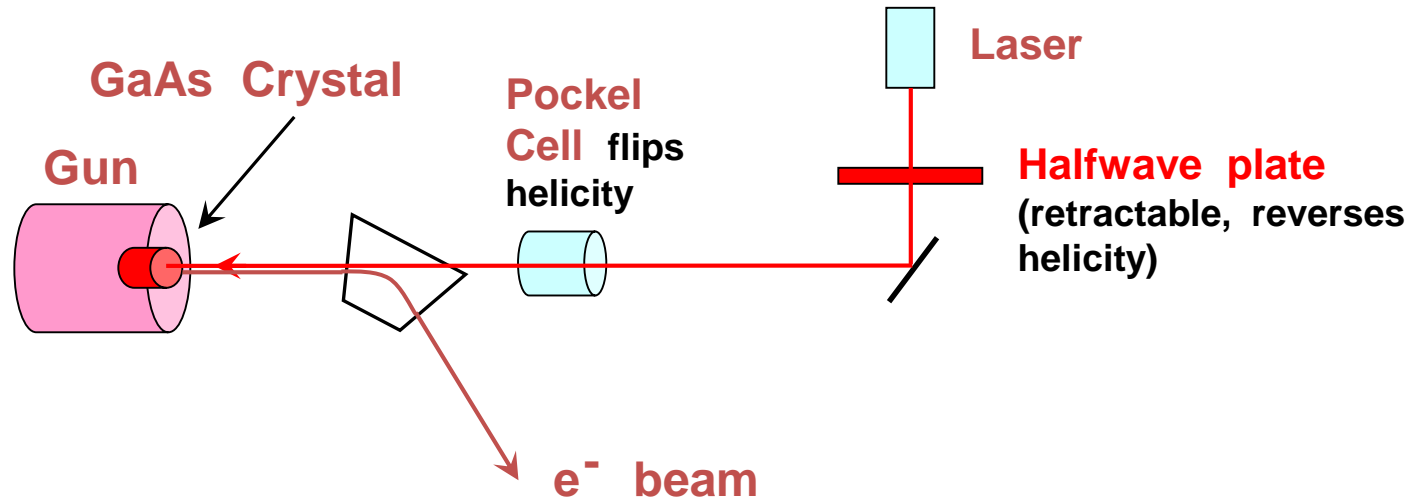


Want to flip spin of electrons and do not flip E, θ or anything else !

Beam Asymmetries

- Want to flip the helicity of the beam and nothing else.
- EM cross section depends on E, θ
- Residual helicity correlations will exist at some level. Need to measure them and correct for them.
- Systematic error = error in the correction.

Polarized Electron Source



- **Based on Photoemission from GaAs Crystal**
- **Polarized electrons from polarized laser**
- **Need :**
 - Rapid, random helicity reversal
 - Electrical isolation from the rest of the lab
 - Feedback on Intensity Asymmetry

Parity Quality Beam : Unique Strength of JLab

Helicity – Correlated Position Differences

$$\langle X_R - X_L \rangle \text{ for helicity } L, R$$

$$A_{\text{raw}} = A_{\text{det}} - A_Q + \alpha \Delta_E + \sum \beta_i \Delta x_i$$

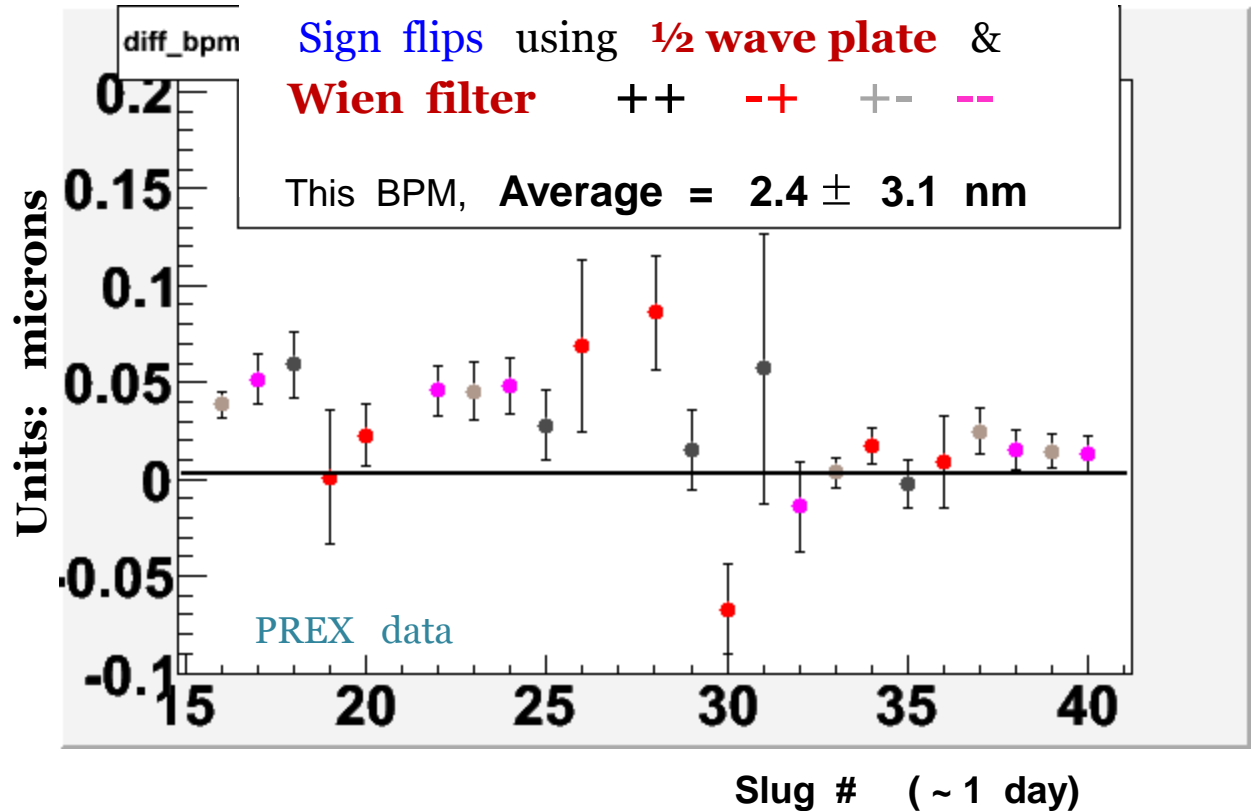
Plotted below

Measured separately

Points: Not sign-corrected.
20-50 nm diffs. with pol.
source setup & feedback

Sign flips provide
further suppression :
Average with signs =
what experiment feels

achieved
< 5 nm



Important Systematic : **P I T A** Effect

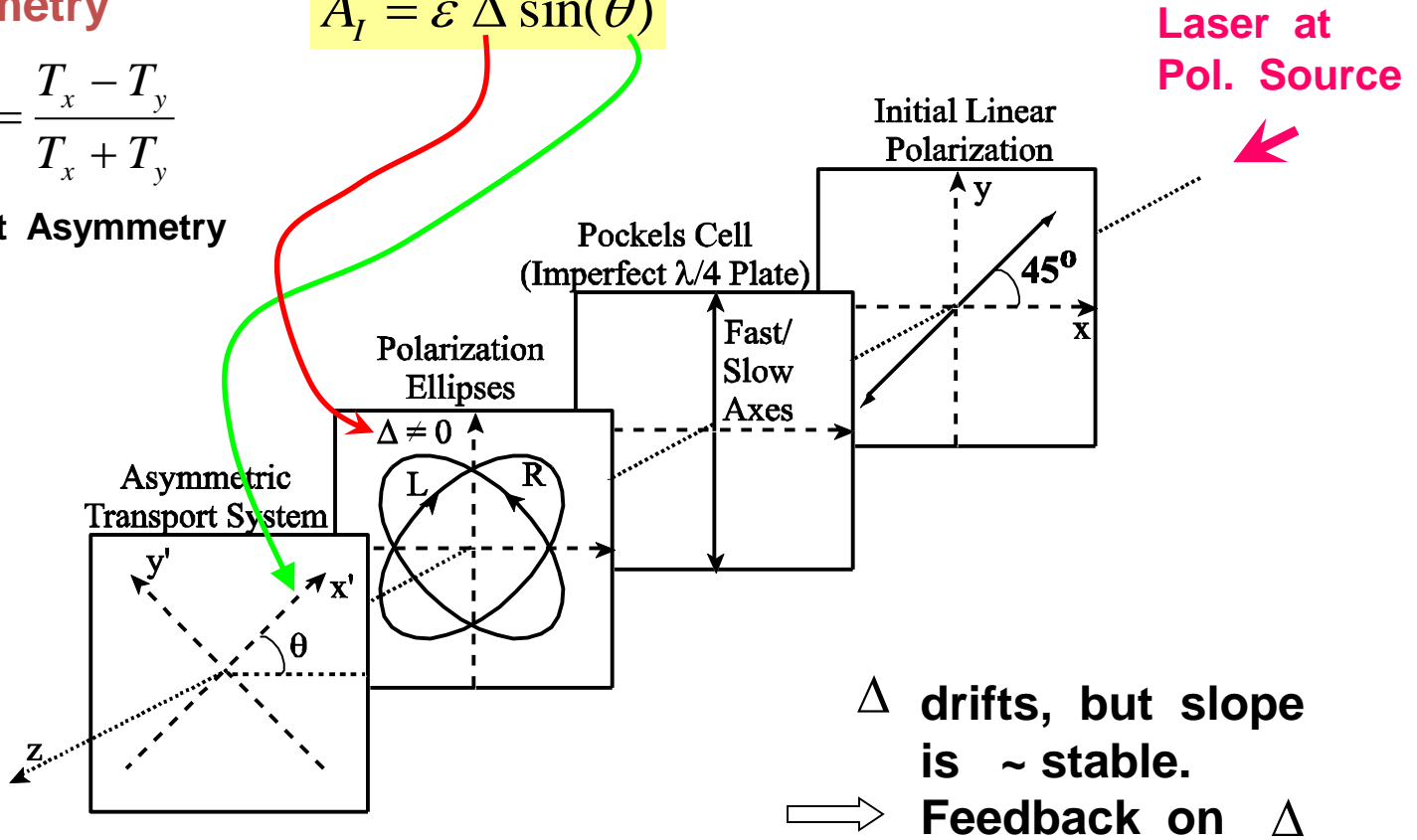
Polarization **I**nduced **T**ransport **A**symmetry

Intensity Asymmetry

$$A_I = \varepsilon \Delta \sin(\theta)$$

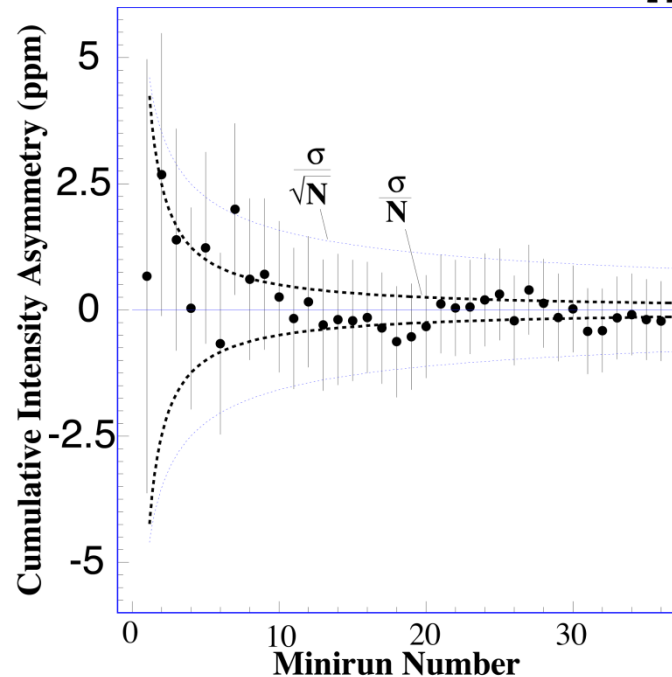
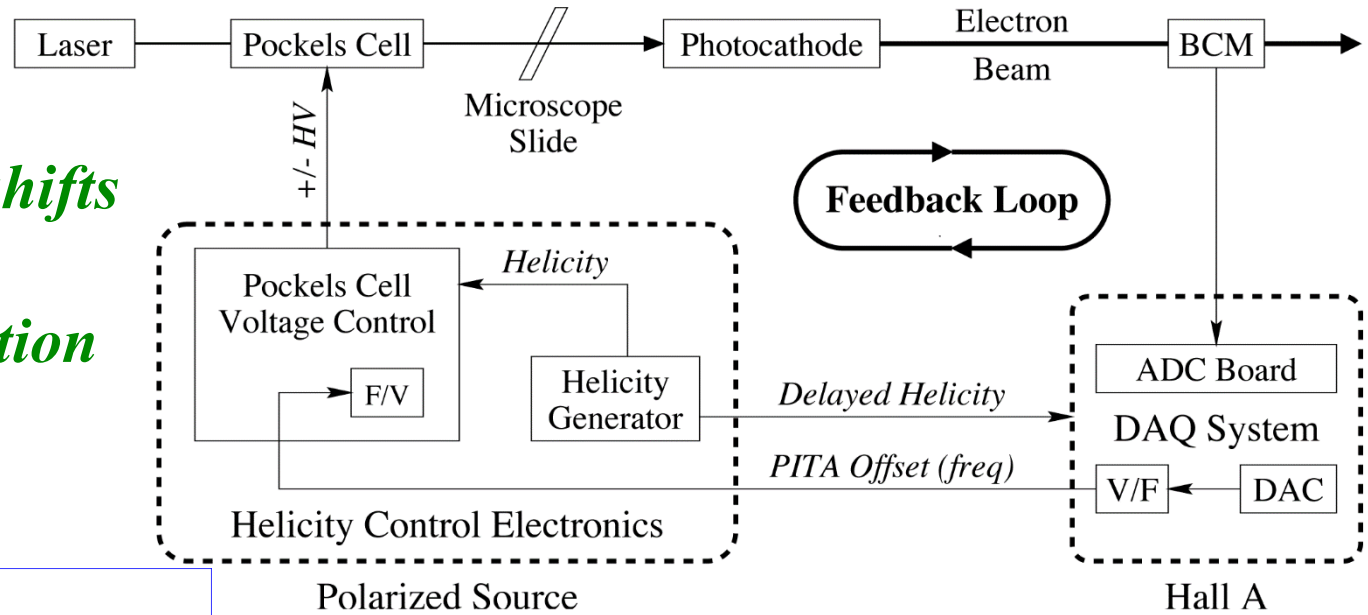
where $\varepsilon = \frac{T_x - T_y}{T_x + T_y}$

Transport Asymmetry



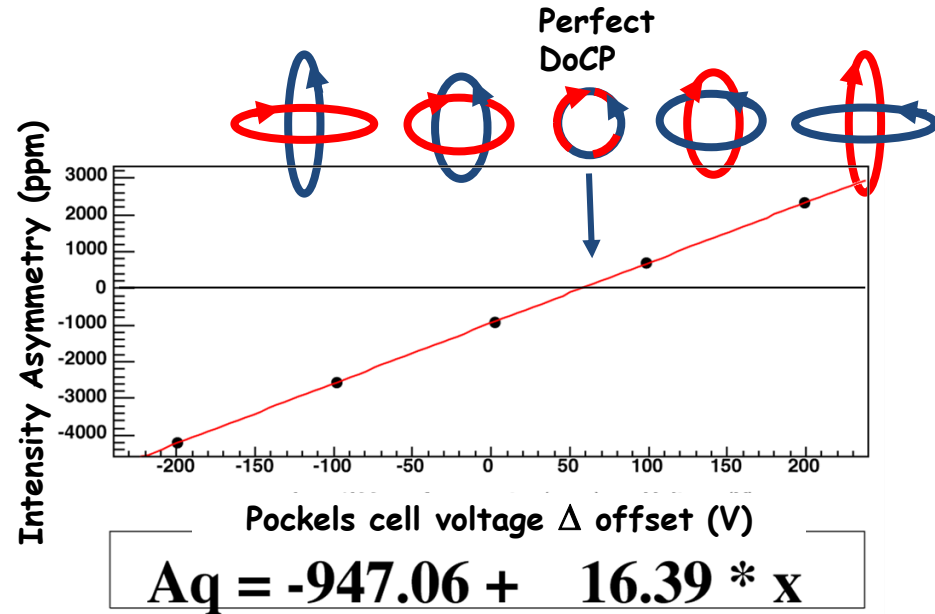
Intensity Feedback

*Adjustments
for small phase shifts
to make close to
circular polarization*



*Low jitter and high accuracy allows sub-ppm
cumulative charge asymmetry in ~ 1 hour*

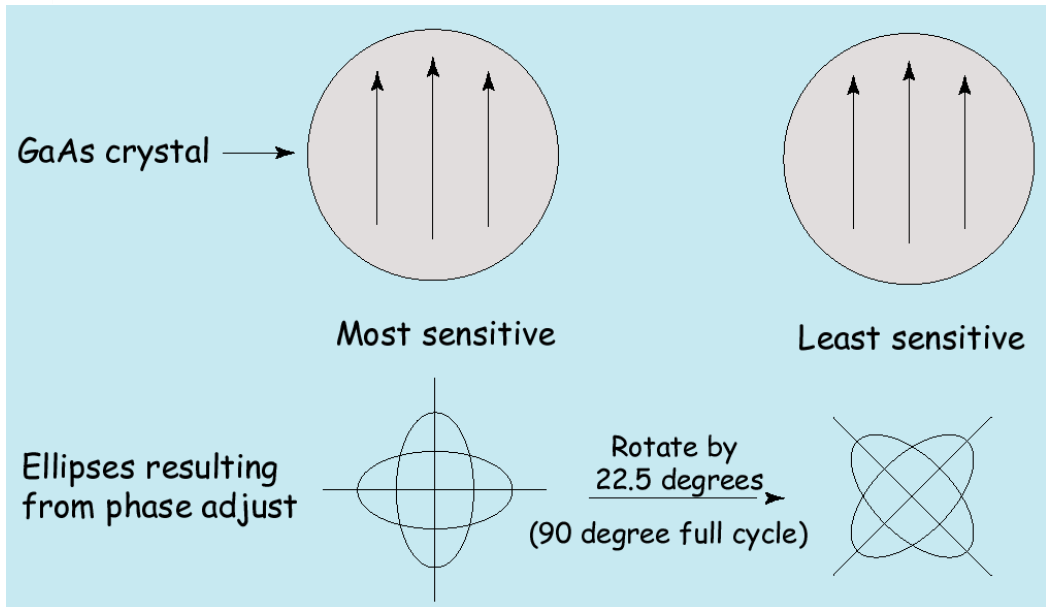
Methods to Reduce Systematics



Scanning the Pockels Cell voltage = scanning the residual linear polarization (DoLP)

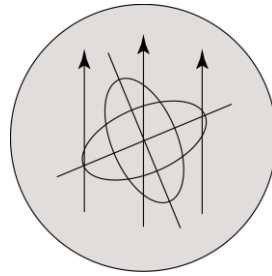
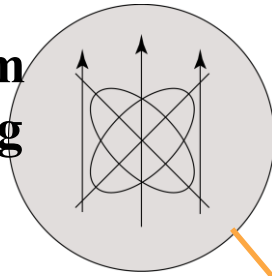
A simplified picture: asymmetry=0 corresponds to minimized DoLP at analyzer

A rotatable $\lambda/2$ waveplate downstream of the P.C. allows arbitrary orientation of the ellipse from DoLP

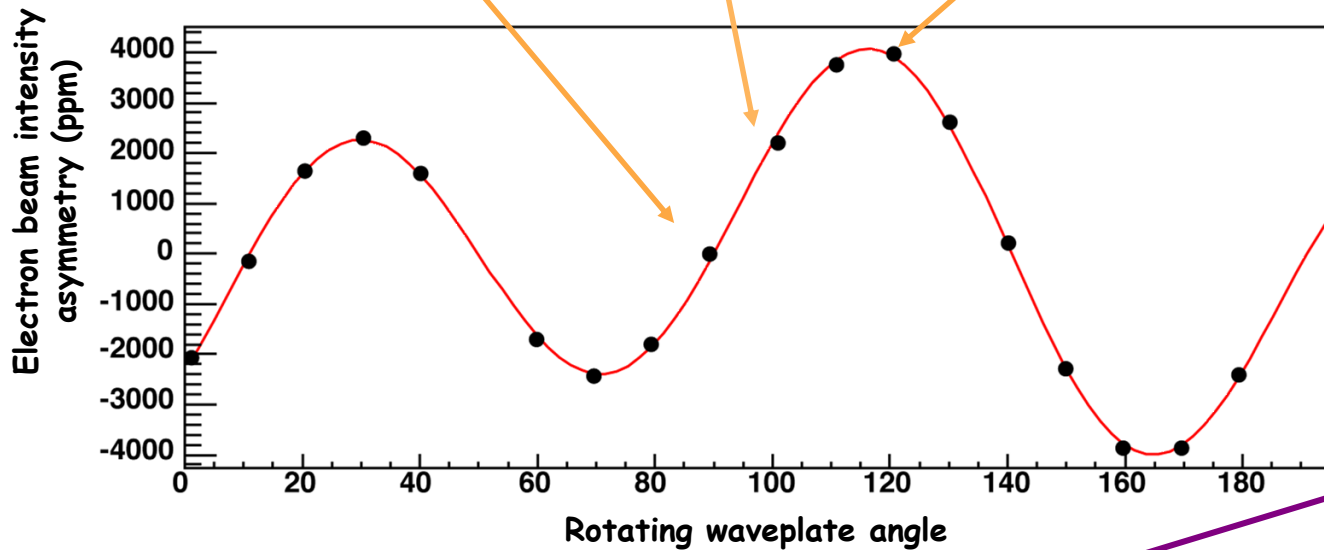
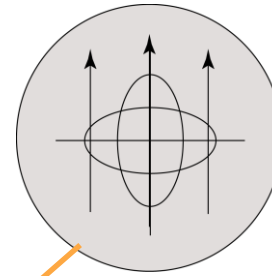


Rotatable Half-Wave Plate

minimum
analyzing
power



maximum
analyzing
power



Add $\lambda/2$ plate
to minimize
analyzing
power

40 term measures
analyzing
power*DoLP (from
Pockels cell)

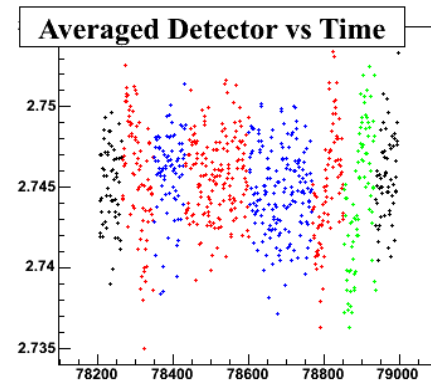
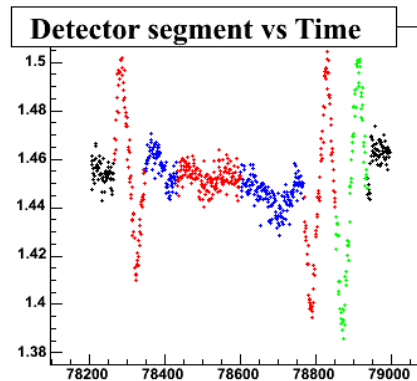
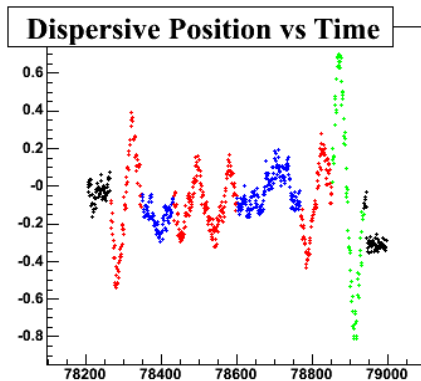
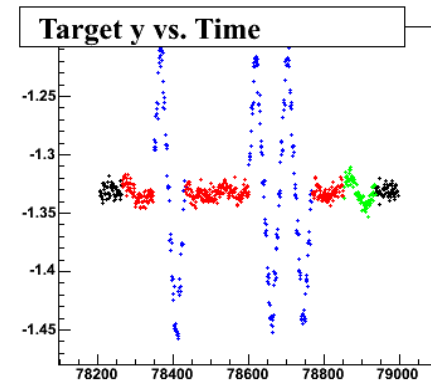
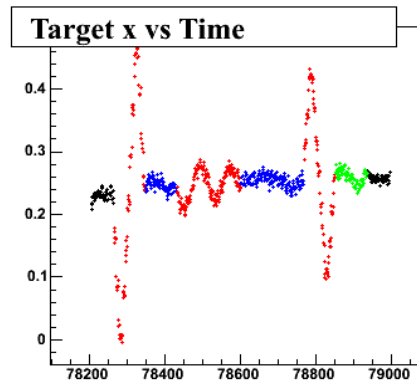
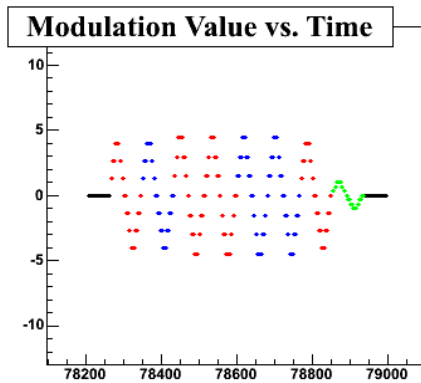
$$Aq = -7.98 + -1211.75 \sin(2x + 75.52) + -3151.04 \sin(4x + 158.47)$$

Beam Asymmetries

$$A_{\text{raw}} = A_{\text{det}} - A_Q + \alpha \Delta_E + \sum \beta_i \Delta x_i$$

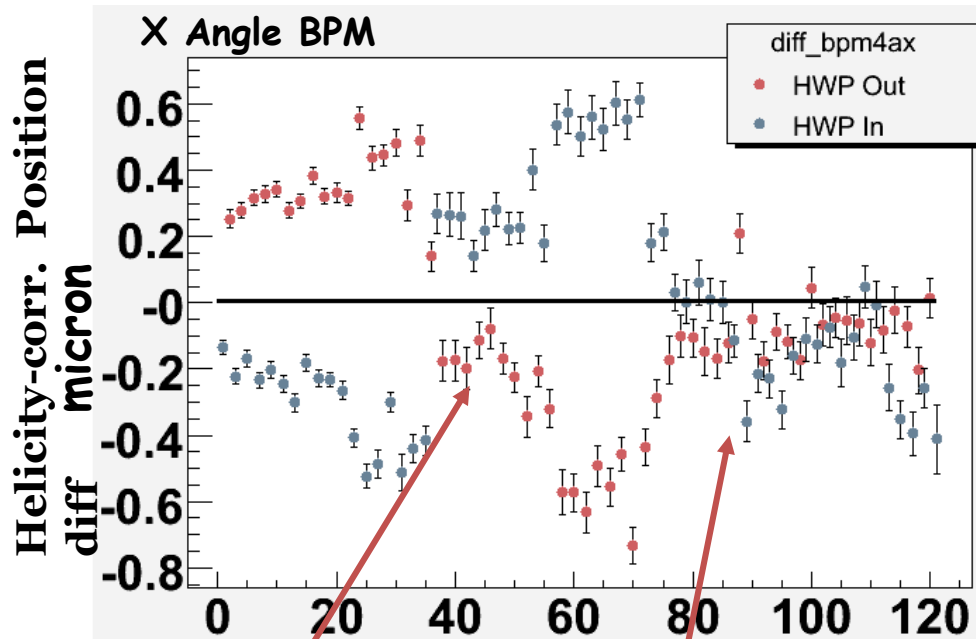
Slopes from

- natural beam jitter (regression)
- beam modulation (dithering)



The Corrections Work !

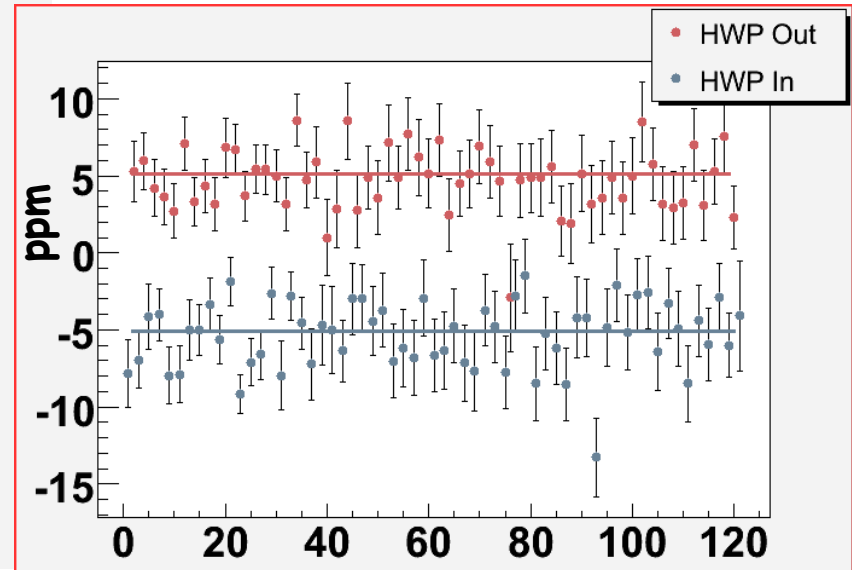
Shown : period of data during HAPPEX (^4He) when beam had a helicity-correlated position due to a mistake* in electronics.



Helicity signal to driver reversed

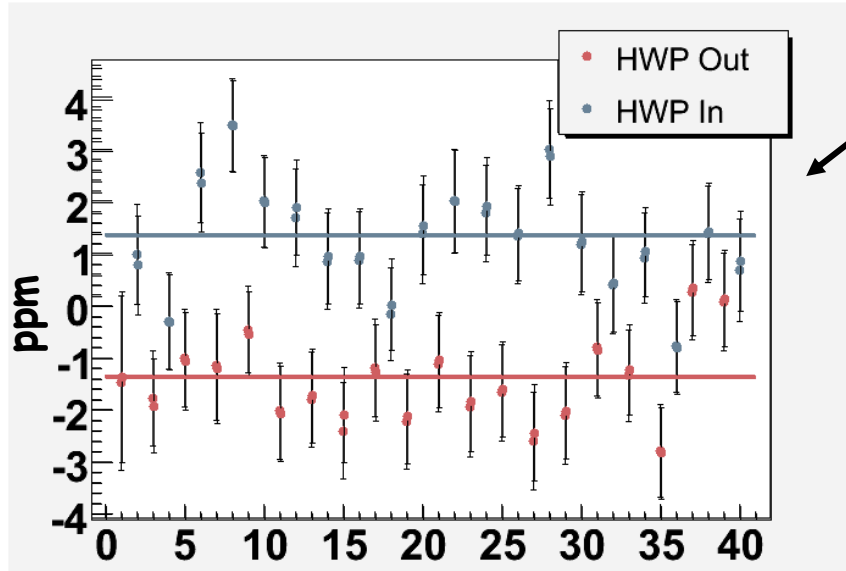
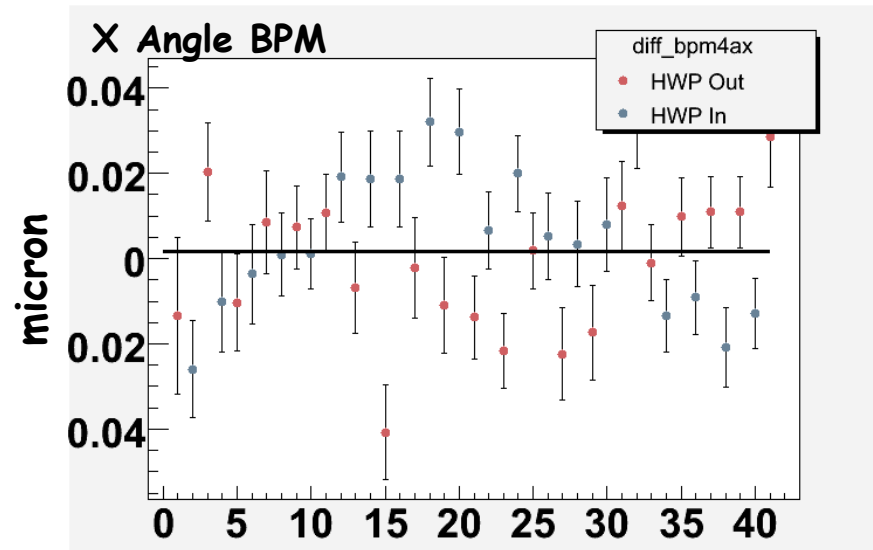
Helicity signal to driver removed

With corrections



* The mistake: Helicity signal deflecting the beam through electronics "pickup"

Final Beam Position Corrections (HAPPEX-H)



Beam Asymmetry Results

Energy: -0.25 ppb

X Target: 1 nm

X Angle: 2 nm

Y Target : 1 nm

Y Angle: <1 nm

Corrected and Raw, Left spectrometer arm alone, Superimposed!

Total correction for beam position asymmetry on Left, Right, or ALL detector: **10 ppb**

Spectacular results from HAPPEX-H showed we could do PREX.

Want multiple, redundant ways to flip the helicity.

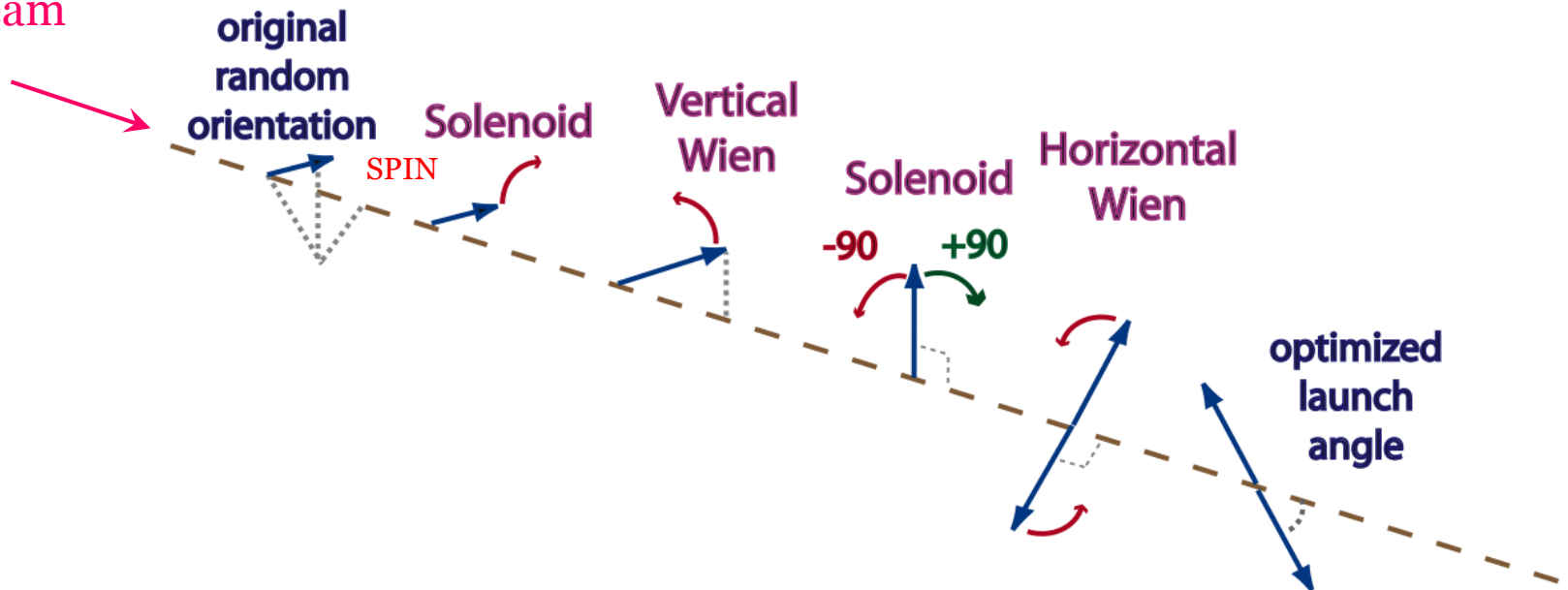
Here, we show the “**Double Wien Filter**”

Crossed E & B fields to rotate the spin

- Two Wien Spin Manipulators in series
- Solenoid rotates spin +/-90 degrees (spin rotation as B but focus as B^2).

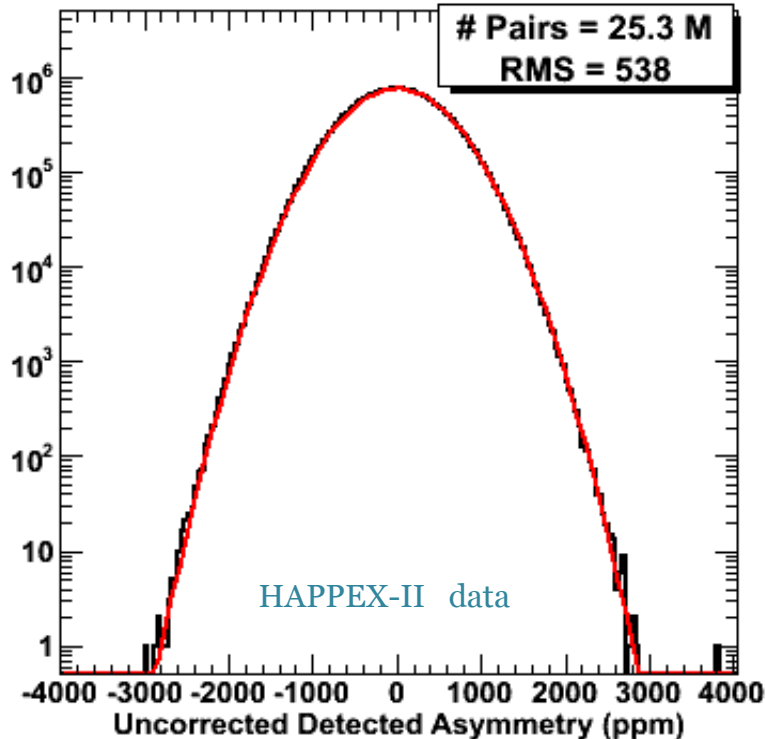
Flips spin without moving the beam !

Electron
Beam

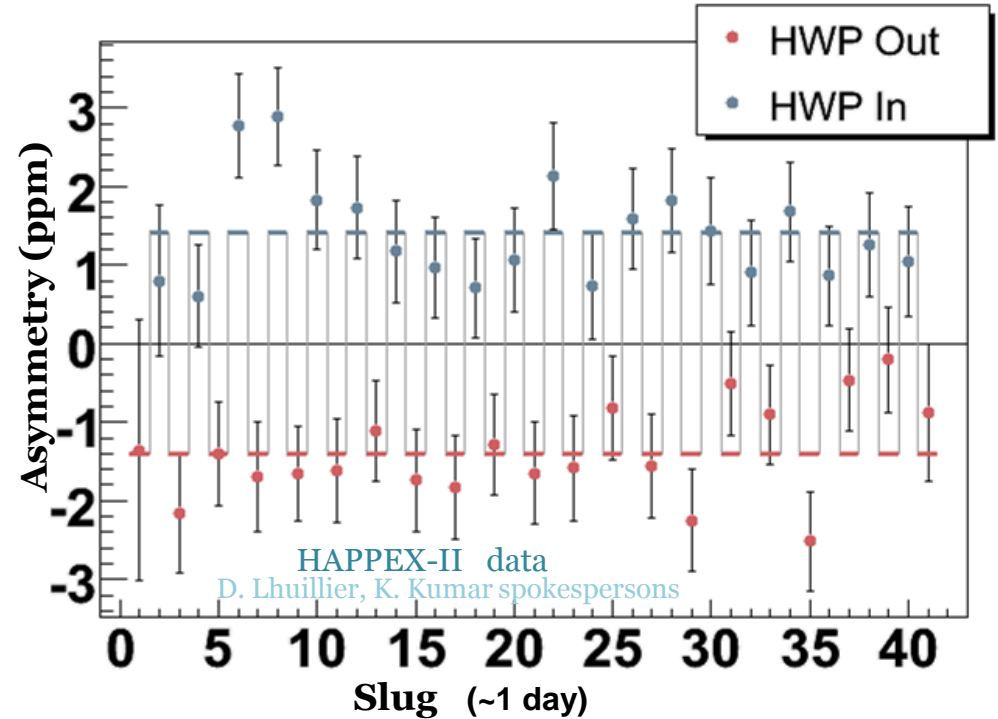


HAPPEX achieved 3 ppb Helicity Correlated Beam Systematics

- Corrections tiny (here, 3 ppb)
- Errors are statistical only



Parity Violating Asymmetry



$$A_{\text{raw}} = -1.58 \text{ ppm} \pm 0.12 \text{ (stat)} \pm 0.04 \text{ (syst)}$$

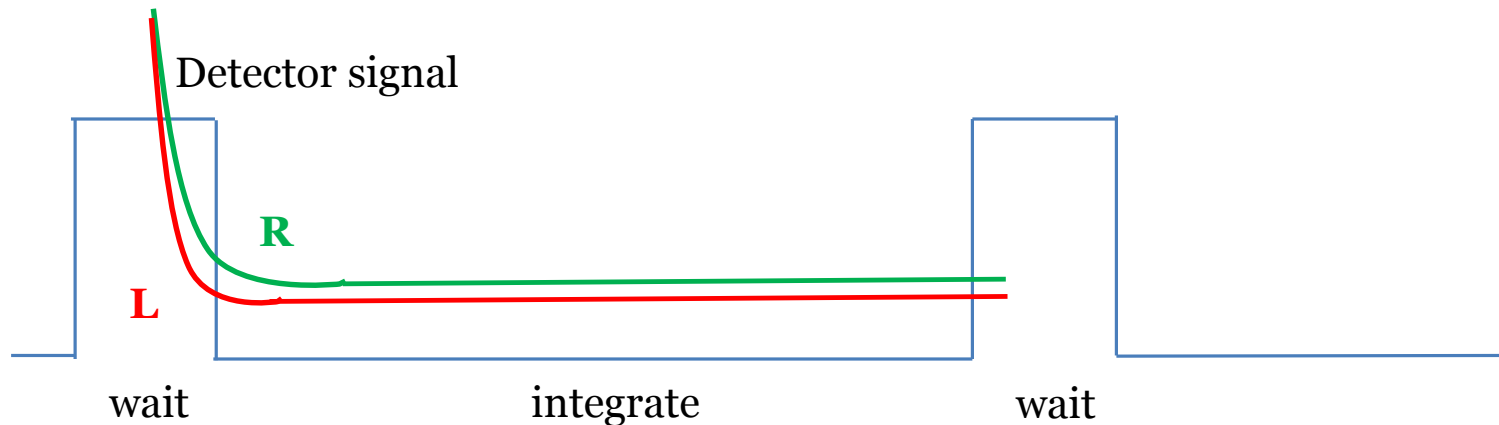
(HWP = optical element used to flip beam helicity, helps suppress some systematics)

$$A = \frac{\left(\frac{D}{Q}\right)_R - \left(\frac{D}{Q}\right)_L}{\left(\frac{D}{Q}\right)_R + \left(\frac{D}{Q}\right)_L}$$

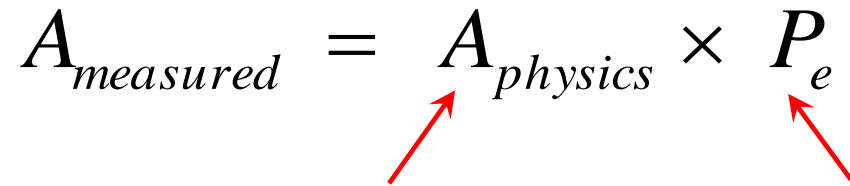
There may be helicity-correlated transients in the beam.

- D & Q may have different time constants in response to transients.
- D may be non-linear in Q.

D = Detector signal Q = Charge in same time interval



Beam Polarization

$$A_{measured} = A_{physics} \times P_e$$


Want to
extract this

Polarization of electrons,
typically 0.9

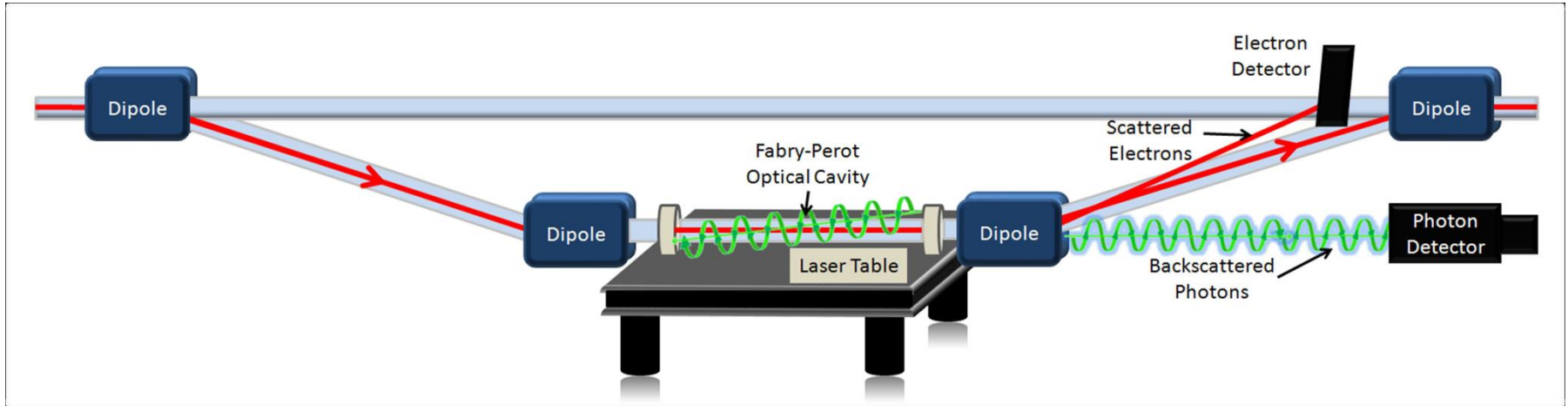
Measure P_e through

- Compton Scattering $\vec{e} - \vec{\gamma}$
- Moller Scattering $\vec{e} - \vec{e}$
- Mott Scattering (not discussed)

Compton Polarimetry

$$\vec{e} - \vec{\gamma}$$

beam laser



1. **Laser system: 1 W drive laser coupled to high gain Fabry-Perot cavity → several kW intracavity power**
2. **Photon detector: GSO crystal for low energies, or lead-tungstate for high energies.**
3. **Electron detector: silicon strip detector, 240 μm pitch, 192 strips/plane**

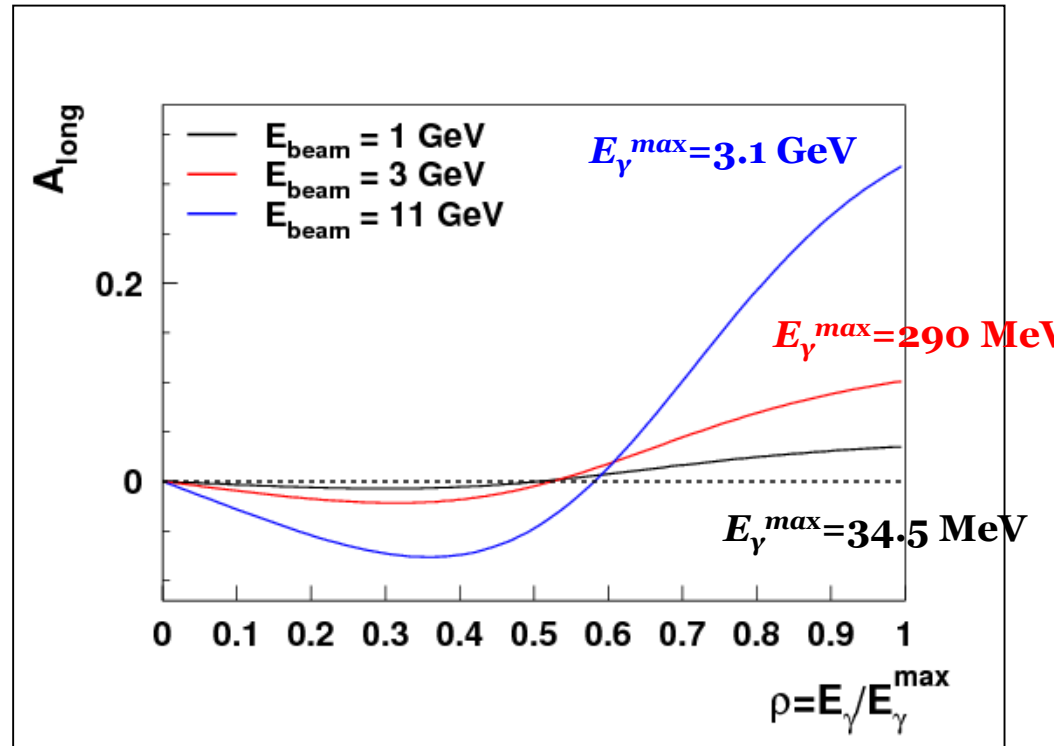
Compton Polarimetry

$$\frac{dP_e}{P_e} < 1\%$$

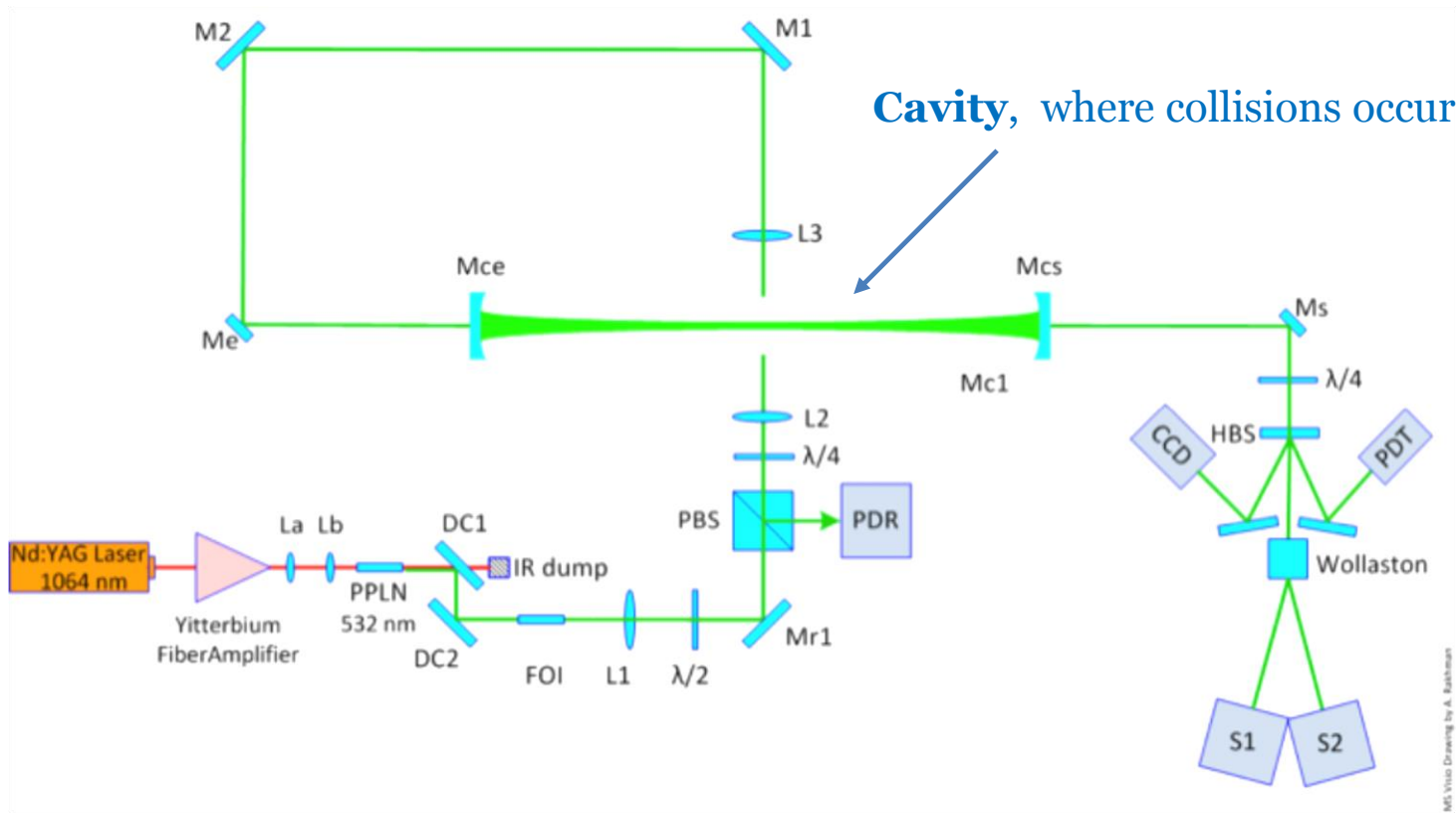
1%-level Compton polarimetry has been achieved with Hall A Compton at ~ 1 GeV (PREX) and ~ 3 GeV (HAPPEX-II)

$$A_{measured} = P_e P_\gamma A_{long}$$

Known exactly from QED



Compton Laser System



Measure P_γ

Recent Compton Data with 11 GeV Beam

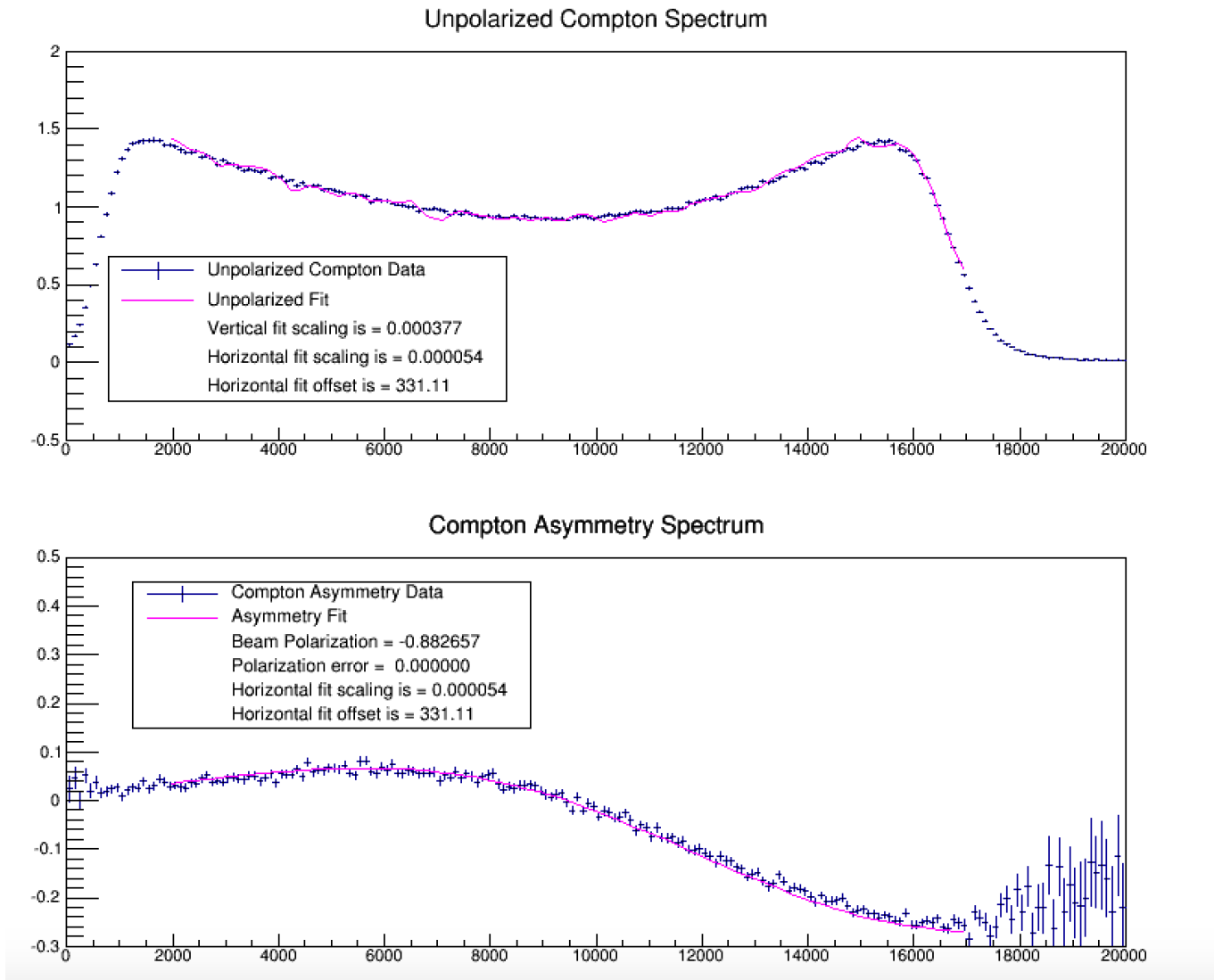
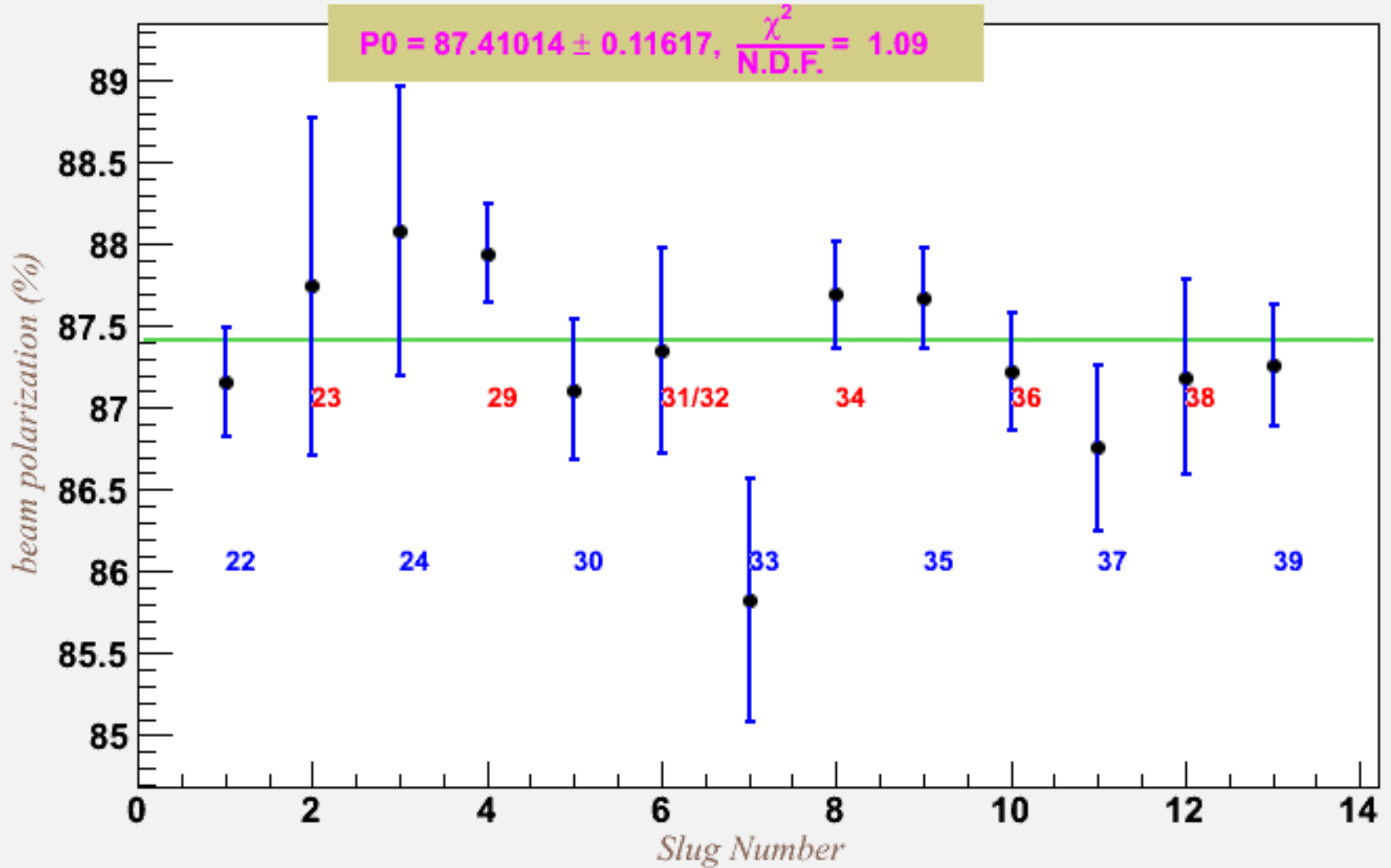


Figure courtesy Alexa Johnson

Compton Polarimeter in PREX-I

the grand average of laser cycle wise beam polarization V.S. slug number



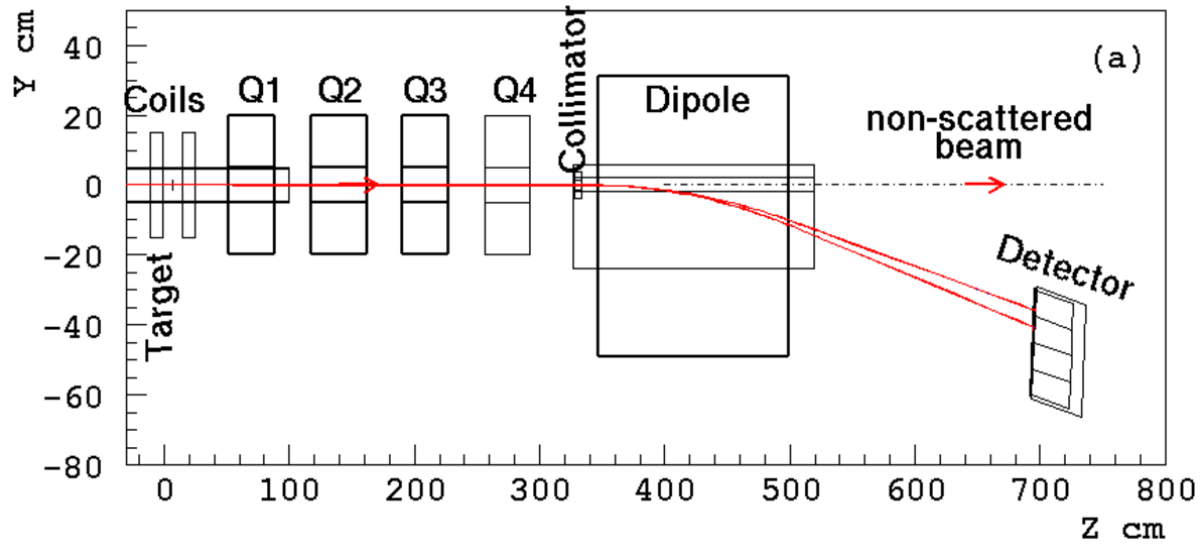
Moller Polarimeter

$$\vec{e} - \vec{e}$$

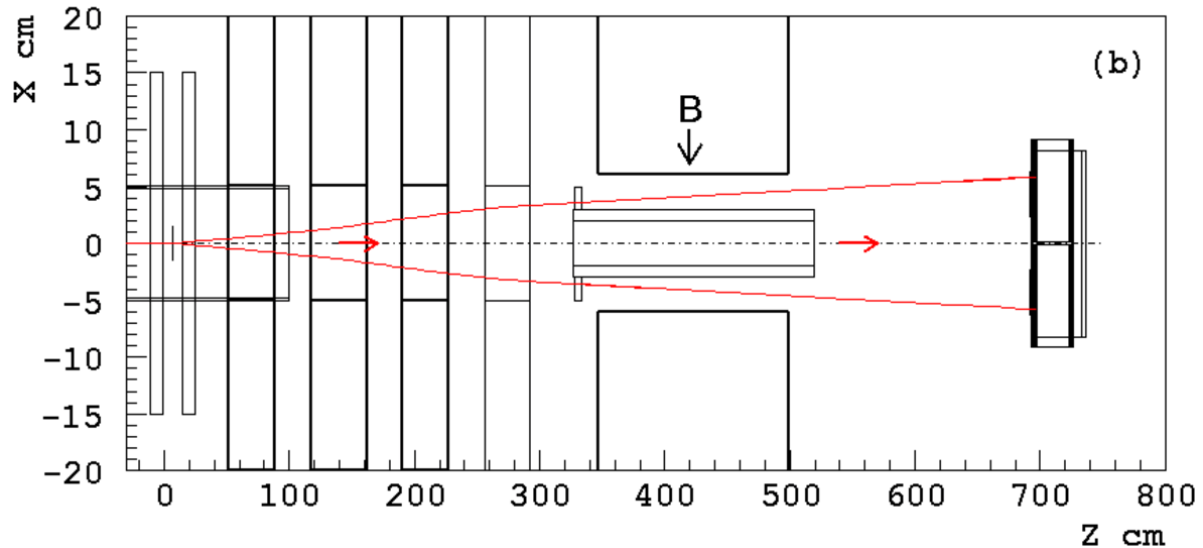
beam

Electrons in
an iron foil
in a strong
B field

Side View



Top View



5 T MAGNET



Power supply



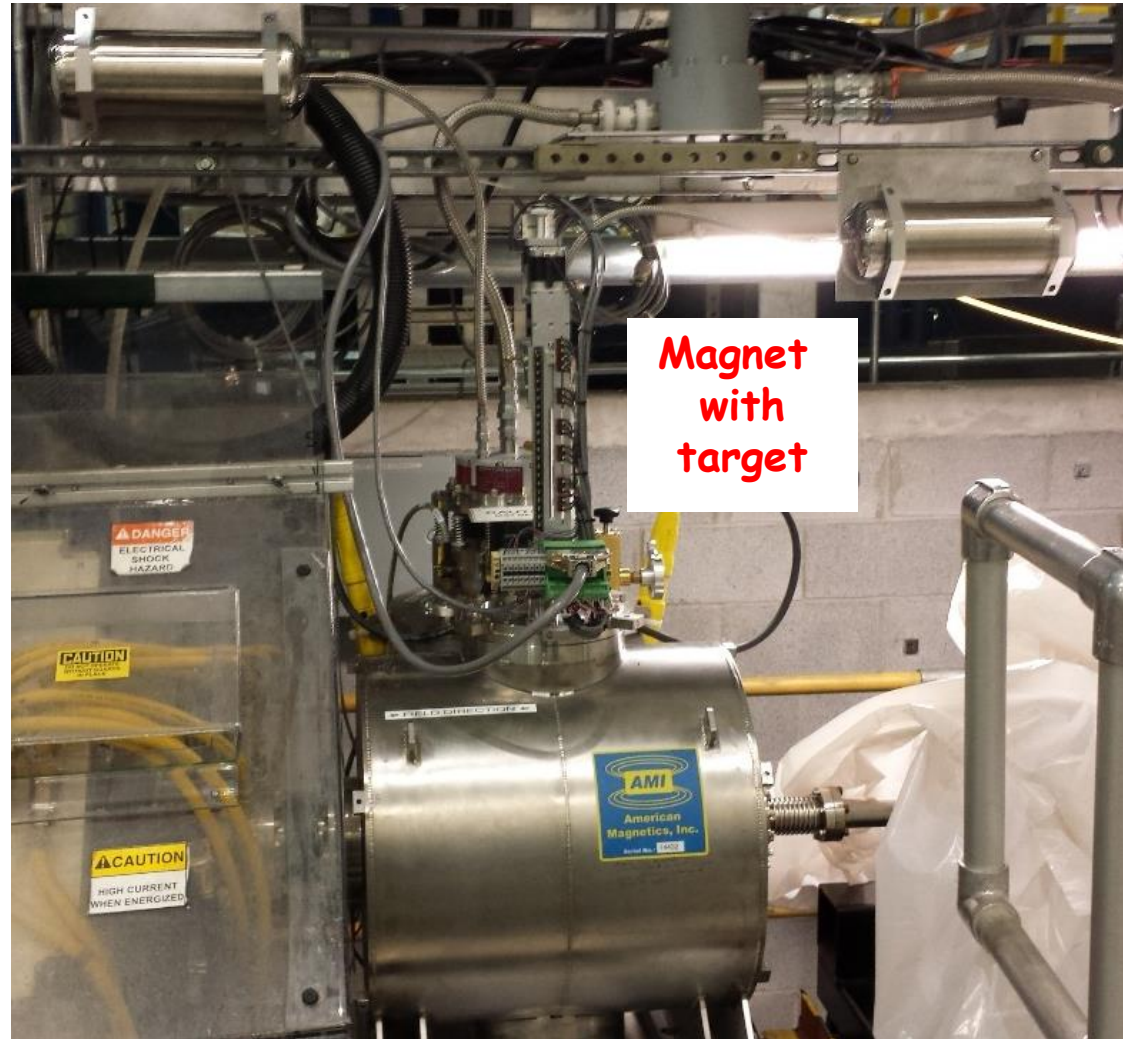
Compressor



Water chiller

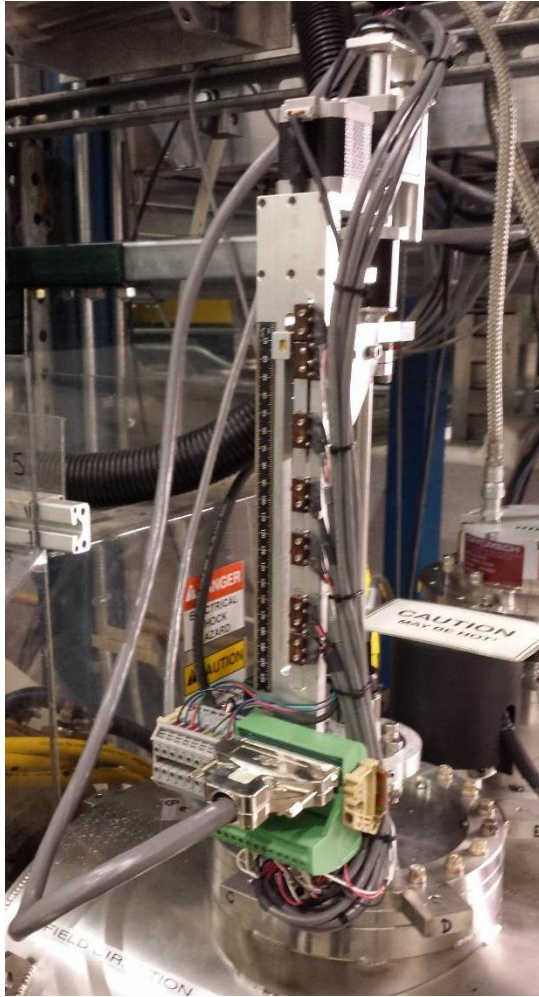
*Superconductive magnet:
Maximal field $\pm 5T$*

The iron foil targets sit in this field and become spin-polarized.



Magnet with target

NEW POLARIZED ELECTRON TARGET



Hall A High-Field Moller Target Motion Control

Linear

Motor Readback (mm) **0.000** Encoder (v) **3.031**

Position (mm)	0	43.444	80.265	117.483	155.298
Encoder (v)	3.025	2.364	1.813	1.251	0.680

Retracted **Park** T1 T2 T3 T4 Extended

P T1 T2 T3 T4 Move

0.000 < 0.000 > STOP

Position Input (mm) Jog (mm)

Rotary

Motor Readback (degree) **-0.112** Encoder (v) **4.568**

Position (degree)	-8.426	-0.200	12.865
Encoder (v)	4.150	4.558	5.206

-Limit **C** +Limit

-L C +L Move

-0.112 < 0.000 > STOP

Position Input (degree) Jog (degree)



Designed and made by Temple University
Rotation in horizontal plane $\pm 10^\circ$
Vertical translation

Targets in holder: pure iron $1\mu\text{m}$, $4\mu\text{m}$, $12\mu\text{m}$, $25\mu\text{m}$

R. Michaels, Jlab
 HUGS Lecture

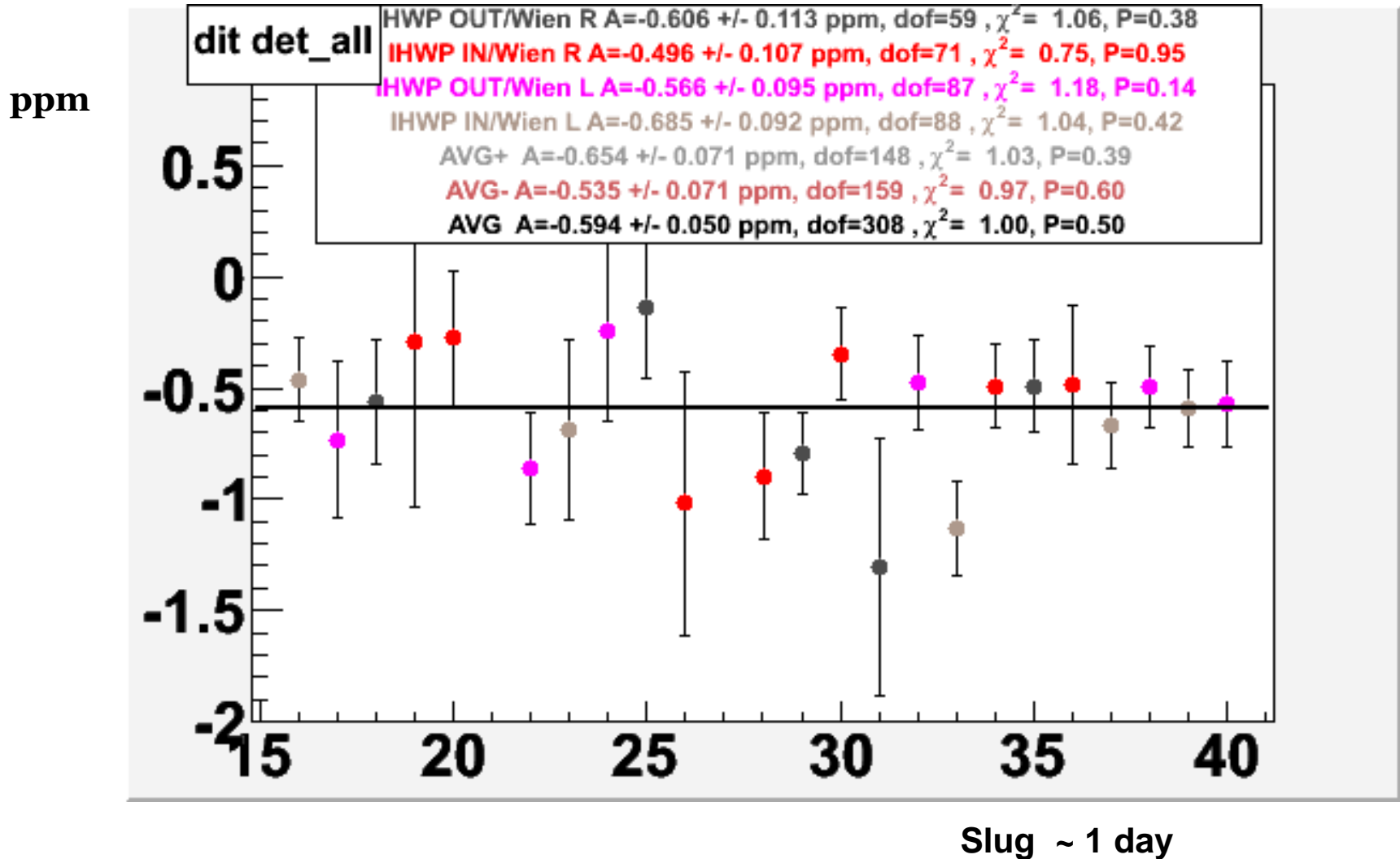


MØLLER SYSTEMATIC ERROR BUDGET

Variable	2010 (3.7T)	PREX (2017)
<i>Target polarization</i>	0.35%	0.25%
<i>Analyzing power</i>	0.3%	0.2%
<i>Levchuk effect</i>	0.3%	0.2%
<i>Target temperature</i>	0.02%	0.02%
<i>Dead time</i>	0.3%	0.05%
<i>Background</i>	0.3%	0.2%
<i>High beam current*</i>	0.2%*	0.2%
<i>Others</i>	0.5%	0.3%
<i>Total</i>	0.87%	0.53%

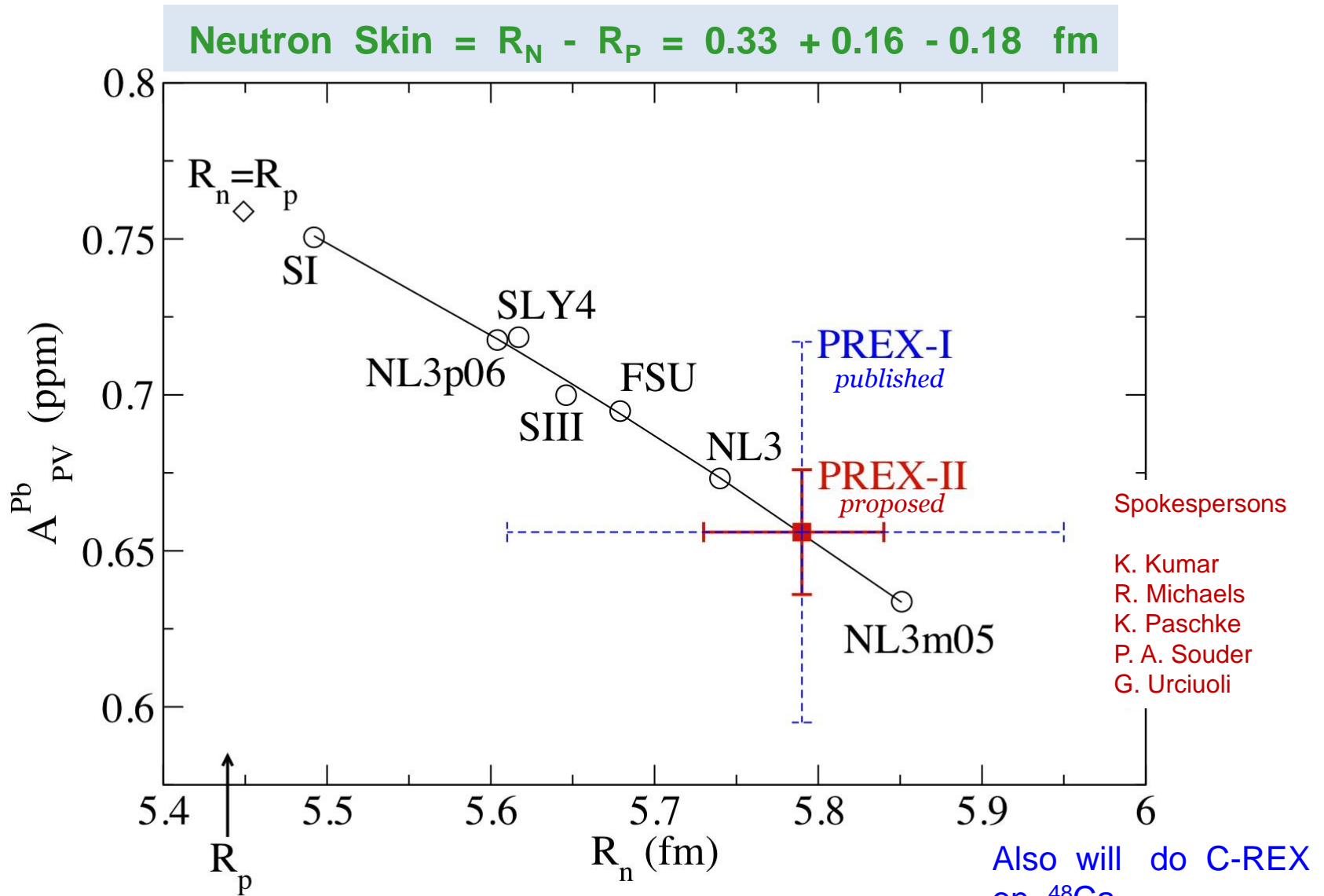
Thanks, Sasha Glamazdin

Returning to ... **PREX Asymmetry** ($P_e \times A$)



Asymmetry leads to R_N

Establishing a neutron skin at ~95% CL

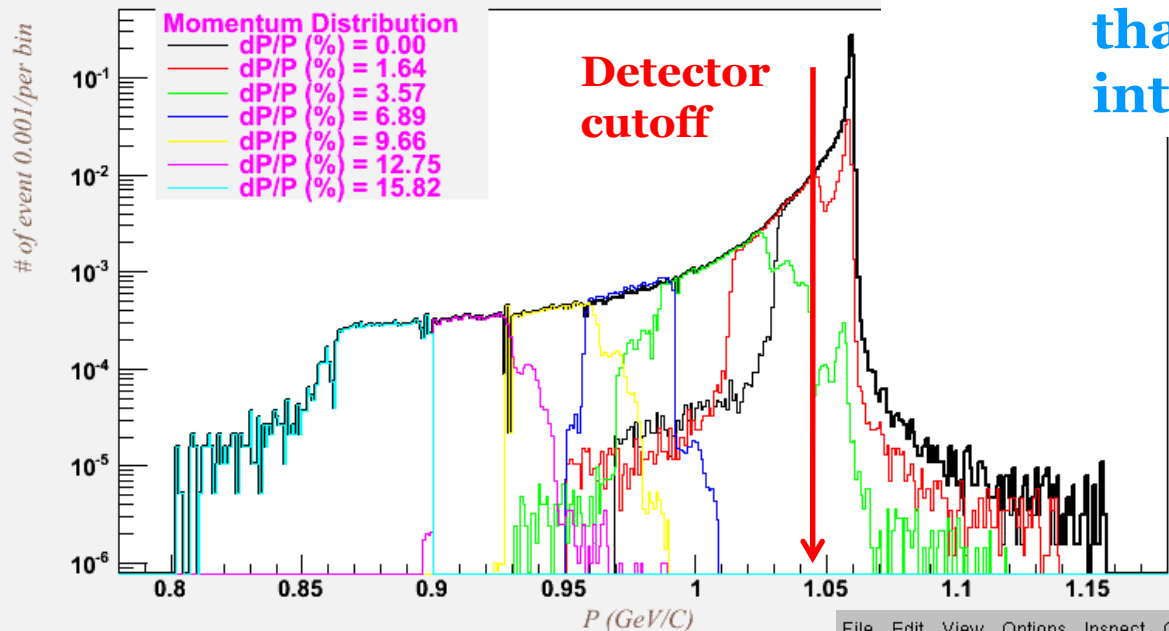


PREX/CREX : Summary

- Fundamental Nuclear Physics with many applications
- PREX-I achieved a 9% stat. error in Asymmetry (original goal : 3 %)
- Systematic error goals already achieved.
- PREX-II and C-REX to run back-to-back, possibly in 2018 (being decided now)

Extra Slides

momentum distribution



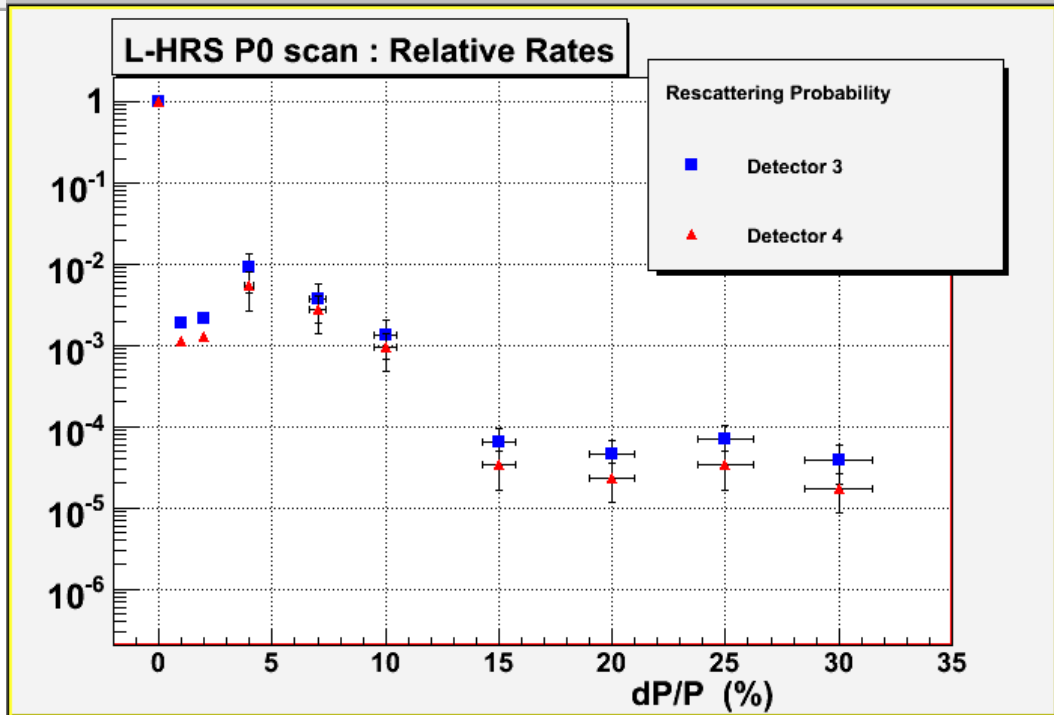
Backgrounds that might re-scatter into the detector ?

Run magnets down: measure inelastic region

Run magnets up: measure probability to rescatter

No inelastics observed on top of radiative tail. Small systematic for tail.

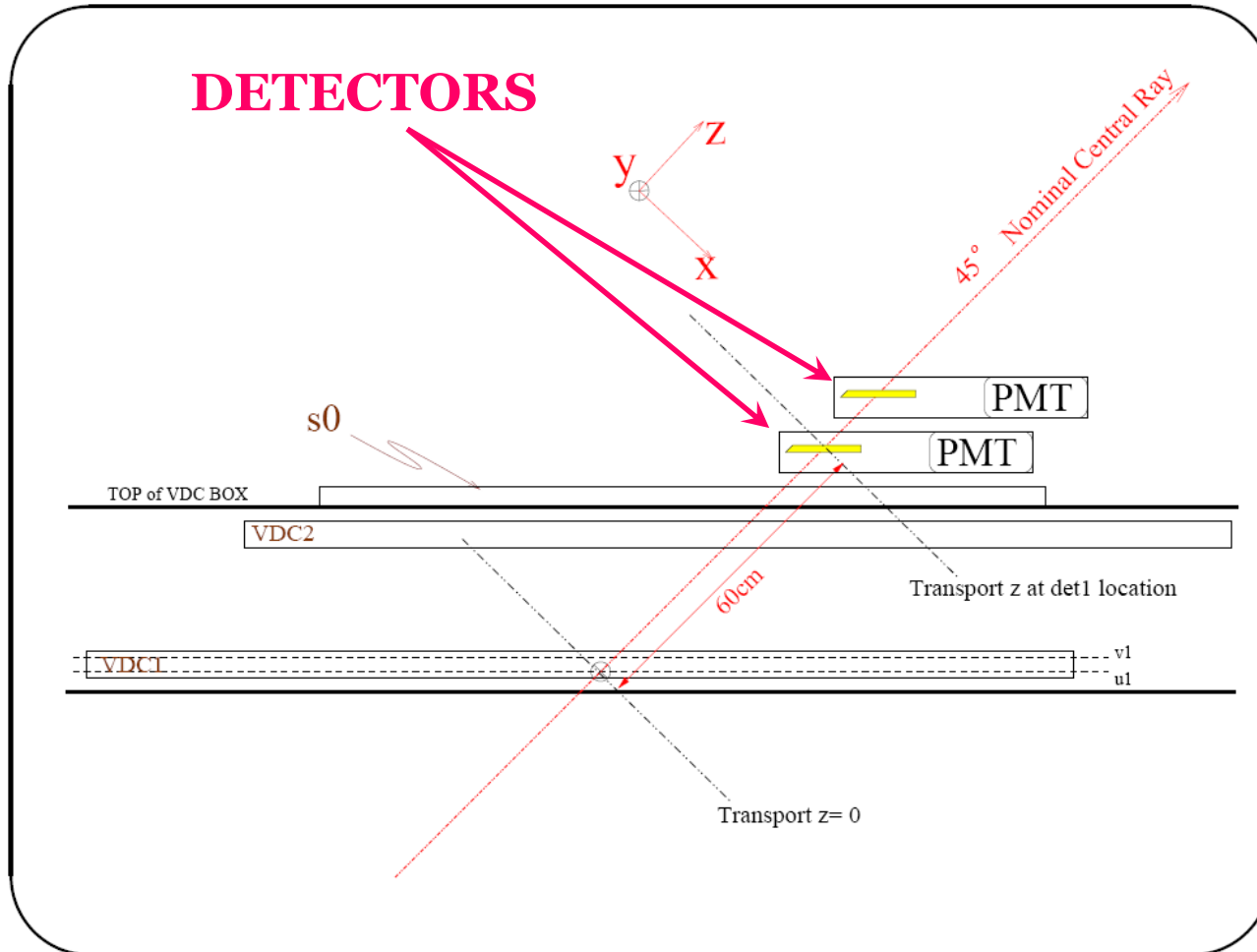
File Edit View Options Inspect Classes Help



Detector Package in HRS

PREX Integrating Detectors

UMass / Smith



Lead / Diamond Target

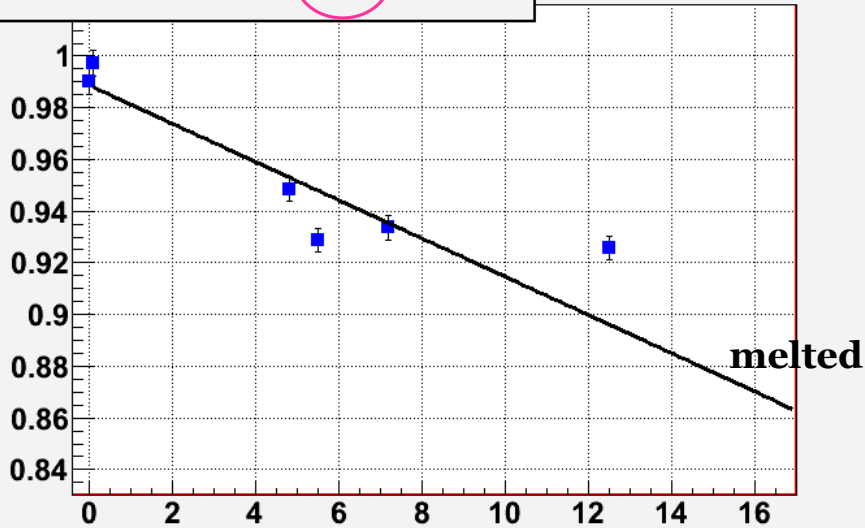
Diamond

LEAD

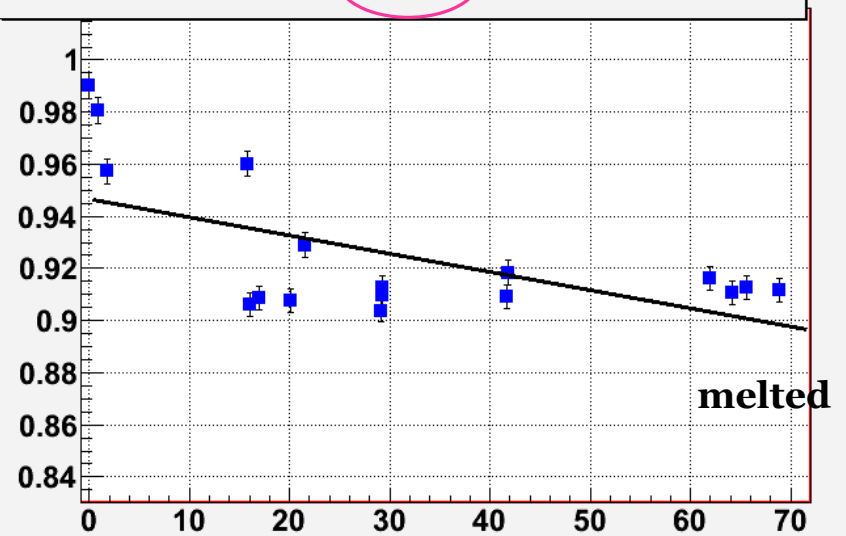
- Three bays
- **Lead** (0.5 mm)
sandwiched by
diamond (0.15 mm)
- Liquid He cooling (30 Watts)

Performance of Lead / Diamond Targets

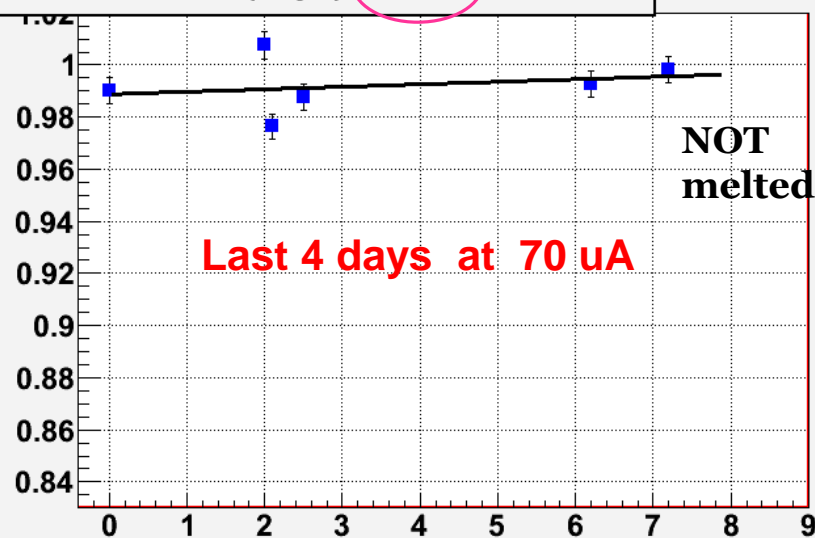
Rate vs Time (days) **THIN** Diamond



Rate vs Time (days) **MEDIUM**-thickness Diamond



Rate vs Time (days) **THICK** Diamond



Targets with thin diamond backing (4.5 % background) degraded fastest.

Thick diamond (8%) ran well and did not melt at 70 uA.

→ Solution: Run with 10 targets.

PREX-I Result

Systematic Errors

Error Source	Absolute (ppm)	Relative (%)
Polarization (1)	0.0083	1.3
Beam Asymmetries (2)	0.0072	1.1
Detector Linearity	0.0076	1.2
BCM Linearity	0.0010	0.2
Rescattering	0.0001	0
Transverse Polarization	0.0012	0.2
Q ² (1)	0.0028	0.4
Target Thickness	0.0005	0.1
¹² C Asymmetry (2)	0.0025	0.4
Inelastic States	0	0
TOTAL	0.0140	2.1

(1) Normalization Correction applied

(2) Nonzero correction (the rest assumed zero)

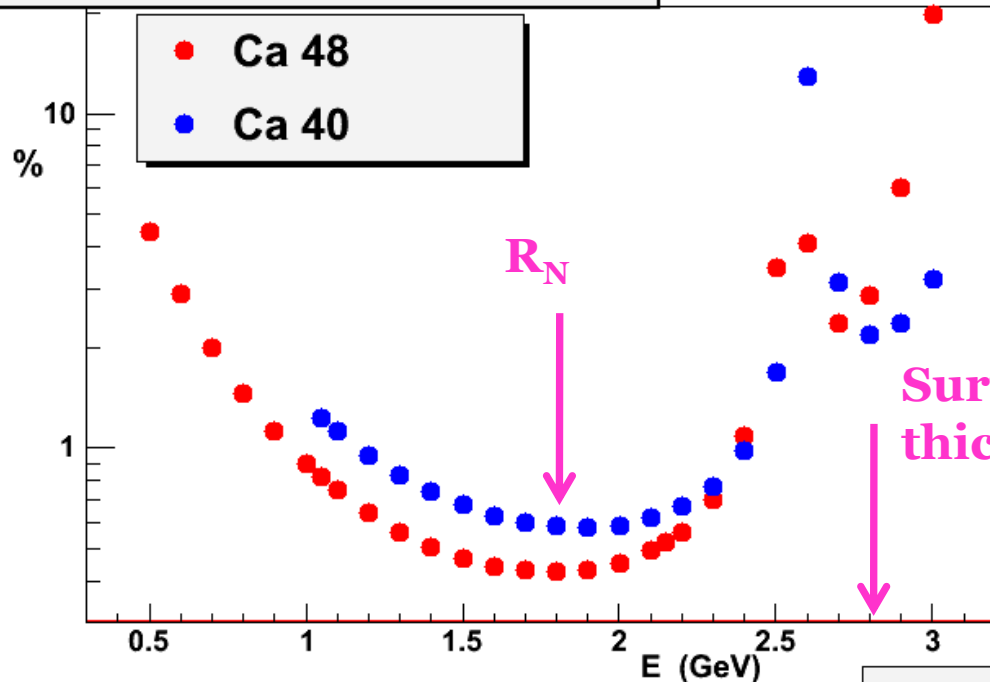
Physics Asymmetry

$$A = 0.656 \text{ ppm} \\ \pm 0.060(\text{stat}) \pm 0.014(\text{syst})$$

→ Statistics limited (9%)

→ Systematic error goal achieved! (2%)

Percent Error in R_N vs Energy (Calcium Isotopes)



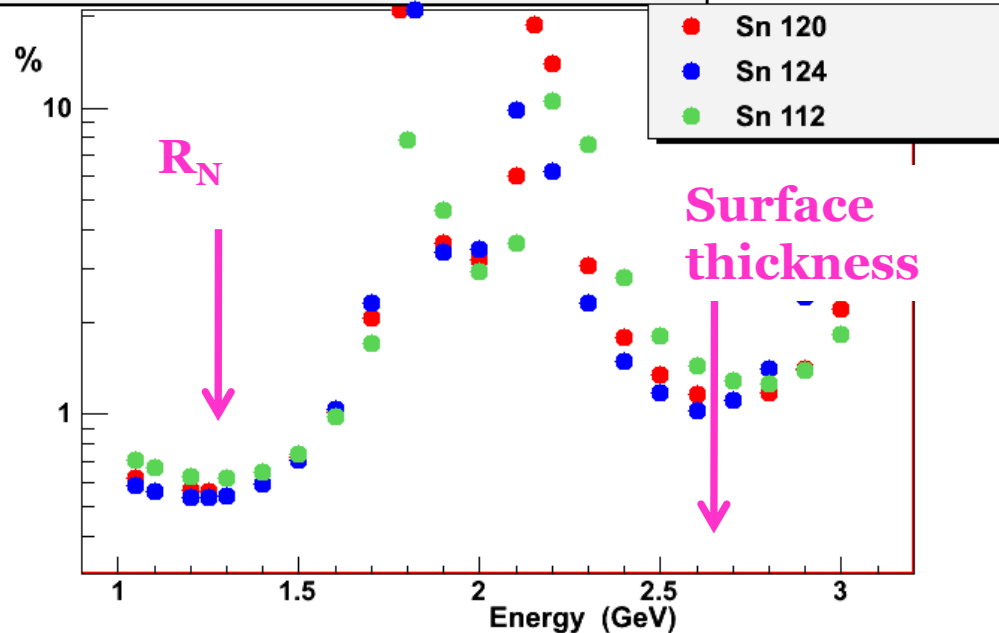
After PREX ...

Other Nuclei ?

and Shape Dependence ?

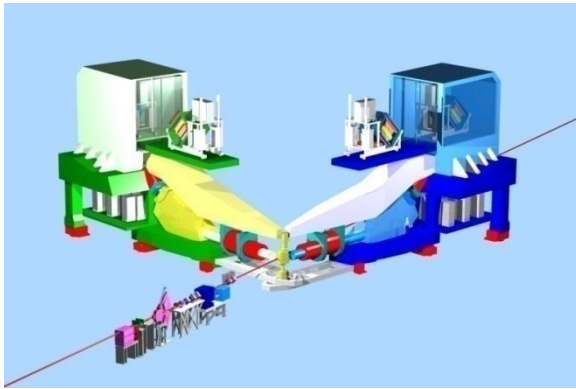
each point 30 days

Percent Error in R_N vs Energy (Tin Isotopes)

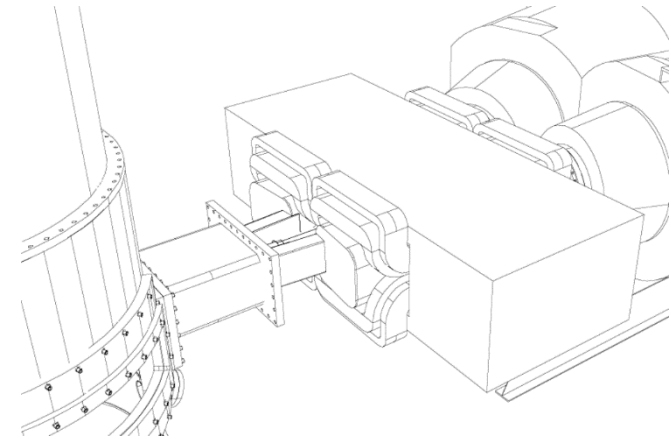


Parity Violating Electron Scattering
Measurements of Neutron Densities
Shufang Ban, C.J. Horowitz, R. Michaels

J. Phys. G39 014104 2012



Possible Future **PREX** Program ?



Each point 30 days

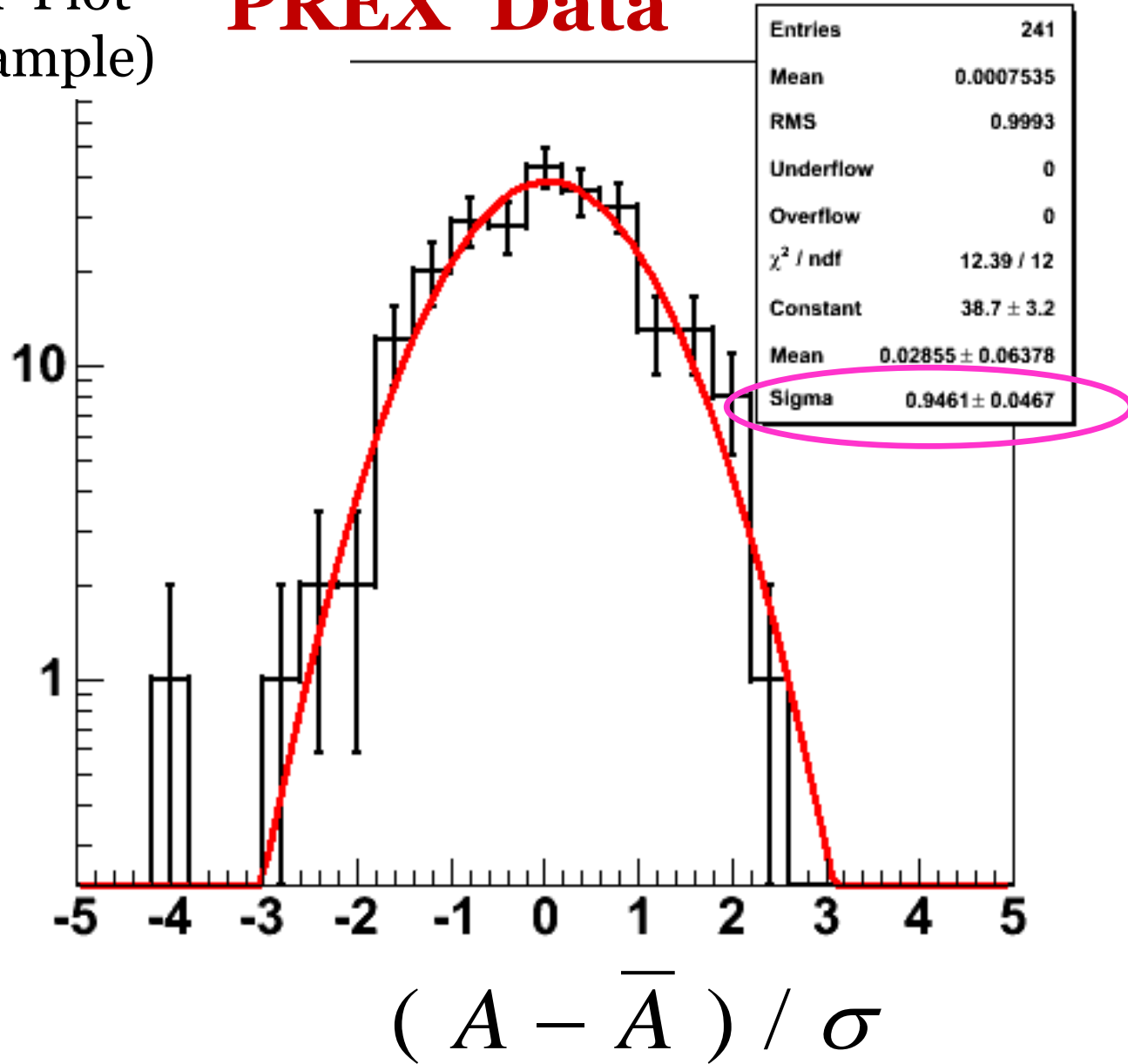
stat. error only

Nucleus	E (GeV)	dR_N / R_N	comment
^{208}Pb	1	1 %	PREX-II (approved by Jlab PAC, A rating)
^{48}Ca	2.2 (1-pass)	0.4 %	natural 12 GeV exp't will propose @ next PAC
^{48}Ca	2.6	2 %	surface thickness
^{40}Ca	2.2 (1-pass)	0.6 %	basic check of theory
tin isotope	1.8	0.6 %	apply to heavy ion
tin isotope	2.6	1.6 %	surface thickness

} Not proposed

Pull Plot
(example)

PREX Data



Corrections to the Asymmetry are Mostly Negligible

- **Coulomb Distortions** ~20% = the biggest correction.
- **Transverse Asymmetry** (to be measured)
- Strangeness
- Electric Form Factor of Neutron
- Parity Admixtures
- Dispersion Corrections
- Meson Exchange Currents
- Shape Dependence
- Isospin Corrections
- Radiative Corrections
- Excited States
- Target Impurities

Horowitz, *et.al.* PRC 63 025501

Optimum Kinematics for Lead Parity: $E = 1$ GeV if $\theta = 5^\circ$

$\langle A \rangle = 0.5$ ppm. Accuracy in Asy 3%

²⁰⁸Pb

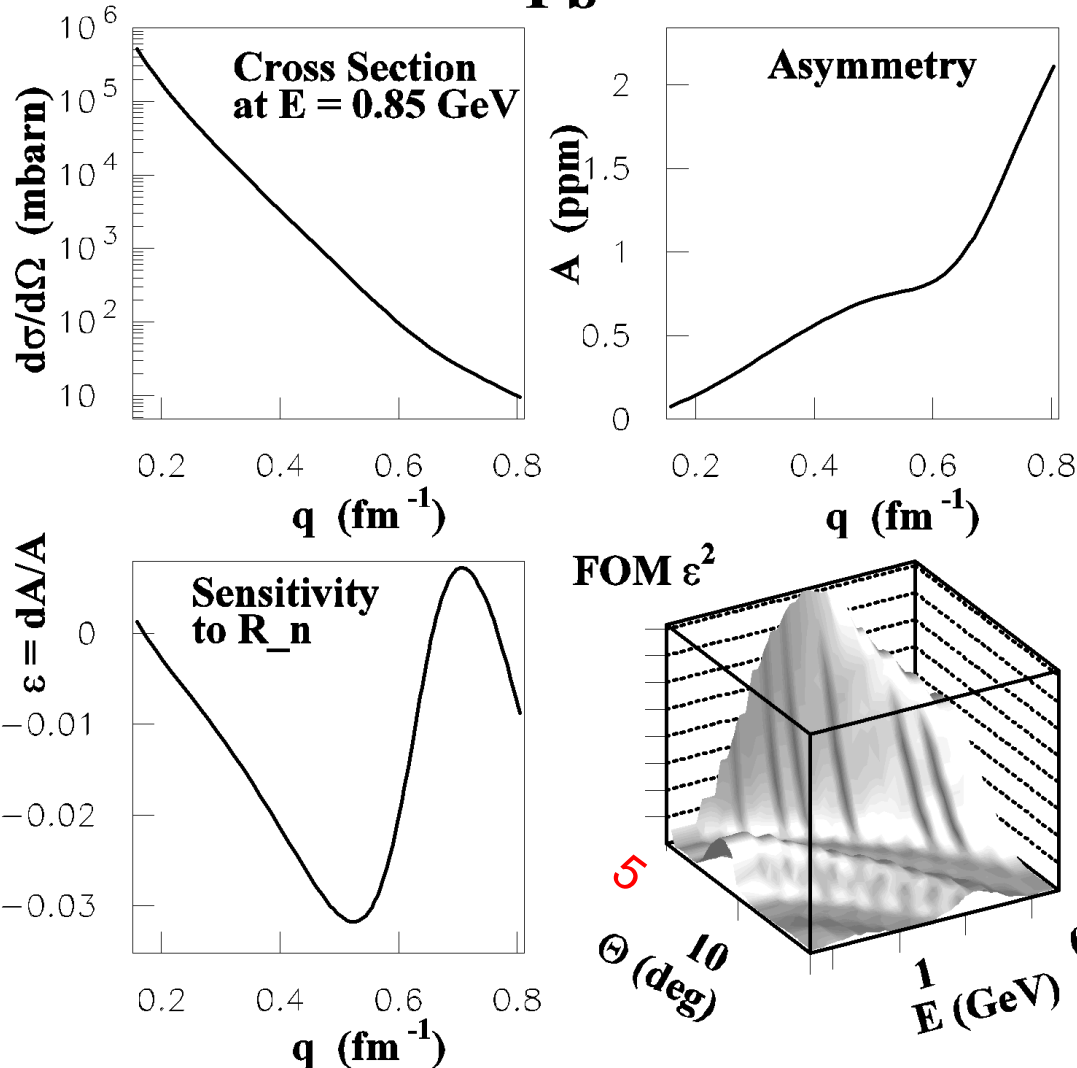


Fig. of merit

$$FOM = \frac{d\sigma}{d\Omega} \times A^2$$

Min. error in R_n

maximize:

$$\rightarrow FOM \times \epsilon^2$$

1 month run

$$\rightarrow 1\% \text{ in } R_n$$

(2 months x
100 uA
 \rightarrow 0.5% if no
systematics)